



U.S. DEPARTMENT OF
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Quasi-steady states and miscodynamics in “once-in-a-lifetime encounter” neutrino models

Anson Kost

University of New Mexico

(In collaboration with Huaiyu Duan and Lucas Johns)

Collective Neutrino Oscillations in Supernovae and Neutron Star Mergers

25 March 2026

The mean-field approximation ignores entanglement

$$\underbrace{i(\partial_t + \vec{v} \cdot \nabla)f = [H_f, f]}_{\text{Mean-field approximation}} + C_f$$

Mean-field approximation

$$f = \begin{pmatrix} f_{ee} & f_{ex} \\ f_{xe} & f_{xx} \end{pmatrix} (t, \vec{r}, \vec{p})$$

- C_f : loss of coherence/purity
 - Randomness (collisions)
 - Entanglement

Side note:

- Engineers: $3 \approx \pi$
- Physicists: $3 \approx 2$

Many-body models

$$|\psi\rangle_{N \text{ neutrinos}} \doteq \begin{pmatrix} c_{(1)} \\ c_{(2)} \\ c_{(3)} \\ \vdots \\ c_{(2^N)} \end{pmatrix}$$

$$\rho_a = \begin{pmatrix} \rho_{ee} & \rho_{ex} \\ \rho_{xe} & \rho_{xx} \end{pmatrix}_a = \text{Tr}_{\text{not } a} (|\psi\rangle\langle\psi|)$$

- “Curse of dimensionality”: $N \lesssim 20$

Side note:

- Engineers: $20.1 \approx 20$
- Physicists: $10^{58} \approx 20$

Is entanglement important?

- Millions–billions of moles of neutrinos per cubic centimeter
- We know they collectively oscillate
- Can neutrinos collectively entangle, on the scale of an entire star?

Two-neutrino forward scattering

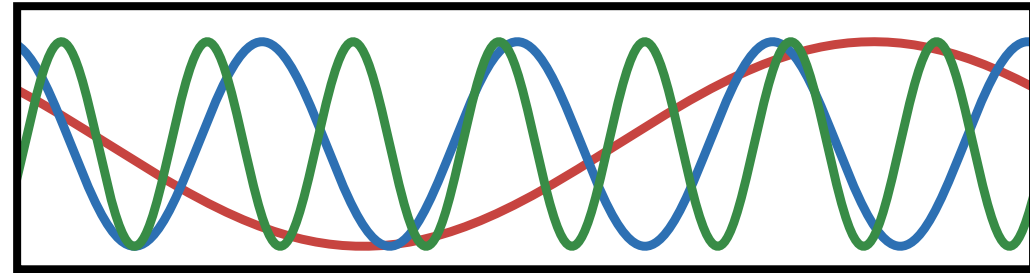
$$H_{12} = \mu(1 - \vec{v}_1 \cdot \vec{v}_2) \boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2$$

$$\boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2 \propto \begin{matrix} & \begin{matrix} ee & ex & xe & xx \end{matrix} \\ \begin{pmatrix} 1 & & & \\ & & 1 & \\ & 1 & & \\ & & & 1 \end{pmatrix} & \begin{matrix} ee \\ ex \\ xe \\ xx \end{matrix} \end{matrix}$$

$$|ex\rangle \rightarrow c_1|ex\rangle + c_2|xe\rangle \text{ (entanglement)}$$

Plane wave models

$$H = \sum_{\text{pairs } i,j} H_{ij}$$

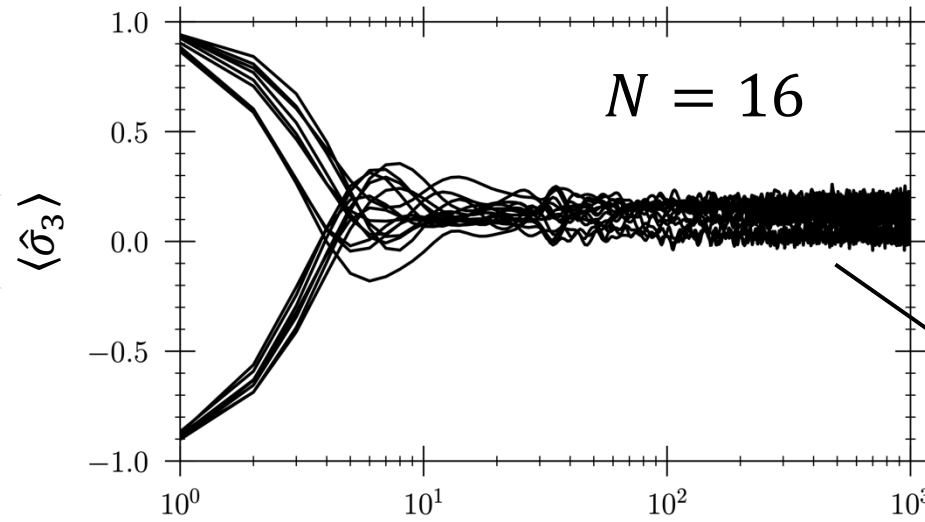


“Plane waves in a (periodic) box”

Unentangled

Entangled

Unentangled



equilibration

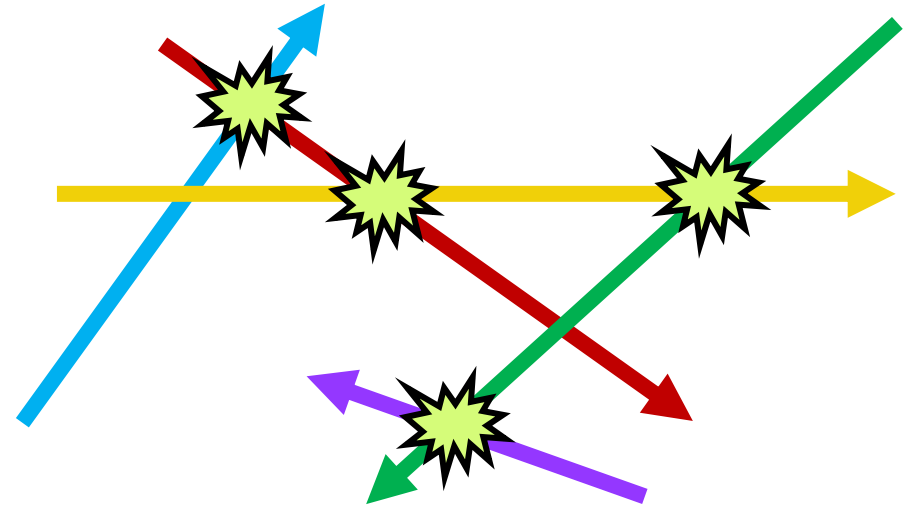
Martin et al.

Dense but dilute gas

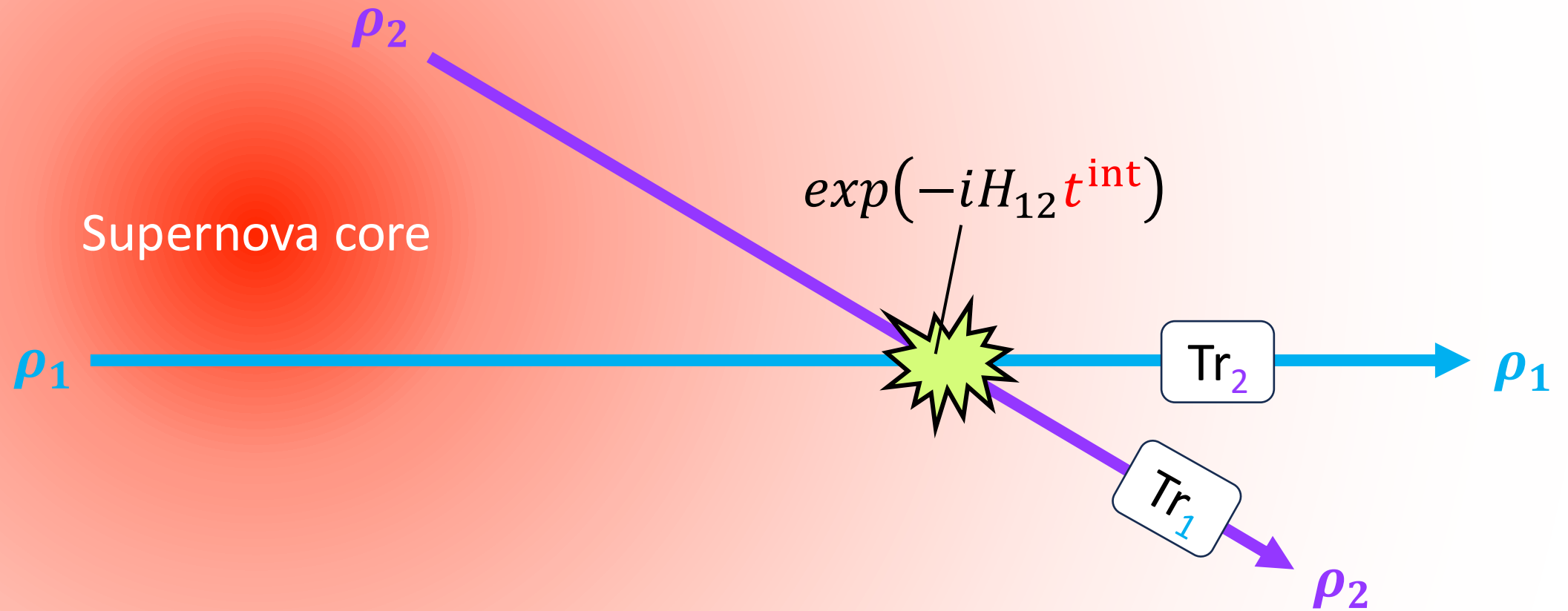
- Dense: Distance between neutrinos ~ 1 pm
- But also **dilute**: Wavepacket size ~ 0.1 pm (?)



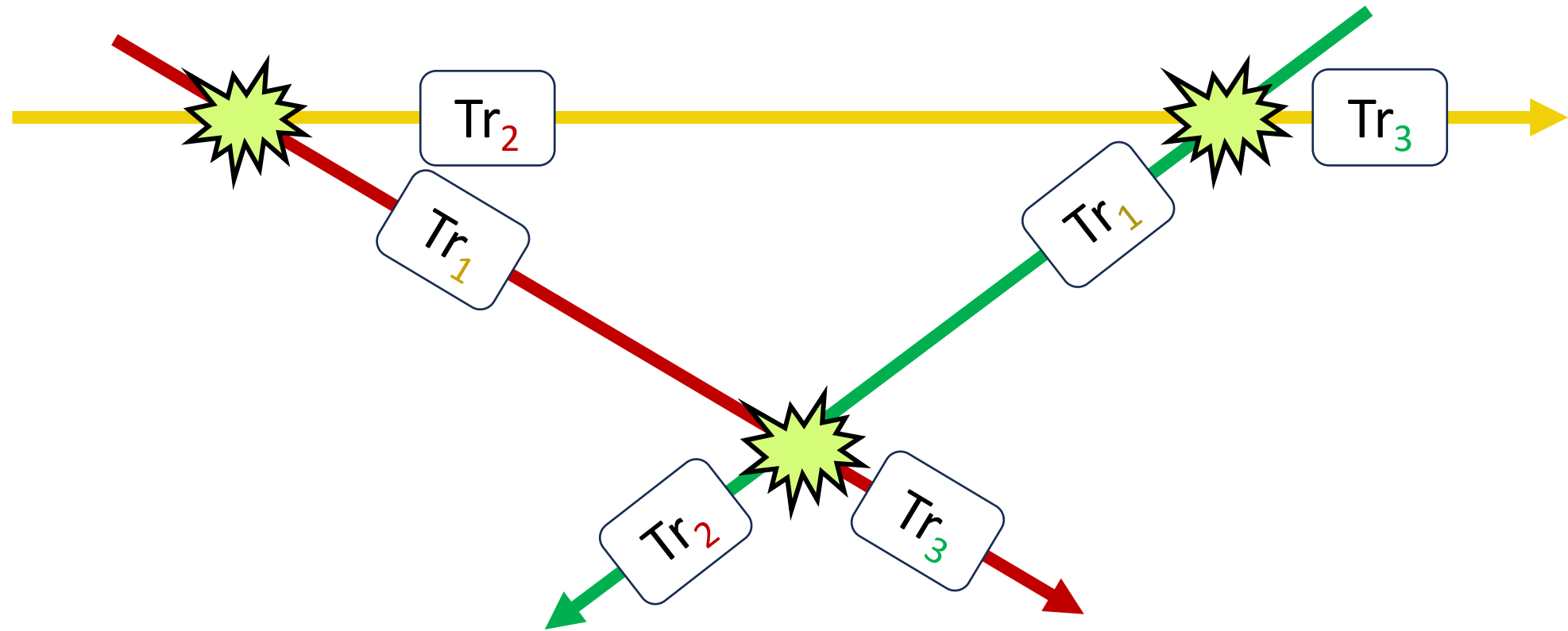
vs.



A once-in-a-lifetime encounter



Once-in-a-lifetime encounter (OILE) models

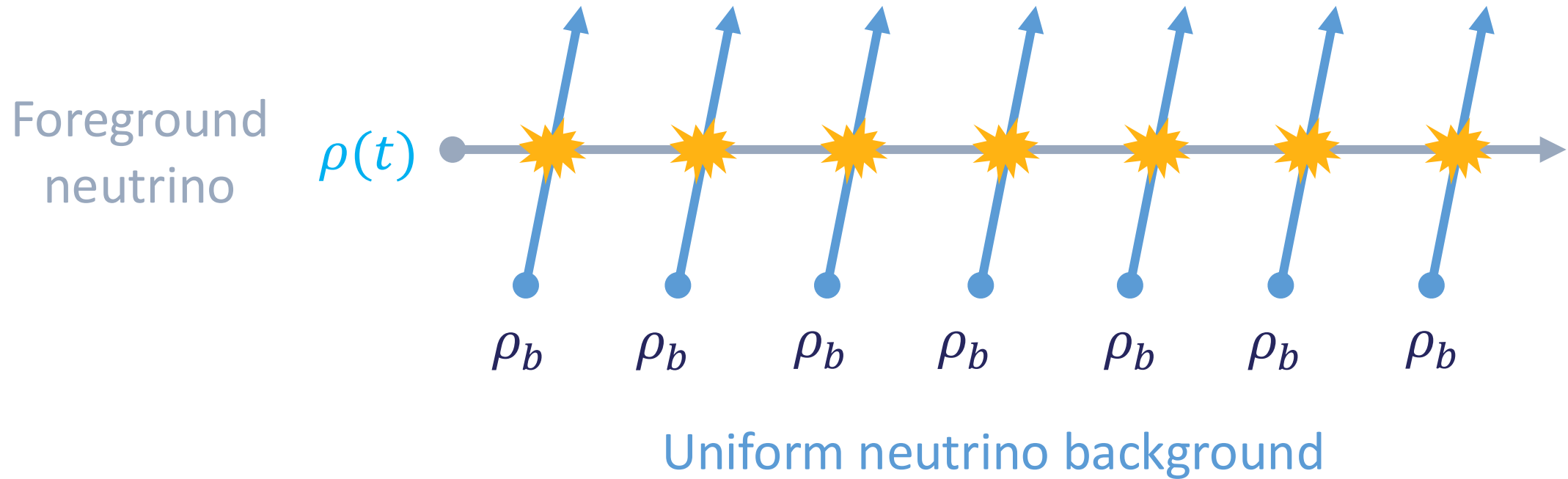


Discreteness parameter \rightarrow cross section?

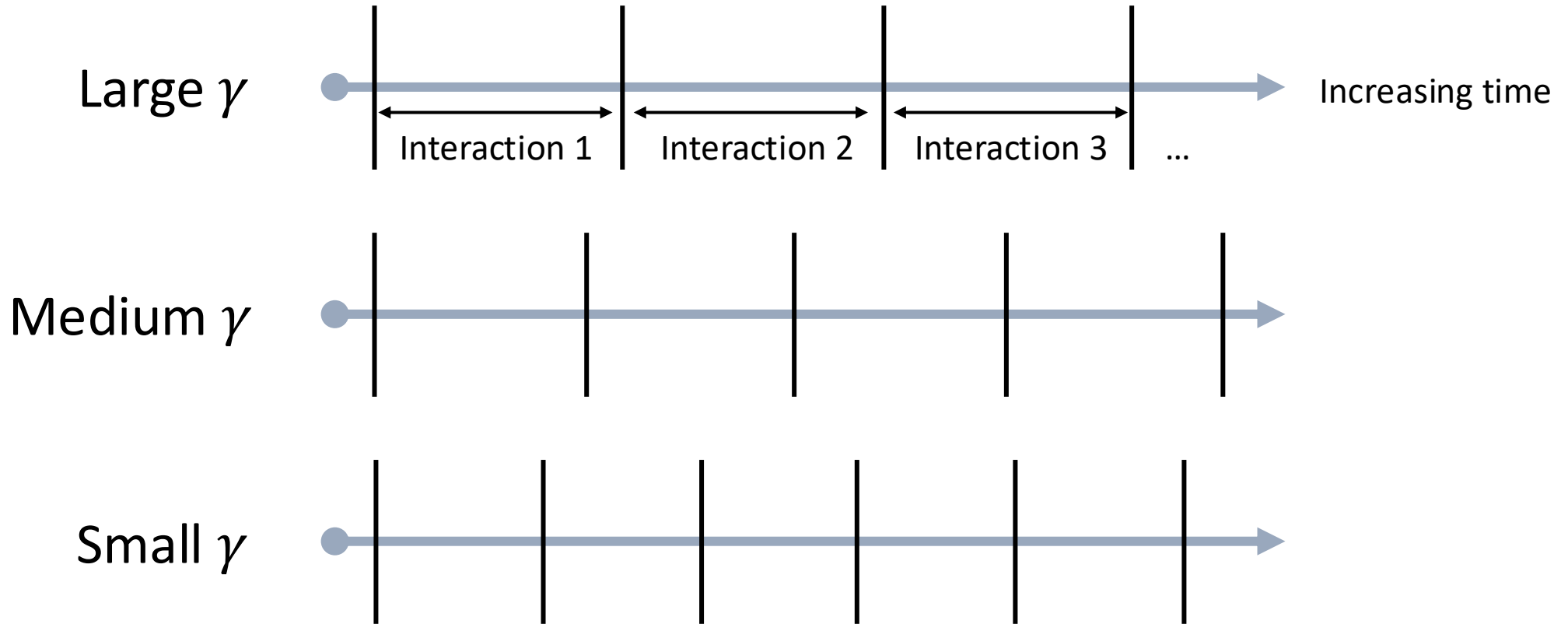
$$\underbrace{\gamma}_{\text{discreteness parameter}} \equiv \mu t^{\text{int}} = \frac{\sqrt{2}G_F}{A \cdot L^{\text{wavepacket}}} t^{\text{int}} = \frac{\sqrt{2}G_F}{\mathbf{A}} \underbrace{\quad}_{\text{cross section}}$$

(using $L^{\text{wavepacket}} = t^{\text{int}}$)

A simple once-in-a-lifetime encounter model



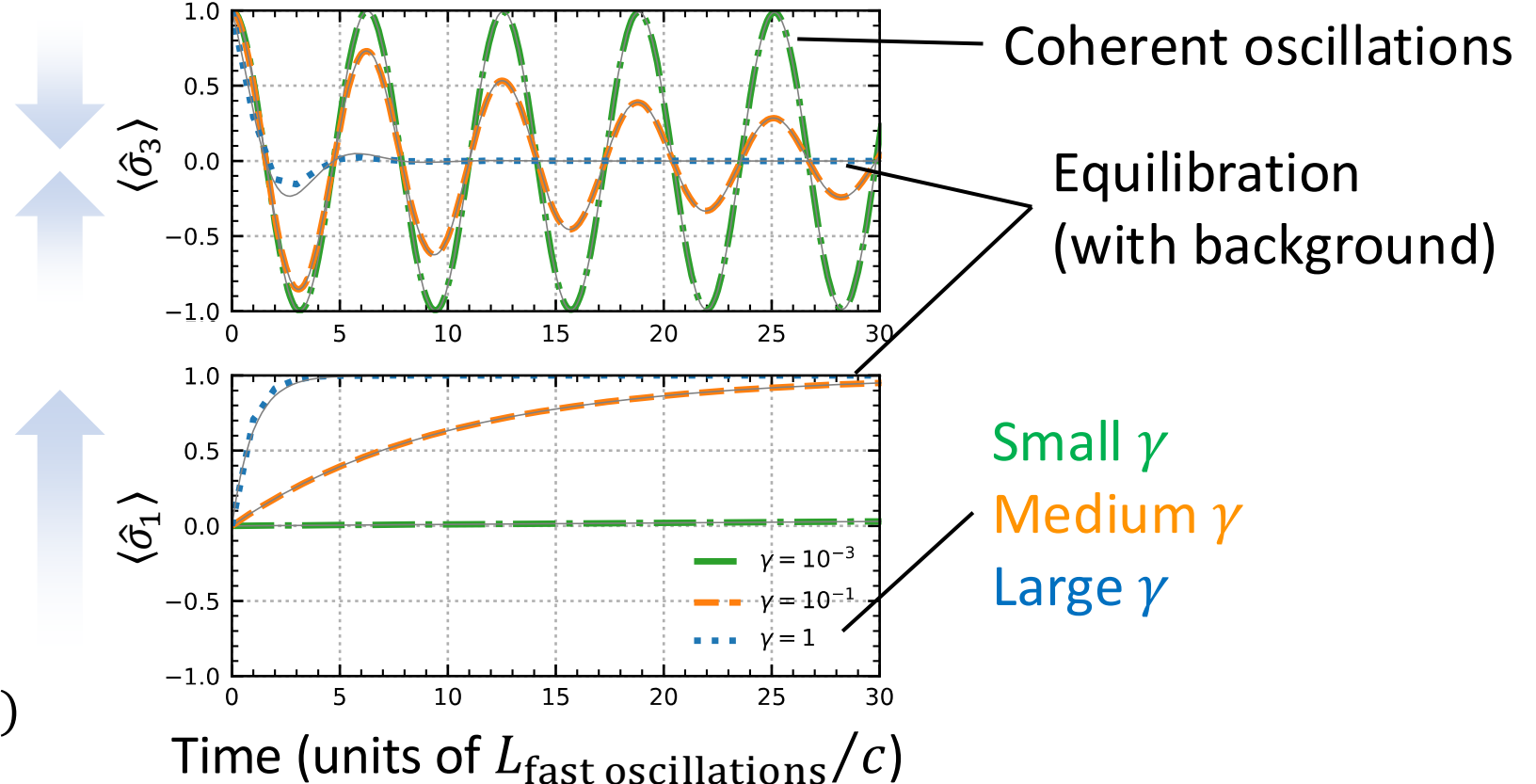
A (new) parameter γ



A numerical experiment with OILE (Result: Rate of entanglement $\propto \gamma$)

Increasing entanglement

Increasing entanglement



$$\dot{\mathbf{P}} = \underbrace{\mu \mathbf{P}_{bg}}_{\text{mean-field}} \times \underbrace{\mathbf{P} + \mu \gamma (\mathbf{P}_{bg} - \mathbf{P})}_{\text{equilibrate to } \mathbf{P}_{bg}}$$

Two limits of once-in-a-lifetime encounter models

Mean-field
approach

Many-body
models



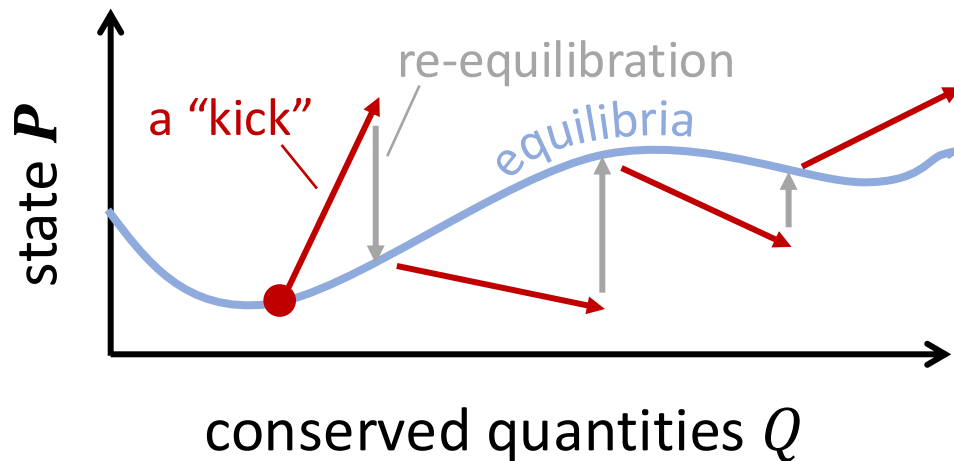
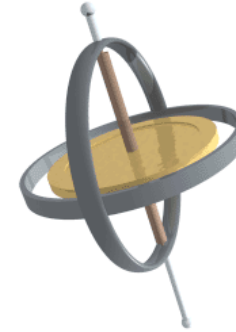
Miscidynamics (a flash talk)

Another perturbative expansion

- Separation of timescales:
 - Possible in OILE models, but not plane wave models
- $$\mu\gamma \text{ (entanglement rate)} \ll \mu \text{ (mean-field rate)}$$
- First order in γ : The adiabatic approximation
 - Like in thermodynamics or quantum mechanics
 - Evolution along a **path of** different mean-field **equilibria**

What path is taken?

- Mean-field limit:
equilibria parameterized by conserved quantities Q
- A series of small “kicks” away from equilibrium
 - Each “kick” changes Q
 - System quickly re-equilibrates with the new Q

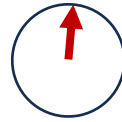


$$\rightarrow \frac{d\mathbf{P}}{dt} = \frac{d\mathbf{P}_{\text{eq}}(Q)}{dQ} \cdot \frac{dQ}{dt}$$

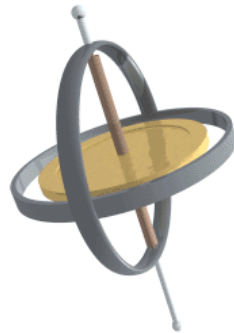
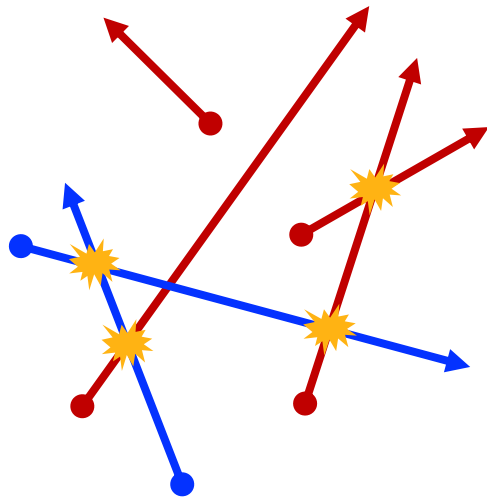
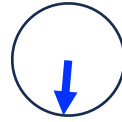
given by terms involving $\boldsymbol{\gamma}$

OILE and miscidynamics agree

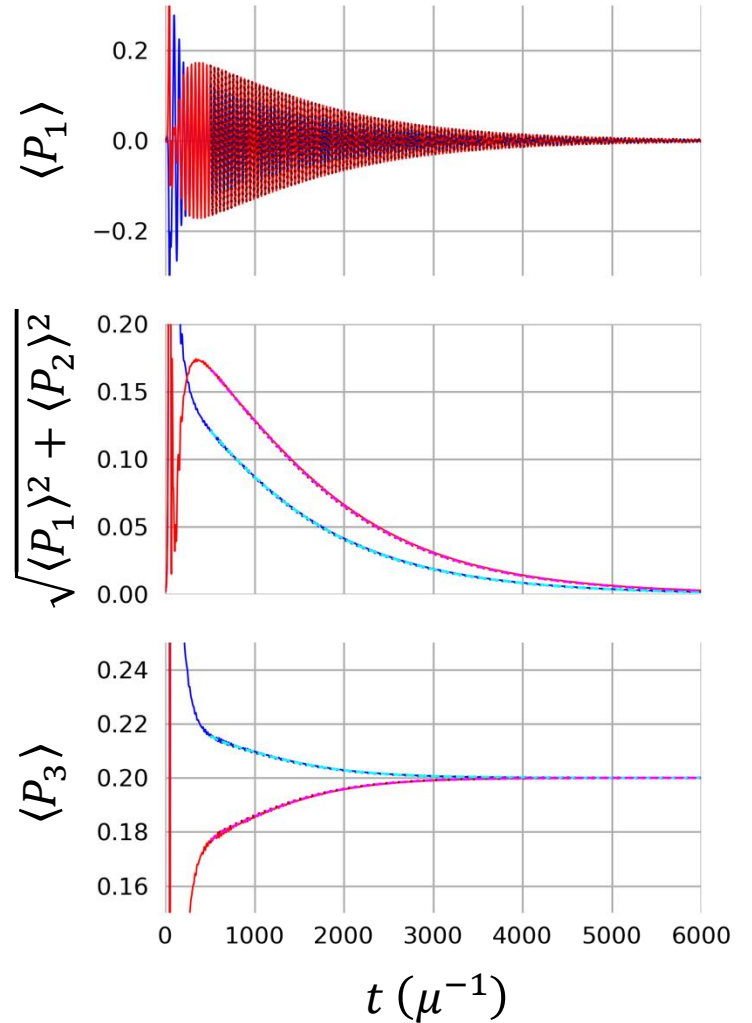
Group 1
60 neutrinos, initially



Group 2
40 neutrinos, initially



Group 1 & 2, \mathbf{P} vs. t
Miscidynamics: dashed curves, 1 & 2



Adiabatic flavor evolution

Phenomenon	Adiabatic evolution of neutrinos driven by:
The MSW effect	Changing matter density
Spectral splits	Changing neutrino density
OILE/Miscidynamics	“Decoherence” through entanglement

Conclusions

- The standard approach to collective neutrino oscillations, the mean-field approach, ignores entanglement
- Once-in-a-lifetime encounter models show that entanglement can depend on the finite size of neutrino wavepackets
 - Small wavepacket size: mean-field approximation might be okay!
- Miscidynamics “factors out” small-scale dynamics by tracking large-scale changes in equilibria
 - Similar to the flavomon approach
 - Preliminary success with once-in-a-lifetime encounter models as a testing ground