

# Kalman-Based Reconstruction in the EIC ZDC



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# Motivations

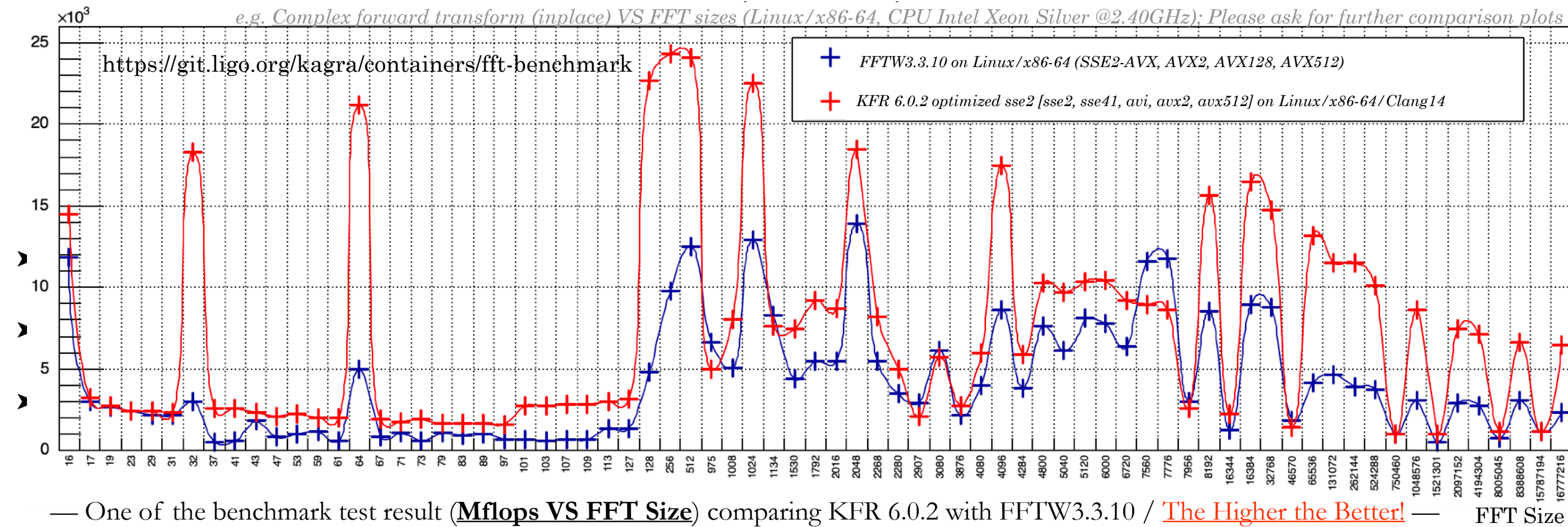
- **Goal:** reconstruct neutral particles (neutron, photons) in the ZDC
  - Physics case:  $\Lambda^0 \rightarrow n + \pi^0 \rightarrow n + \gamma\gamma$
  - Assess detector performances, prior to further tuning of the geometry
- **Challenge:**
  - only calorimetric information (position along x,y and energy deposits)
  - possible multi-particle overlap
  - noise possibly underestimated

## My group's interest:

- Explore and elaborate AI technics based on our past expertise either in pure AI (or Astrophysics)
- Streaming readout & DAQ (Echelon 0-1) for event filtering; ZDC detector for study of meson structure

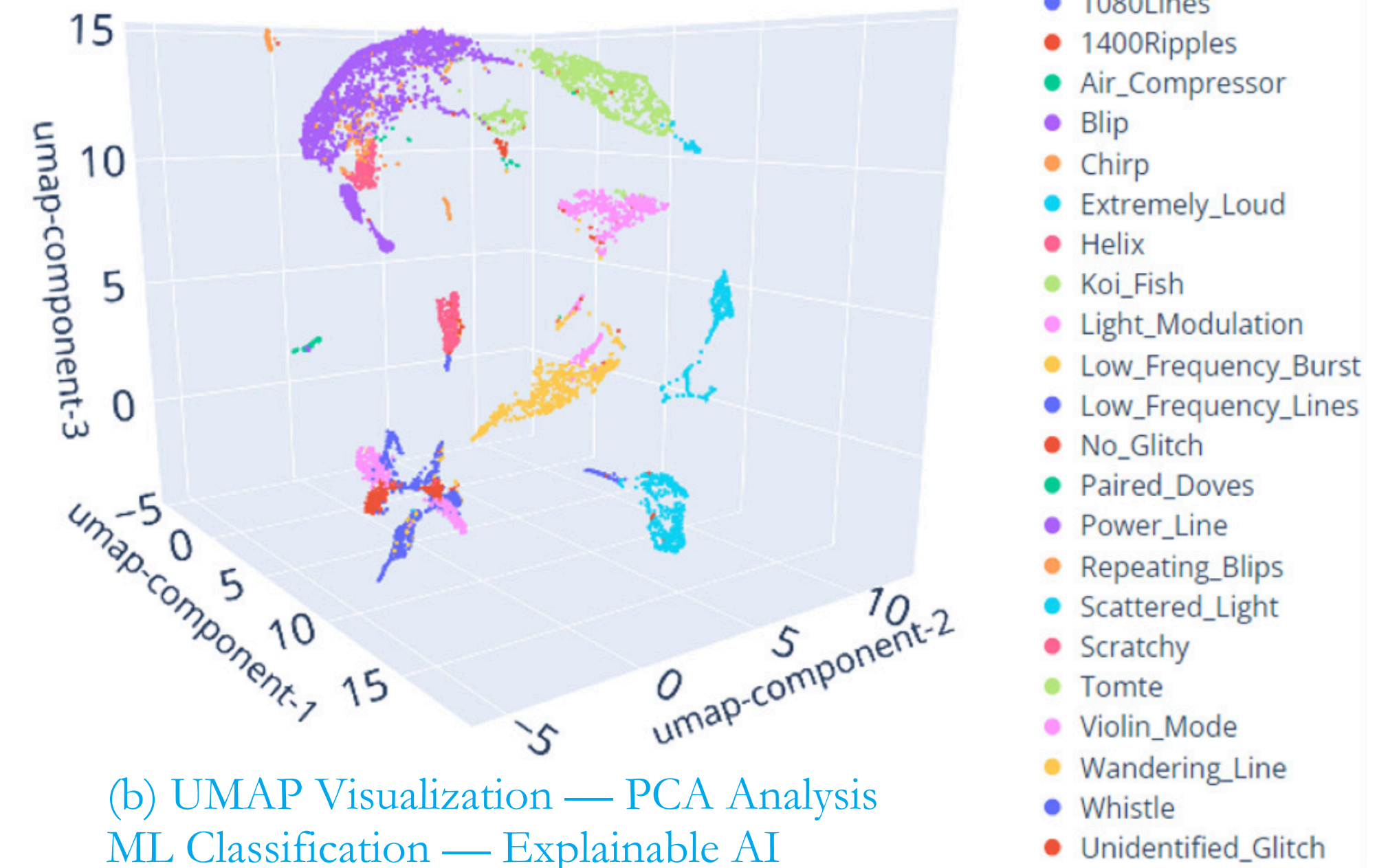
# Introduction

Tokyo City University: "Data Science" Laboratory



(a) Optimization (in red) & Benchmarking of Signal Transformation (FFT) — Extending Python/C++ ROOT Framework toward Multidimensional Tensors / TimeSeries Analysis at CERN  
arXiv:2503.14292v1 [astro-ph.IM] - ACAT2024 Invited Talk.

## Pattern Recognition in Vector Space



(b) UMAP Visualization — PCA Analysis  
ML Classification — Explainable AI

<https://nature.com/articles/s41598-022-13329-4> [Scientific Reports]

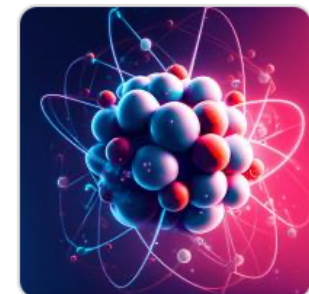
(d) First Gravitational Wave Extraction using ROOT Experimental Signal Processing (e) + **Kalman Filtering** <https://link.aps.org/doi/10.1103/jpr4-m3ck>



### conda-forge

<https://github.com/hep-packaging-coordination>  
dedicated HEP initiative aiming to compile all sort of arch. and machines (powerpc, aarch64, window, macos etc.)

\* (no link with <http://hepforge.org>)



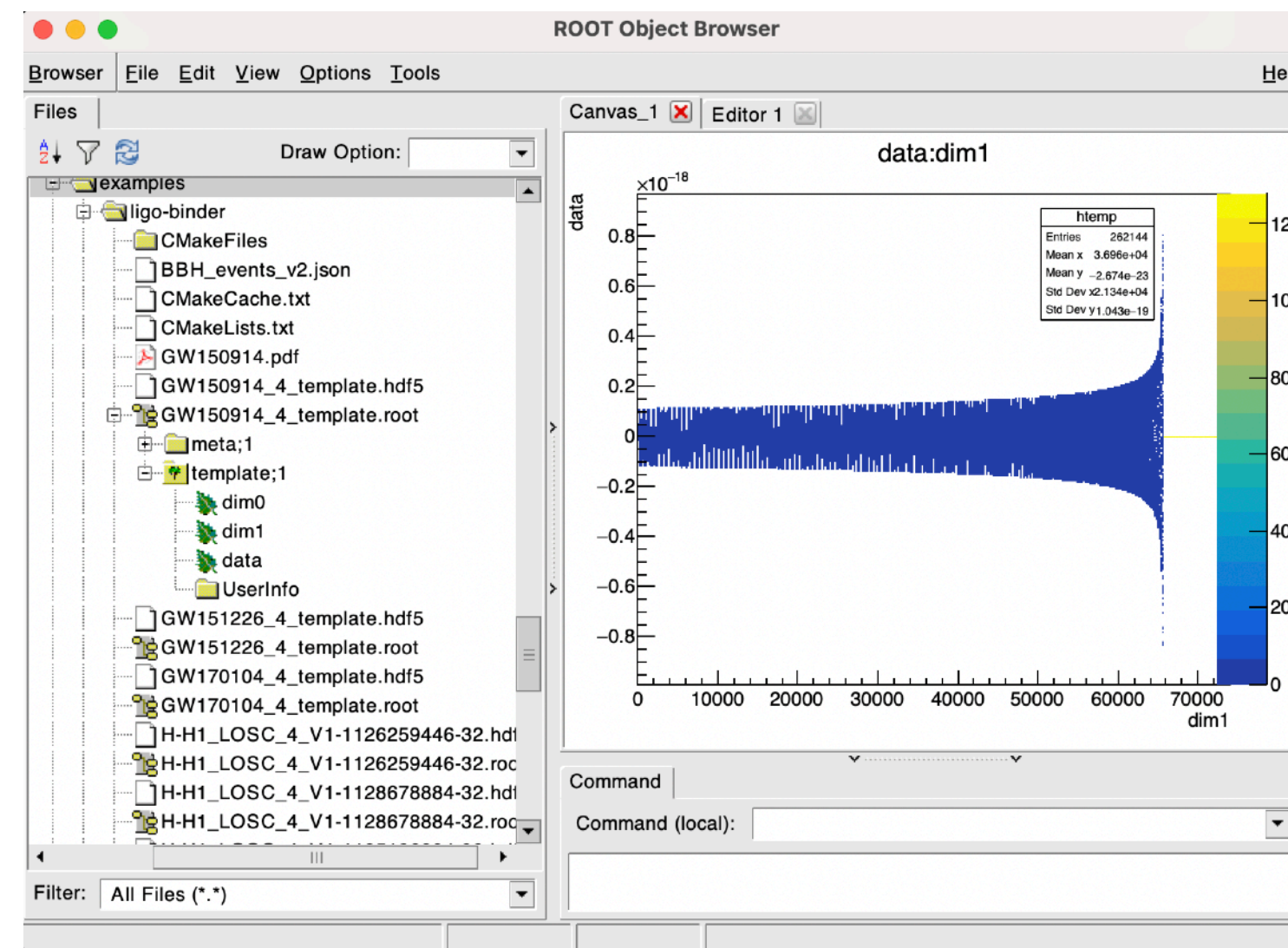
### HEP Forge

<https://github.com/hep-forge>  
<http://anaconda.org/hep-forge>

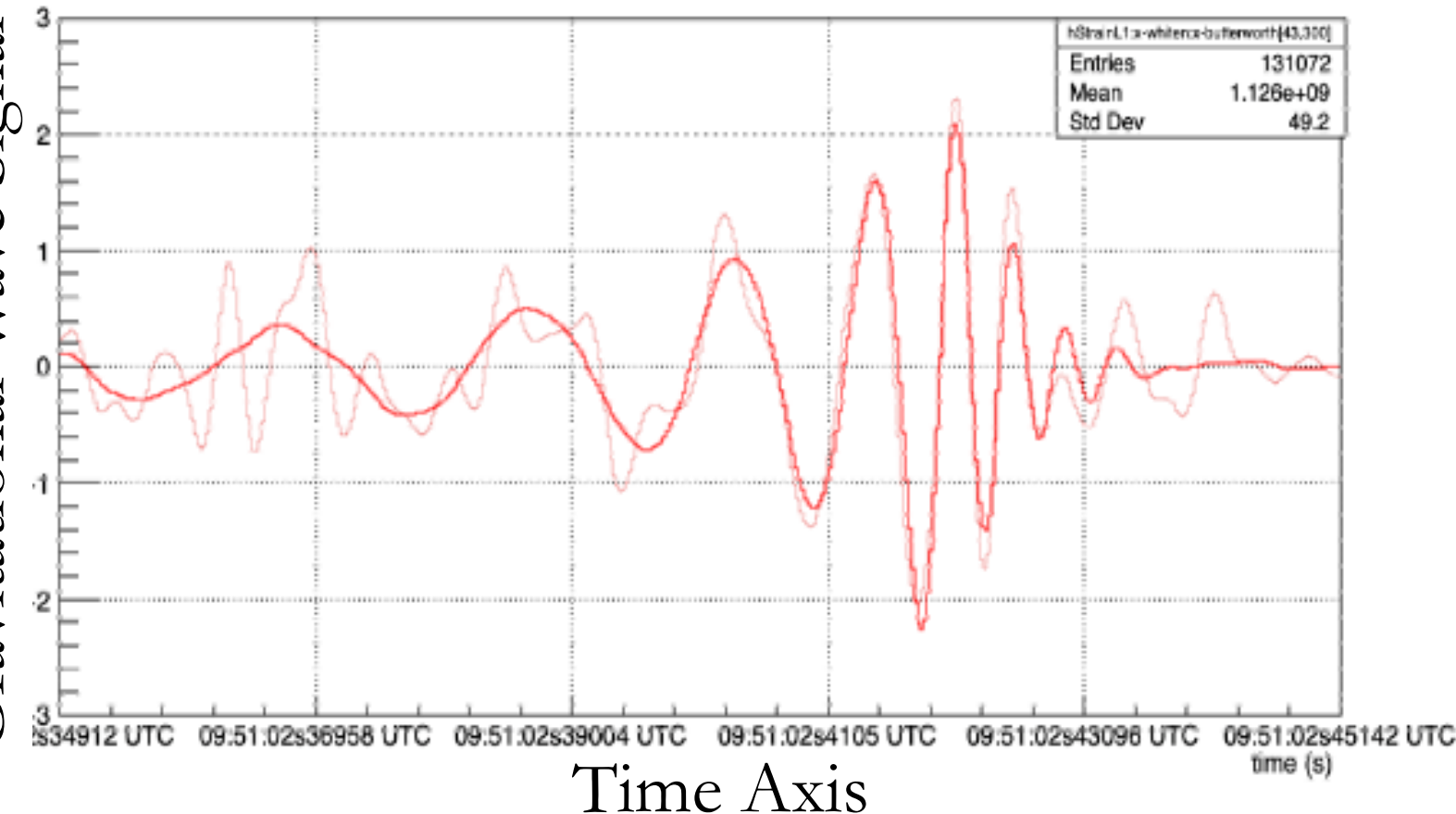
repository with *no strict requirements* to publish (linux only, AMD64/ARM64)



(c) Coordination & Maintenance of the HEP Packaging Software on Conda Repositories (pre-compiled software installation & legacy support)

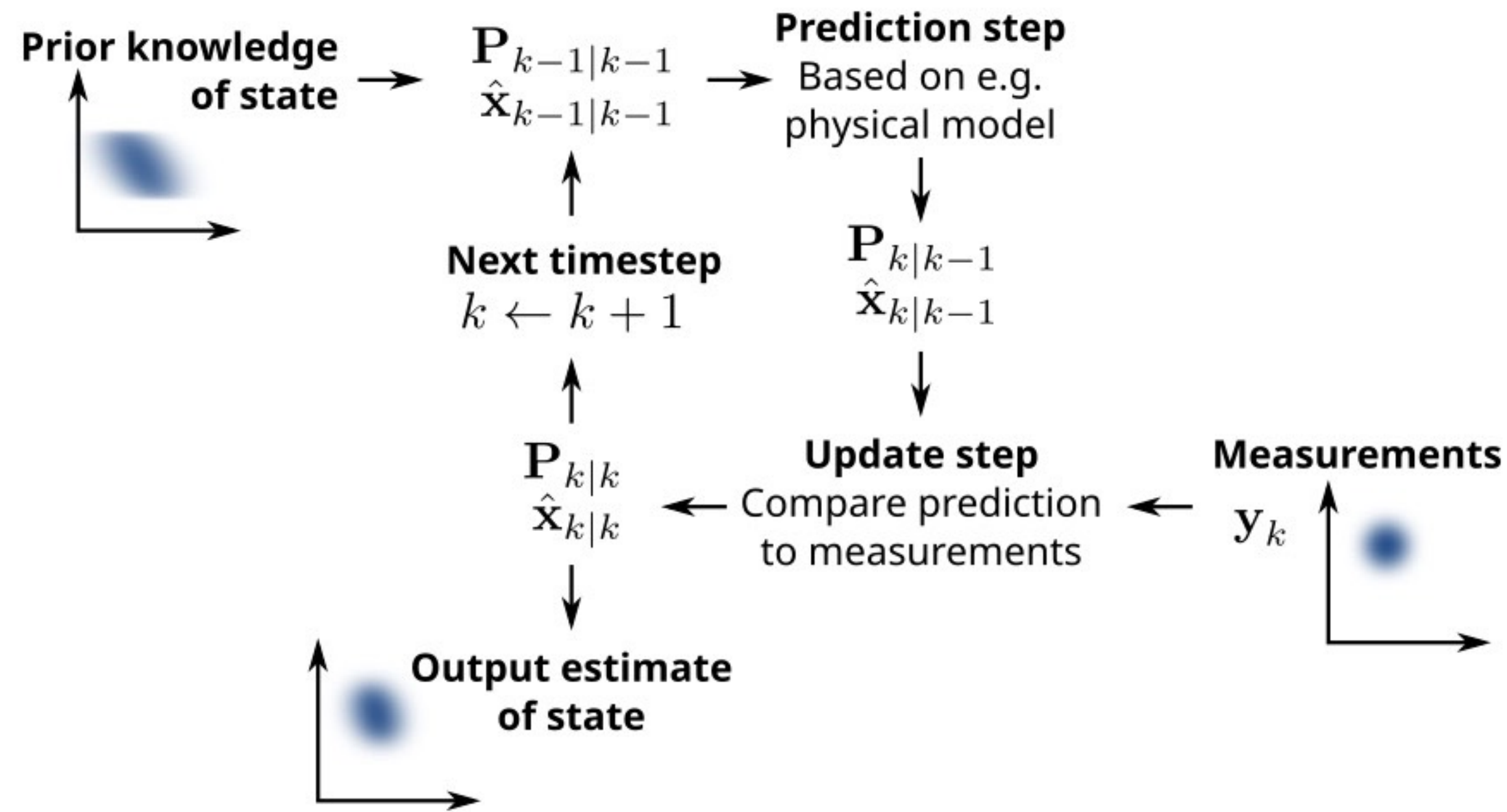


Gravitational Wave Signal



# Traditional Kalman example

Ballistic model



Developed by Rudolf Kalman in 1960.

First use in Apollo guidance computer.

Equations of Kalman

$$\hat{\mathbf{x}}_{k|k-1} = F_{k-1} \hat{\mathbf{x}}_{k-1|k-1},$$

$$P_{k|k-1} = F_{k-1} P_{k-1|k-1} F_{k-1}^T + Q_{k-1}$$

$$\mathbf{w}_k \sim \mathcal{N}(\mathbf{0}, \mathbf{Q})$$

(process noise)

$$\mathbf{v}_k \sim \mathcal{N}(\mathbf{0}, \mathbf{R})$$

(Measurement noise)

From State Space  $\longrightarrow$  To Observing (Measurement) Space

$$\mathbf{x}_{k+1} = F_k \mathbf{x}_k + \mathbf{w}_k$$

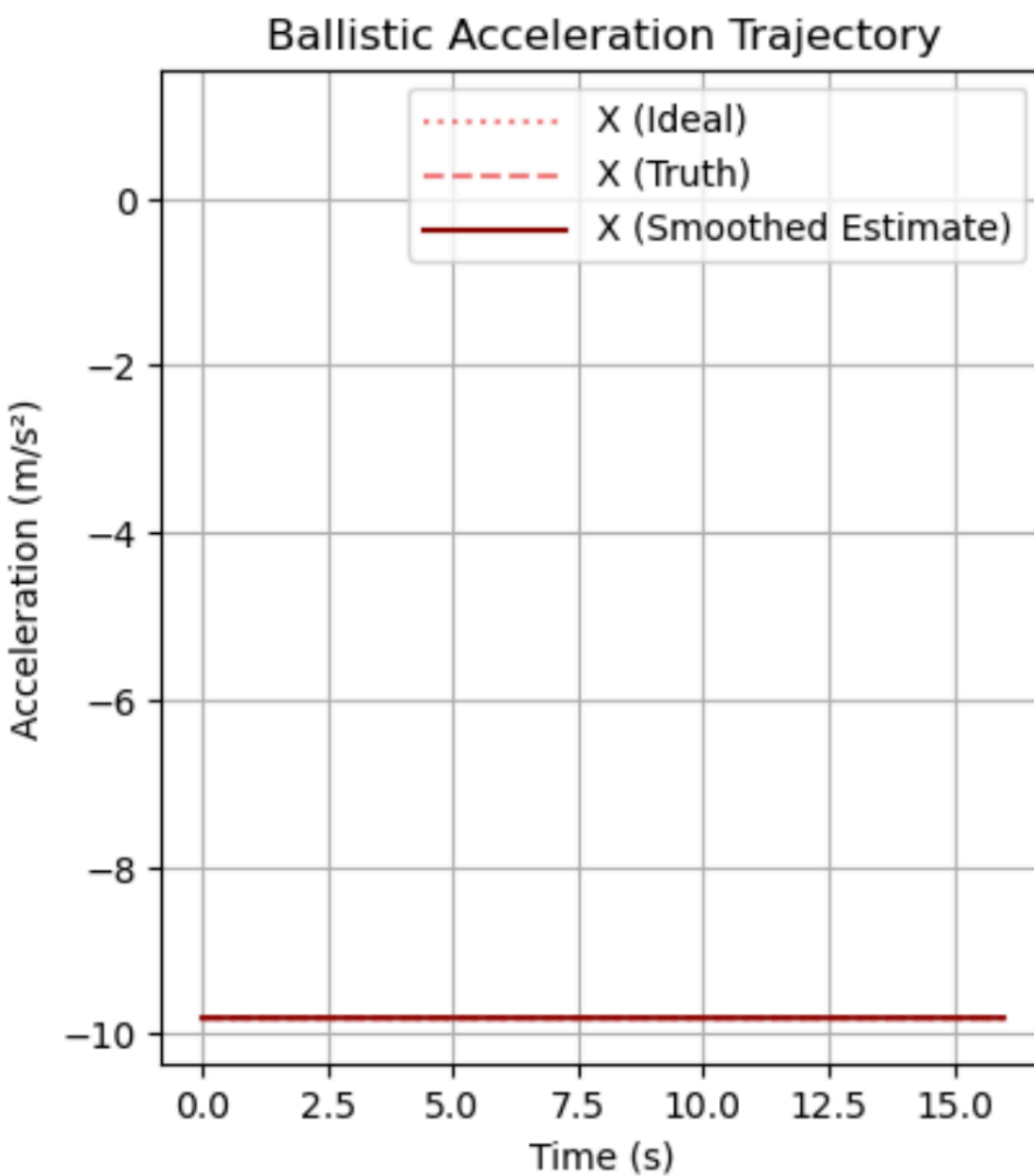
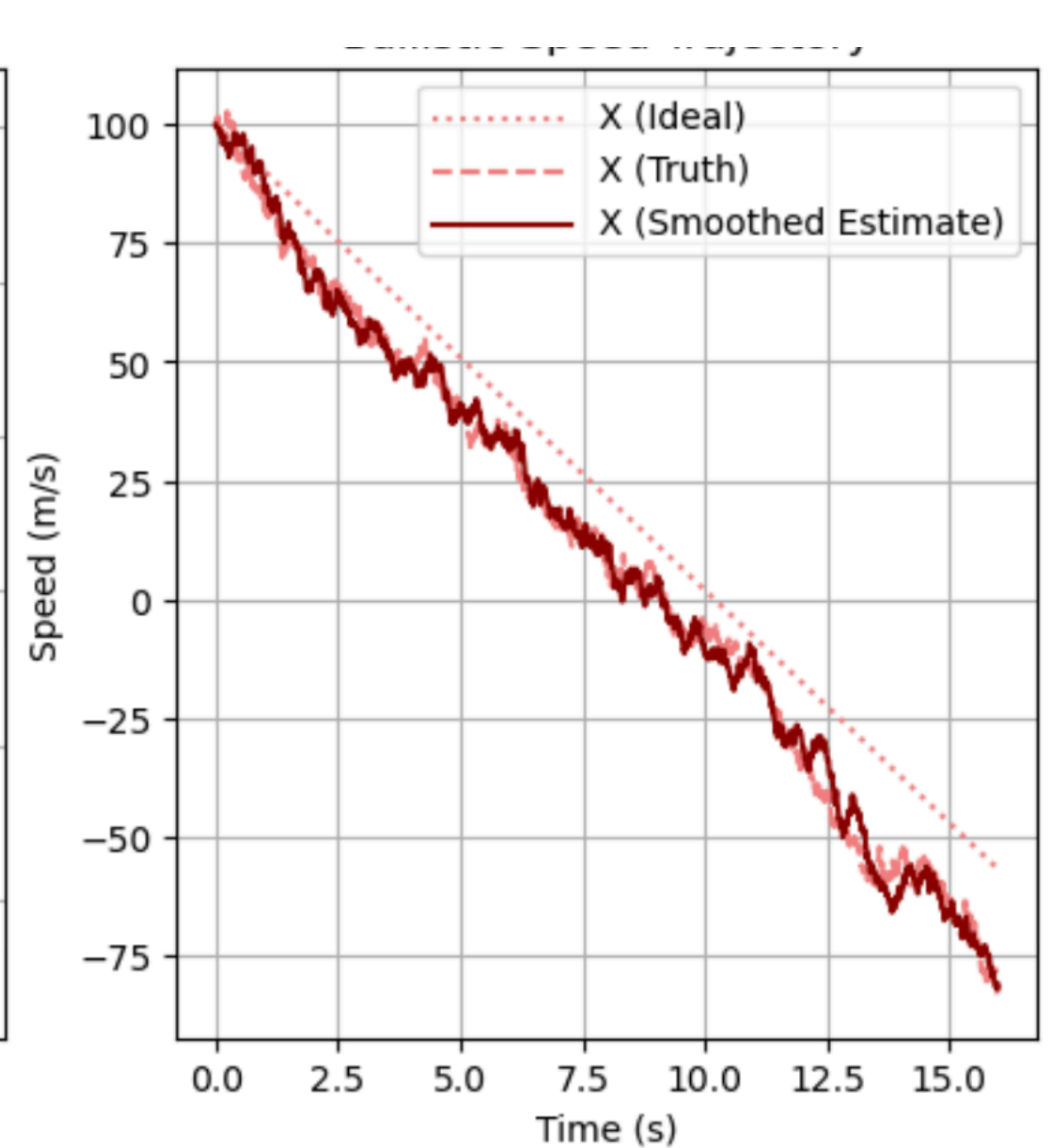
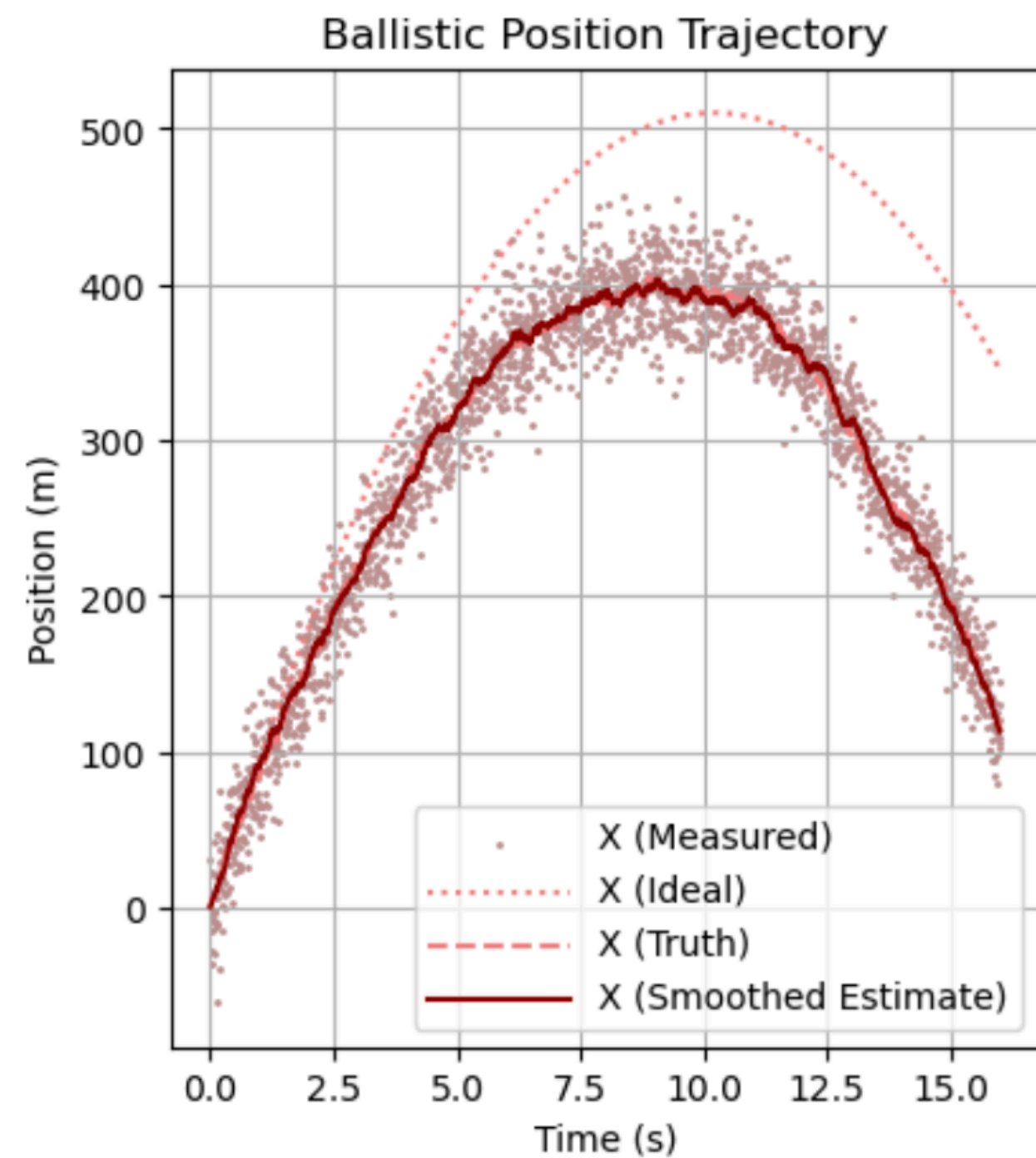
$$\mathbf{z}_k = H \mathbf{x}_k + \mathbf{v}_k,$$

# Traditional Ballistic Model

- Simple model for practical introduction of the concept (constant acceleration)

$$x(t) = \frac{1}{2}gt^2 + v_0t + x_0 \longrightarrow \mathbf{x} = [z, \dot{z}, \ddot{z}]^T \text{ (state vector)}$$

$$F = \begin{pmatrix} 1 & \Delta t & 0 \\ 0 & 1 & \Delta t \\ 0 & 0 & 1 \end{pmatrix}, \quad H = (1 \ 0 \ 0) \longrightarrow \mathbf{z} \text{ (measurement vector)}$$



- Noise model ? Process noise VS Measurement Noise. (See next slide)

# Transition from Ballistic Model to ZDC

By analogy

	<b>Projectile model (toy)</b>	<b>ZDC EIC</b>
System dynamics	Ballistic trajectory	Track position / angle
(Physics induced-) Process noise	Friction, wind	EM shower, stochastic fluctuation
Measurement noise	GPS or Camera uncertainty	Detector resolution (per slice); position and energy deposit

# Transition to ZDC Reconstruction

## Why Kalman Here?

- **Detector Characteristics**

- layered calorimeter  $\rightarrow$  natural discretization along  $z$
- 4D information available:  $(x, y, "z", E)$

- **It naturally handles**

- sequential layers (ZDC sampling): transition matrix  $(k \rightarrow k+1)$
- process noise (fluctuating showers)
- partial information

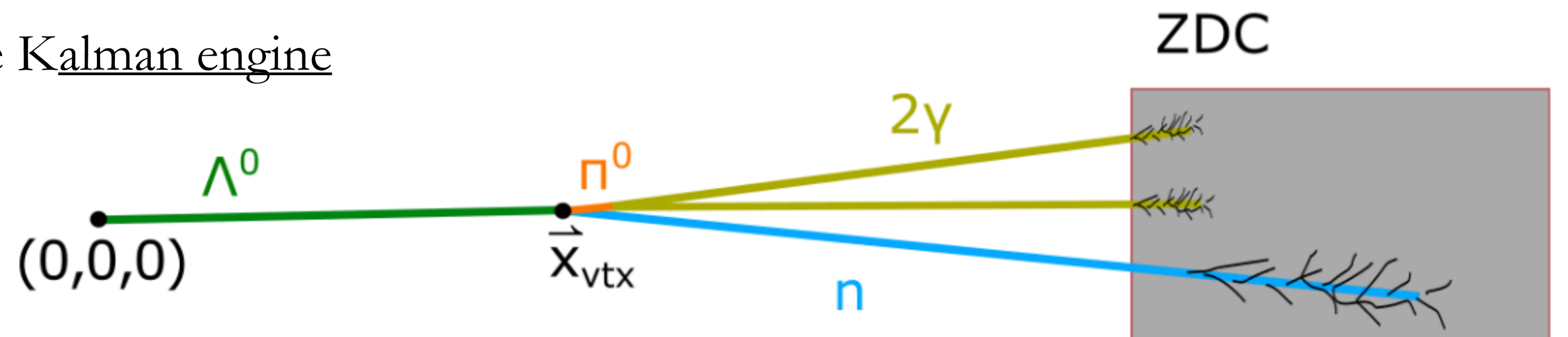
- One Kalman component per track, but one Kalman engine

$$\mathbf{A} = \text{diag}(\mathbf{A}_1, \dots, \mathbf{A}_N),$$

- **Track reconstruction engine:**

- Linear Kalman Filter
- Extended Kalman Filter (non linear) for including energy deposit (later)
- Unscented Kalman Filter for non-linear noise distribution.

- Rauch-Tung-Striebel smoother for taking into account a backward pass on the measurement



(Taken from: SJ Paul *et al.*)

# Kalman Filter per Component

One Component = One Candidate Track

$$\mathbf{w}_k \sim \mathcal{N}(\mathbf{0}, \mathbf{Q})$$

(process noise)

$$\mathbf{v}_k \sim \mathcal{N}(\mathbf{0}, \mathbf{R})$$

(Measurement noise)

- **State Definition:**

$$\mathbf{x}_k = (x_k, y_k, t_{x,k}, t_{y,k})^T, \quad \mathbf{x}_{k+1} = F_k \mathbf{x}_k + \mathbf{w}_k, \quad F_k = \begin{pmatrix} 1 & 0 & \Delta z_k & 0 \\ 0 & 1 & 0 & \Delta z_k \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}, \quad \mathbf{z}_k = H \mathbf{x}_k + \mathbf{v}_k, \quad H = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \end{pmatrix}$$

- **Measurement** : traditional clustering, possibility for machine learning (Kalman is still flexible)

$$\mathbf{z}_k = \begin{pmatrix} x_k \\ y_k \end{pmatrix}, \quad R_k^{(0)} = \begin{pmatrix} \sigma_{x,k}^2 & \sigma_{xy,k} \\ \sigma_{xy,k} & \sigma_{y,k}^2 \end{pmatrix}, \quad \text{with} \quad x_k = \frac{\sum_{i \in \mathcal{H}_j^{(k)}} E_i x_i}{\sum_{i \in \mathcal{H}_j^{(k)}} E_i}, \quad y_k = \frac{\sum_{i \in \mathcal{H}_j^{(k)}} E_i y_i}{\sum_{i \in \mathcal{H}_j^{(k)}} E_i}$$

$+\Delta E_k$

- **Noise Modeling**

- Process noise (Q) accounts for shower fluctuations & scattering

Q to be determined with CWNA model by calculating  $t_x(z_k + \Delta z_k)$

R. Frühwirth, “App. of Kalman filtering to track and vertex fitting”, Nucl. Instrument. Meth. A 262, 444 (1987)

$$\frac{dt_x}{dz} = \eta_x(z), \quad \langle \eta_x(z) \eta_x(z') \rangle = \sigma_t^2 \delta(z - z') \quad \text{(stochastic diffusion: } \sigma_t^2 \sim \text{strength diffusion of the shower axis; typically including Molière radius+radiation length for spread of the EM \& hadronic showers)}$$

- Measurement noise (R) depends on cluster spread, detector granularity & energy deposition

Detector characteristics  $\sim$  position & energy resolutions

# Vertex Reconstruction for assessing precision

Performance and detector reconstruction

Backward extrapolation of reconstructed tracks to find common intersection along  $z$

## Method:

- Pairwise closest approach from gamma and neutron
- $\chi^2$  minimization
- Weighted intersection using uncertainties from Kalman filter ( $P_k$ )

## Resolution considerations and advantages of Kalman

- Limited by angular resolution, shower fluctuations, detector segmentation
- Assess detector performance on the fly  
comparing generated vs reconstructed information
- Compute pull distribution for each track, using covariance matrix  $P_k$ :  
any deviation  $\sim$  badly conditioned (R,Q) noise



# Summary

**Kalman filtering is a swift and adaptable approach for describing a system's dynamic from differential equations.**

1. **3D Clustering:** Separate particles are created.
2. **Layer Integration:** Measurements are incorporated into the layers.
3. **One Kalman per Component:** Trajectories are calculated for each component.
4. **Backward Extrapolation:** Vertex positions are determined.
5. **Combine Particles:** The particles are combined to obtain physics observables.

## **Extensions:**

- **Incorporating Energy E:** Energy is included in the state vector  $x_k$ .
- **Detailed Model Study:** A comprehensive study of the electromagnetic and neutron showers is conducted.
- **Testing Systematics:** Systematics are tested using the resulting covariance matrix,  $\chi^2$ , and by changing the detector design on-the-fly.
- **Dynamic energy reconstruction:** implementing Longo–Sestili model for describing more precisely the energy deposit VS. nuclear interaction length.

## **Extension (for fun):**

- **Neural Networks for Clustering:** Neural networks are used for clustering.
- **Comparison with Traditional Graph Neural Networks (GNN)**