



中國科學院高能物理研究所
Institute of High Energy Physics
Chinese Academy of Sciences

AMS-02 Layer0 Tracker Upgrade

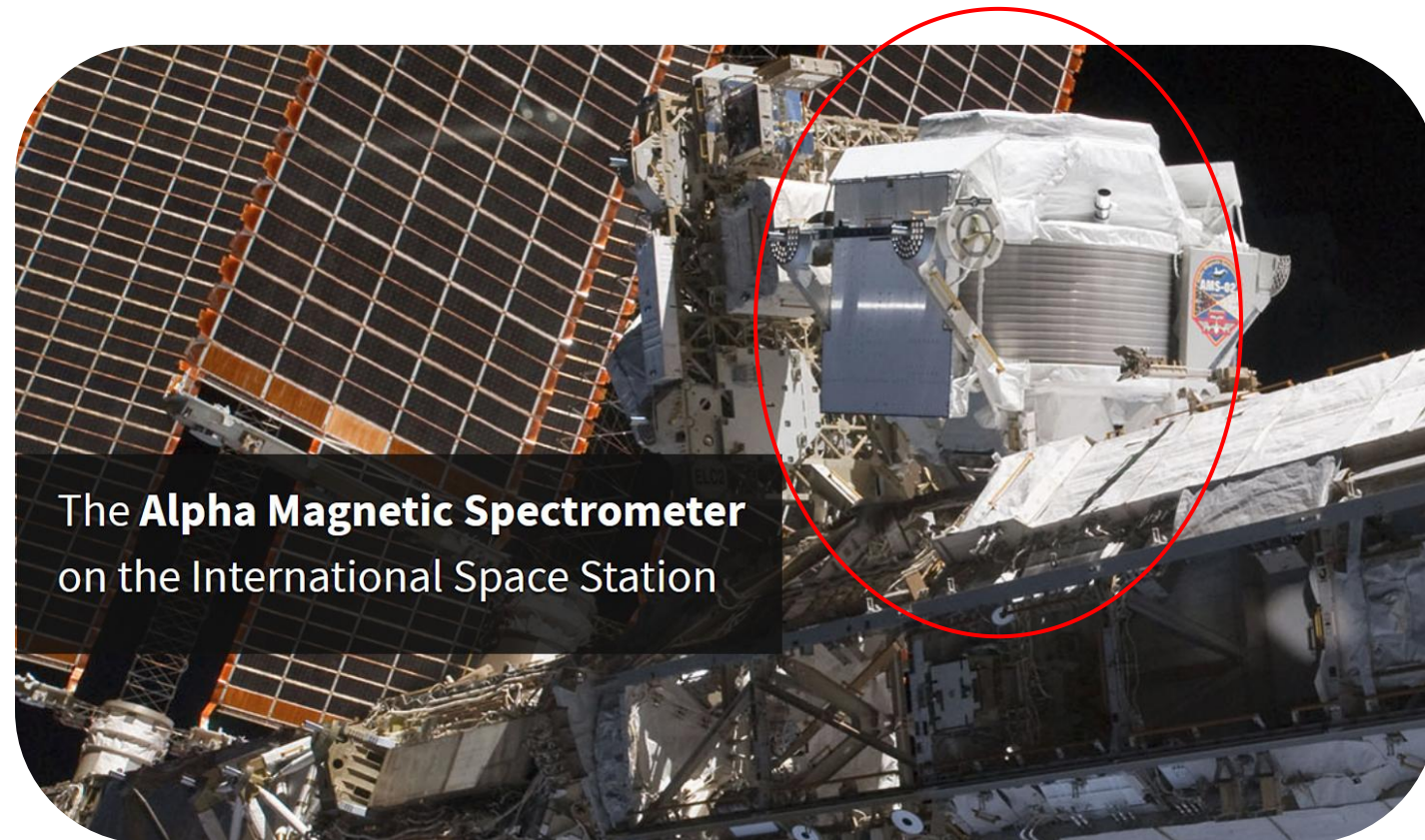
徐子駿 (IHEP, CAS)

海峽兩岸尖端探測器與技術研討會

2024年6月17-19日

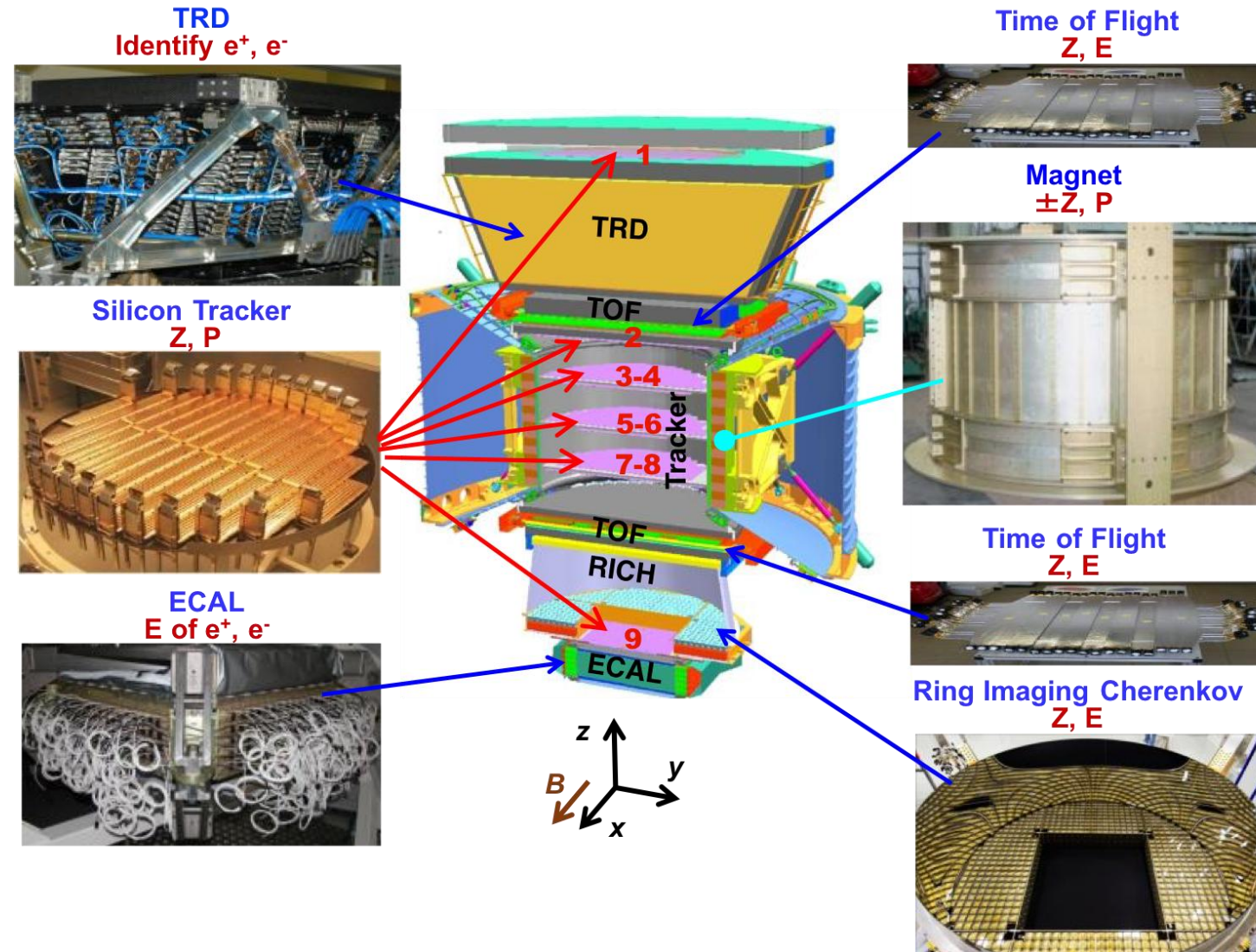
AMS-02

- launched to the ISS and installed on May 19, 2011
- taking data for the whole life of the ISS
- track deflection within its magnetic field => unique capability of distinguishing matter from anti-matter
- Main objectives:
 - search for Primordial Antimatter by direct detection of antinuclei
 - search for indirect Dark Matter signals
 - study of production, acceleration and propagation of Cosmic-Rays
 - study of Solar Modulation



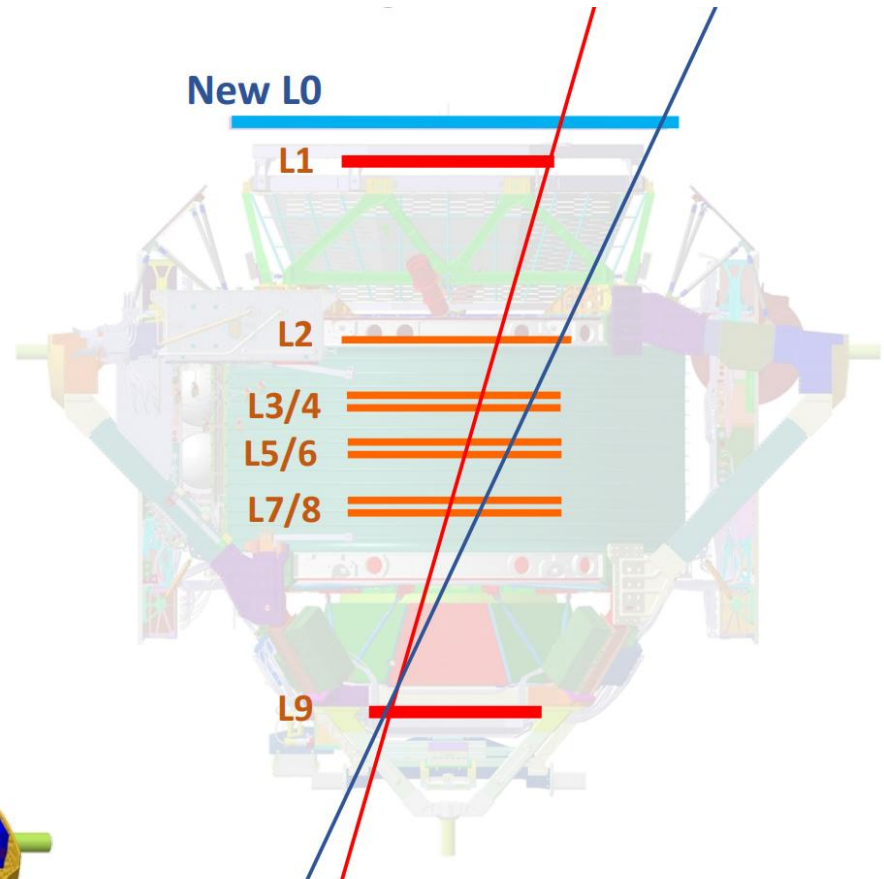
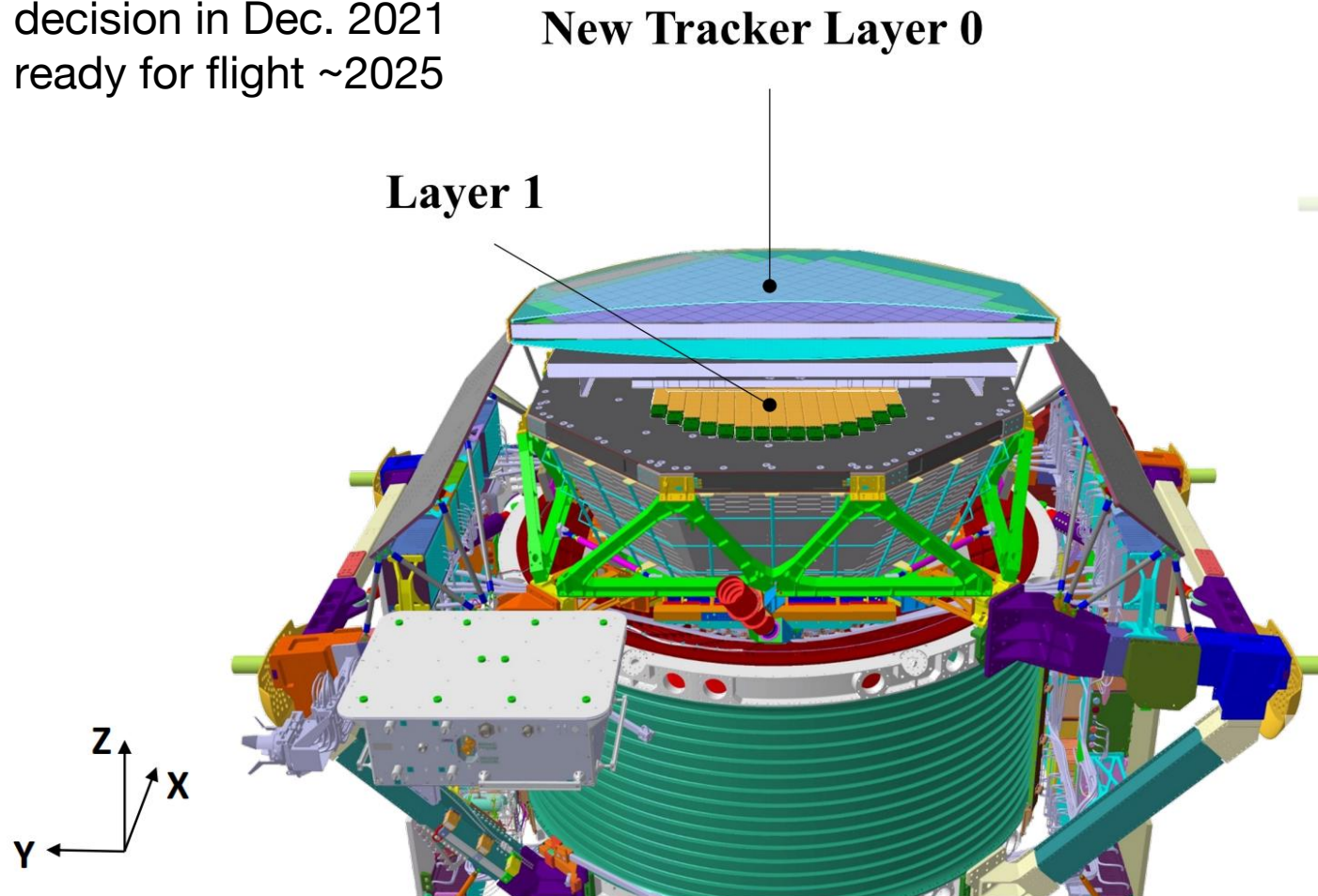
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New Tracker Layer 0 Added

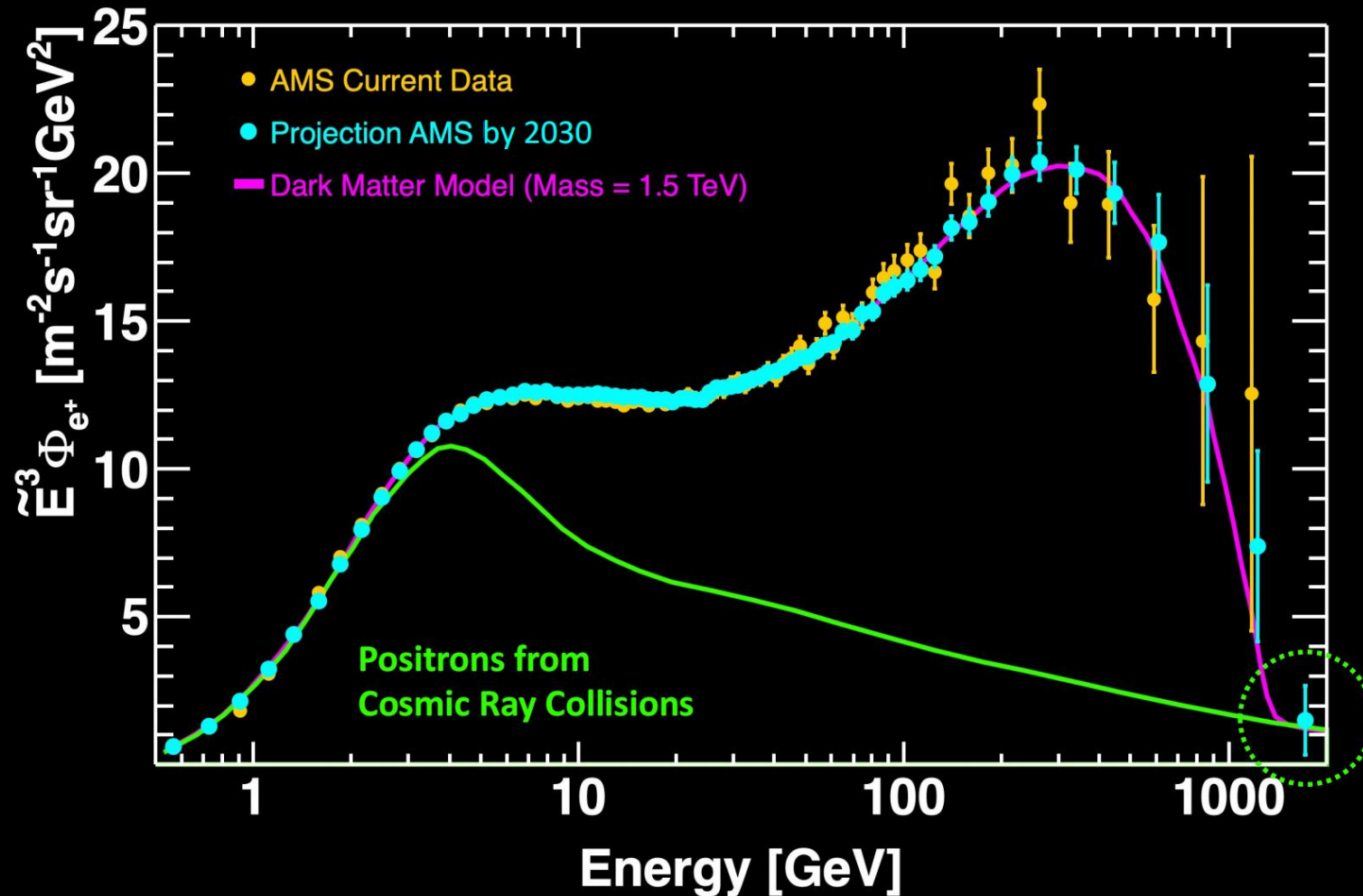
- decision in Dec. 2021
- ready for flight ~2025



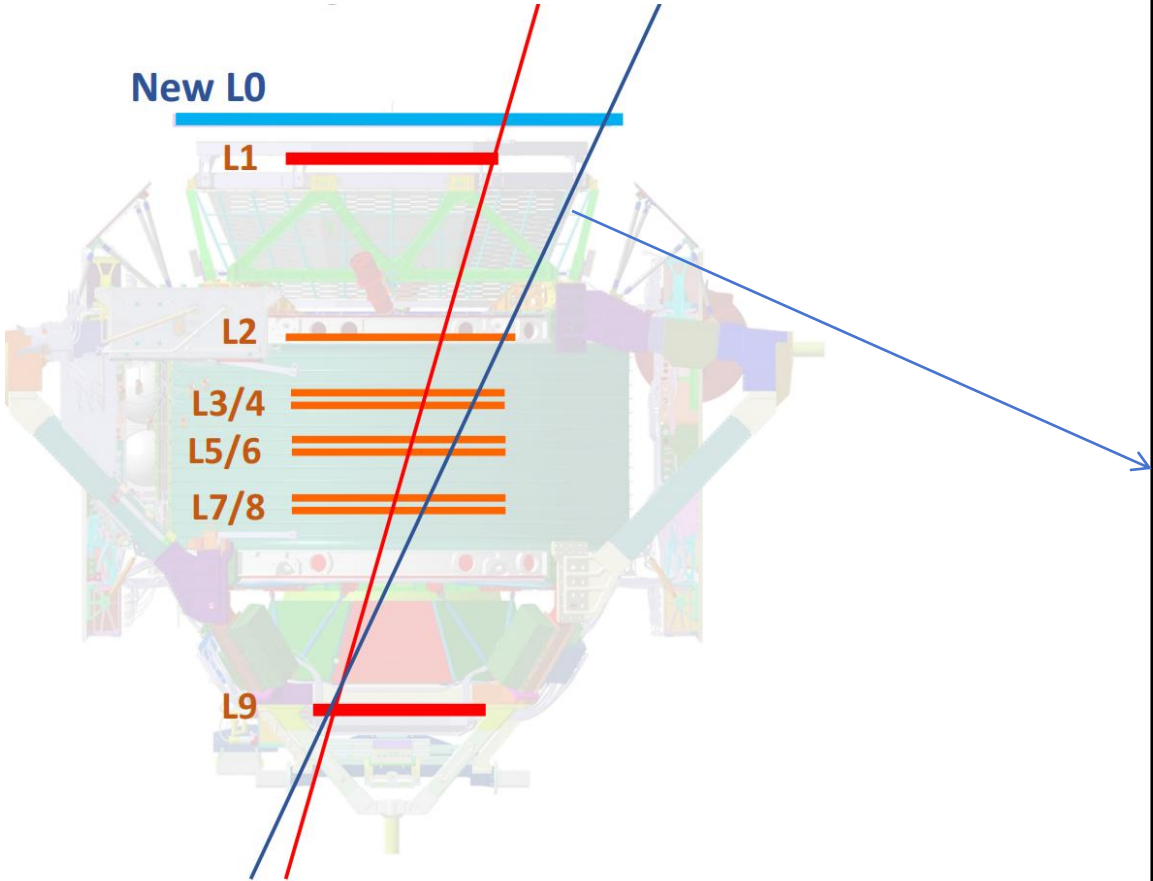
Increase of the detector acceptance by 300%

Determination of the Origin of Cosmic Positrons by 2030

AMS will ensure that the measured high energy positron spectrum indeed drops off quickly and, at the highest energies, the positrons only come from cosmic ray collisions as predicted by dark matter models



Layer 0



L0: ~8 m² of silicon detector
L1+L2+...+L9: ~6 m²

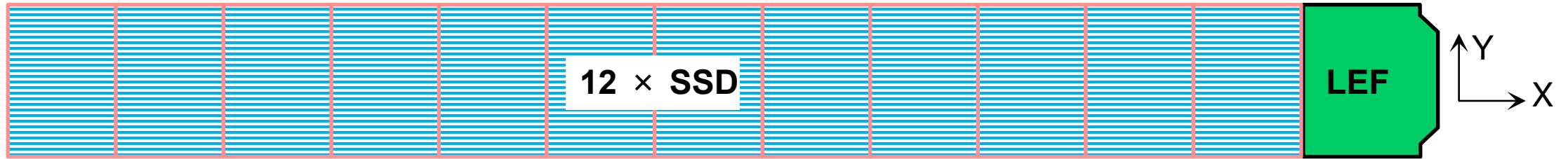
Detector Ladder

L0 tracker	
2	planes
72	ladders
768	detectors

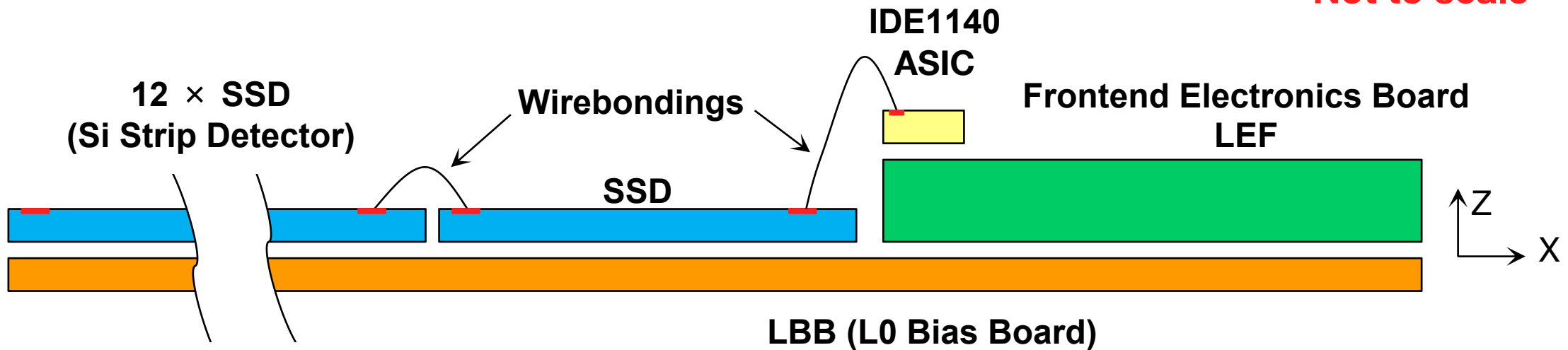
Top View

Bottom View

Illustration of A L0 Detector Ladder



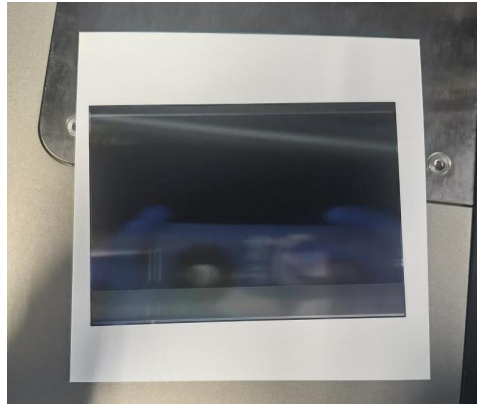
Not to scale



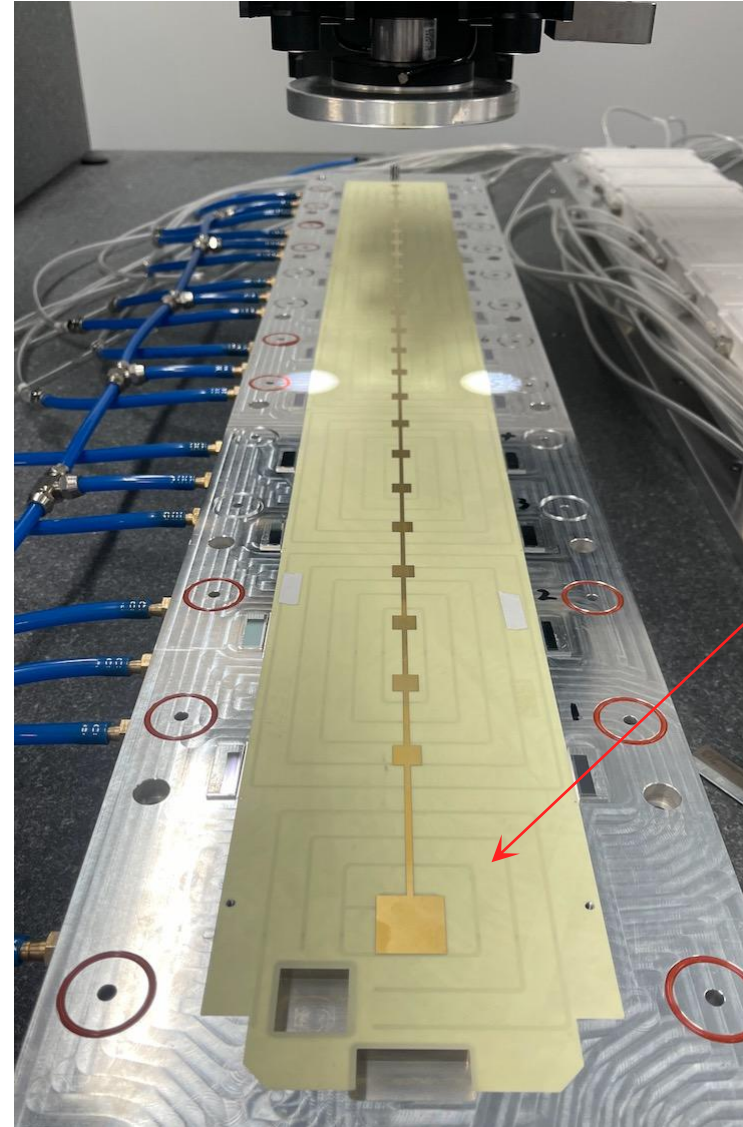
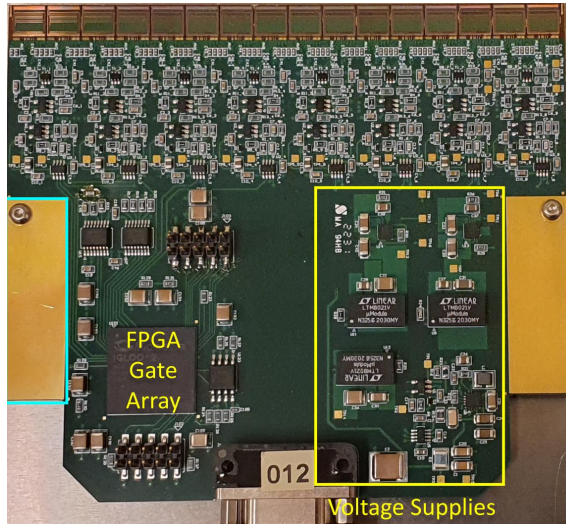
- ❑ Coupling and noise level due to long strips impose big challenges
- ❑ Precise placement of SSDs on a LBB affects the performance. We aim for $< 5 \mu\text{m}$ precision
- ❑ Highly efficient and reliable wire-bonding ($> 12\text{K}$ wires per ladder)

L0 Ladder Components

SSD



LEF



LBB

AMS L0 SiDet Laboratory at IHEP

Visual 3D Measuring System
(not visible from this angle)

Wire Bonding Machine

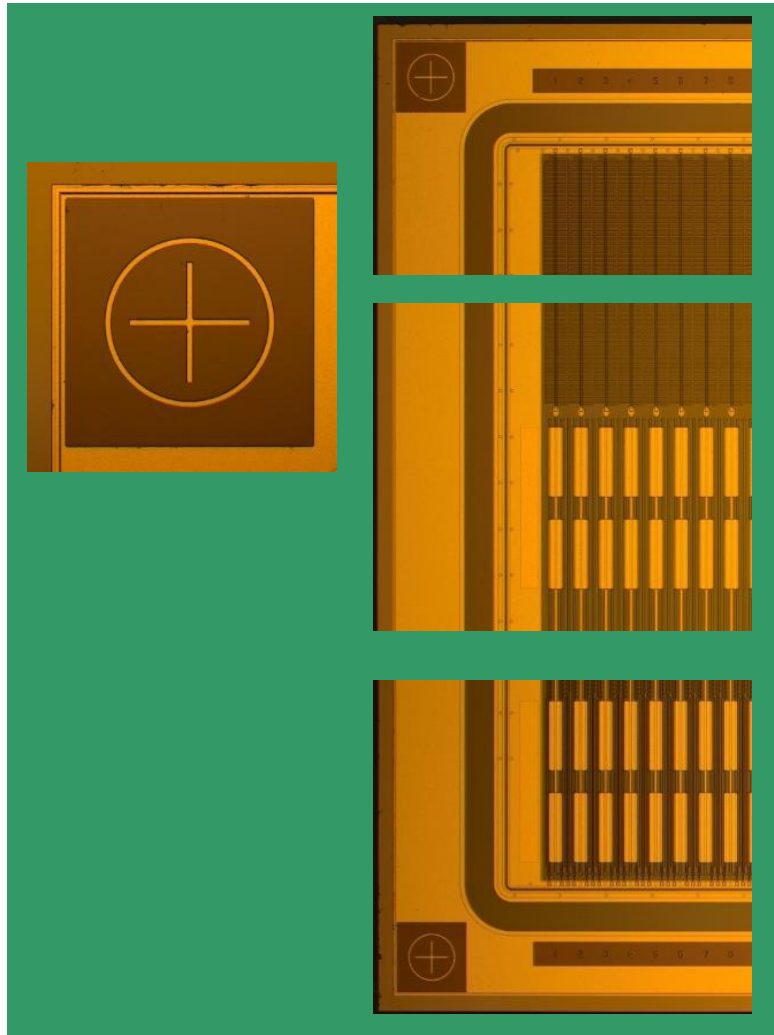


Pull Tester
for Q/A of wire-bonding

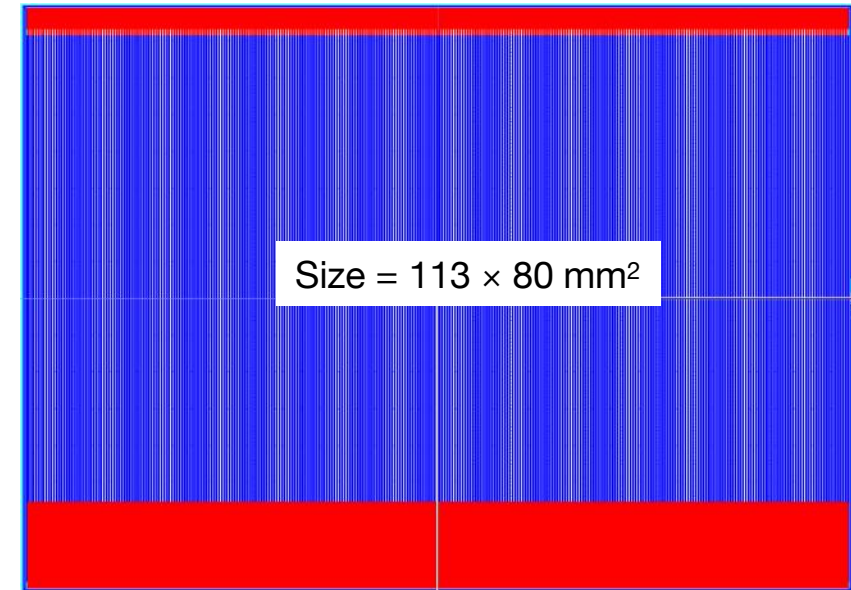
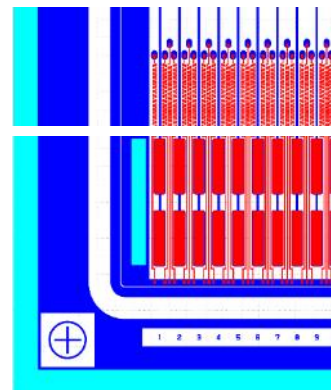
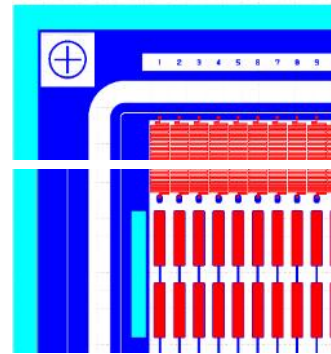
Gantry
for ladder assembling

Probe Station
for Silicon Strip Detector Q/A

The Silicon Strip Detector



Designed by IHEP+Perugia+HPK Swiss

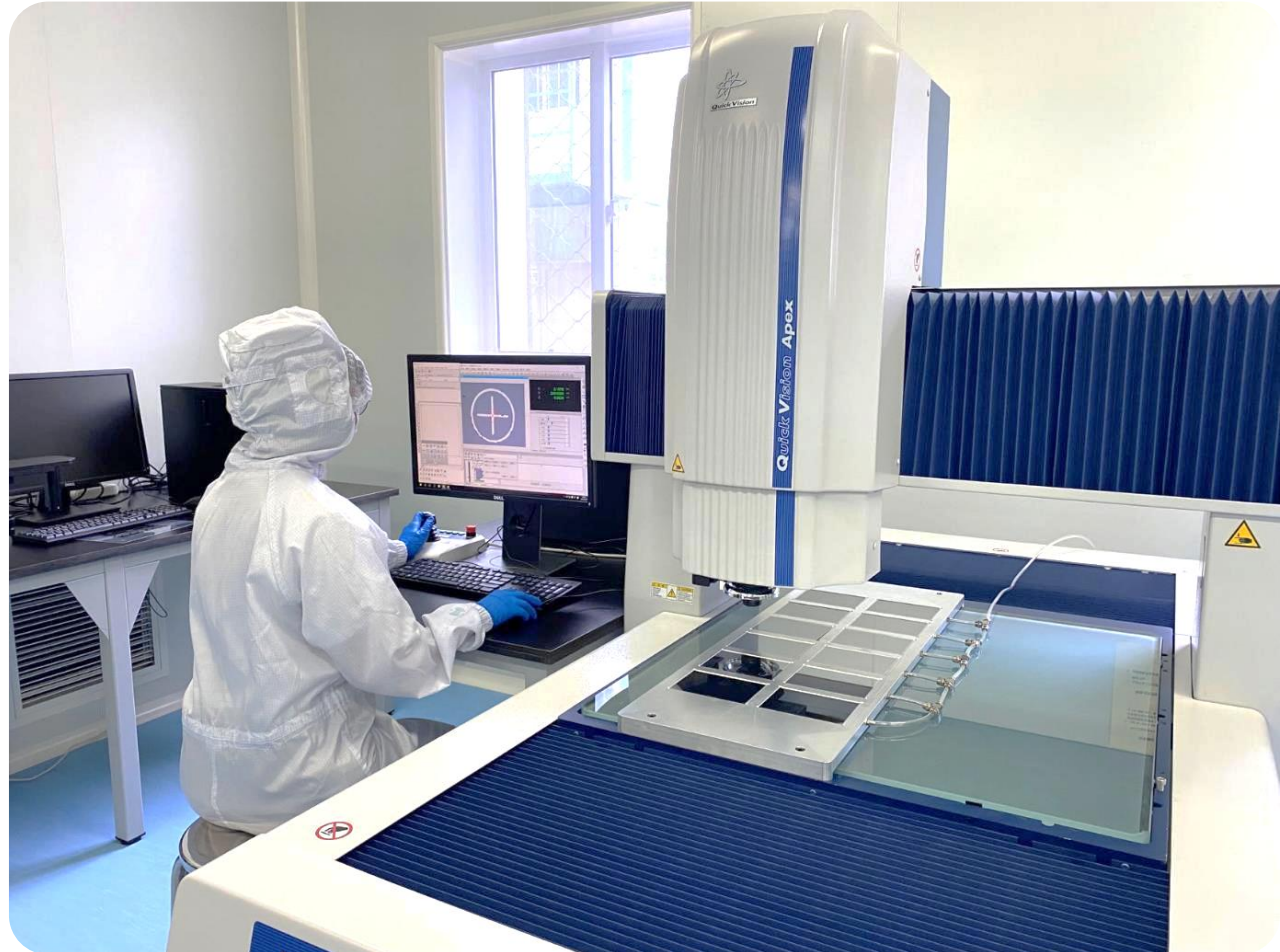
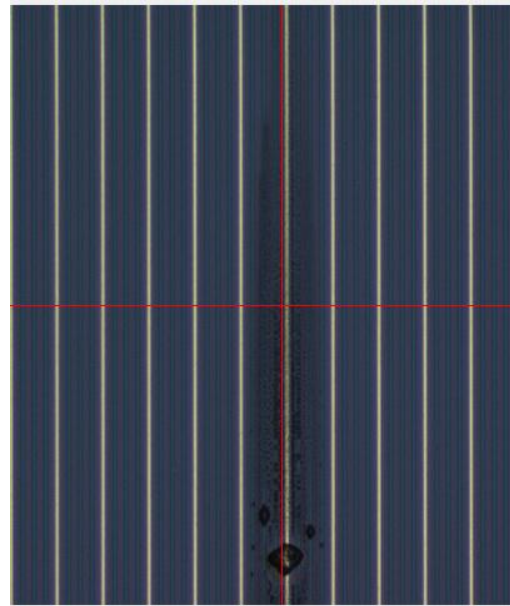
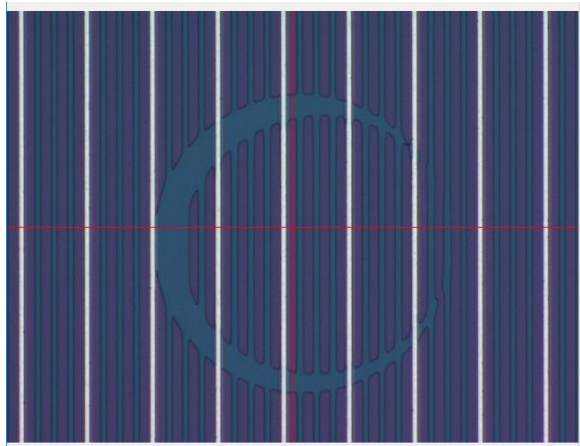


Highlights:

- Special design of readout pads / bias resistors
- Alignment marks for precise placement
- strip pitch 27.25 μm , readout pitch 109 μm

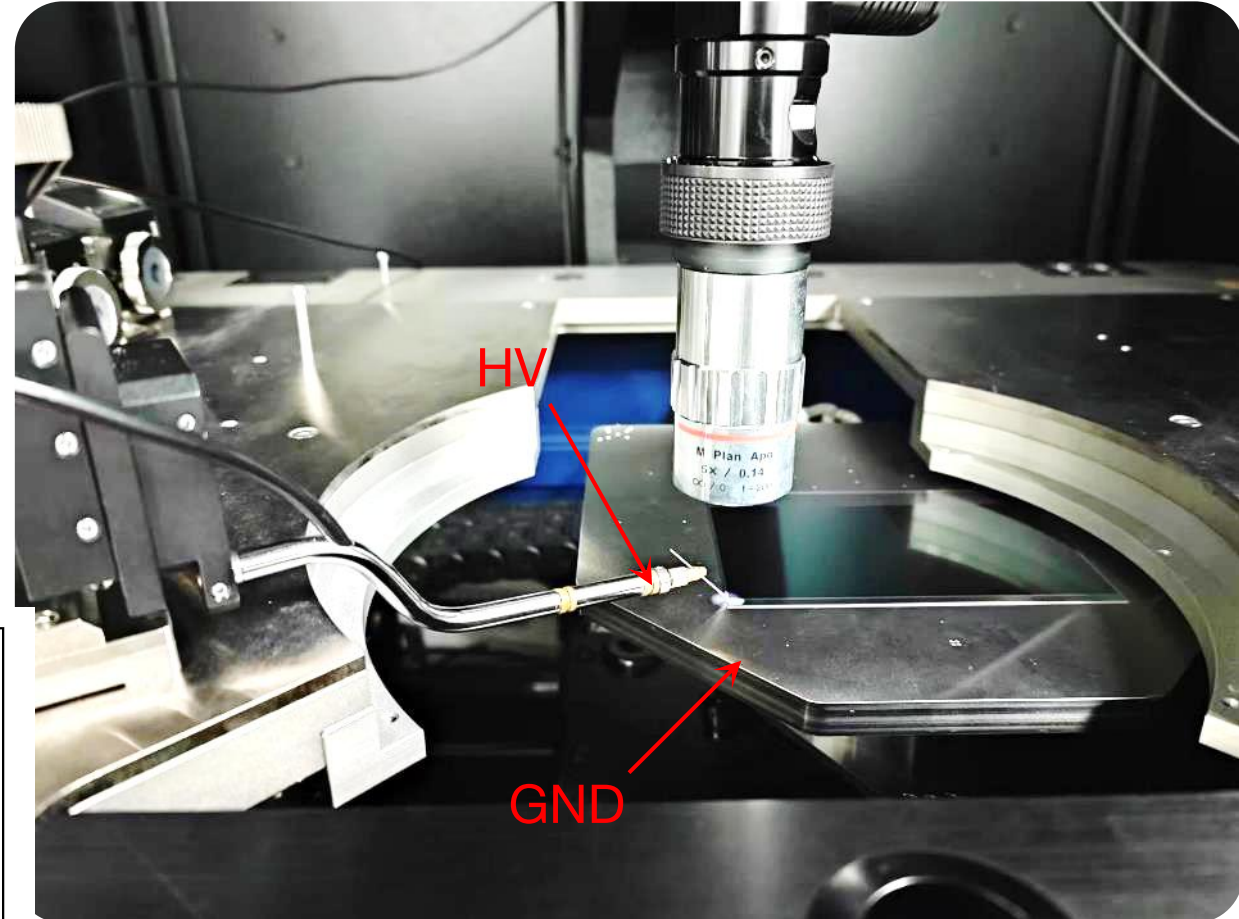
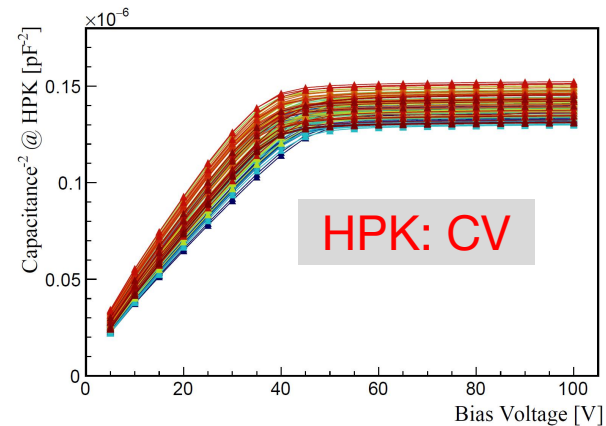
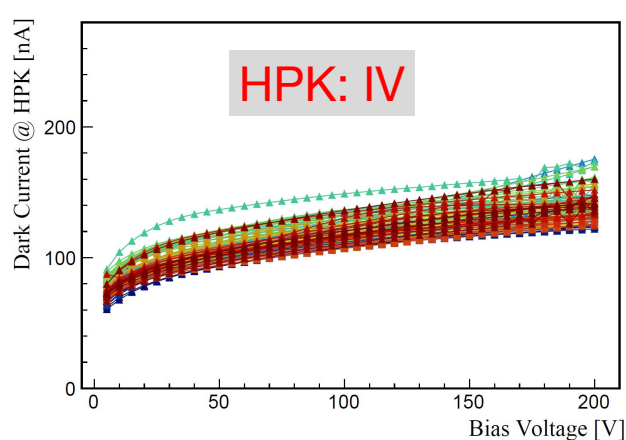
SSD Visual Inspection

- Identify debris, scratch or any other defect
- Generally there is no surprise



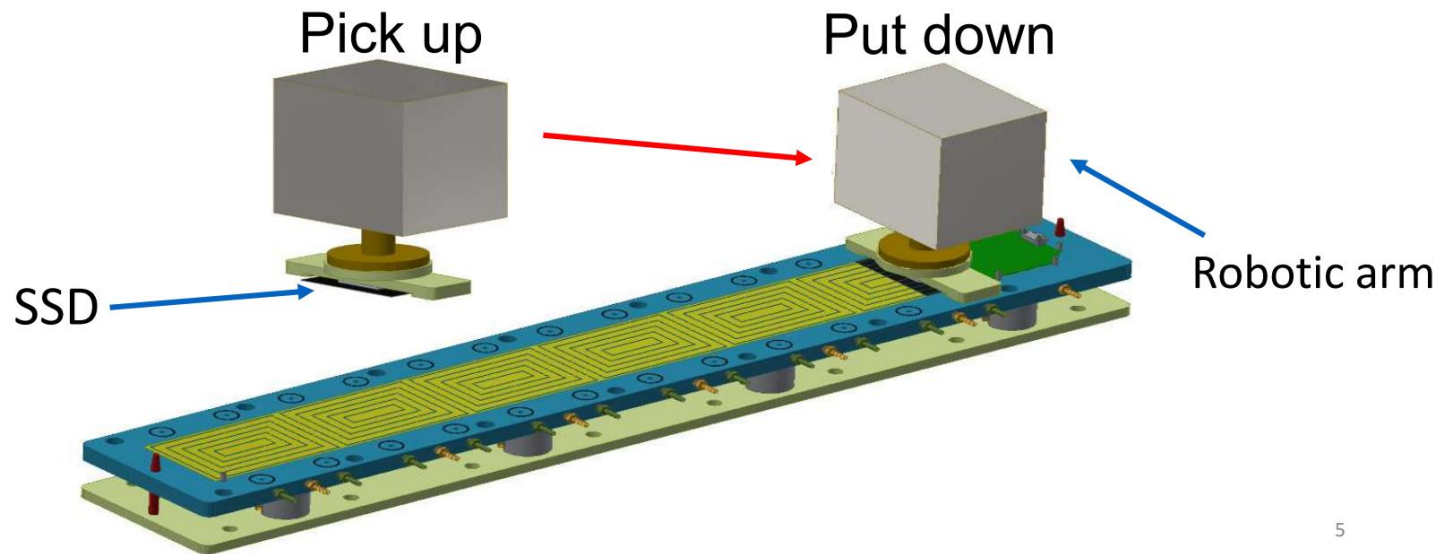
SSD IV & CV Scan

- All SSDs were characterized by HPK before the deliveries
- ~10% of all SSDs were re-tested at IHEP
 - consistent results between IHEP and HPK

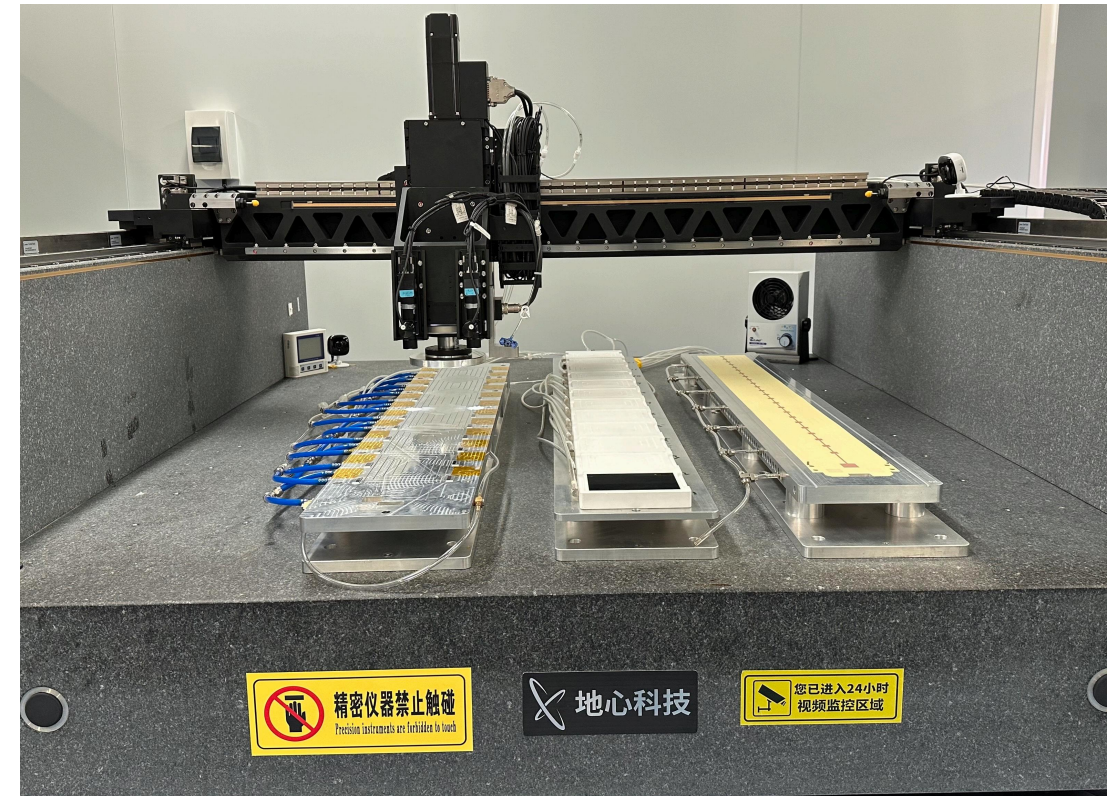


Ladder Production Procedure

- high-precision ($\sim 1\mu\text{m}$) gantry system for assembly

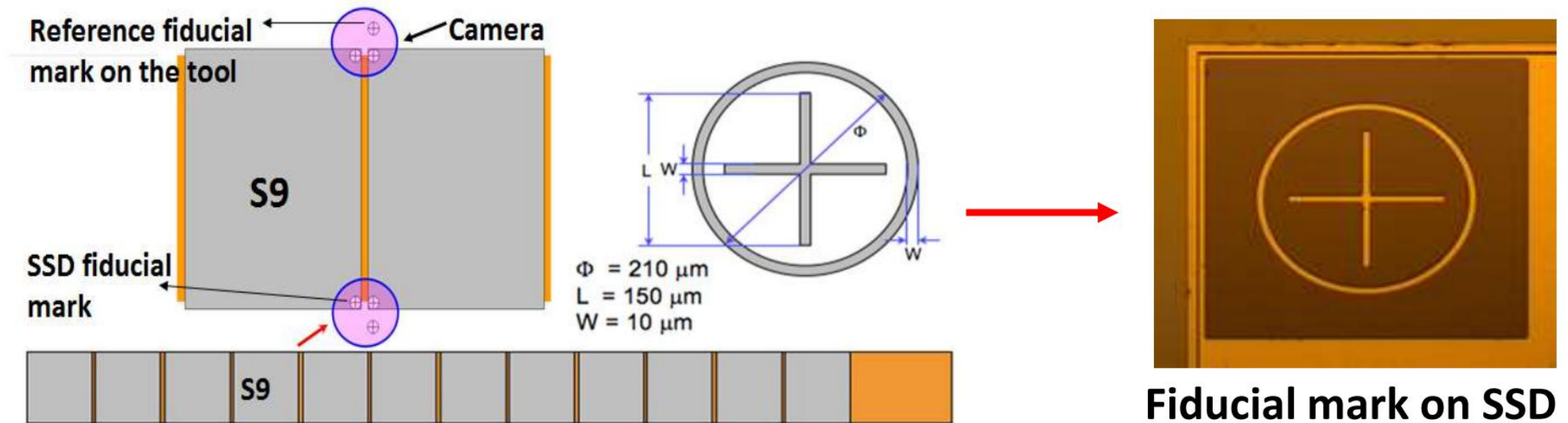


5



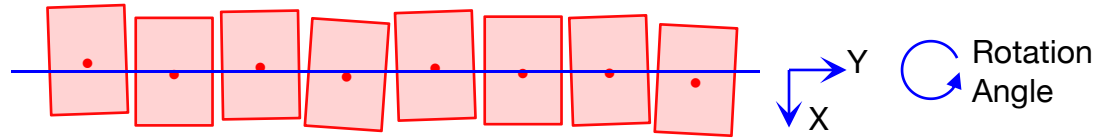
Ladder Production Procedure

- in-situ align the fiducial marks on SSDs during assembly

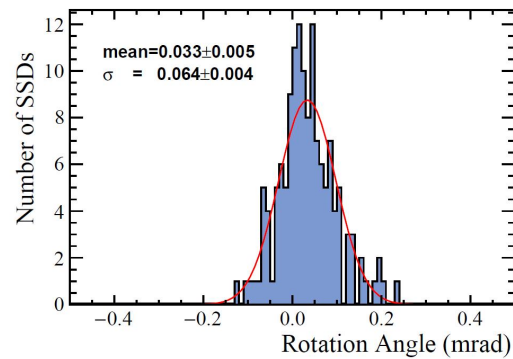
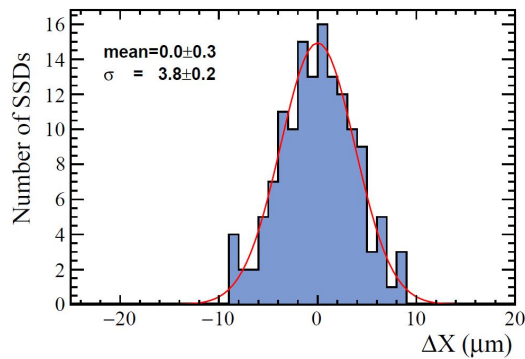


Metrology Using A Visual 3D Measuring System

Relative positions of all SSDs are precisely measured



8 dummy ladders & 7 QM ladders

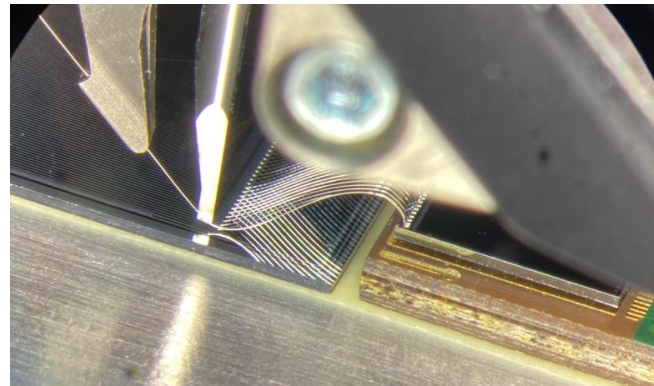
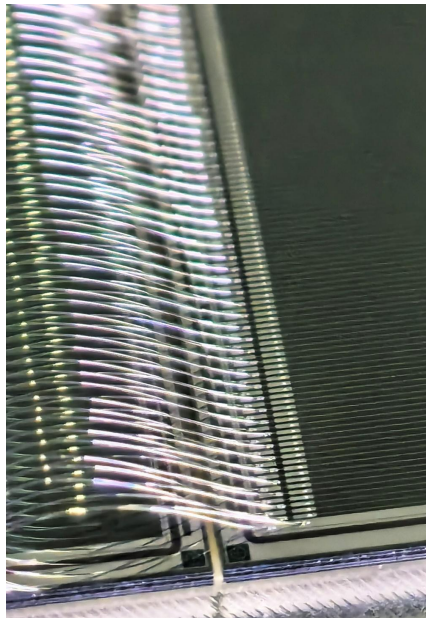
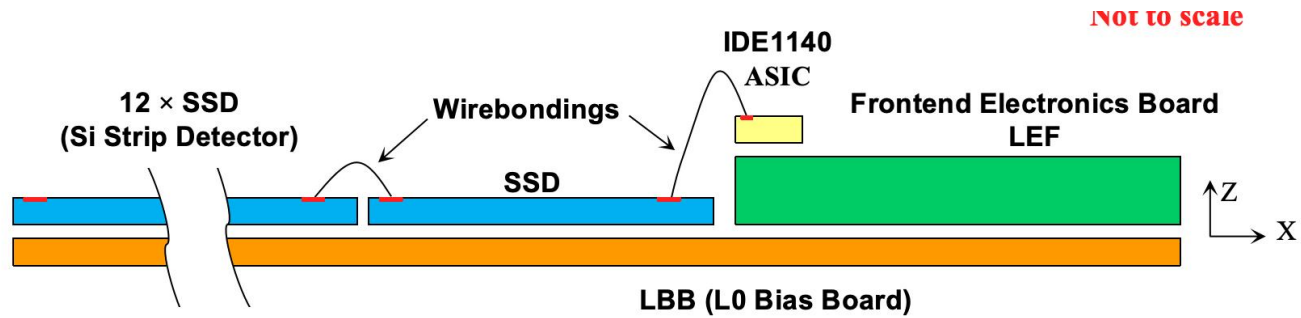


The **5 μm** precision goal has been achieved



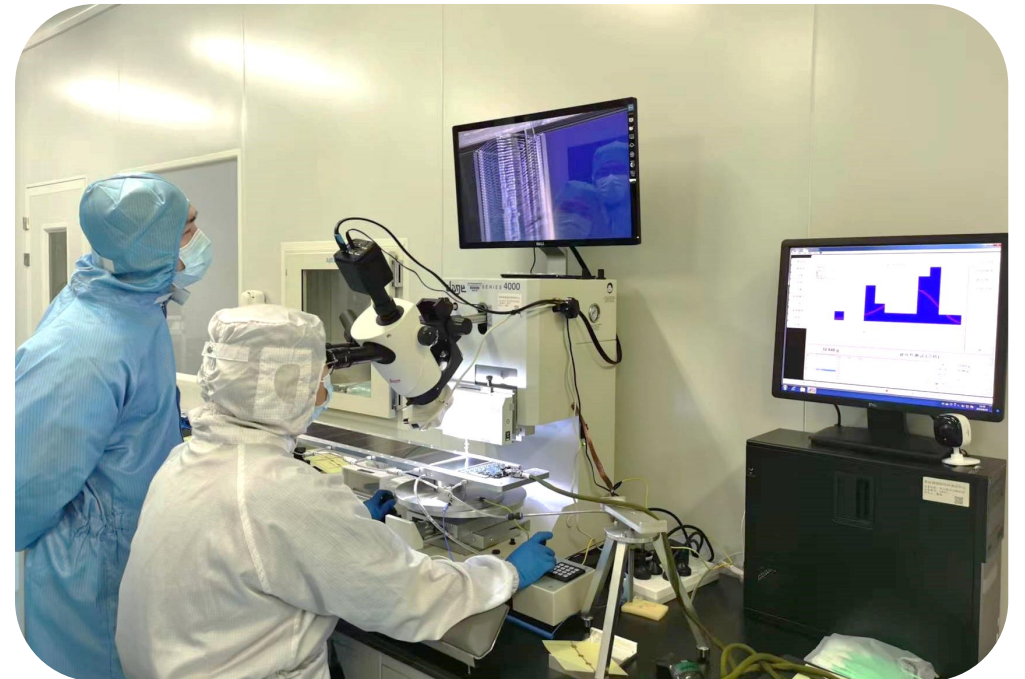
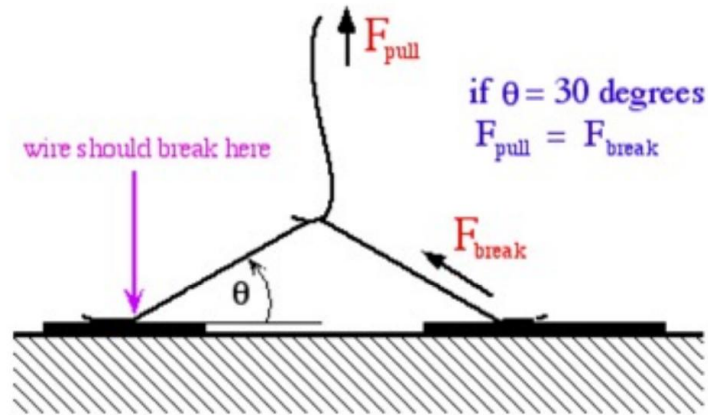
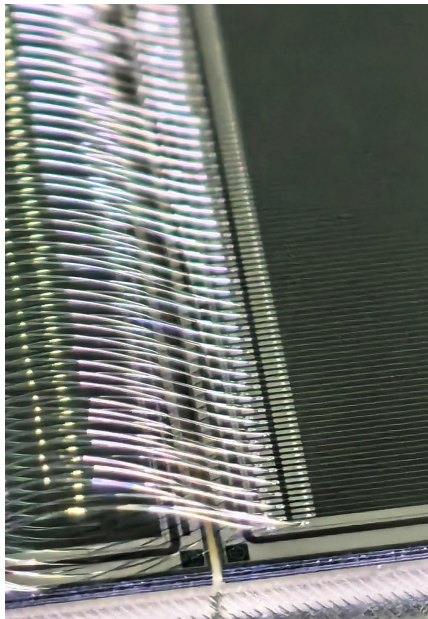
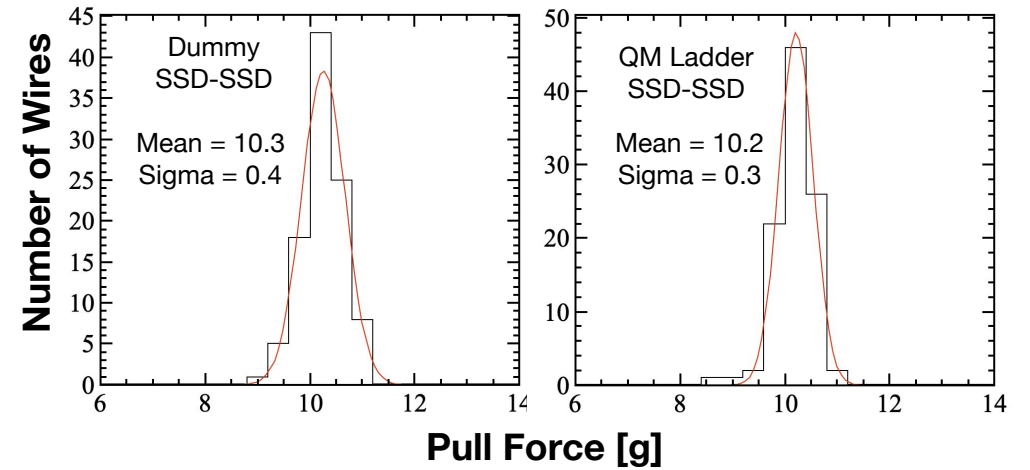
Wire Bonding Process

- Highly efficient and reliable wire-bonding (>12K wires per ladder)



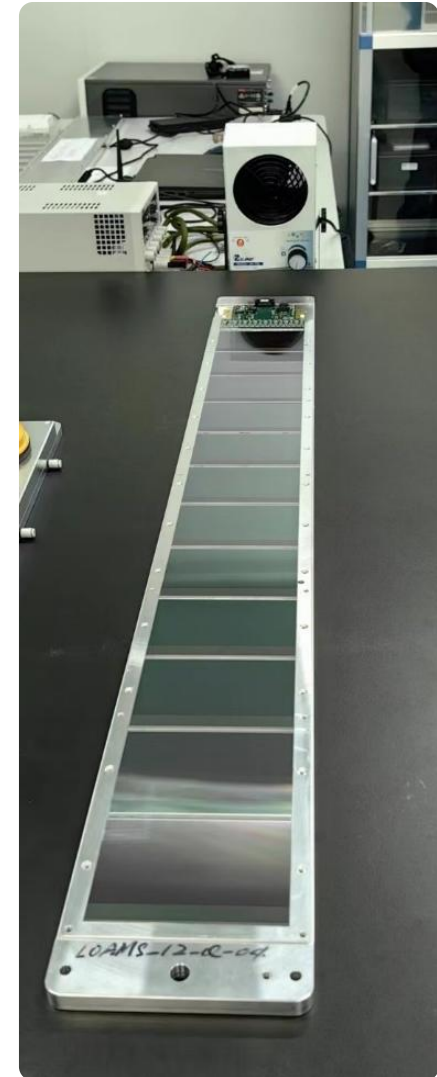
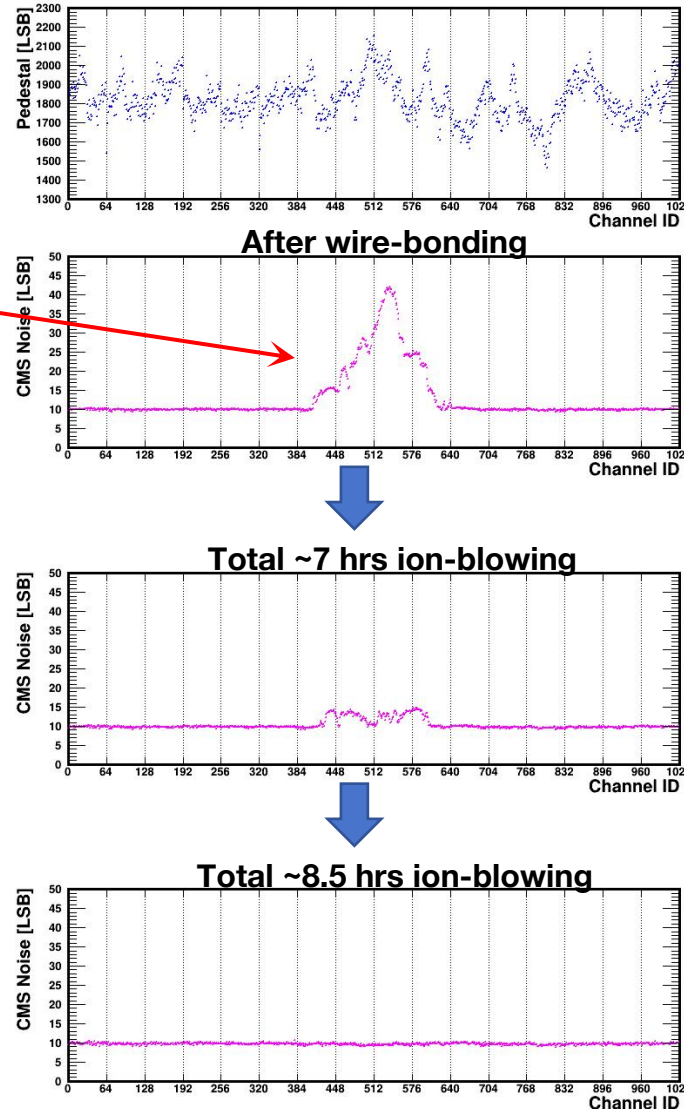
Pull Test

- ❑ High F_{break} means strong and reliable
- ❑ Regulate the bond force & ultrasonic power to optimize the F_{pull} (F_{break})



Ladder Electronics-Test: noise

- Large abnormal noise happend around the central region of the sensors
- Recovering after very long time ionizing-air blowing
- static charge from the vaccum pen when moving sensor:
“cleaning” the vaccum pen can avoid this issue

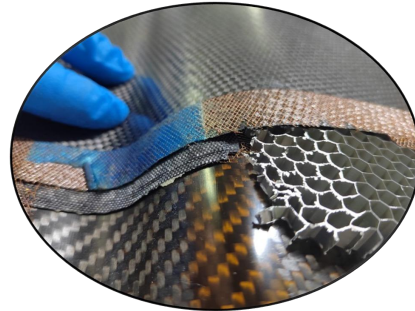
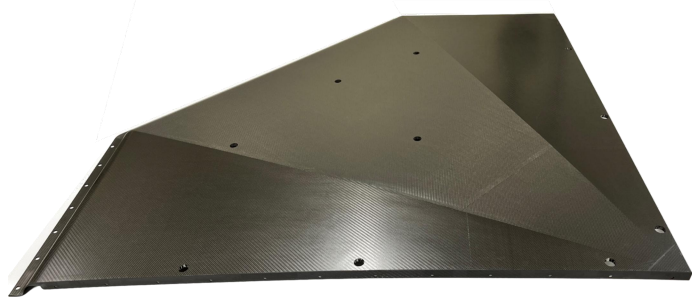


L0 Ladder Production

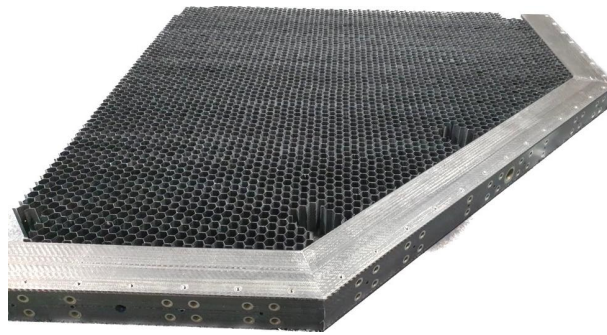
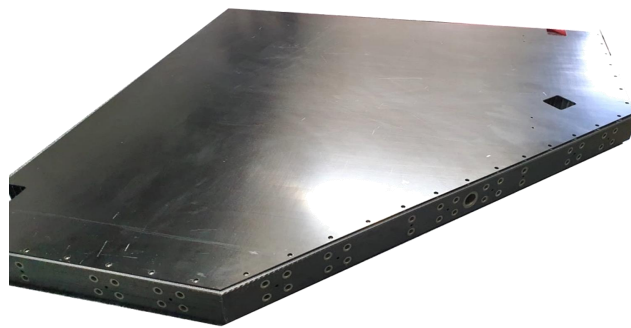


Layer0 Mechanics

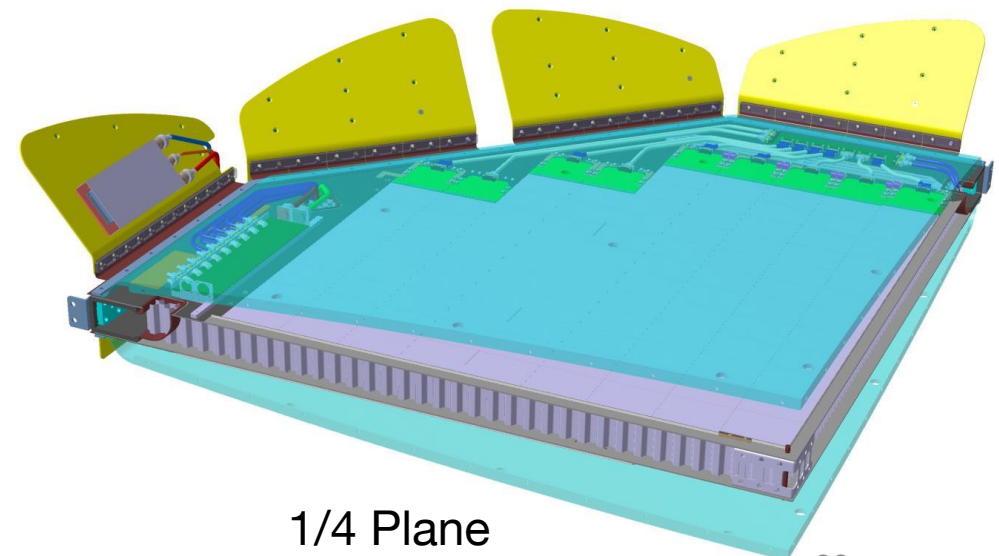
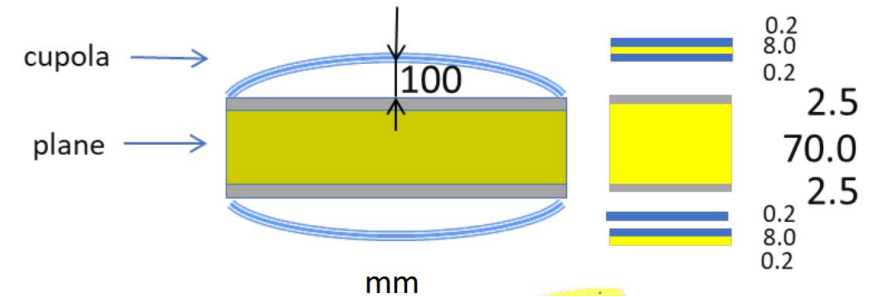
- Cupola Top & Bottom
 - EMI Shielding, Thermal radiator, Light shielding



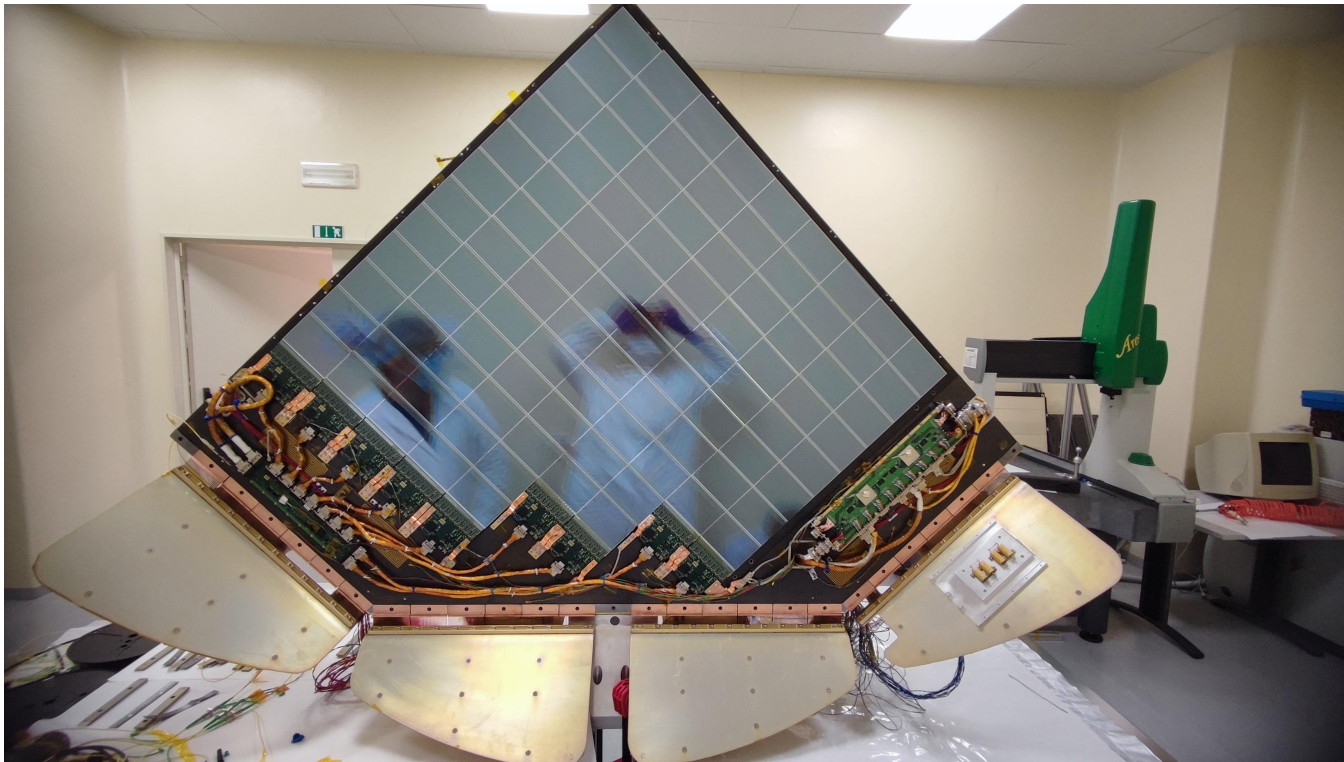
- Carbon-plane and carbon-frame realize a thick stiff plane



Dimensions
~2.6m diameter
~ 30 cm thick
~ 250 kg



1/4 Plane Integration

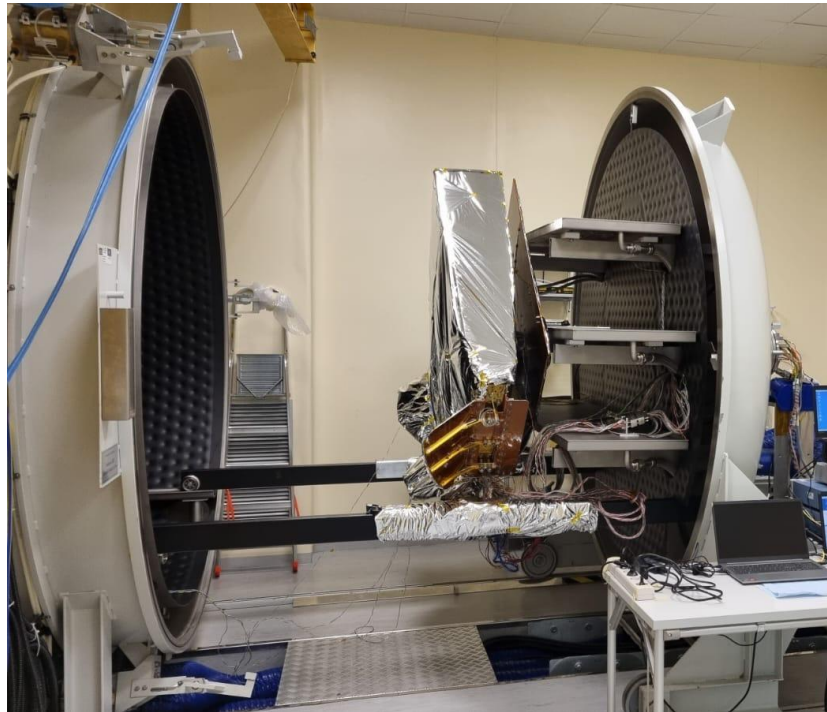


1/4 Plane with ladders and electronics



1/4 Plane with Cupola

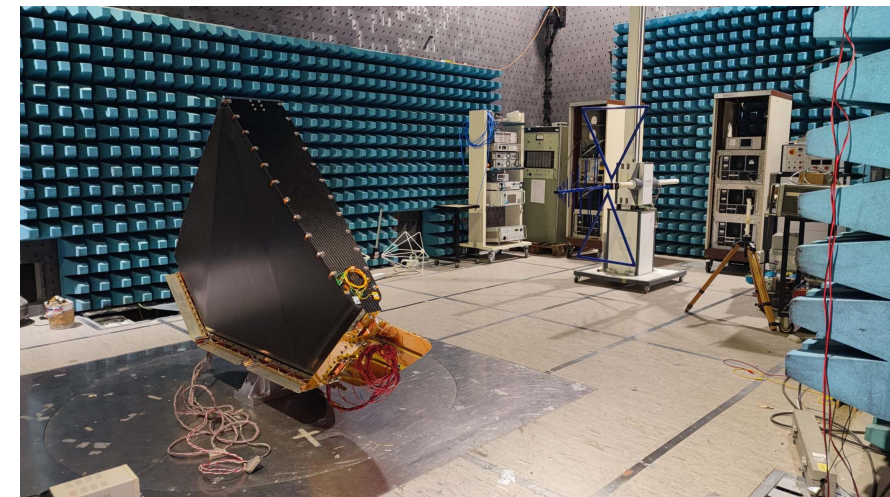
1/4 Plane Testing



Thermal Vacuum test

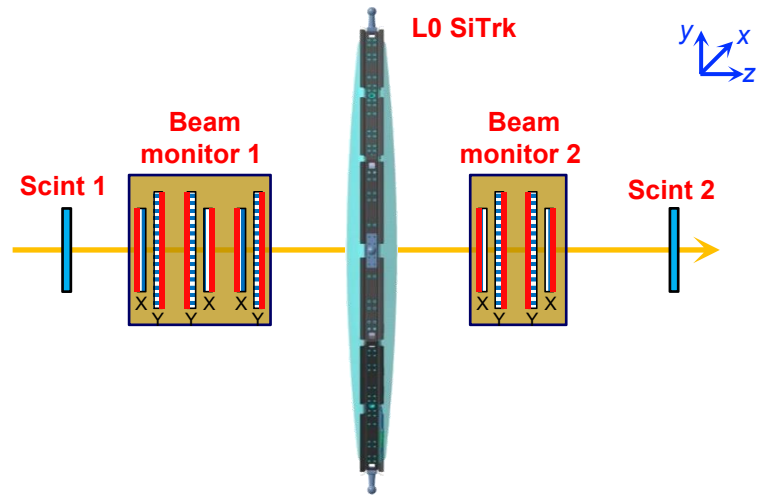


vibration test: no bonding-wire failure

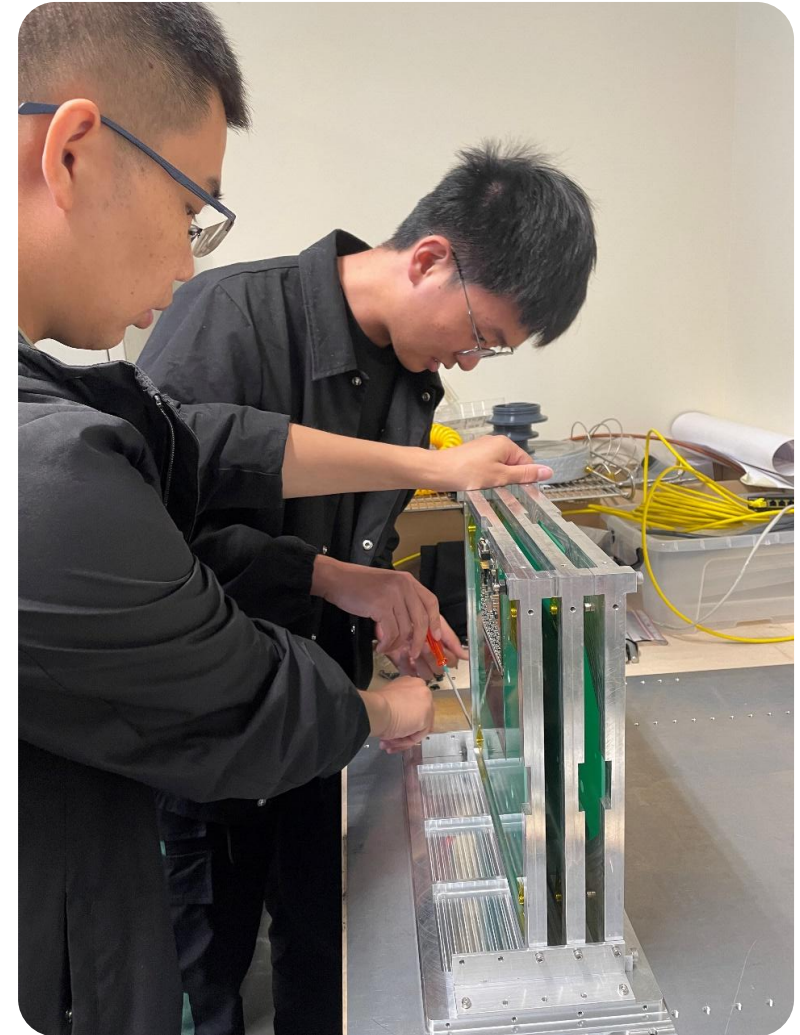


EMI test

Beam Monitors

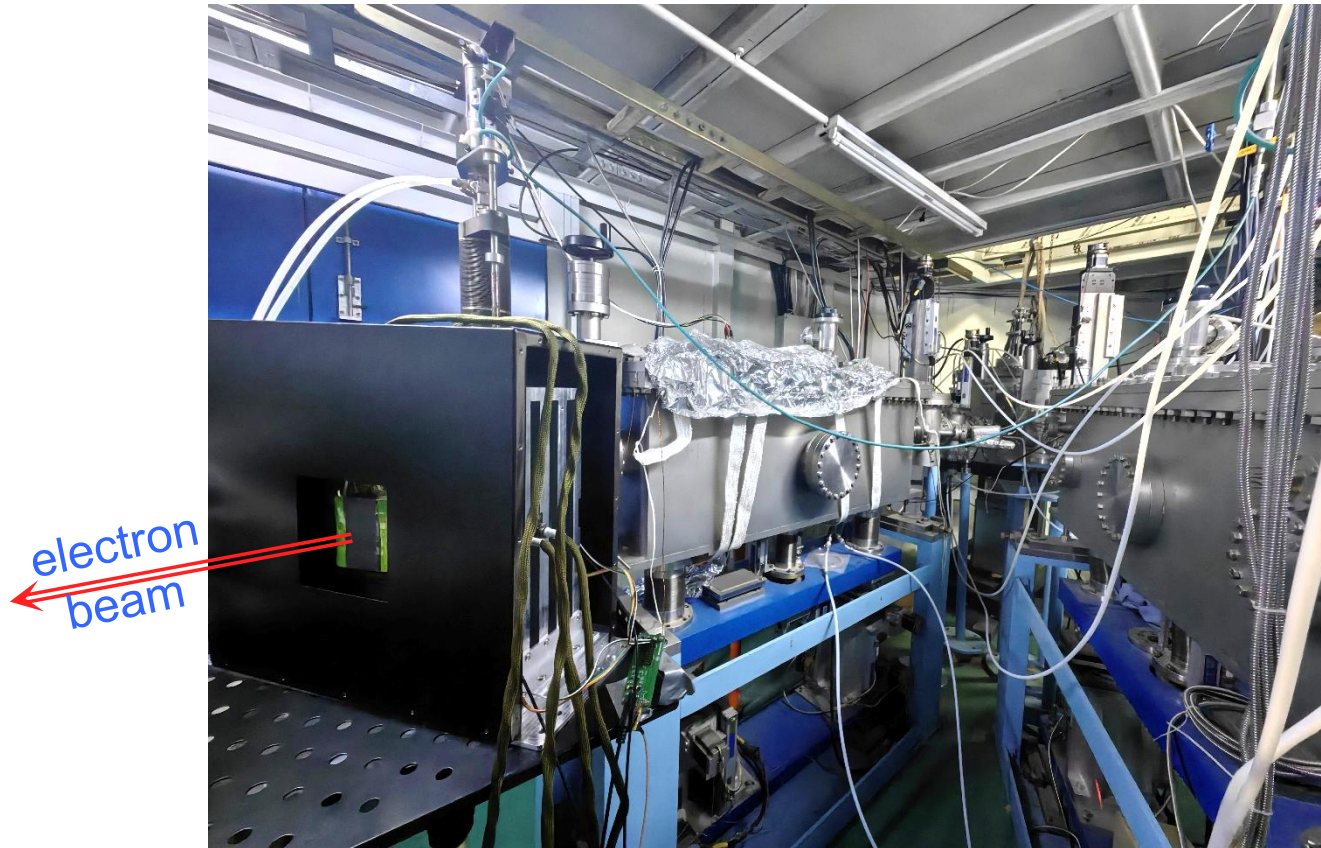
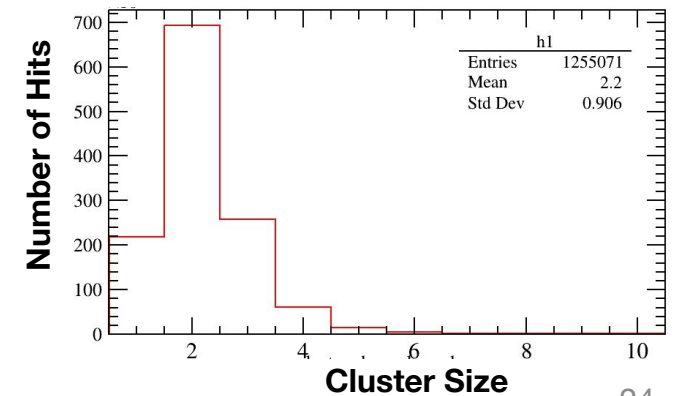
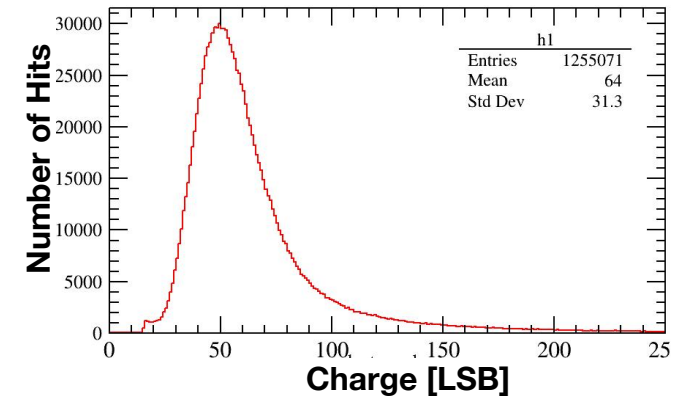
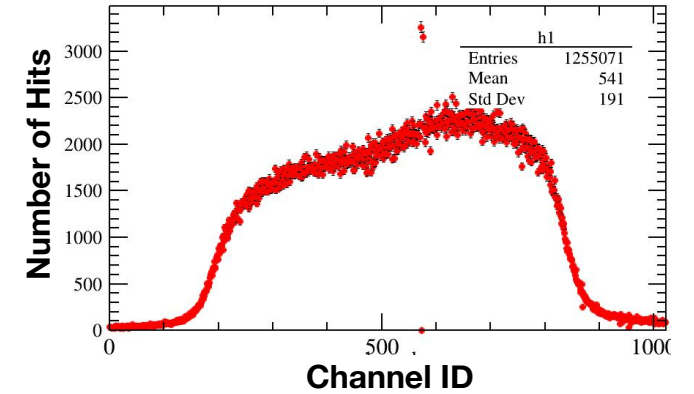


- ❑ high energy beam for characterizing the SSD/Ladder/Layer0 performance
- ❑ same SSD and LEF used for beam monitor



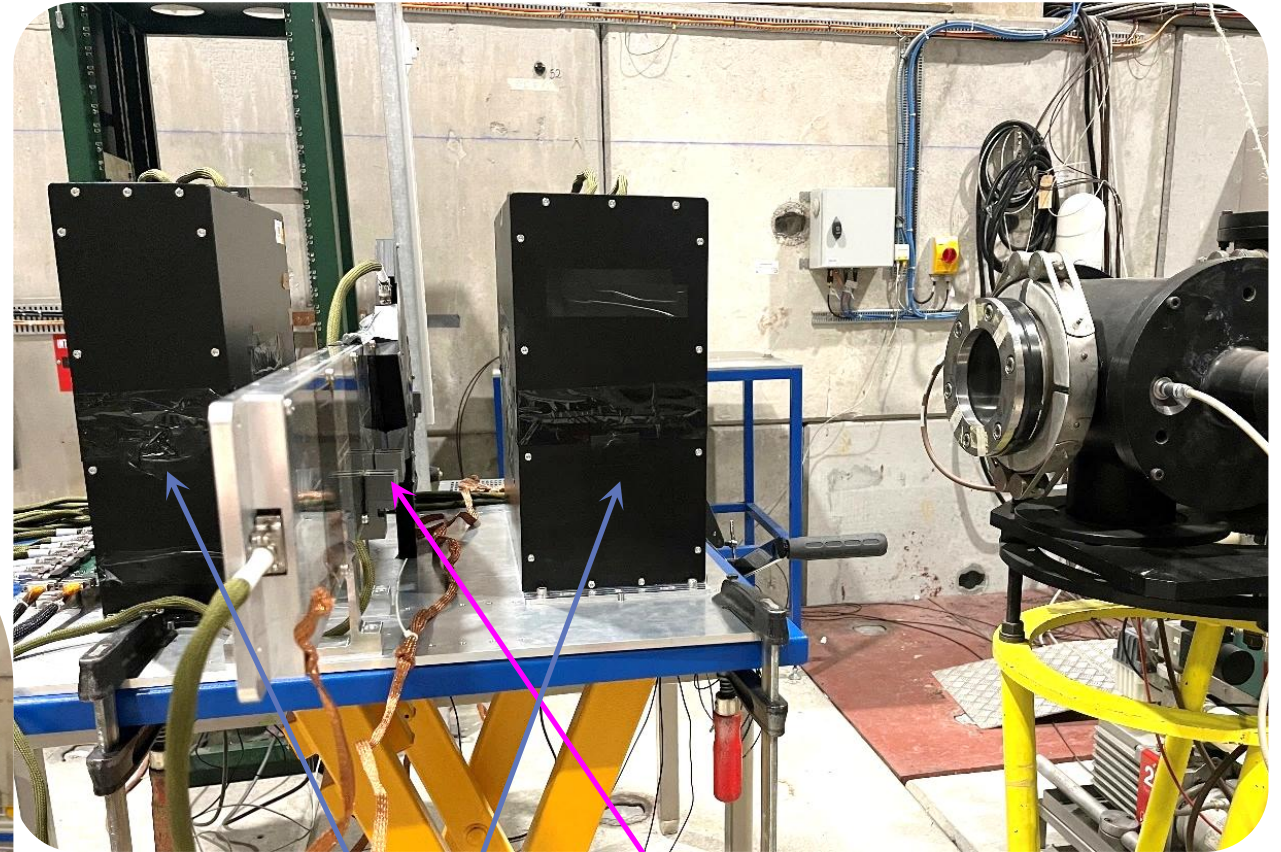
Testbeam at IHEP

- ❑ Beam monitors were tested in an electron beam at IHEP
- ❑ To validate the data acquisition system, and reveal issues



Testbeam at CERN: 2023, 2024

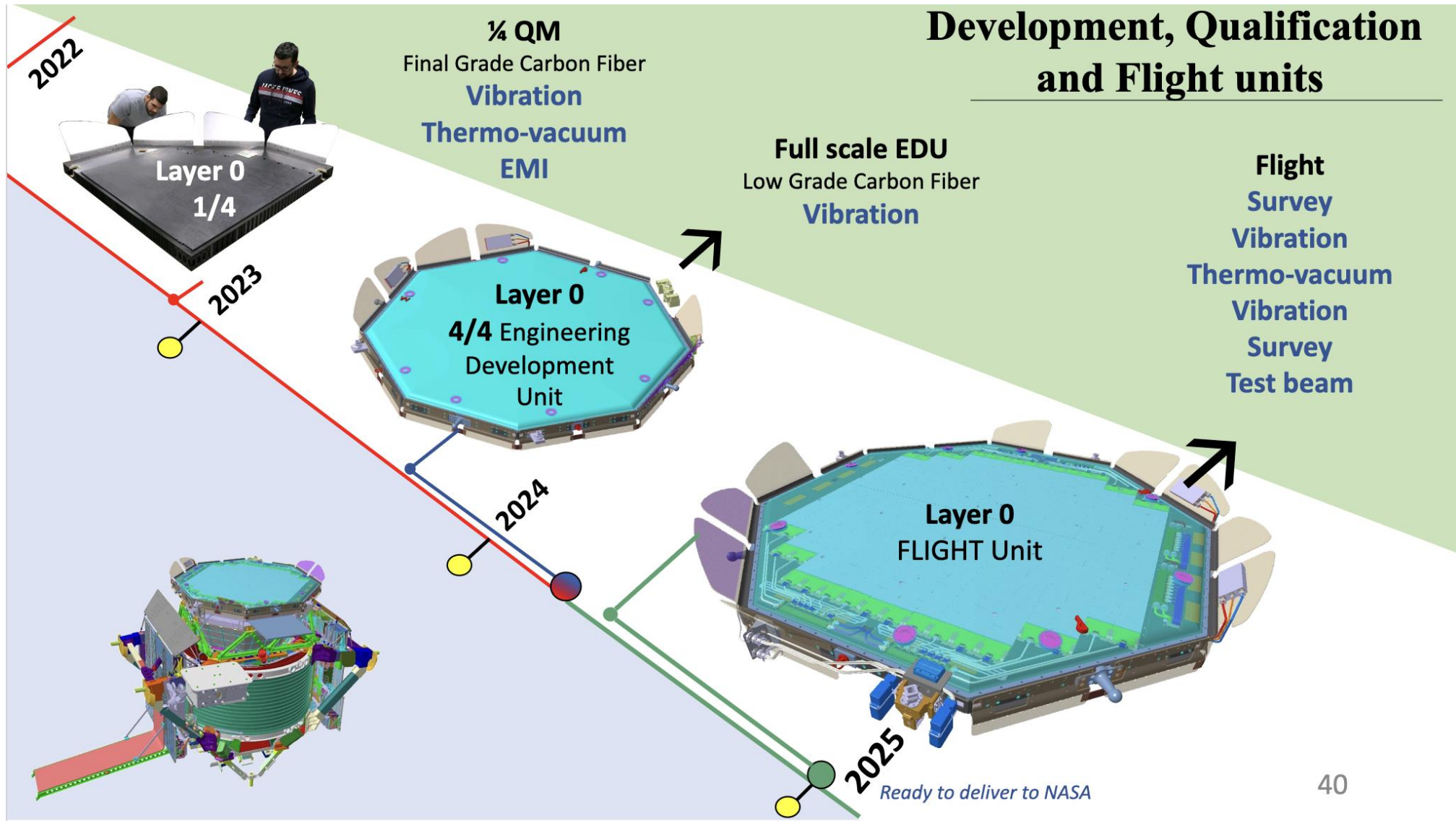
- ❑ Prototype detector ladders were tested in muon, proton, and ion beams at CERN
- ❑ Beam monitors were used as tracking
- ❑ Data are being analyzed
- ❑ test-beam planed in 2024



Prototype ladders under test

Beam monitor boards and trigger scintillators inside the boxes

Roadmap



Backup

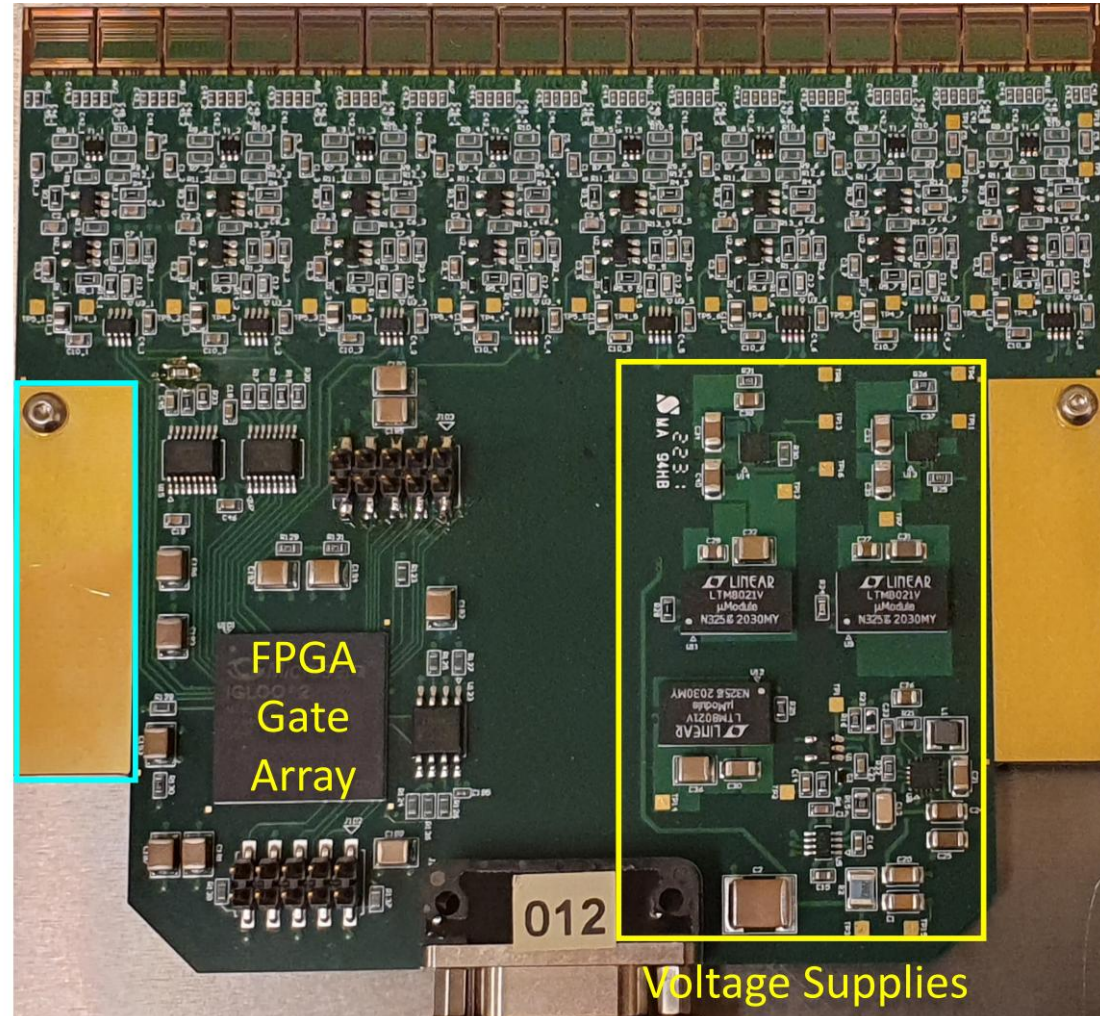
L0 Electronic Front End (LEF)

8 Amplifiers

8 14-Bit, 2.5Msps, Serial Sampling ADCs, 4096mV

Thermal Strips

1. Heat path to radiators.
2. Ground path to chassis

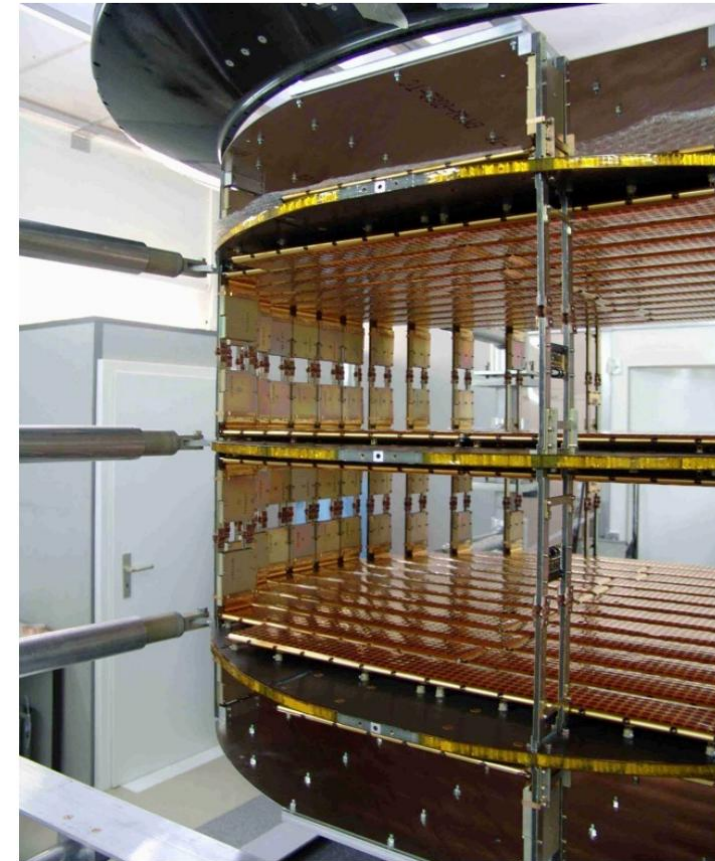
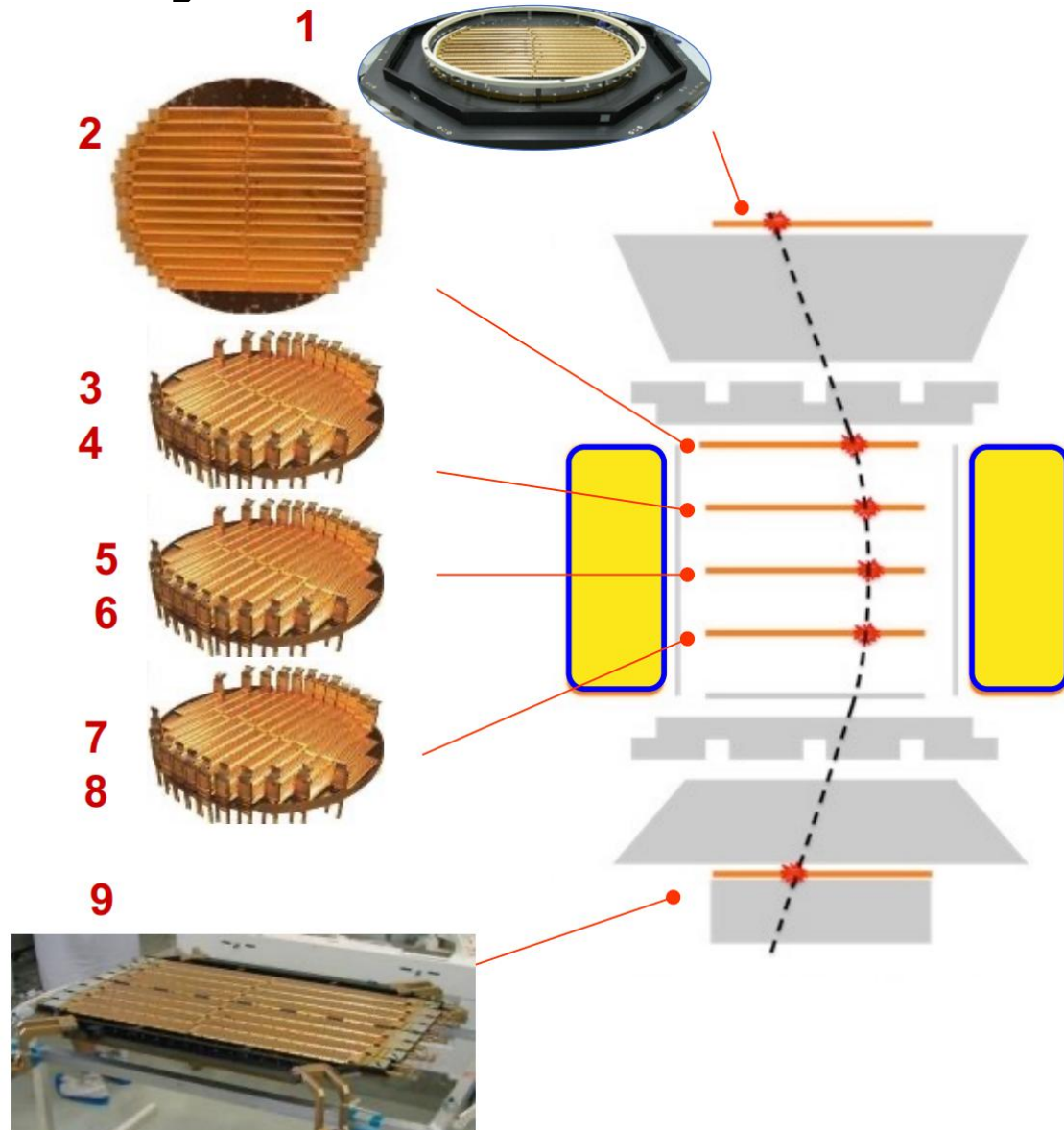


16 IDE1140, 1024 strips

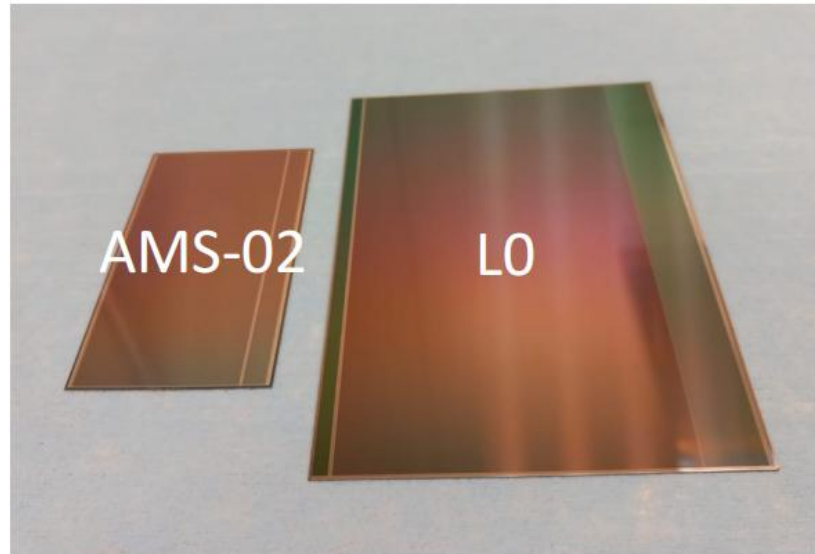
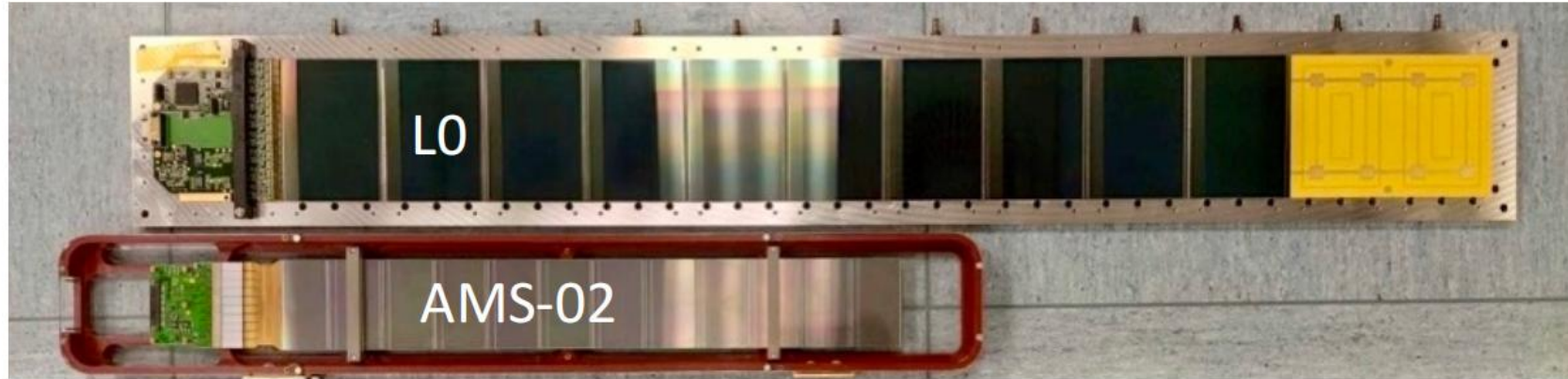
Each IDE1140 (“VA”)

- 64 channels charge amplifier/ shaper.
- Sample and hold.
- 64 channels analog multiplexor.
- 2.6 uA per 1 fC differential current output

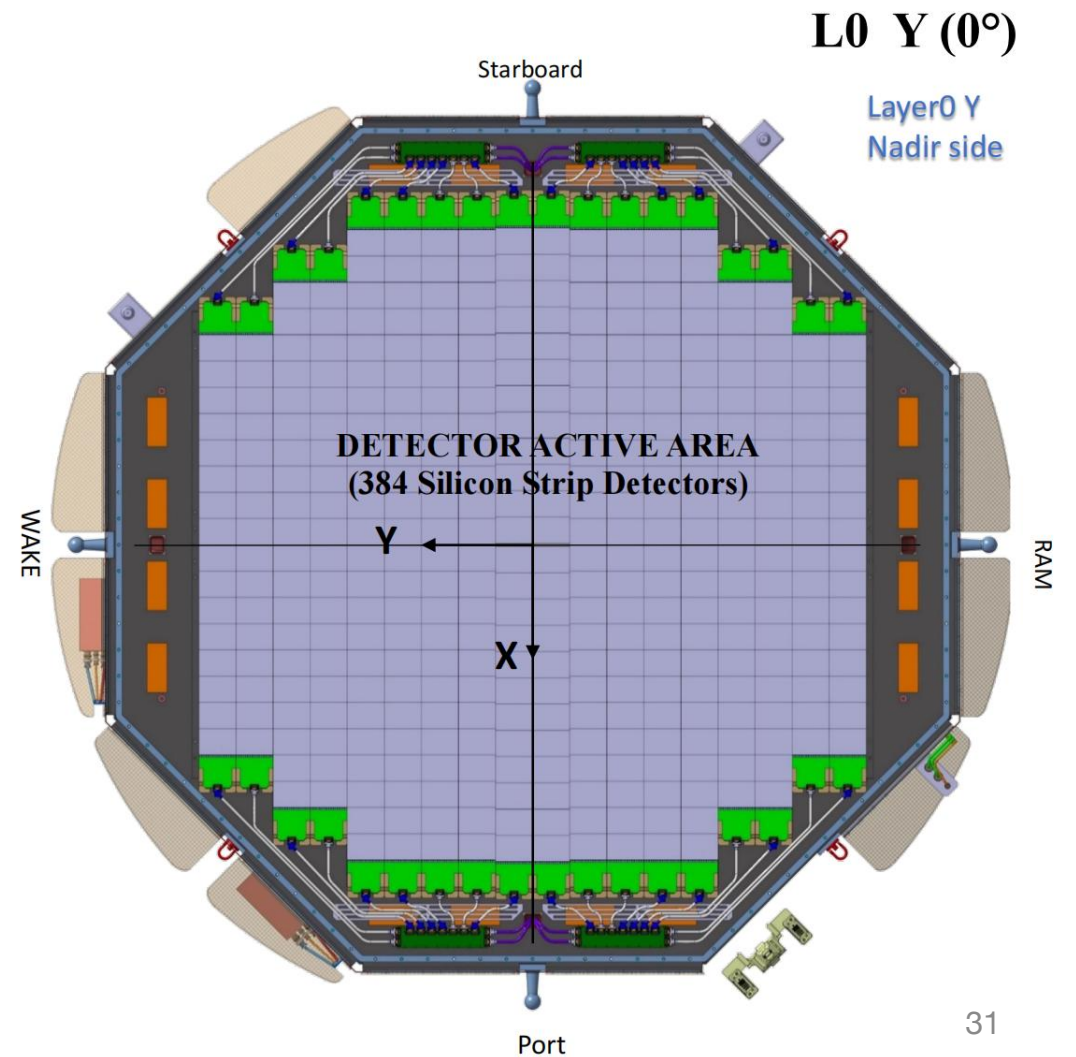
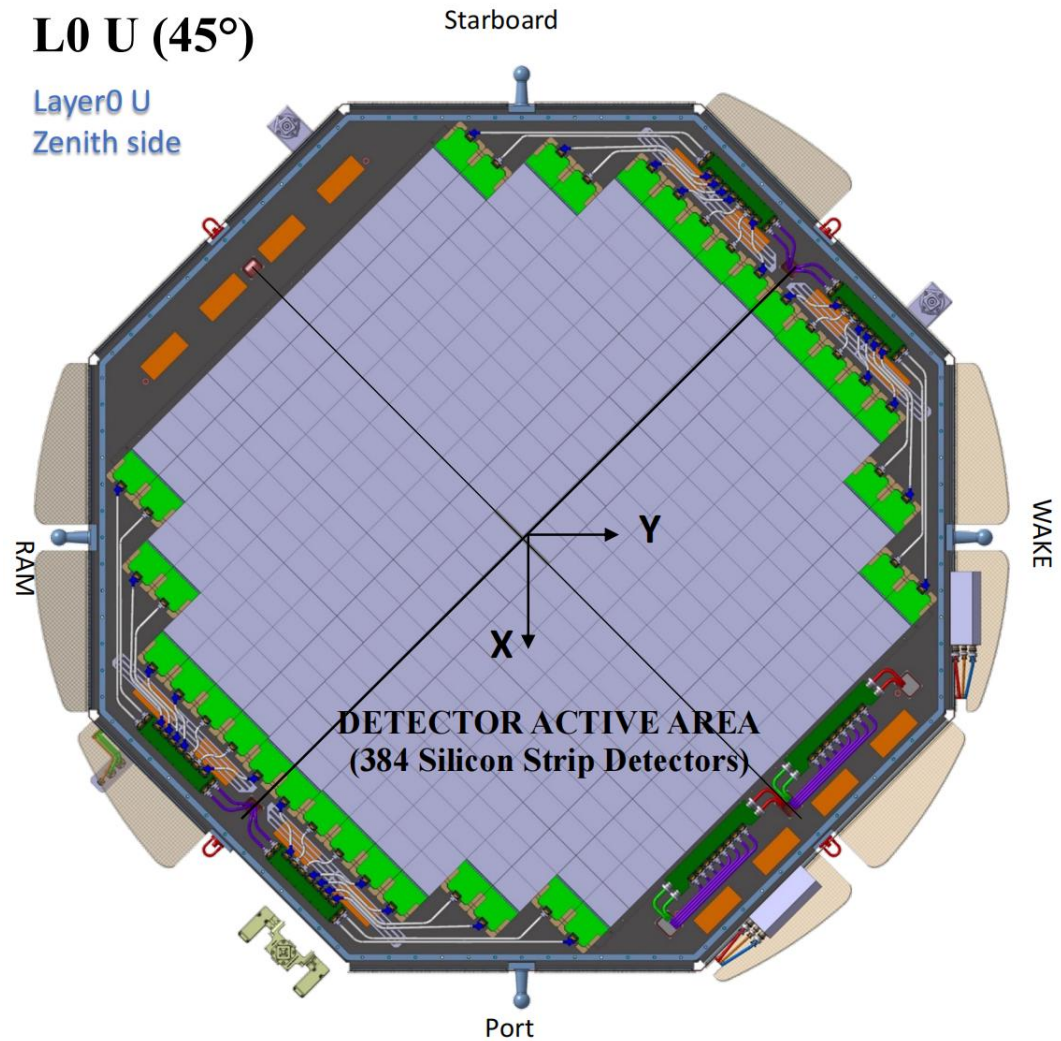
The layout of the AMS-02 Tracker



AMS-02 vs L0 ladders



Plane layout



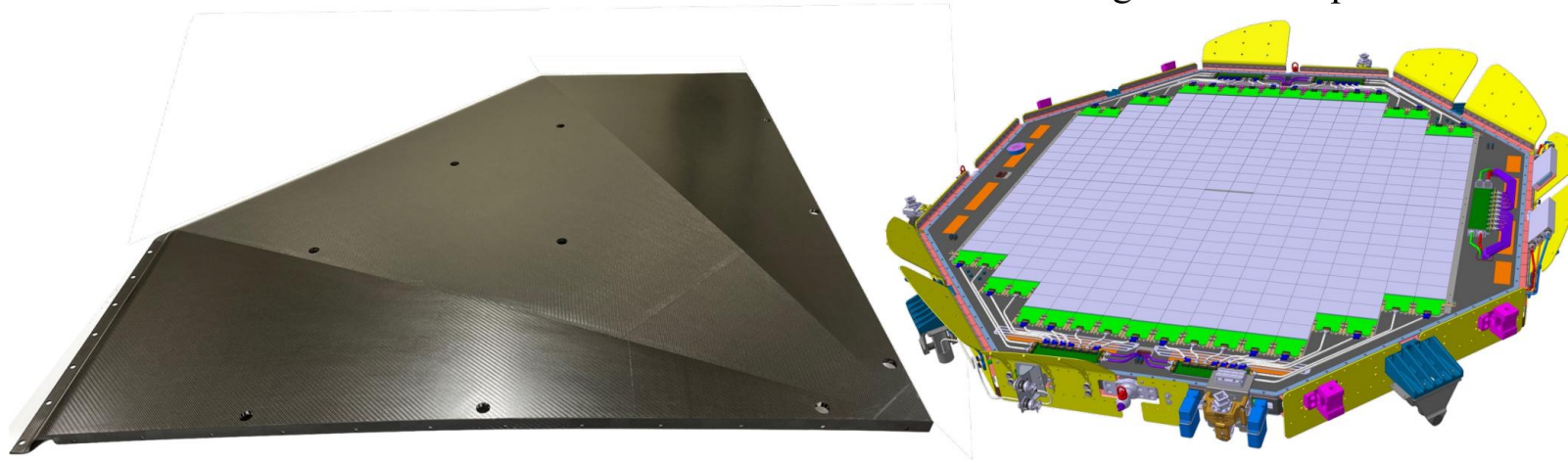
ASAPP 2023

AMS-L0: upgrade status and prospects

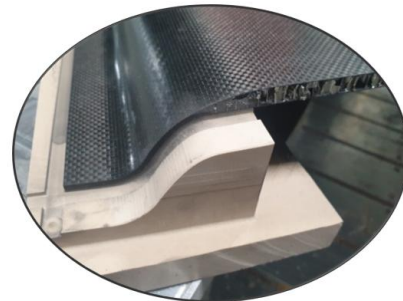
<https://indico.cern.ch/event/1208314/contributions/5283387/>

Mechanics

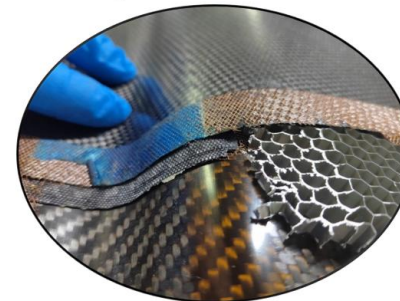
Light carbon cupola



Transition from sandwich to bolted interface



C. Gargiulo



Aluminum honeycomb
Carbon skin
Copper net for EMI

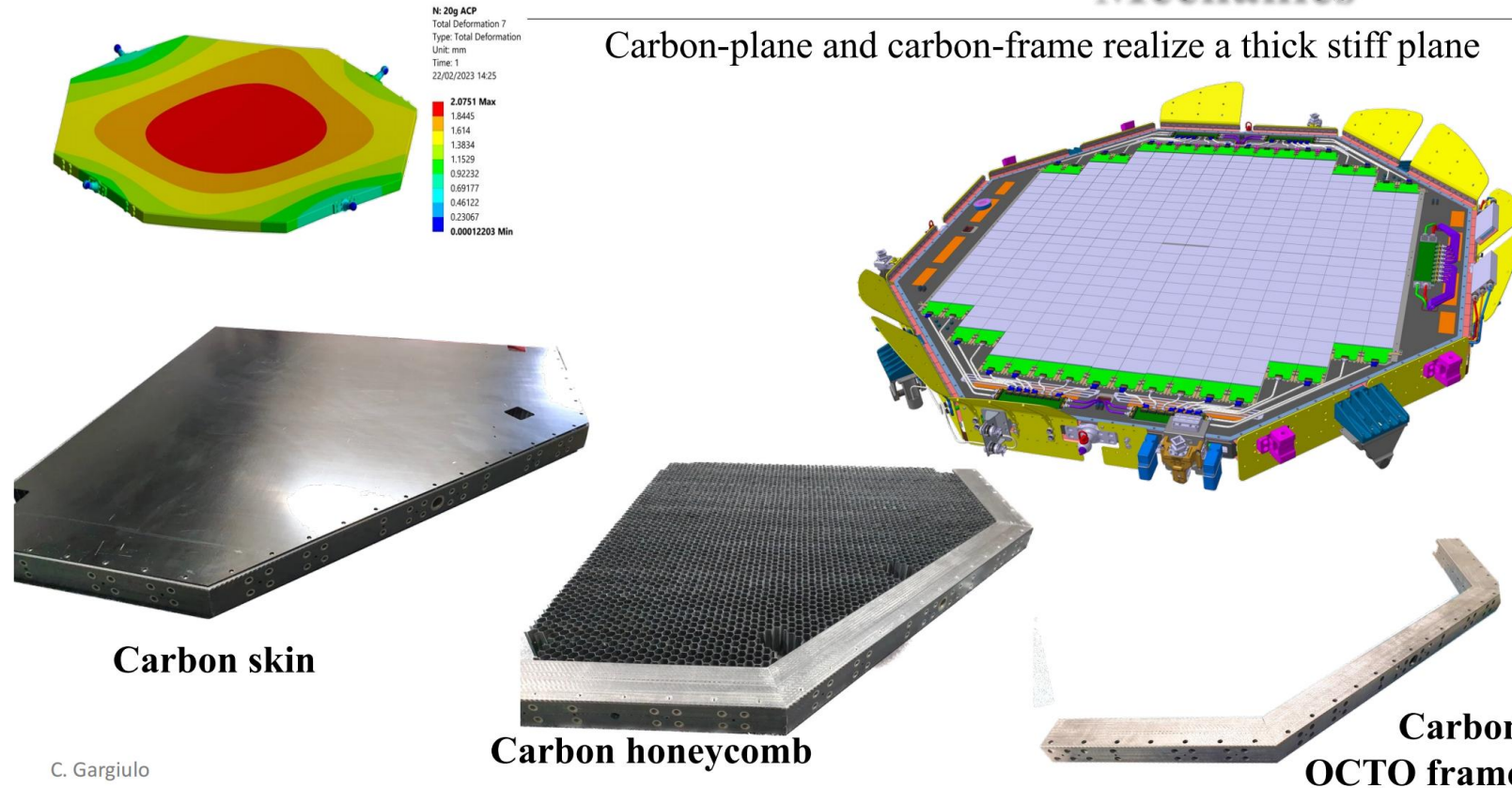
ASAPP 2023

AMS-L0: upgrade status and prospects

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Mechanics

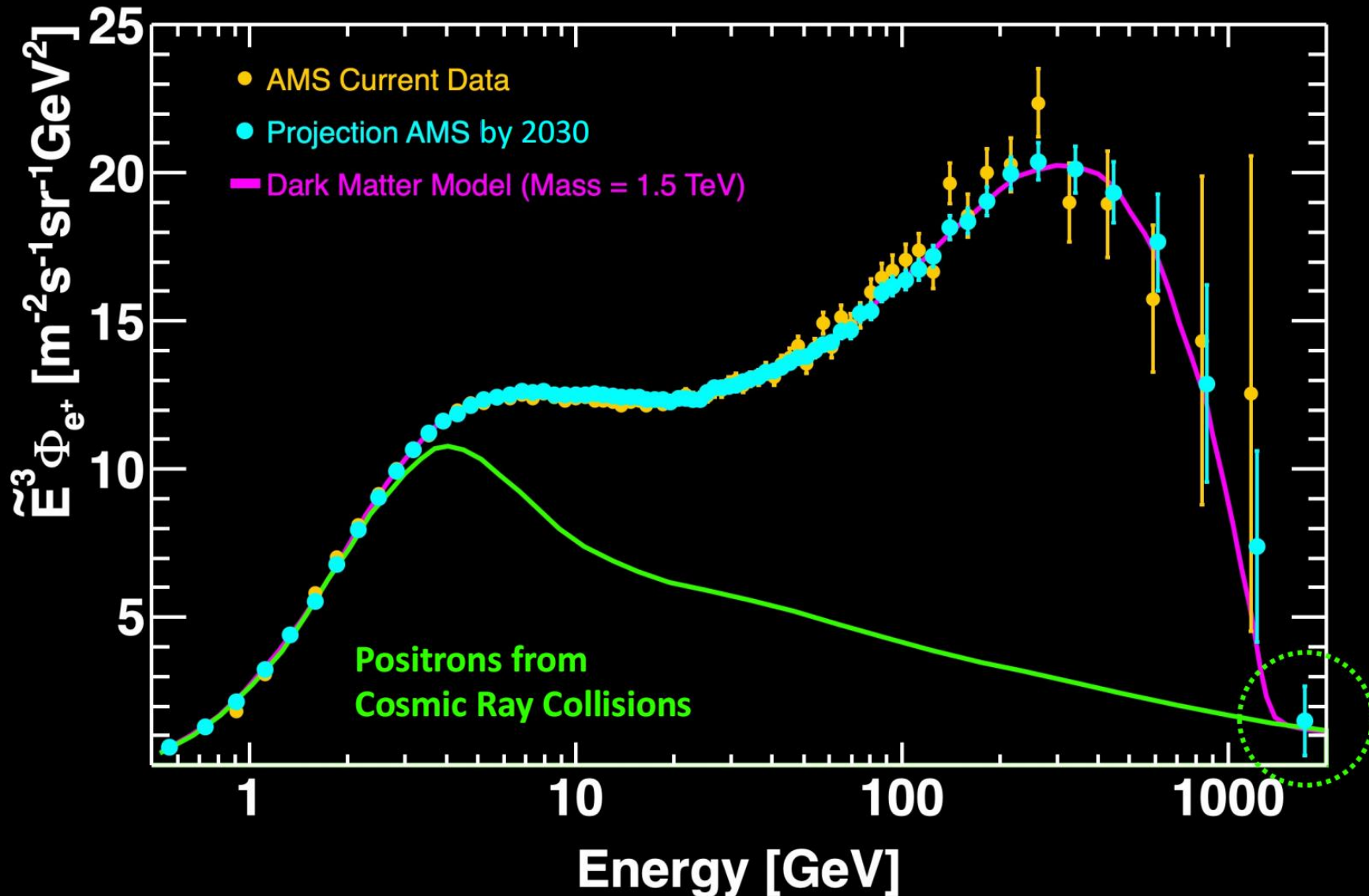
Carbon-plane and carbon-frame realize a thick stiff plane



C. Gargiulo

Determination of the Origin of Cosmic Positrons by 2030

AMS will ensure that the measured high energy positron spectrum indeed drops off quickly and, at the highest energies, the positrons only come from cosmic ray collisions as predicted by dark matter models



By 2030, AMS will extend the energy range of the positron flux measurement from 1.4 to 2 TeV and reduce the error by a factor of two compared to current data

