

ANPhA 2025

Recent Highlights from the STAR Experiment at RHIC

(nuclear structure across energy scales)

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Fudan University

Nov 27-29, 2025, Institute of Physics, Academia Sinica

Outline

I. Nuclear structure across energy scales

II. Nuclear deformation in ground-state ^{238}U and ^{96}Zr nuclei

STAR, Nature 635, 67-72 (2024)

STAR, Rep. Prog. Phys. 88, 108601 (2025)

STAR, Preliminary results

III. Nucleonic cluster pattern in light ^{16}O nucleus

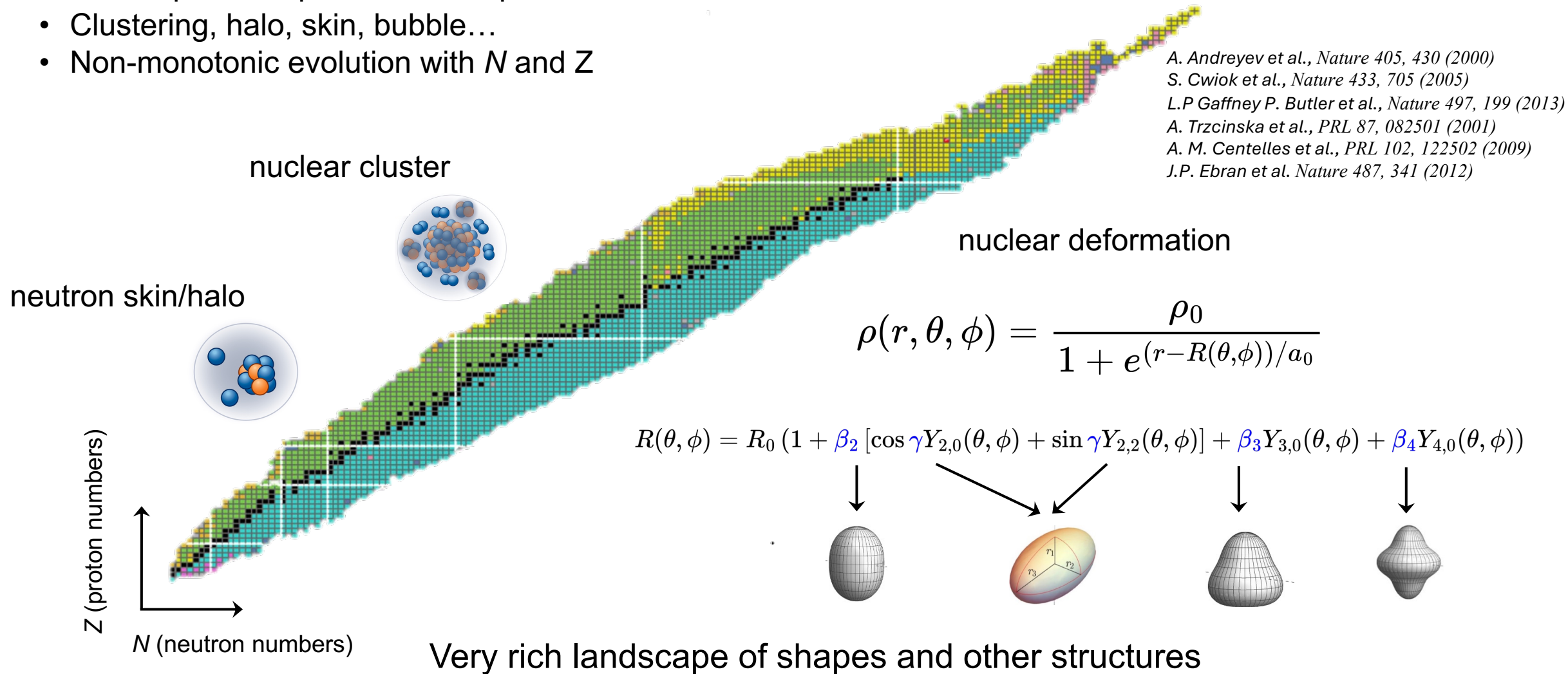
STAR, arXiv: 2510.19645

IV. Conclusions and outlooks

Shape of atomic nuclei

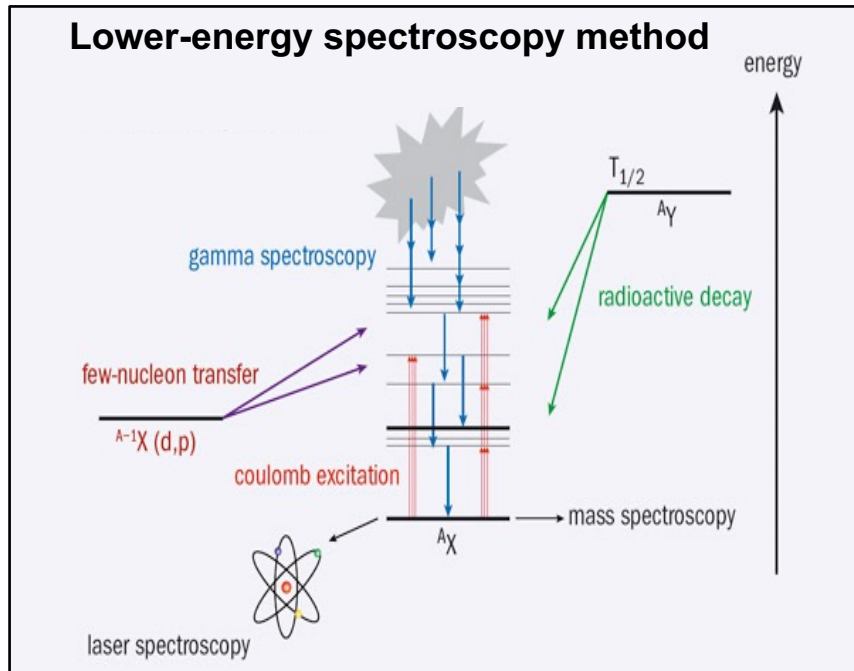
Emergent phenomena of the many-body quantum system, govern by short-range strong nuclear force

- Quadrupole/octupole/hexadecapole deformations
- Clustering, halo, skin, bubble...
- Non-monotonic evolution with N and Z



A. Andreyev et al., *Nature* 405, 430 (2000)
 S. Cwiok et al., *Nature* 433, 705 (2005)
 L.P Gaffney P. Butler et al., *Nature* 497, 199 (2013)
 A. Trzcinska et al., *PRL* 87, 082501 (2001)
 A. M. Centelles et al., *PRL* 102, 122502 (2009)
 J.P. Ebran et al. *Nature* 487, 341 (2012)

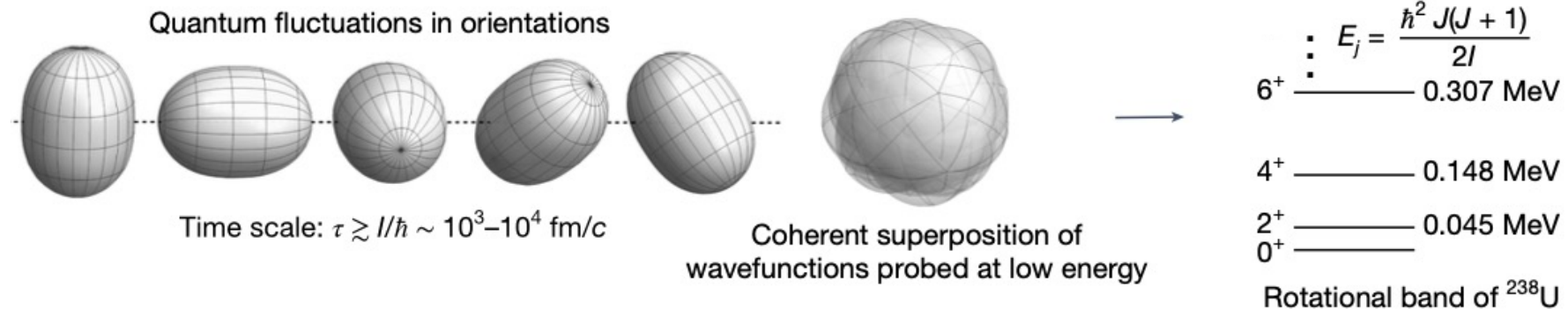
Nuclear shape at low energy: long exposure



Traditional imaging method taken before destruction

- Low energy spectroscopic methods probe a superposition of these fluctuations.
- Instantaneous shapes not directly seen, but inferred from model comparison.

Each DOF has zero-point fluctuations within certain timescales.



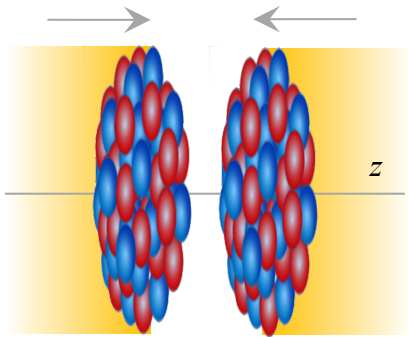
Imaging by smashing method

Take a snapshot

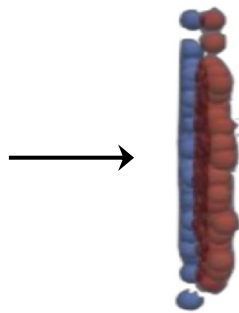
Evolution

Measurement

Nuclei collisions



$$T_{\mu\nu}(\tau = 0)$$

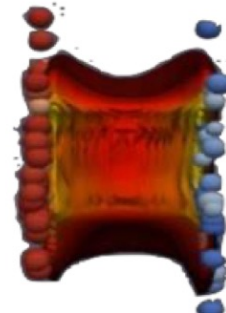


$$\tau \sim 2R_0/\Gamma \sim 0.1\text{fm}/c$$

exposure

Pressure-driven expansion of
Quark-gluon plasma (QGP)

$$\partial_\mu T^{\mu\nu} = 0, \text{ EOS, viscosity...}$$

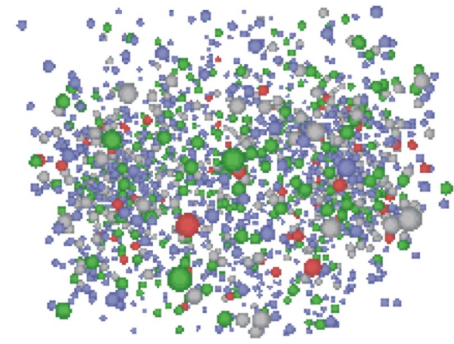


$$\tau \sim 10\text{fm}/c$$

expansion

free-streaming

$$T_{\mu\nu}(\tau = \infty)$$



$$\tau \sim 10^{15}\text{fm}/c$$

detection

$$\rho(r, \theta, \phi) = \frac{\rho_0}{1 + e^{(r-R(\theta, \phi))/a_0}}$$

$\beta_2 \rightarrow$ quadrupole deformation

$\beta_3 \rightarrow$ octupole deformation

$\gamma \rightarrow$ triaxiality

$a_0 \rightarrow$ surface diffuseness

$R_0 \rightarrow$ nuclear size

ab initio theory/shell model/DFT

J. Jia et al., Nucl. Sci. Tech. 35, 220 (2024)

Imaging by smashing method

Take a snapshot

Evolution

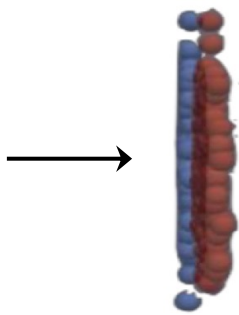
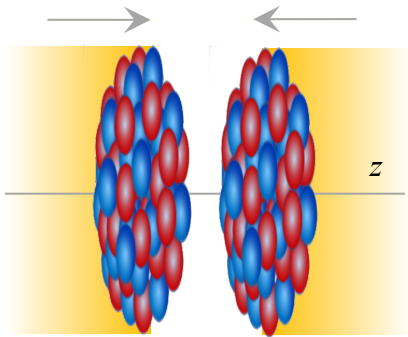
Measurement

Nuclei collisions

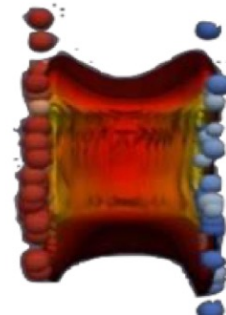
$$T_{\mu\nu}(\tau = 0)$$

$$\partial_\mu T^{\mu\nu} = 0, \text{ EOS, viscosity...}$$

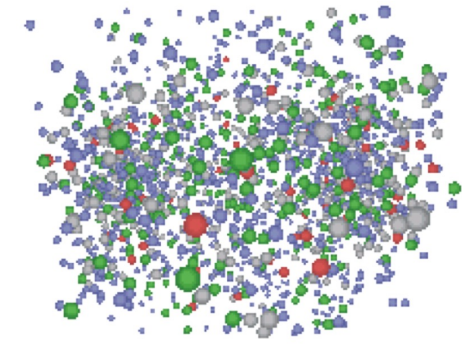
$$T_{\mu\nu}(\tau = \infty)$$



Pressure-driven expansion of
Quark-gluon plasma (QGP)



free-streaming



$\tau \sim 2R_0/\Gamma \sim 0.1\text{fm}/c$
exposure

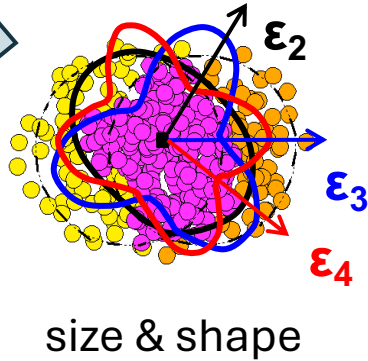
$\tau \sim 10\text{fm}/c$
expansion

$\tau \sim 10^{15}\text{fm}/c$
detection

$$\rho(r, \theta, \phi) = \frac{\rho_0}{1 + e^{(r-R(\theta, \phi))/a_0}}$$

- $\beta_2 \rightarrow$ quadrupole deformation
- $\beta_3 \rightarrow$ octupole deformation
- $\gamma \rightarrow$ triaxiality
- $a_0 \rightarrow$ surface diffuseness
- $R_0 \rightarrow$ nuclear size

ab initio theory/shell model/DFT



size & shape

J. Jia et al., Nucl. Sci. Tech. 35, 220 (2024)

$$R_\perp^2 \propto \langle r_\perp^2 \rangle \quad \mathcal{E}_n \propto \langle r_\perp^n e^{in\phi} \rangle$$

Diagram showing the relationship between R_0 , a_0 , and β_n with arrows indicating dependencies.

$$\text{Observables } \frac{d^2 N}{d\phi dp_T} = N(p_T) \left(\sum_n V_n e^{-in\phi} \right)$$

Event-by-event linear responses:

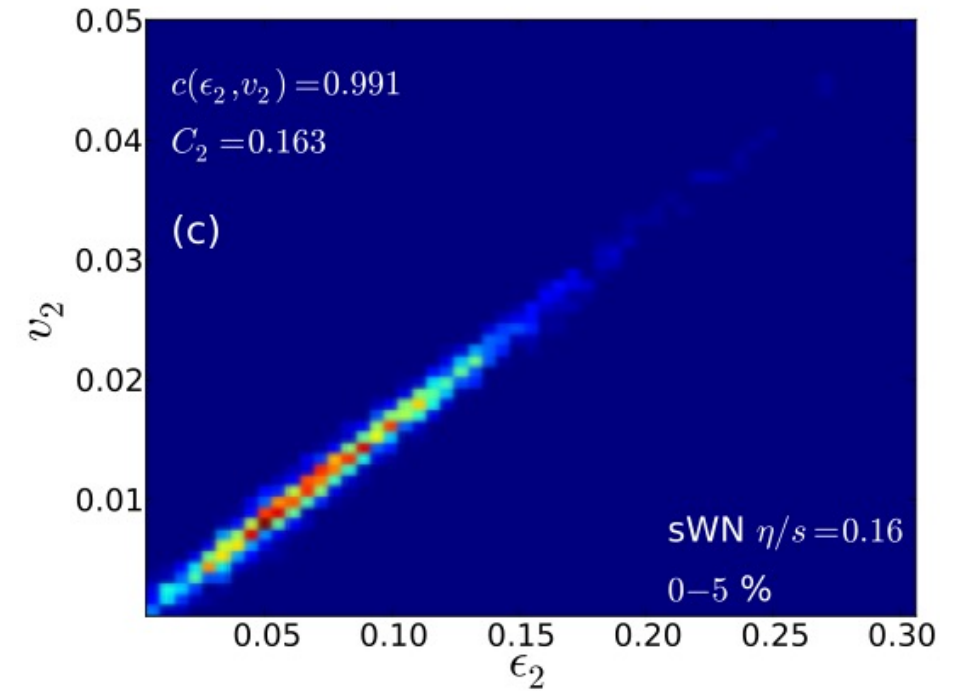
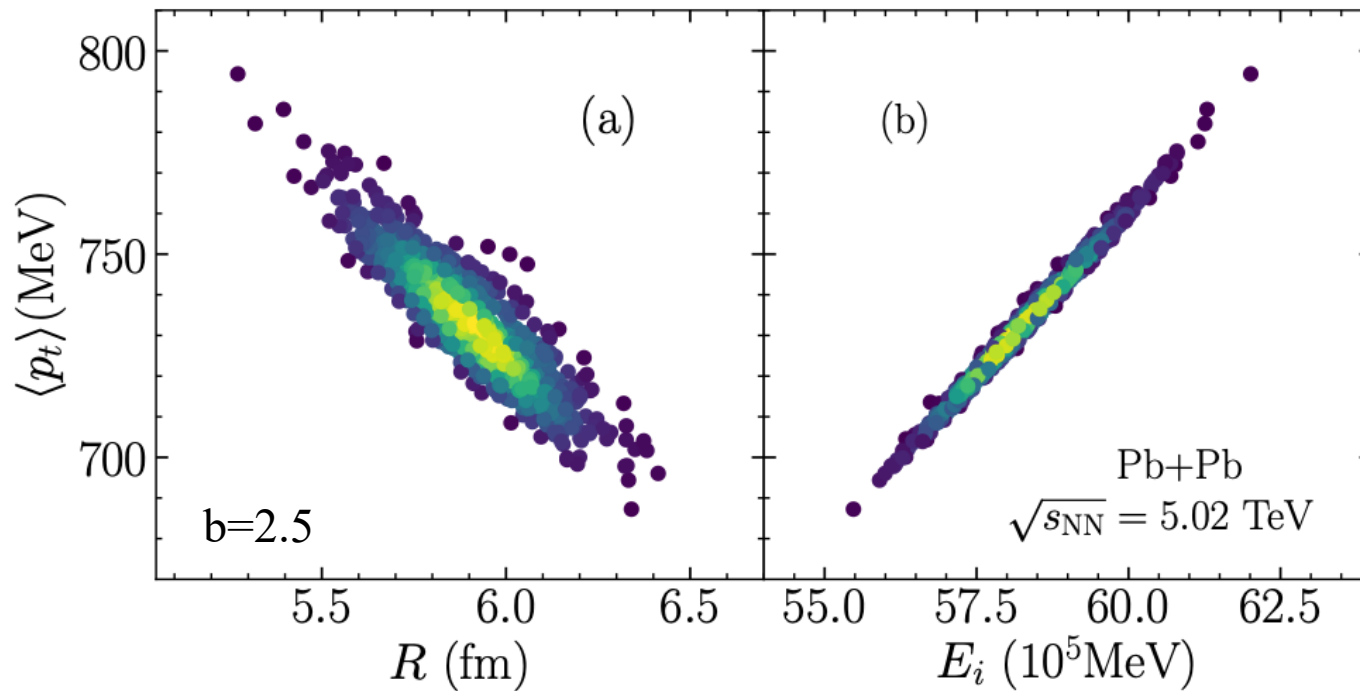
$$\frac{\delta[p_T]}{[p_T]} \propto -\frac{\delta R_\perp}{R_\perp} \quad V_n \propto \mathcal{E}_n$$

Key: 1) fast snapshot, 2) linear response, 3) large multiplicity for many-body correlation

Linear response in ultra-central collisions

$$\frac{\delta[p_T]}{[p_T]} \propto -\frac{\delta R_{\perp}}{R_{\perp}}$$

$$V_n \propto \mathcal{E}_n$$

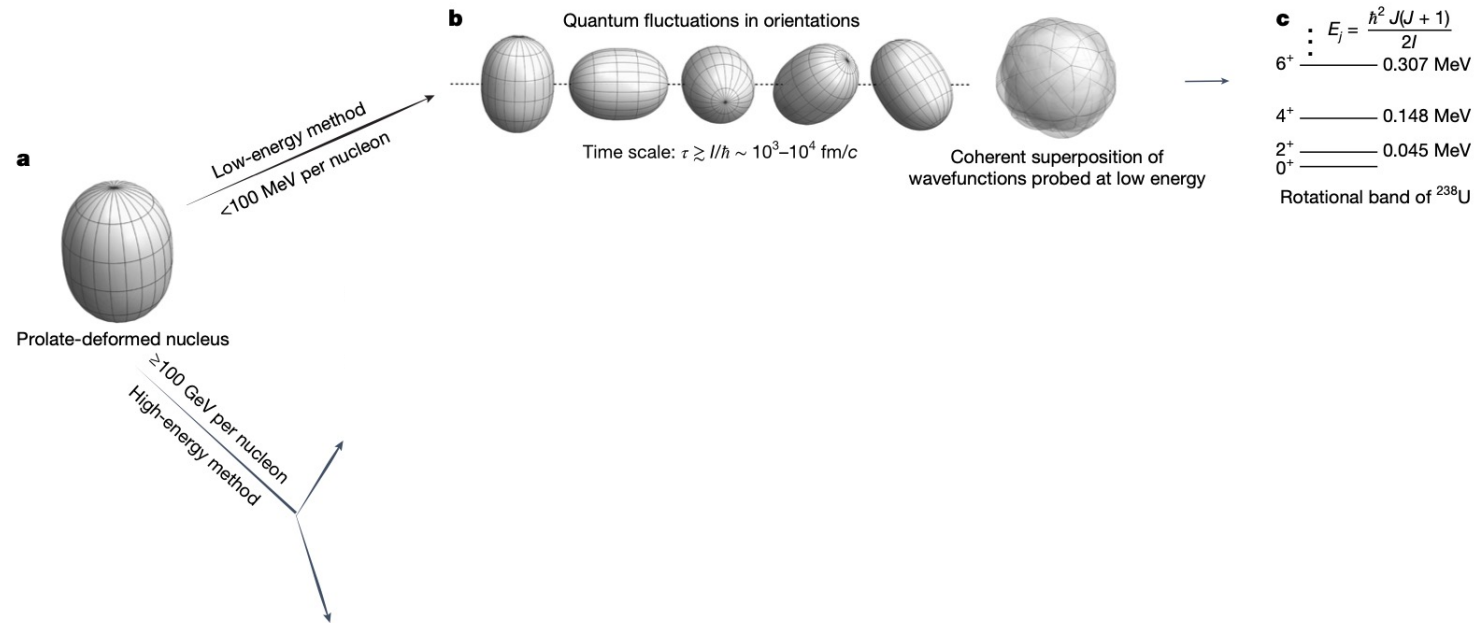


G. Giacalone, F. Gardim, J. Hostler, J. Ollitrault, et al, PRC 103, 024909 (2021)

H. Niemi, G. Denicol, H. Holopainen, P. Huovinen, PRC 87, 054901 (2013)

Imaging nuclear shape in high-energy snapshot

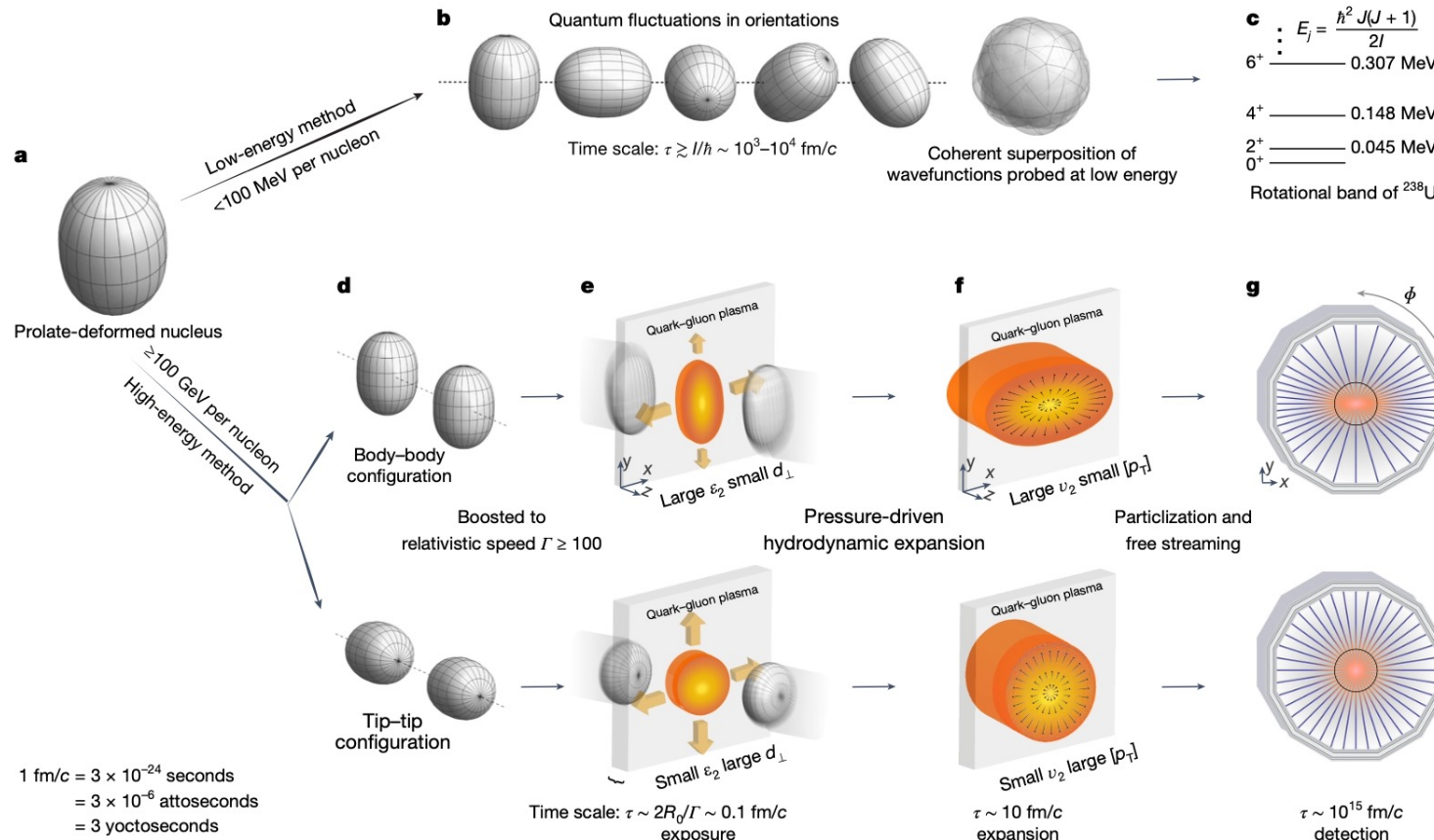
- Nuclear shape in intrinsic (body-fixed) frame not directly visible in the lab frame
--Mainly inferred from non-invasive spectroscopy methods.



$1 \text{ fm/c} = 3 \times 10^{-24} \text{ seconds}$
 $= 3 \times 10^{-6} \text{ attoseconds}$
 $= 3 \text{ yoctoseconds}$

Imaging nuclear shape in high-energy snapshot

- Nuclear shape in intrinsic (body-fixed) frame not directly visible in the lab frame
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STAR, Nature 635, 67-72 (2024)

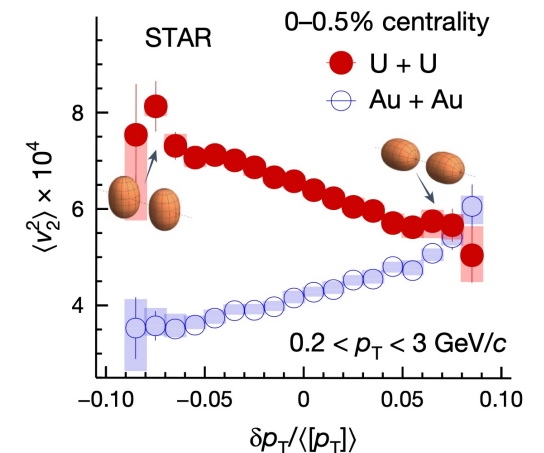
<https://www.nature.com/articles/s41586-024-08097-2>

Body-body: large-eccentricity large-size

$v_2 \nearrow$ $p_T \searrow$

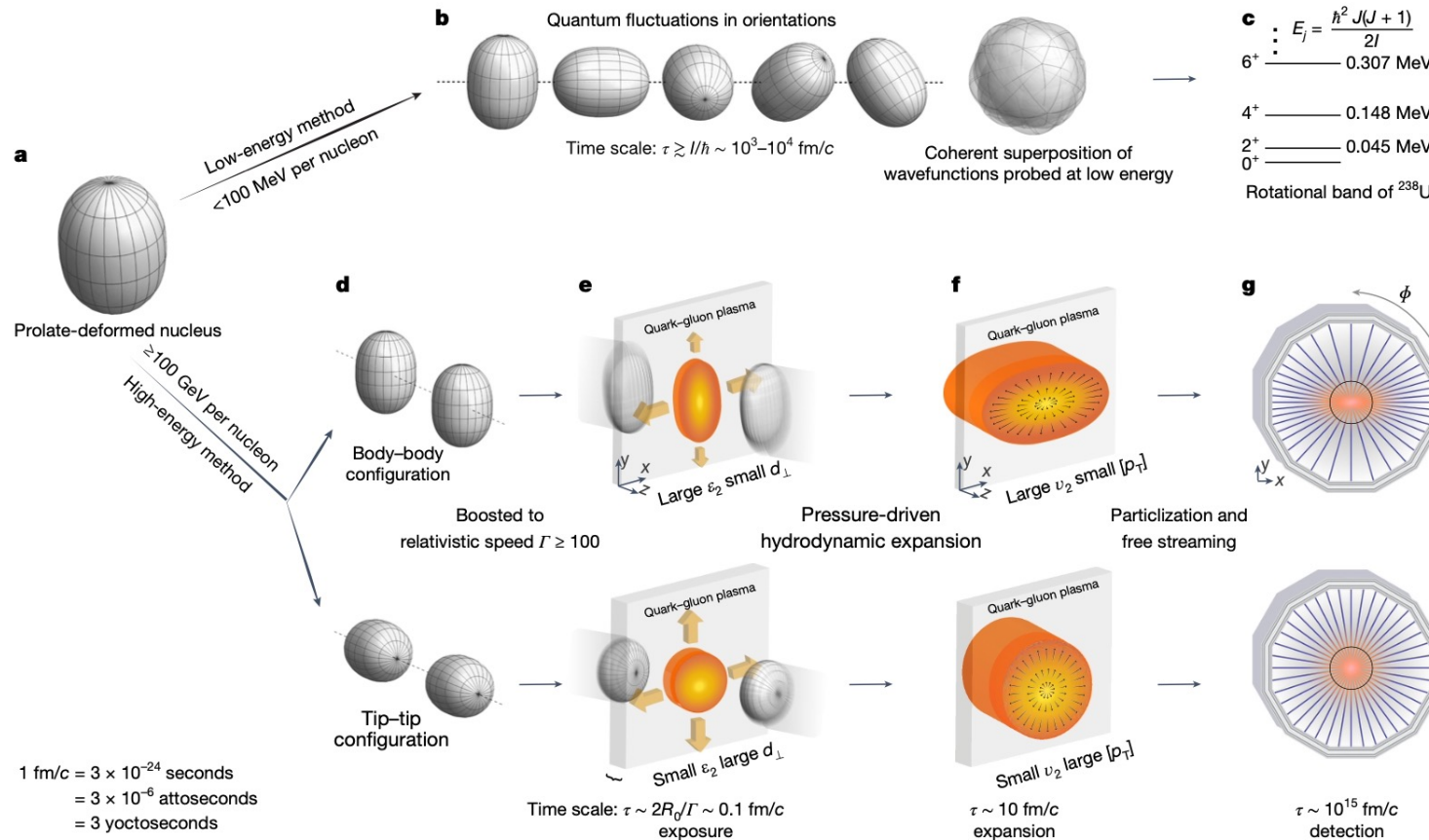
Tip-tip: small-eccentricity small-size

$v_2 \searrow$ $p_T \nearrow$



Imaging nuclear shape in high-energy snapshot

- Nuclear shape in intrinsic (body-fixed) frame not directly visible in the lab frame
--Mainly inferred from non-invasive spectroscopy methods.



STAR, *Nature* 635, 67-72 (2024)

<https://www.nature.com/articles/s41586-024-08097-2>

across energy scales

Body-body: large-eccentricity large-size

$v_2 \nearrow$ $p_T \searrow$

Tip-tip : small-eccentricity small-size

$v_2 \searrow$ $p_T \nearrow$

$$\langle v_2^2 \rangle = a_1 + b_1 \beta_2^2,$$

$$\langle (\delta p_T)^2 \rangle = a_2 + b_2 \beta_2^2,$$

$$\langle v_2^2 \delta p_T \rangle = a_3 - b_3 \beta_2^3 \cos(3\gamma).$$

G. Giacalone, J. Jia, C. Zhang, *PRL* 127, 242301(2021)

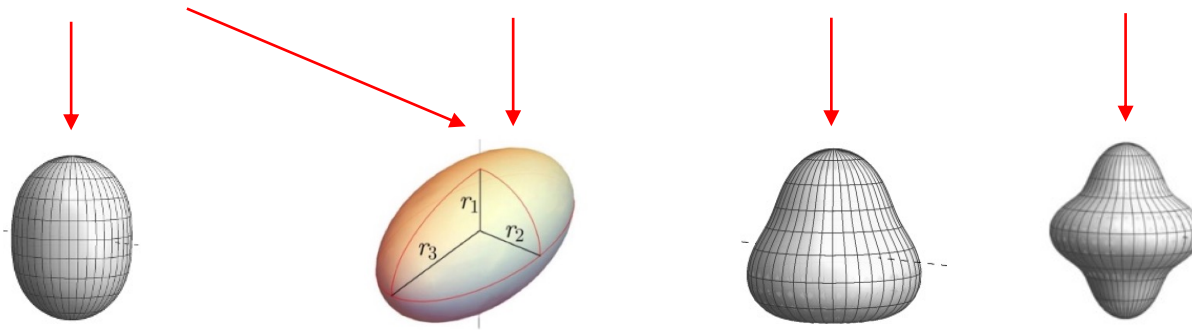
Shape-frozen like a snapshot during nuclear crossing ($10^{-25}\text{s} \ll$ rotational time scale 10^{-21}s)

probe entire mass distribution in the intrinsic frame via multi-point correlations

II. Nuclear deformation in heavy ^{238}U and ^{96}Zr nucleus

$$\rho(r, \theta, \phi) = \frac{\rho_0}{1 + e^{(r-R(\theta, \phi))/a_0}}$$

$$R(\theta, \phi) = R_0 (1 + \beta_2 [\cos \gamma Y_{2,0}(\theta, \phi) + \sin \gamma Y_{2,2}(\theta, \phi)] + \beta_3 Y_{3,0}(\theta, \phi) + \beta_4 Y_{4,0}(\theta, \phi))$$



W. Ryssens, G. Giacalone, B. Schenke, C. Shen, *PRL*130, 212302(2023)

DFT calculations predict a slightly small WS deformation

$$\beta_{2\text{U}} \approx 0.28 \rightarrow \beta_{2\text{U,WS}} \approx 0.25$$

corresponding to a larger volume deformation in presence of $\beta_{4\text{U}} \sim 0.1$ $\beta_{2,\text{body}} = \frac{4\pi}{3R_0^2 A} \int d^3r \rho(\mathbf{r}) r^2 Y_{20}$

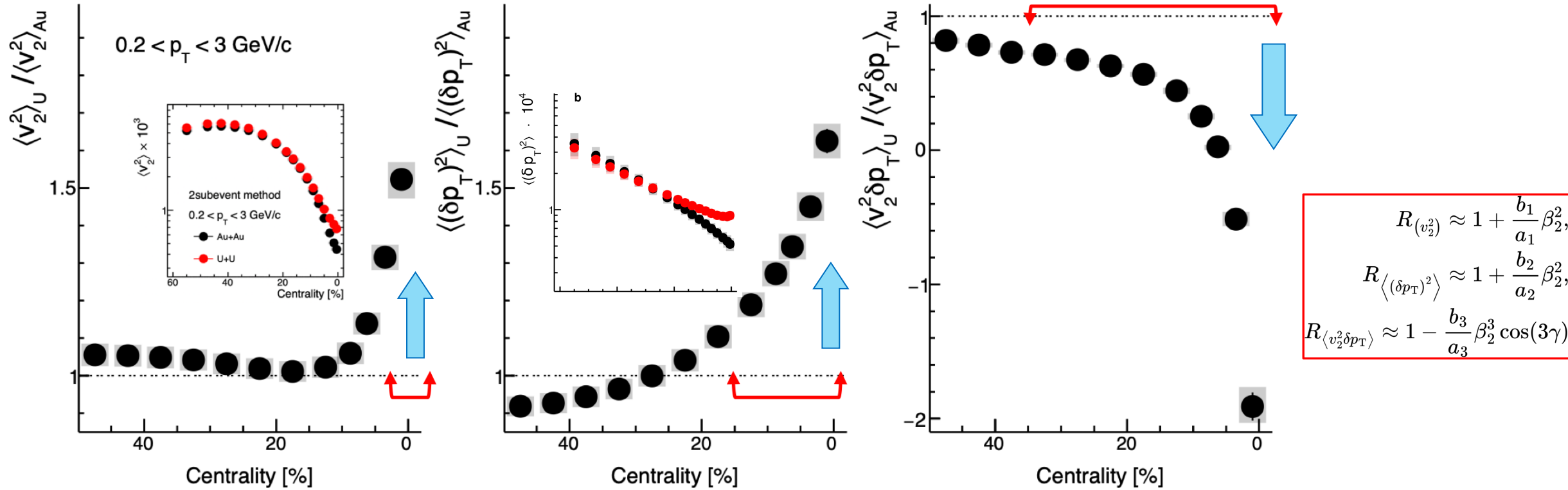
Low-energy estimate with rigid rotor assumption from B(E2) data $\beta_{2,\text{LD}} = \frac{4\pi}{5R_0^2 Z} \sqrt{\frac{B(\text{E2})}{e^2}}$

$$\beta_{2\text{U,LD}} = 0.287 \pm 0.007 \quad \gamma_{\text{U,LD}} = 6^\circ - 8^\circ$$

$$\beta_{3\text{U}} \sim \beta_{4\text{U}} \sim 0.1$$

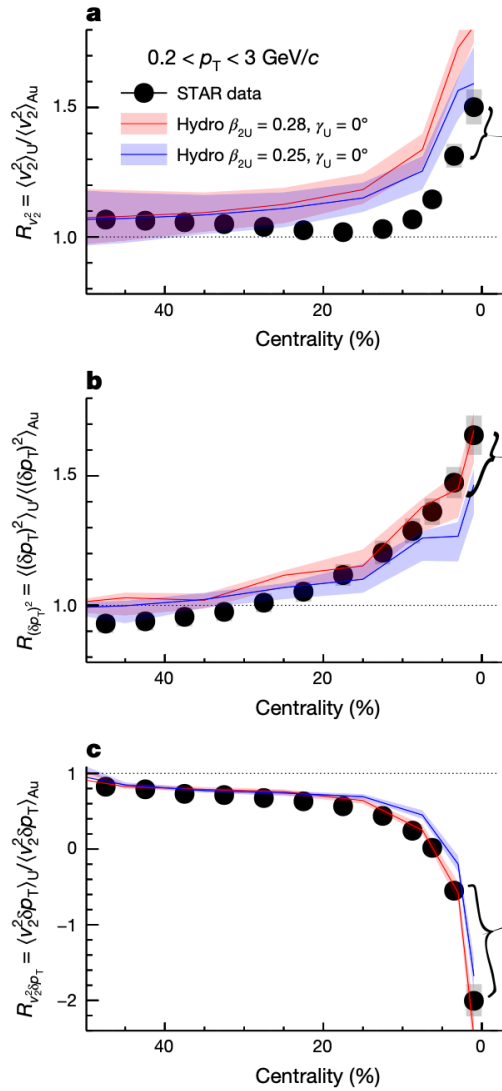
B. Pritychenko et al., *J.ADT*107, 1(2016)
C. Y. Wu et al., *PRC*54, 2356(1996)

Ratio of observables

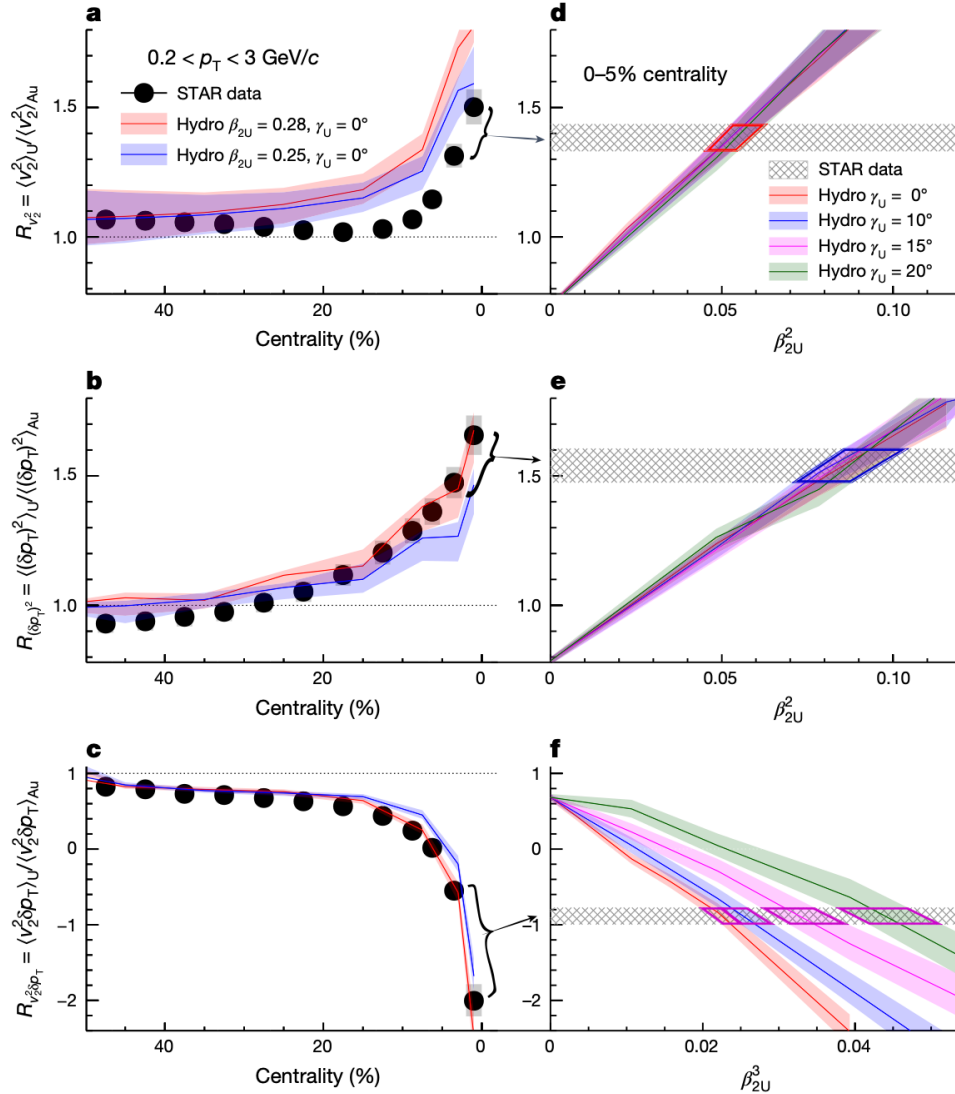


- Elliptic flow and size fluctuation are enhanced by the nuclear deformation effect.
- Ratios cancel final state effects and isolate the effects of initial state/nuclear structures.
→ U deformation dominates the ultra-central collisions (UCC)

Constraining the ground-state ^{238}U : $\beta_{2,U}$ and γ_U



Constraining the ground-state ^{238}U : $\beta_{2,U}$ and γ_U



Sufficient precision is achieved from ratios in ultra-central collisions

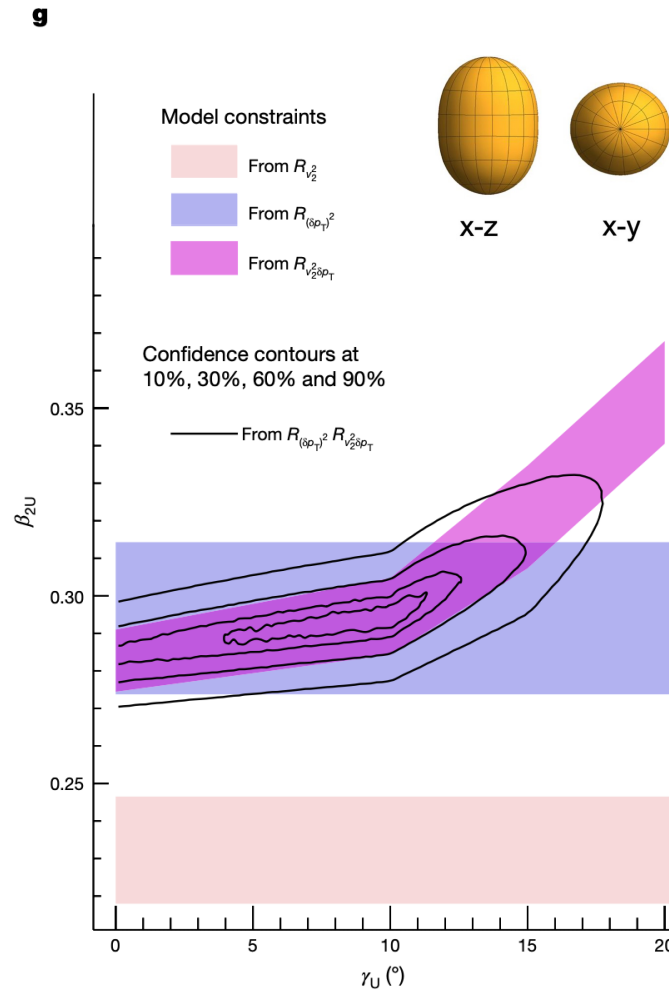
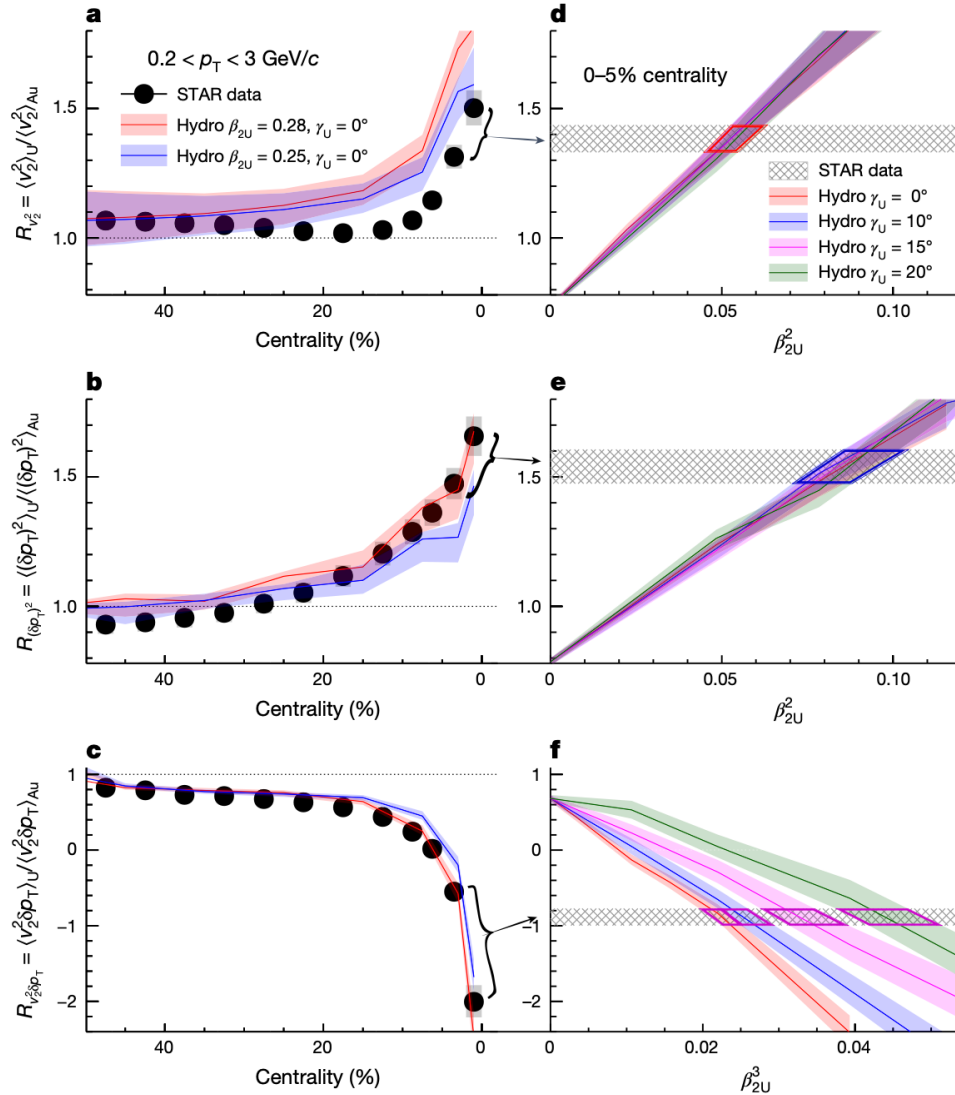
Relation confirmed from hydro

$$R_{(v_2^2)} \approx 1 + \frac{b_1}{a_1} \beta_{2U}^2,$$

$$R_{\langle (\delta p_T)^2 \rangle} \approx 1 + \frac{b_2}{a_2} \beta_{2U}^2,$$

$$R_{\langle v_2^2 \delta p_T \rangle} \approx 1 - \frac{b_3}{a_3} \beta_{2U}^3 \cos(3\gamma_U)$$

Constraining the ground-state ^{238}U : $\beta_{2,U}$ and γ_U



Sufficient precision is achieved from ratios in ultra-central collisions

Relation confirmed from hydro

$$R_{(v_2^2)} \approx 1 + \frac{b_1}{a_1} \beta_{2,U}^2,$$

$$R_{\langle (\delta p_T)^2 \rangle} \approx 1 + \frac{b_2}{a_2} \beta_{2,U}^2,$$

$$R_{\langle v_2^2 \delta p_T \rangle} \approx 1 - \frac{b_3}{a_3} \beta_{2,U}^3 \cos(3\gamma)$$

High-energy estimate
(Ip-glasma+MUSIC and Trajectum)

$$\beta_{2,U} = 0.286 \pm 0.025$$

$$\gamma_U = 8.5^\circ \pm 4.8^\circ$$

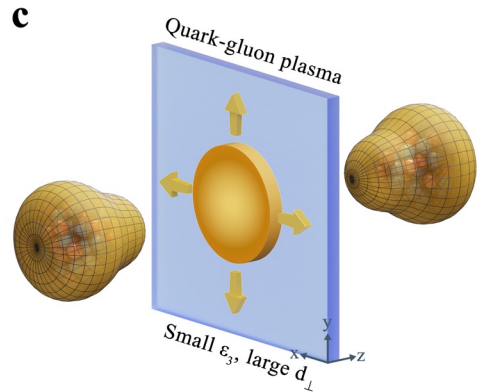
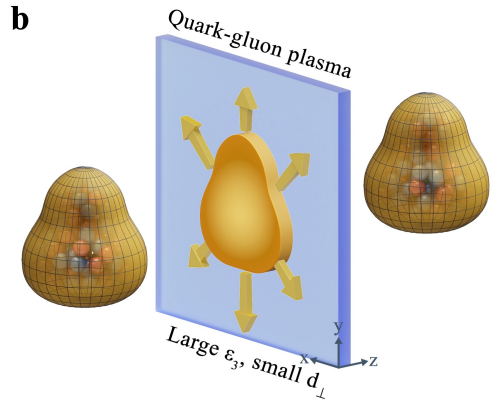
low-energy estimate:

$$\beta_{2,U} = 0.287 \pm 0.007$$

$$\gamma_U = 6^\circ - 8^\circ$$

A large deformation with a slight deviation from axial symmetry in the nuclear ground-state

Evidence of octupole deformation $\beta_{3,U}$



However, v_3 is fluctuation driven, expect in central

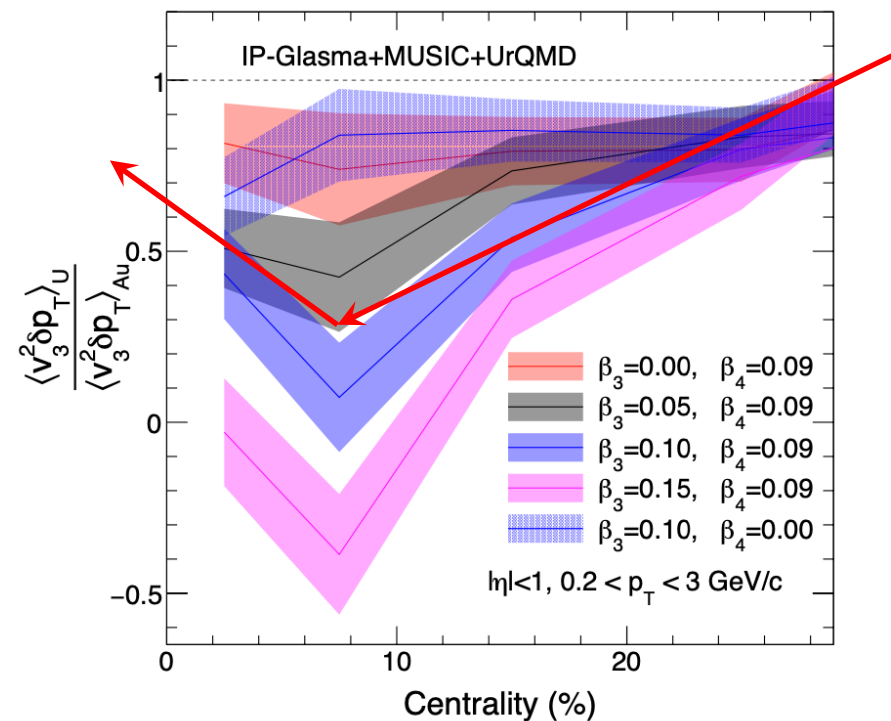
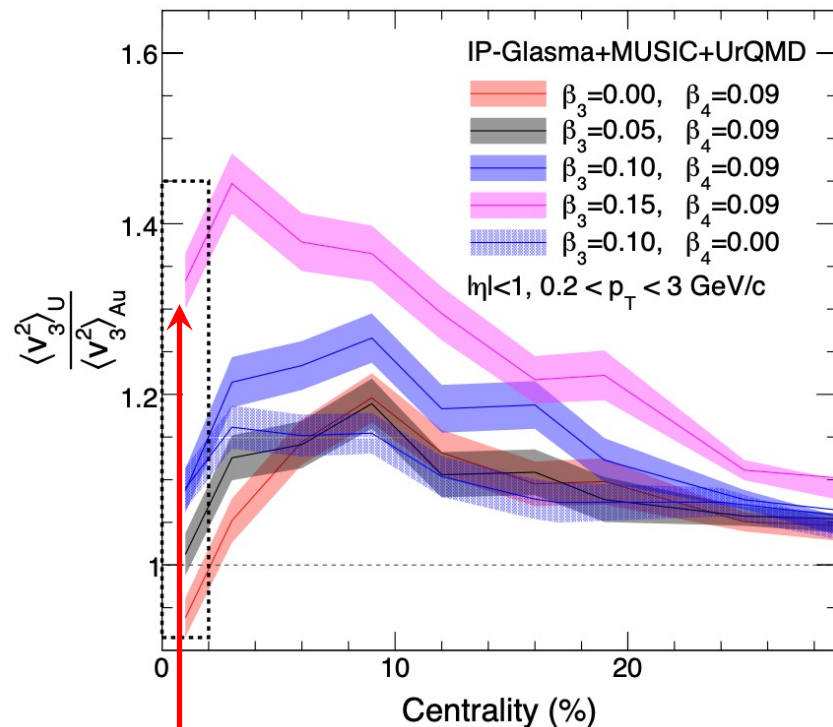
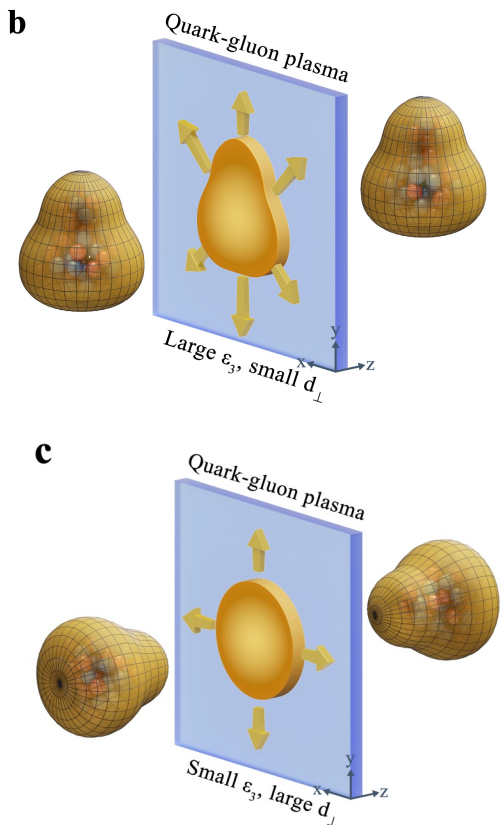
$$\langle v_3^2 \rangle \propto \langle \epsilon_3^2 \rangle \sim 1/A$$

mass number $\nearrow$

Evidence of octupole deformation $\beta_{3,U}$

IP-Glasma+MUSIC calculations

C. Zhang, J. Jia, J. Chen, C. Shen, L. Liu, 2504.15245



However, v_3 is fluctuation driven, expect in central

$$\langle v_3^2 \rangle \propto \langle \epsilon_3^2 \rangle \sim 1/A$$

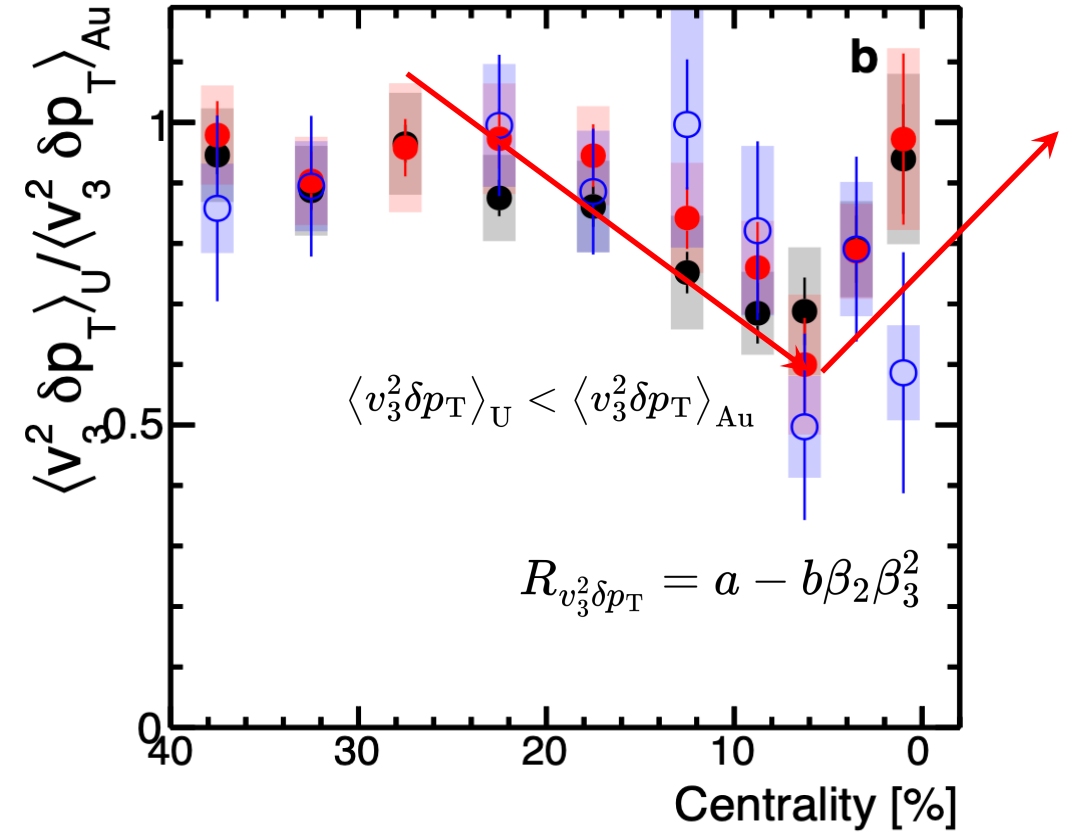
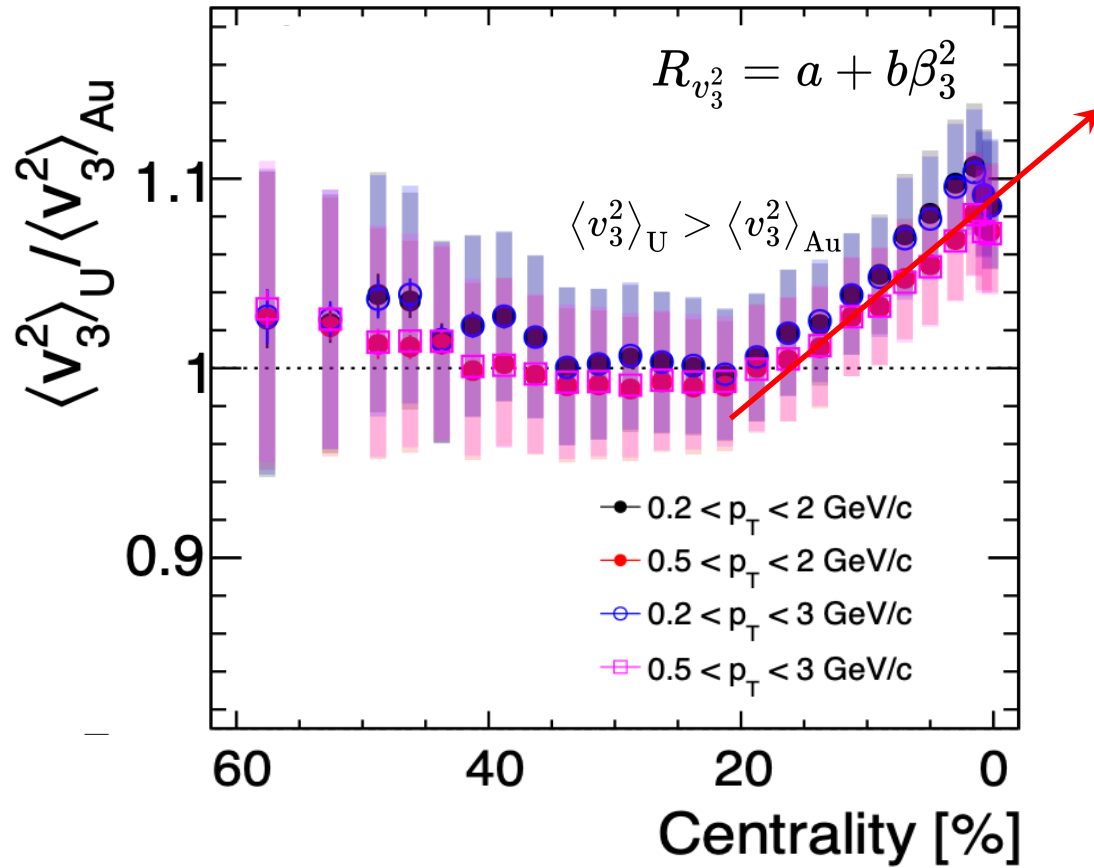
mass number

$\langle v_3^2 \rangle$ follows a linear increase with β_3^2 ,

Characteristic anticorrelation in $\langle v_3^2 \delta p_T \rangle$ shows a pronounced β_3 -dependent suppression.

Evidence of octupole deformation $\beta_{3,U}$

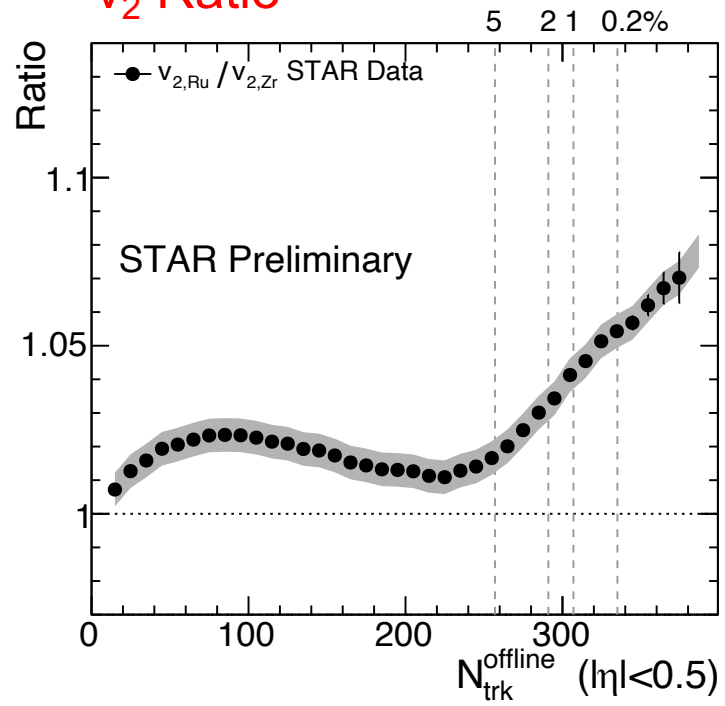
STAR, Rep. Prog. Phys. **88**, 108601 (2025)



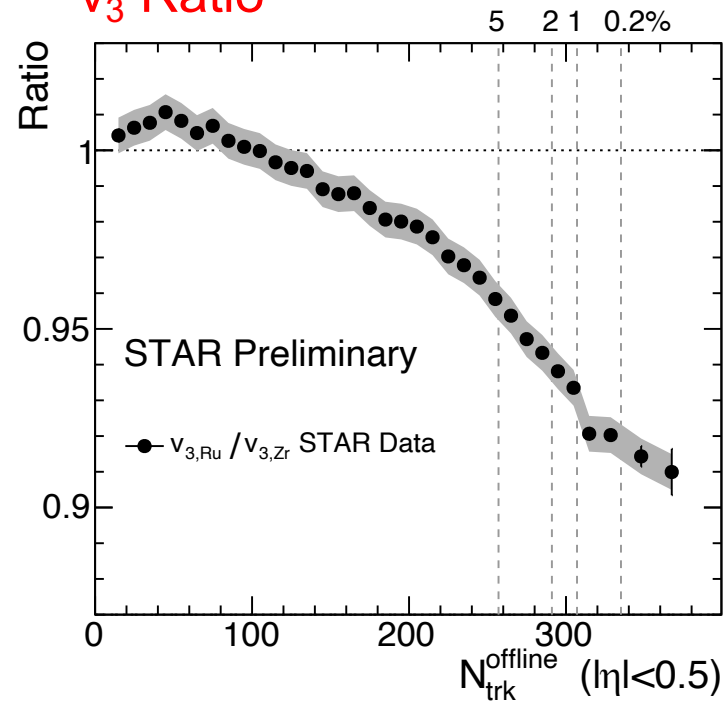
- No apparent p_T dependence
- Order of v_3 and v_3 - p_T reversed by considering non-zero $\beta_{3,U}$, $\beta_{4,U}$.
- An evidence and modest $\beta_{3,U} \sim 0.08$ - 0.10 are confirmed and $\beta_{3U} \sim \beta_{4U}$

Nuclear structure via collectivity v_n ratio

v_2 Ratio

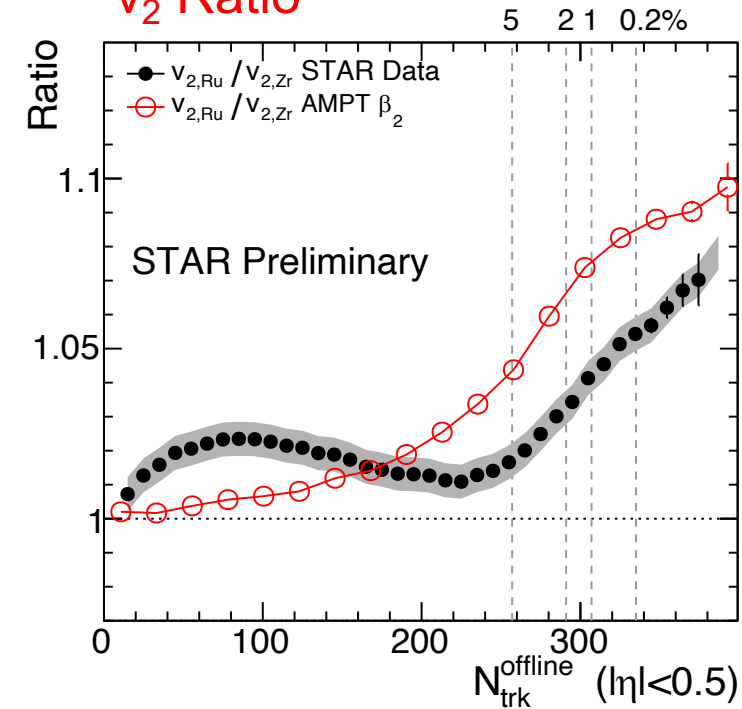


v_3 Ratio

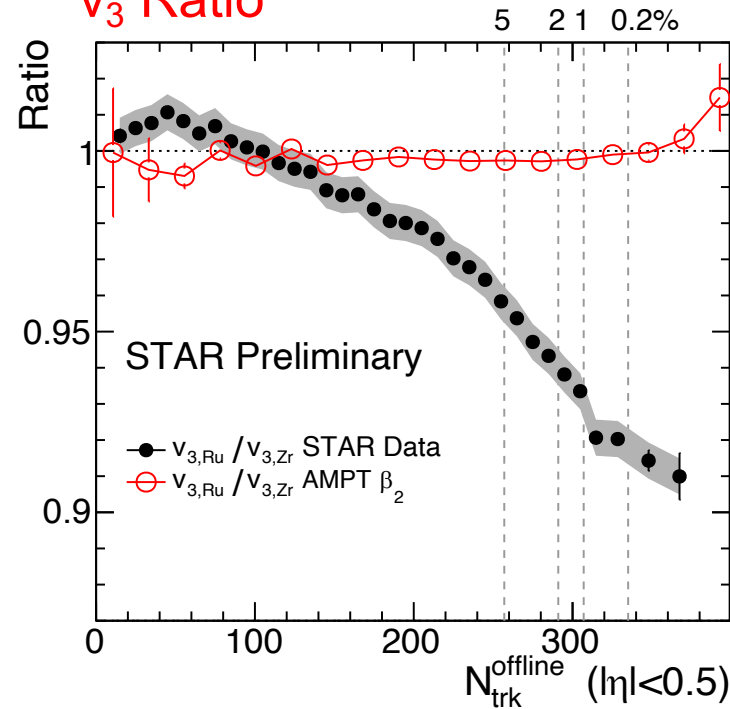


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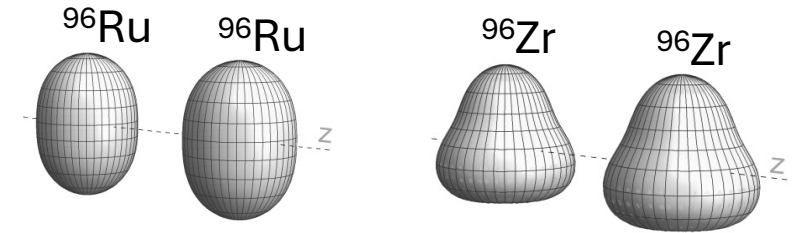
v_2 Ratio



v_3 Ratio



$\beta_{2,Ru} \sim 0.16$ increase v_2 , no influence on v_3 ratio



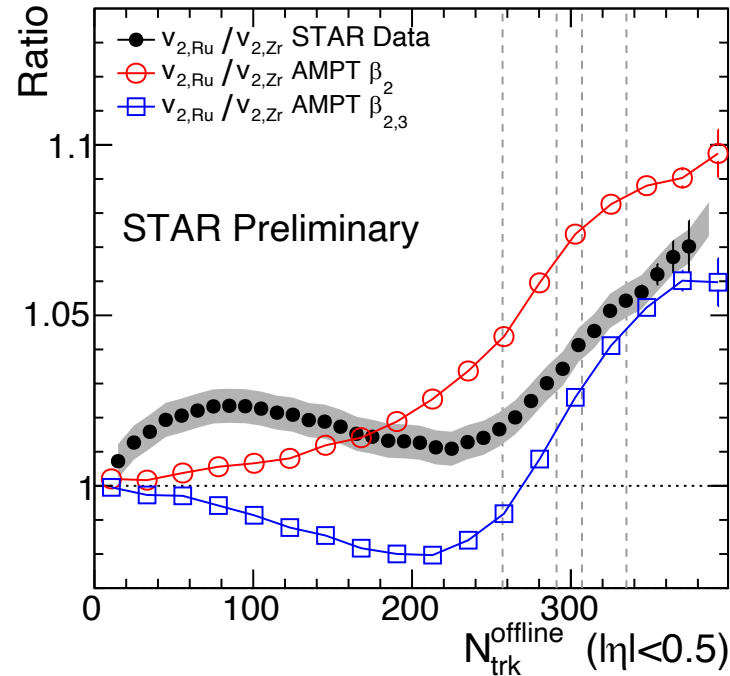
$$\beta_{2,Ru} = 0.16 \pm 0.02$$

difference	$\Delta\beta_2^2$	$\Delta\beta_3^2$	Δa_0	ΔR_0
	0.0226	-0.04	-0.06 fm	0.07 fm

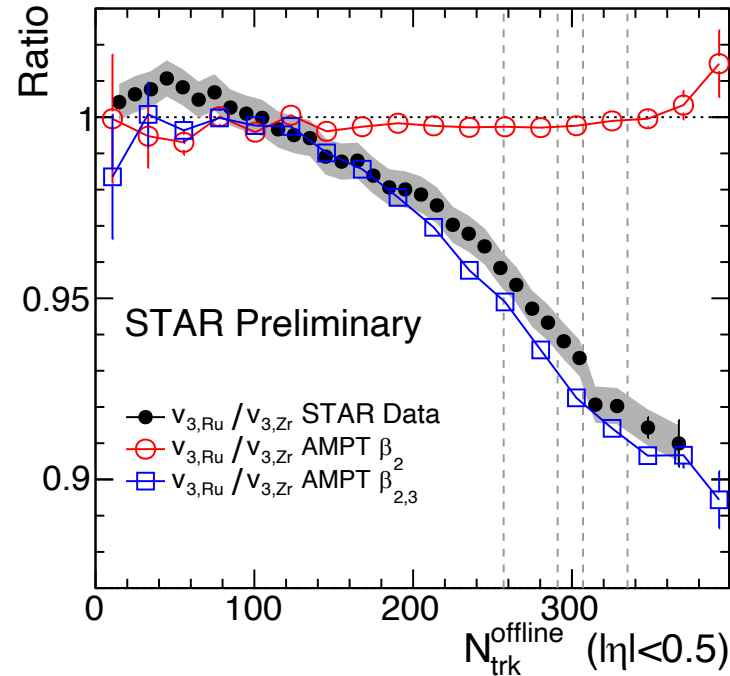
Current estimation is from transport model

Nuclear structure via collectivity v_n ratio

v_2 Ratio

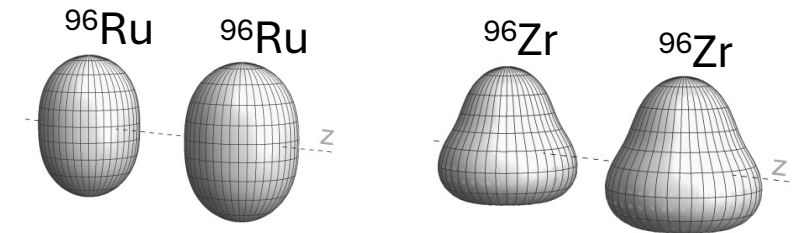


v_3 Ratio



$\beta_{2,Ru} \sim 0.16$ increase v_2 , no influence on v_3 ratio

$\beta_{3,Zr} \sim 0.2$ decrease v_2 in mid-central, decrease v_3 ratio



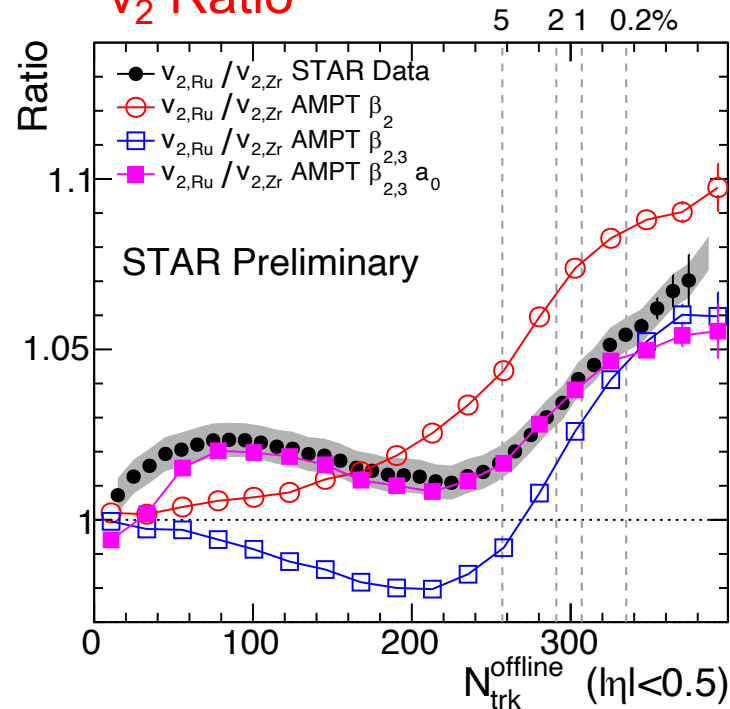
$$\beta_{2,Ru} = 0.16 \pm 0.02 \quad \beta_{3,Zr} = 0.20 \pm 0.02$$

difference	$\Delta\beta_2^2$	$\Delta\beta_3^2$	Δa_0	ΔR_0
	0.0226	-0.04	-0.06 fm	0.07 fm

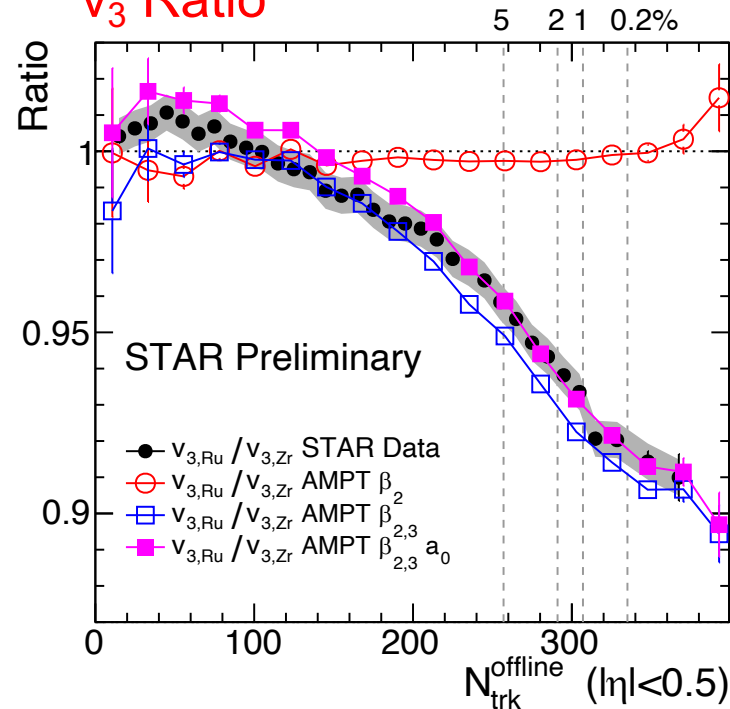
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Nuclear structure via collectivity v_n ratio

v_2 Ratio



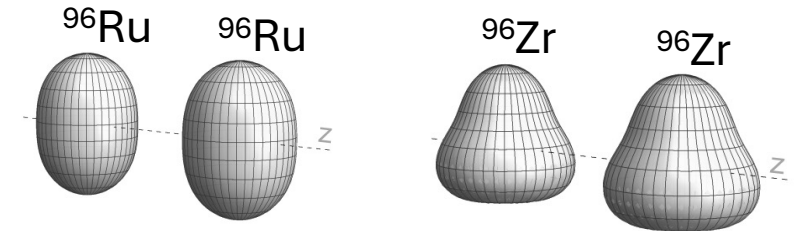
v_3 Ratio



$\beta_{2,Ru} \sim 0.16$ increase v_2 , no influence on v_3 ratio

$\beta_{3,Zr} \sim 0.2$ decrease v_2 in mid-central, decrease v_3 ratio

$\Delta a_0 = -0.06$ fm increase v_2 mid-central, small impact on v_3



$$\beta_{2,Ru} = 0.16 \pm 0.02$$

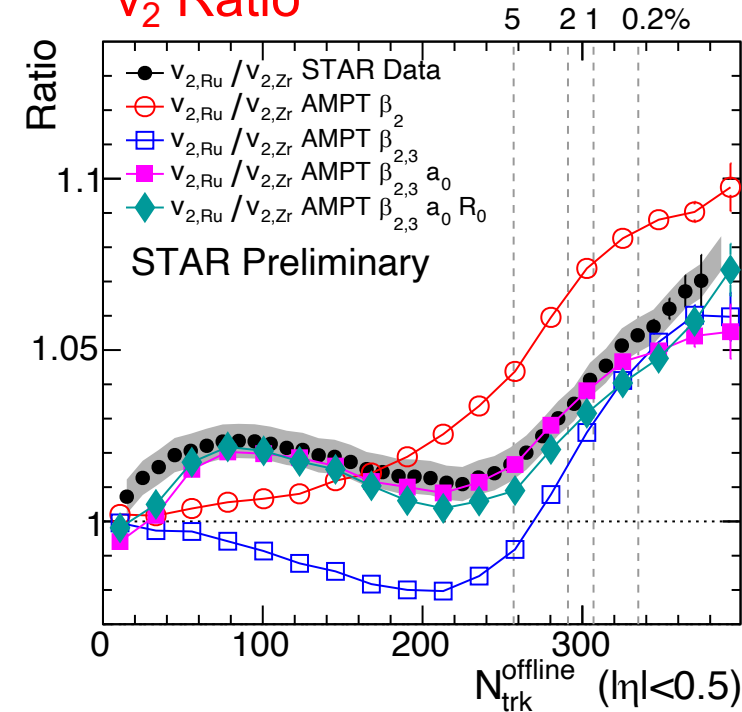
$$\beta_{3,Zr} = 0.20 \pm 0.02$$

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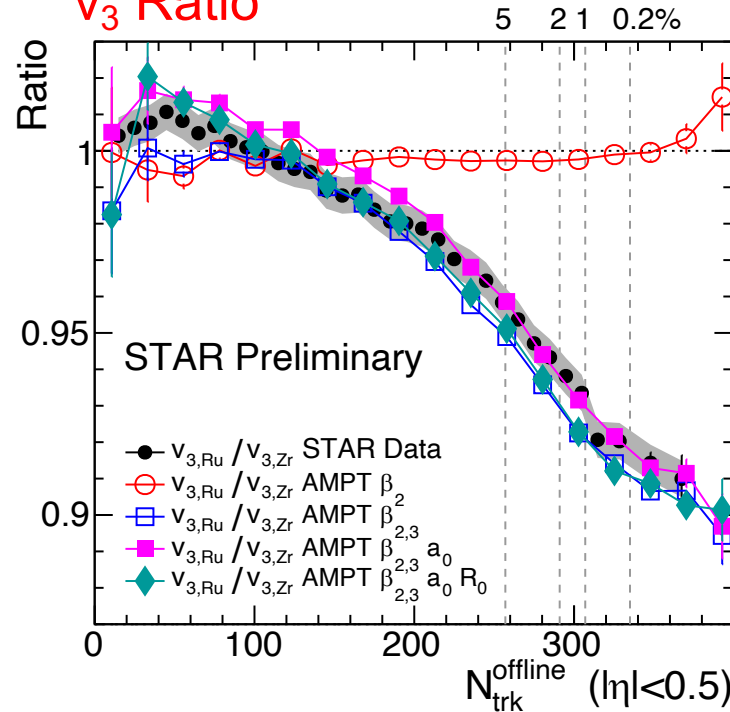
Current estimation is from transport model

Nuclear structure via collectivity v_n ratio

v_2 Ratio



v_3 Ratio

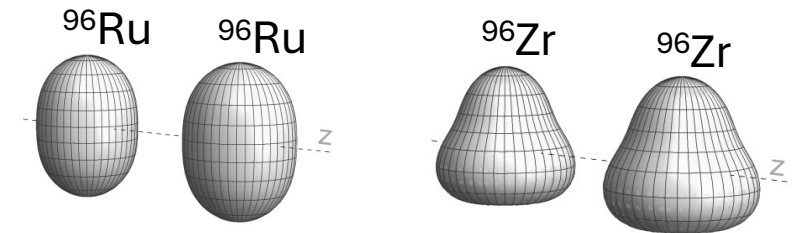


$\beta_{2,Ru} \sim 0.16$ increase v_2 , no influence on v_3 ratio

$\beta_{3,Zr} \sim 0.2$ decrease v_2 in mid-central, decrease v_3 ratio

$\Delta a_0 = -0.06$ fm increase v_2 mid-central, small impact on v_3

Radius $\Delta R_0 = 0.07$ fm only slightly affects v_2 and v_3 ratio.



$$\beta_{2,Ru} = 0.16 \pm 0.02$$

$$\beta_{3,Zr} = 0.20 \pm 0.02$$

difference	$\Delta\beta_2^2$	$\Delta\beta_3^2$	Δa_0	ΔR_0
	0.0226	-0.04	-0.06 fm	0.07 fm

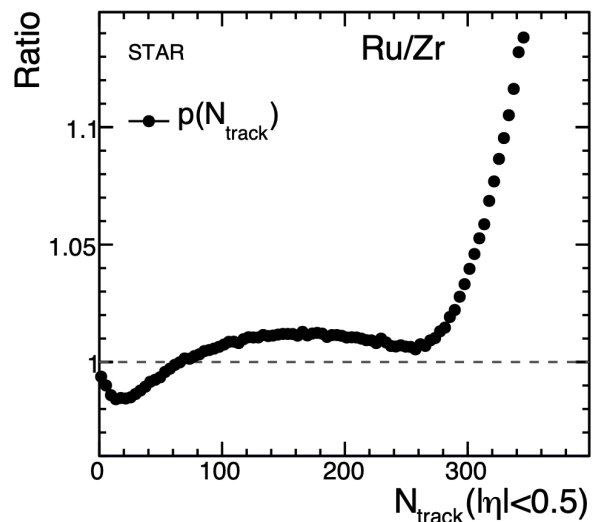
Current estimation is from transport model

- Direct observation of octupole deformation in ^{96}Zr nucleus
- Imply the neutron skin difference between ^{96}Ru and ^{96}Zr
- Simultaneously constrain parameters using Bayesian analysis

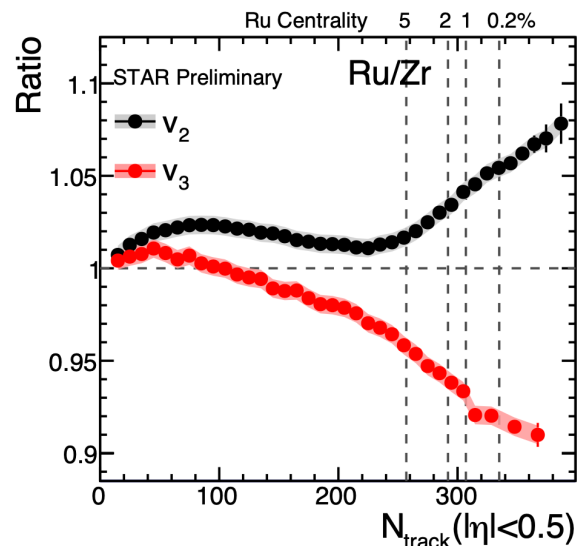
$$R_O \equiv \frac{O_{Ru}}{O_{Zr}} \approx 1 + c_1 \Delta\beta_2^2 + c_2 \Delta\beta_3^2 + c_3 \Delta R_0 + c_4 \Delta a$$

Nuclear structure is inherent of heavy-ion probes

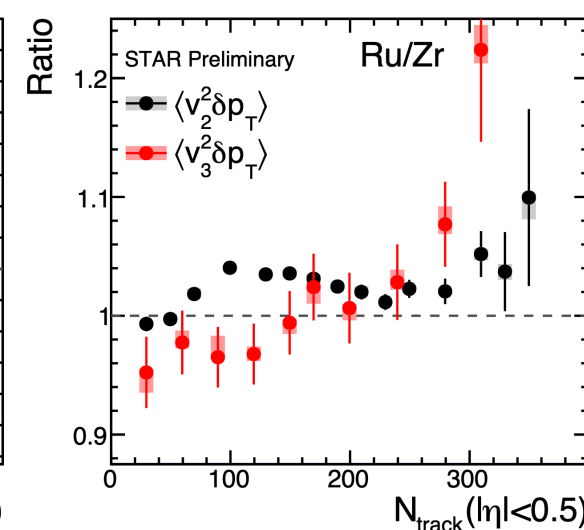
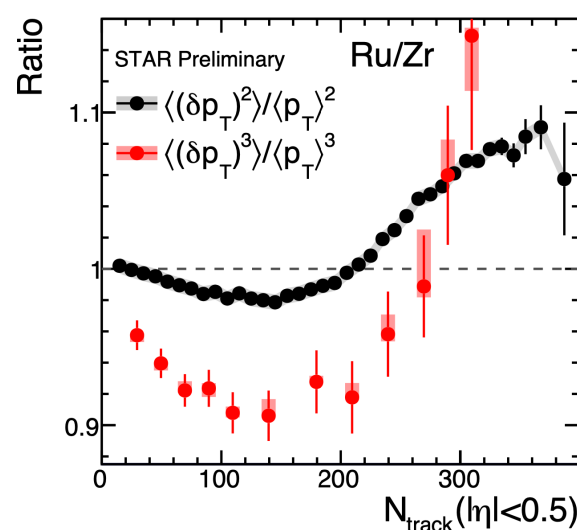
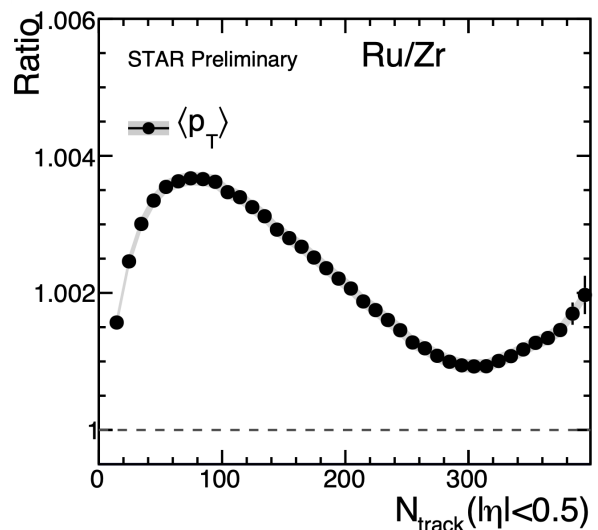
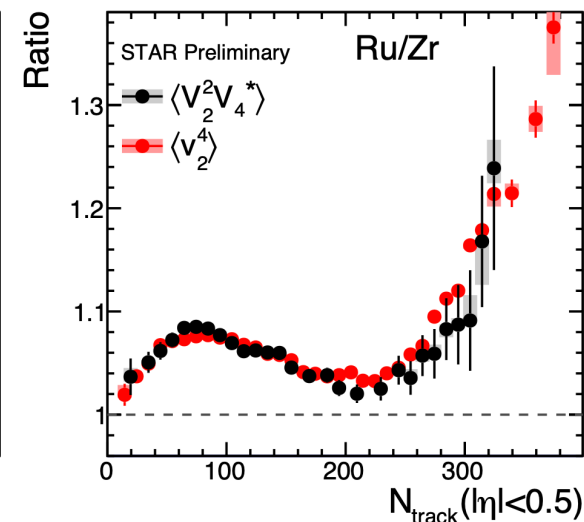
One-body distribution



Two-body correlations



Three-body correlations



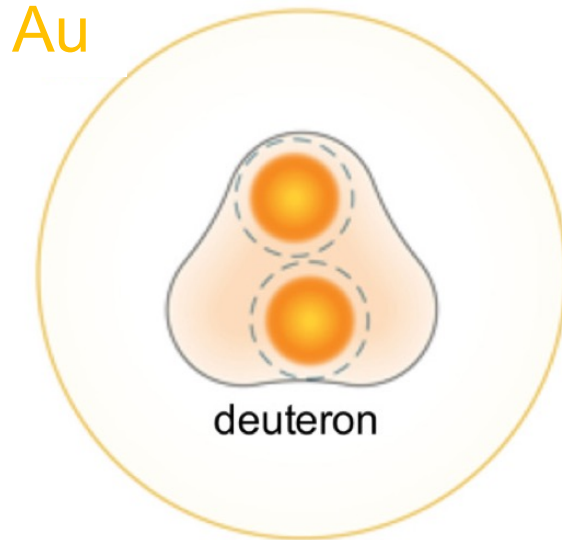
III. Nucleonic cluster pattern in light ^{16}O nucleus

--- from one-body distribution to many-body nucleon correlations

$$\rho(r) \propto \frac{1 + w(r^2/R^2)}{1 + e^{(r-R)/a_0}} \longrightarrow \text{Modern } ab \text{ initio first-principle calculations}$$

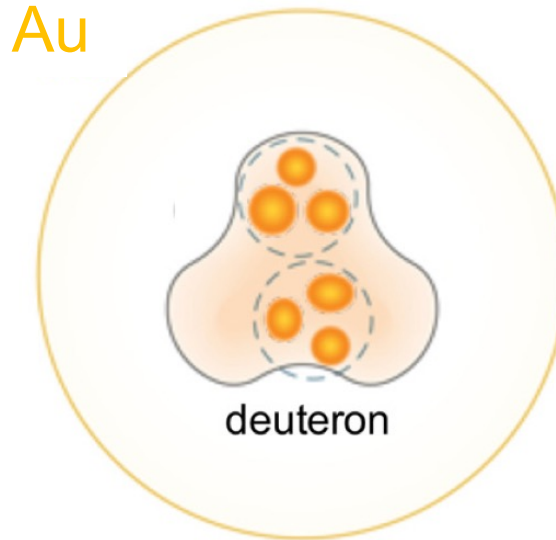
Light nuclear geometry and nucleon-nucleon correlations

Nucleon fluctuation

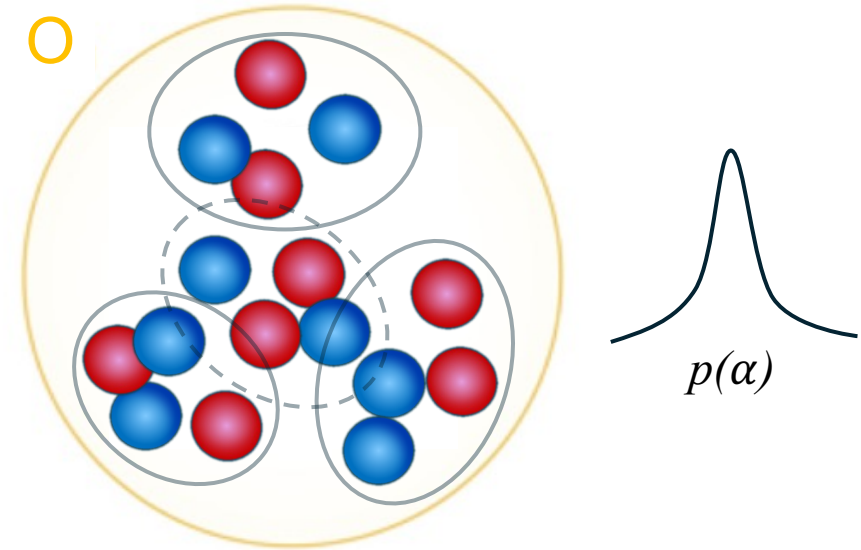


STAR, PRL 130, 242301 (2023)
STAR, PRC 110, 064902 (2024)

subnucleon fluctuation



NN correlation (cluster pattern)



ab initio calculations

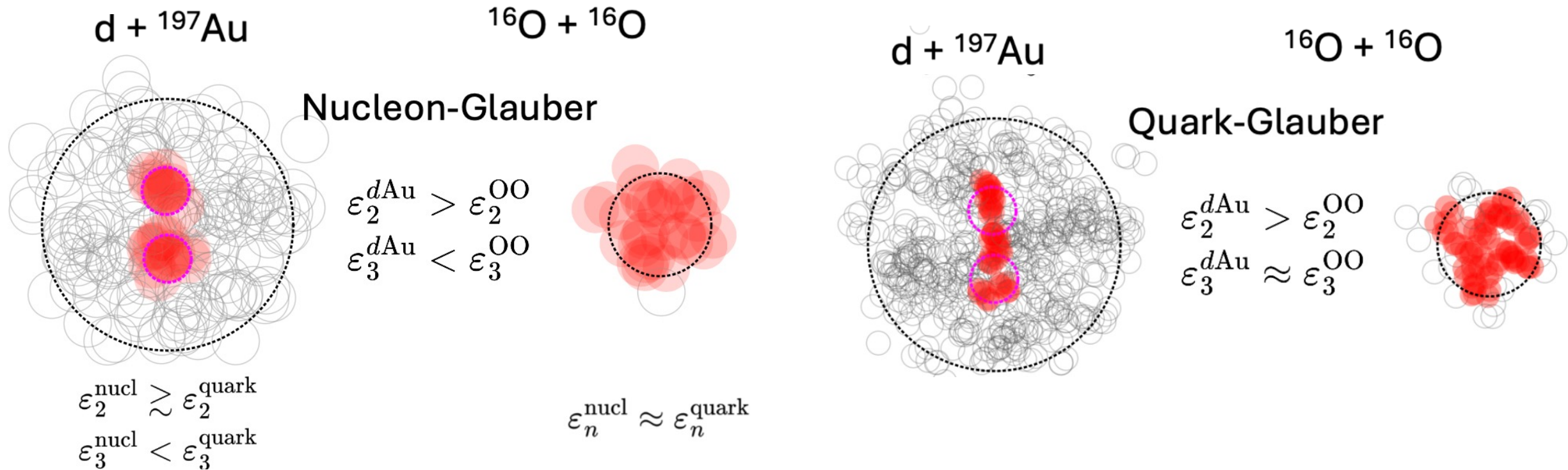
W. He, Y.-G. Ma et al., PRL 113, 032506 (2014)
Y.-G. Ma, S. Zhang, Handbook of Nuclear Physics (2022)
C. Zhang, J. Chen, G. Giacalone, S. Huang, J. Jia, Y.-G. Ma, PLB 862, 139322 (2025)
S. Huang, J. Jia, C. Zhang, PLB 870, 139926 (2025)

d+Au collision, 200 GeV

Nearly same N_{ch}

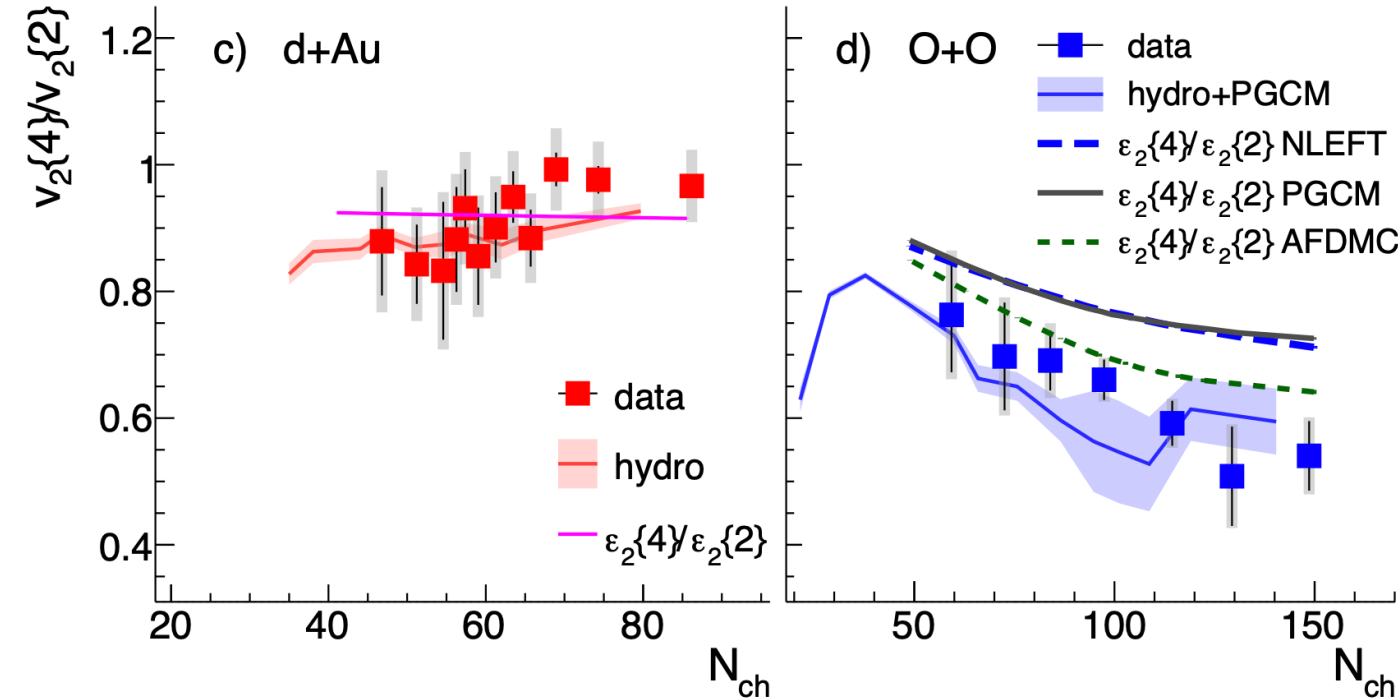
O+O collision, 200 GeV

STAR, 2510.19645



Nucleon-Glauber and quark-Glauber yield different triangular v_3

STAR, 2510.19645



$\varepsilon_2\{4\}/\varepsilon_2\{2\}$ from three models:

AFDMC vs. EFT/PGCM have a visible difference.

Can many-nucleon correlations significantly impact the eccentricity fluctuations? YES!

C. Zhang et al., PLB 862, 139322 (2025)

The interplay between sub-nucleon fluctuation and many-nucleon correlation.

STAR, PRL 130, 242301 (2023)

S. Huang, J. Jia, C. Zhang, PLB 870, 139926 (2025)

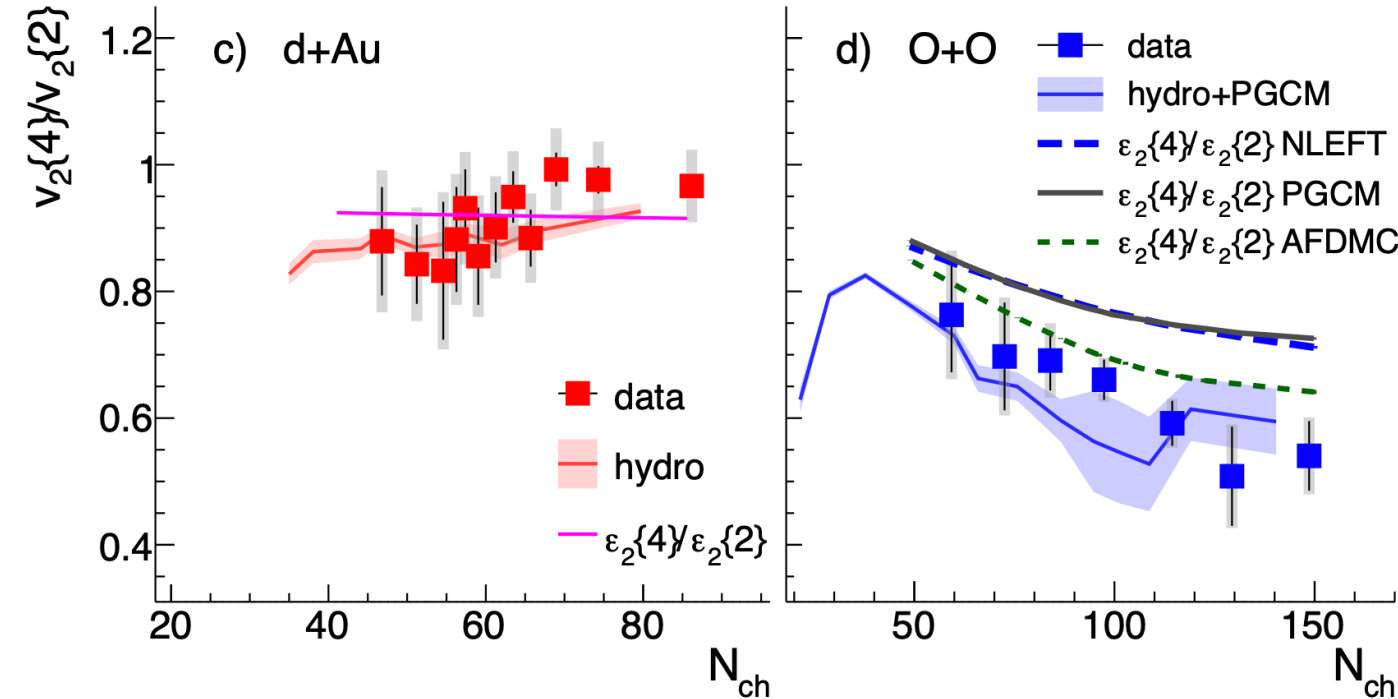
$$\begin{aligned} (v_n\{2\})^2 &= c_n\{2\} = \langle v_n^2 \rangle & \varepsilon_2\{2\}^2 &= \langle \varepsilon_2^2 \rangle \\ (v_n\{4\})^4 &= -c_n\{4\} = 2\langle v_n^2 \rangle^2 - \langle v_n^4 \rangle & \varepsilon_2\{4\}^4 &= 2\langle \varepsilon_2^2 \rangle^2 - \langle \varepsilon_2^4 \rangle \end{aligned}$$

Y.-G. Ma, S. Zhang, Handbook of Nuclear Physics (2022); W. He, Y.-G. Ma et al, PRL 113, 032506 (2014)

G. Giacalone, G. Nijs et al., PRL135, 012302 (2025), G. Giacalone, W. Zhao et al., PRL134, 082301 (2025); P. Li, B. Zhou, G.-L. Ma, 2504.04688

Y. Wang, S. Zhao, B. Cao, H. Xu, H. Song, PRC 109, L051904 (2024); X. Zhao, G.-L. Ma. Y. Zhou, Z. Lin, C. Zhang, 2404.09780; S. Jahan, Roch, C. Shen, 2507. 11394

STAR, 2510.19645



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STAR, PRL 130, 242301 (2023)

S. Huang, J. Jia, C. Zhang, PLB 870, 139926 (2025)

Geometric scan elucidates nuclear tomography and strong nuclear force?

more low-energy model inputs & revisit

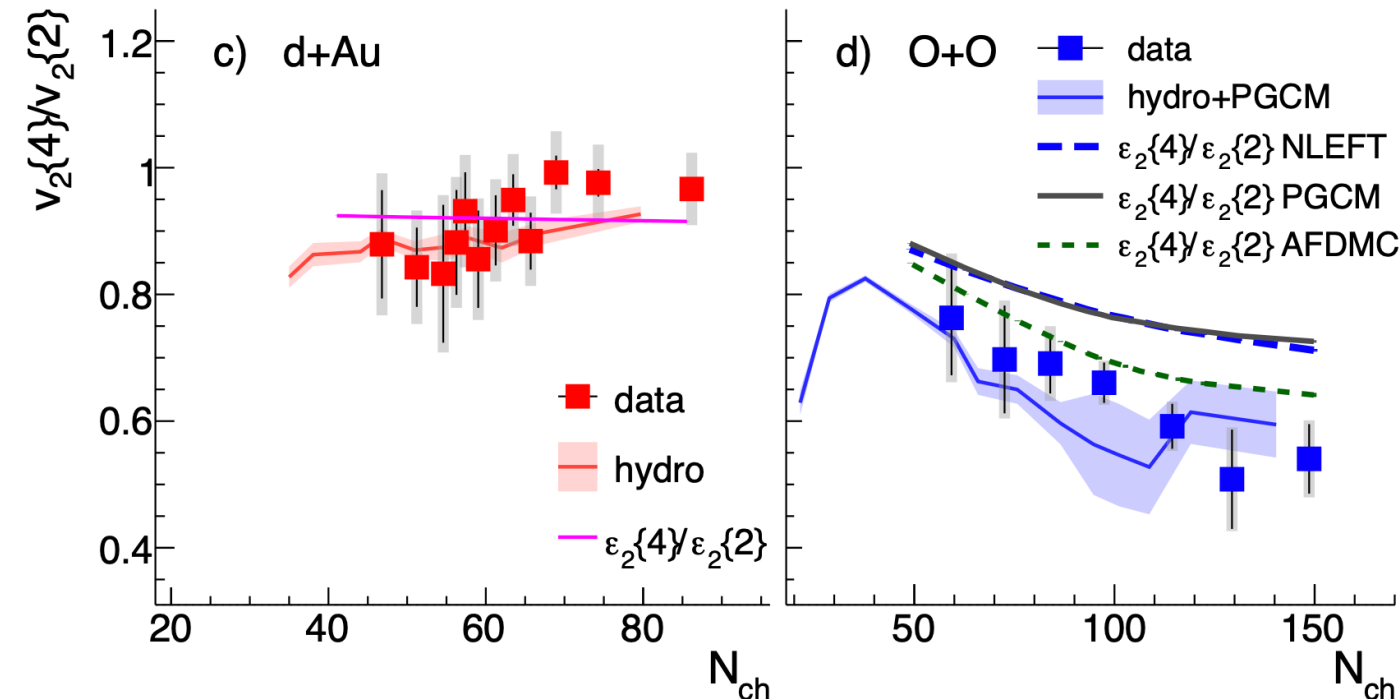
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$$\varepsilon_2\{4\}^4 = 2\langle \varepsilon_2^2 \rangle^2 - \langle \varepsilon_2^4 \rangle$$

LHC (ALICE, ATLAS, CMS, LHCb) also released the O+O/Ne+Ne.
Shape imaging method for light ions (New avenue)

Y.-G. Ma, S. Zhang, Handbook of Nuclear Physics (2022); W. He, Y.-G. Ma et al, PRL 113, 032506 (2014)

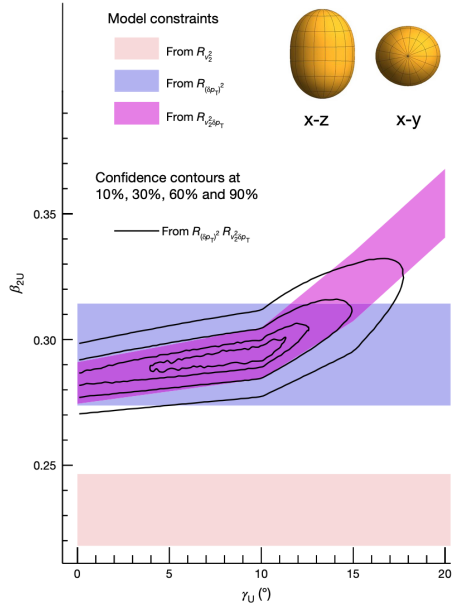
G. Giacalone, G. Nijs et al., PRL135, 012302 (2025), G. Giacalone, W. Zhao et al., PRL134, 082301 (2025); P. Li, B. Zhou, G.-L. Ma, 2504.04688

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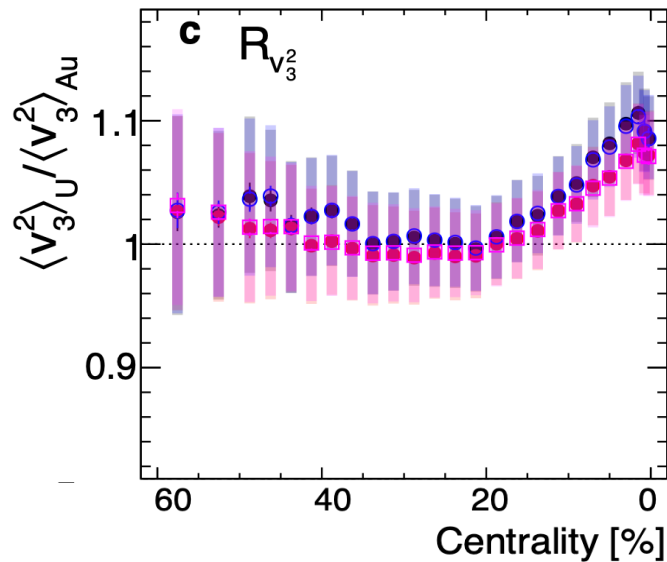
IV. Conclusions and Outlooks

1. The signatures of nuclear structure in nuclear collisions are ubiquitous:

STAR, *Nature* 635, 67-72 (2024)

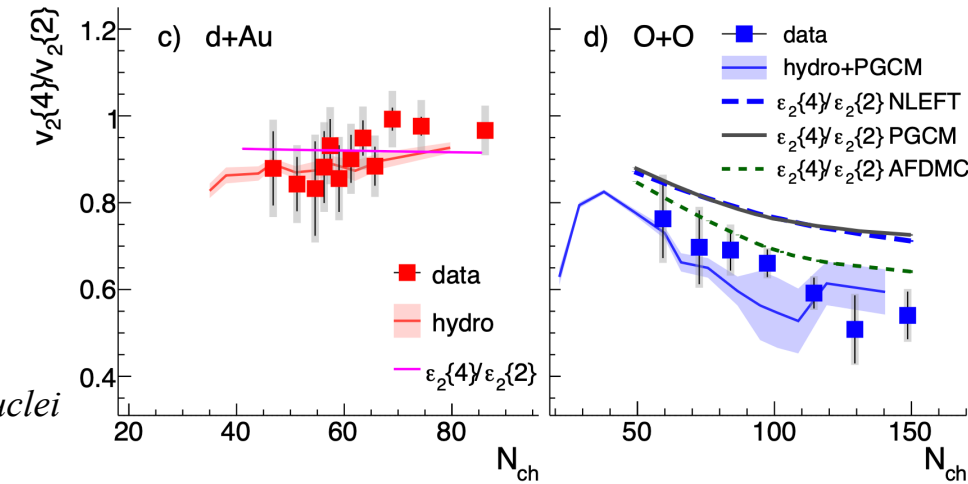


STAR, *Rep. Prog. Phys.* **88**, 108601 (2025)



from *large* to *small* nuclei

STAR, 2510.19645

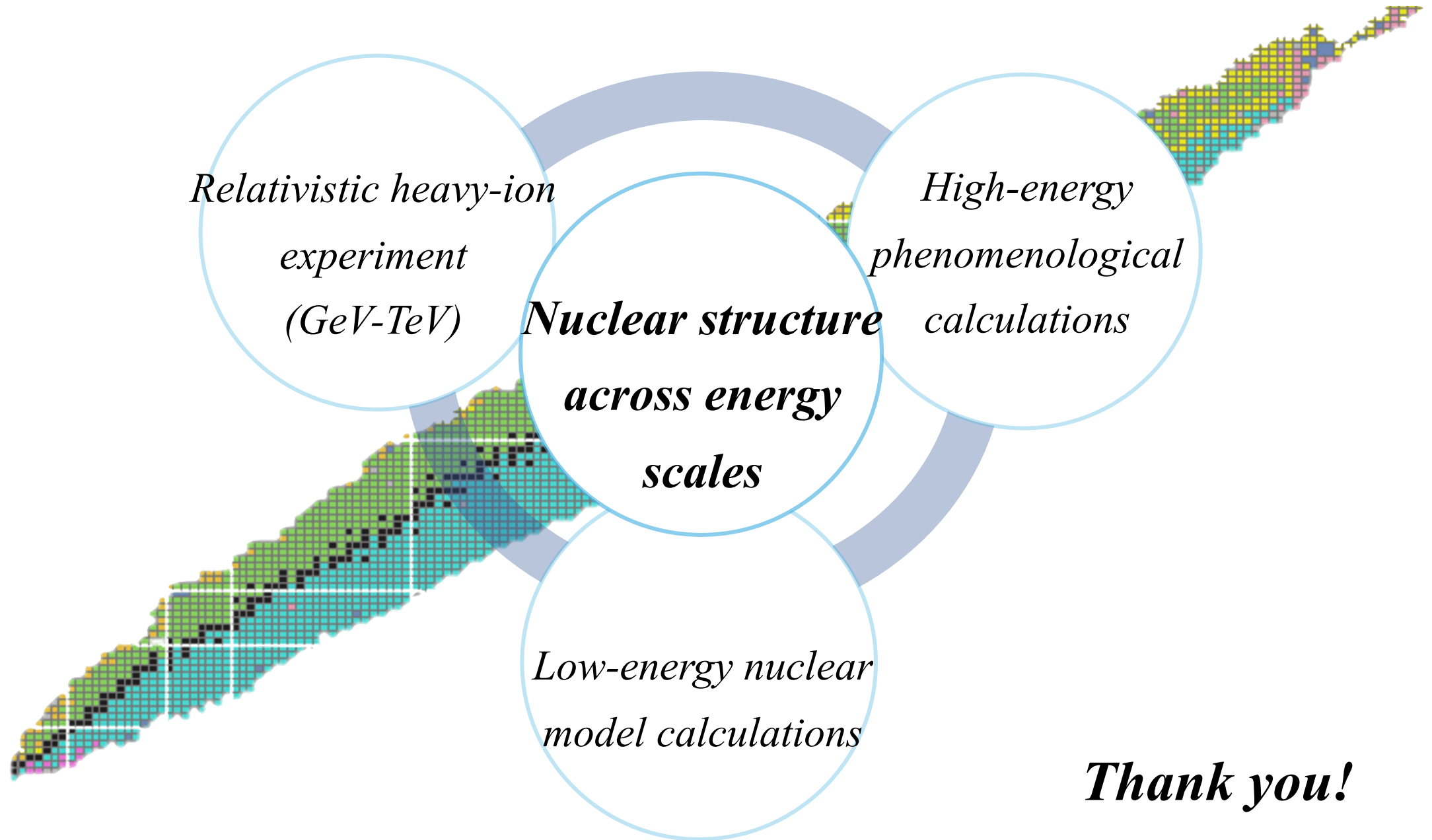


NN correlations is remarkable

2. Many potential applications from large to small collision systems :

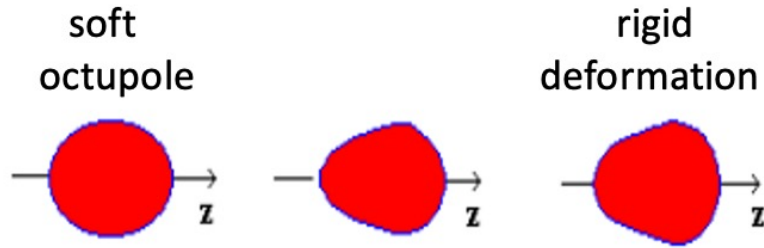
- Rigid and soft β_n and γ (shape fluctuations/coexistence)
- Neutron skin & symmetry energy constraints
- Energy evolution between GeV/TeV and MeV in even-and odd-A nuclei
- Confronted with ab initio calculations and more light-ion studies
- Intersection with other research: Nucleosynthesis, nuclear fission, $0\nu\beta\beta$

IV. Conclusions and Outlooks



Probe $\beta_{3,U}$ and its fluctuation

Octupole collectivity



$$\langle \beta_3^2 \rangle = \bar{\beta}_3^2 + \sigma_{\beta_3}^2$$

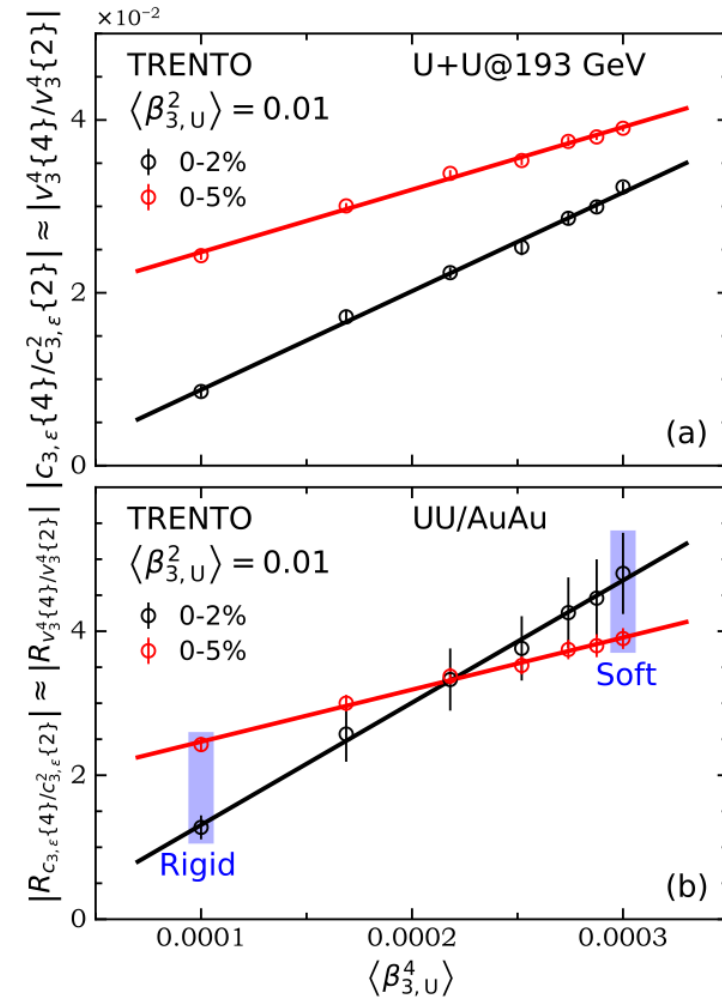
$$c_{n,\varepsilon}\{2\} = \langle \varepsilon_n^2 \rangle \approx \langle \varepsilon_{n,0}^2 \rangle + \langle p_n p_n^* \rangle \langle \beta_n^2 \rangle$$

$$\begin{aligned} c_{n,\varepsilon}\{4\} &= \langle \varepsilon_n^4 \rangle - 2\langle \varepsilon_n^2 \rangle^2 \\ &\approx \langle \varepsilon_{n,0}^4 \rangle - 2\langle \varepsilon_{n,0}^2 \rangle^2 + \langle p_n^2 p_n^{2*} \rangle \langle \beta_n^4 \rangle - 2\langle p_n p_n^* \rangle^2 \langle \beta_n^2 \rangle^2 \end{aligned}$$

Four-particle correlation is linearly scaled to $\langle \beta_{3,U}^4 \rangle$.

A way to discriminate between static and dynamic collective modes in high-energy nuclear collisions

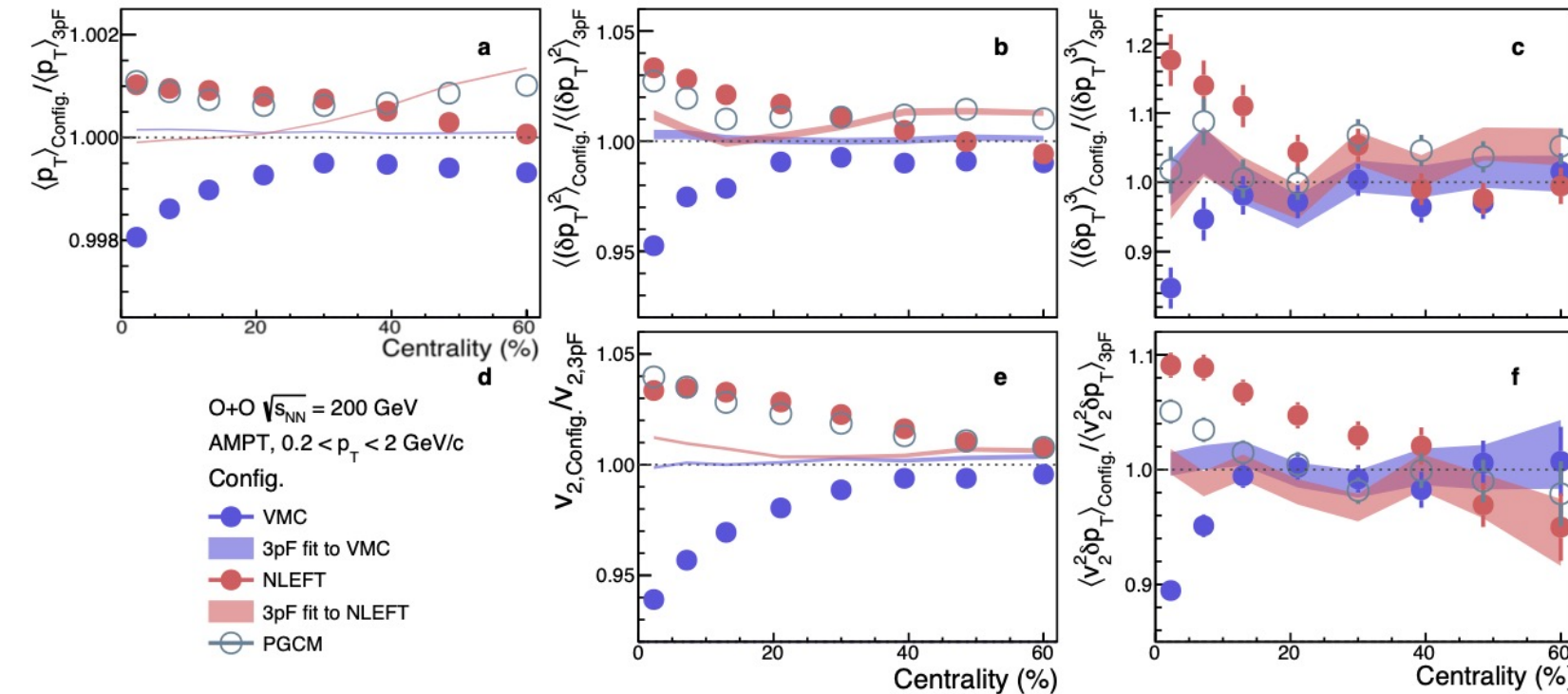
L. Liu, C. Zhang, J. Chen, J. Jia, X. Huang, Y.-G. Ma, 2509.09376



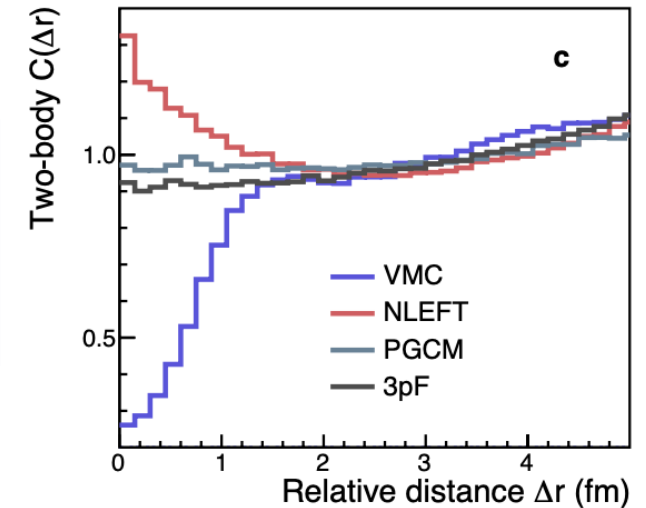
Ab-initio nucleon-nucleon correlations and their impact on high-energy

C. Zhang, J. Chen, G. Giacalone, S. Huang, J. Jia, Y.-G. Ma, *PLB* 862, 139322 (2025)

Multiple final-state observables



Two-body correlations in coordinate space



$$C(\Delta r) = \frac{\int d^3s f(\mathbf{s} + \frac{\mathbf{r}}{2}, \mathbf{s} - \frac{\mathbf{r}}{2})}{\int d^3s f(\mathbf{s} + \frac{\mathbf{r}}{2}) f(\mathbf{s} - \frac{\mathbf{r}}{2})}$$

- 3pF as the baseline can NOT capture the NN correlations.
- Distinct differences are observed between VMC and NLEFT/PGCM.
- Central collision is prominent.

Data analysis is ongoing and will be presented in *sQM*.