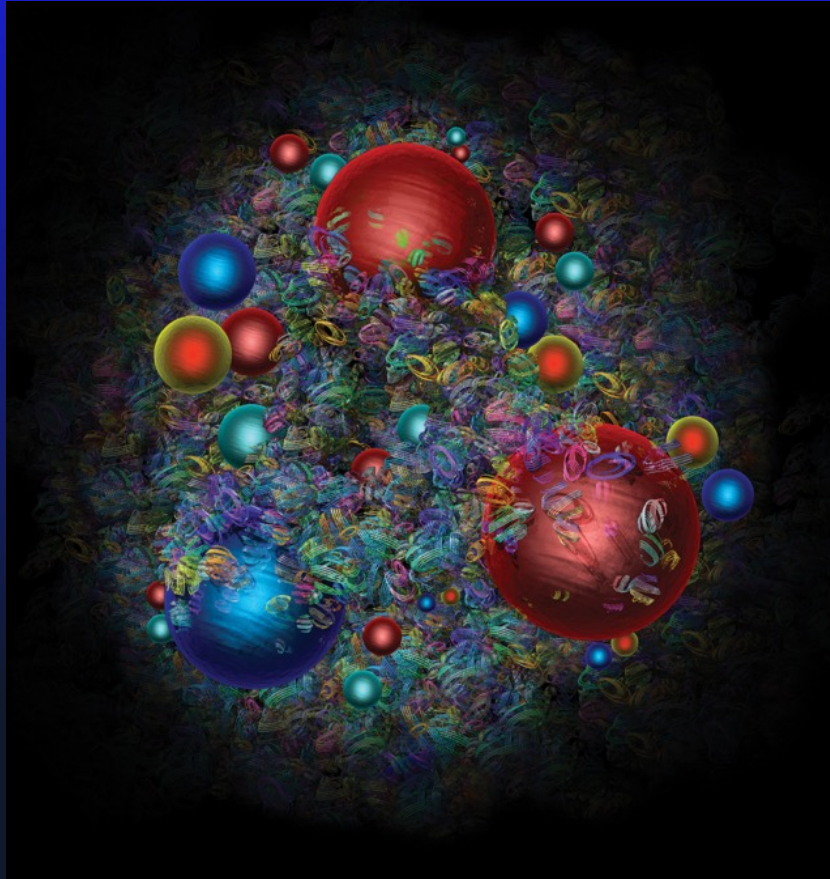


The Electron-Ion Collider: the ultimate electron microscope



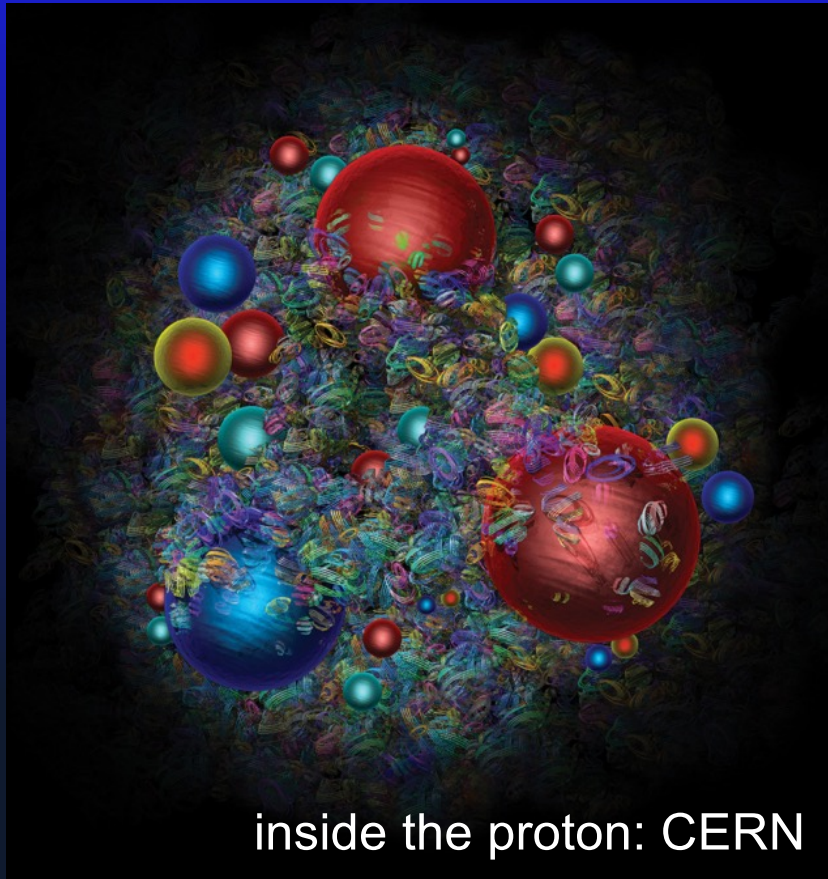
Gordon Baym
University of Illinois
Urbana, Illinois



TIDC 2nd EIC Meeting
Institute of Physics, Academia Sinica
January 3, 2023



The Electron-Ion Collider: the ultimate electron microscope

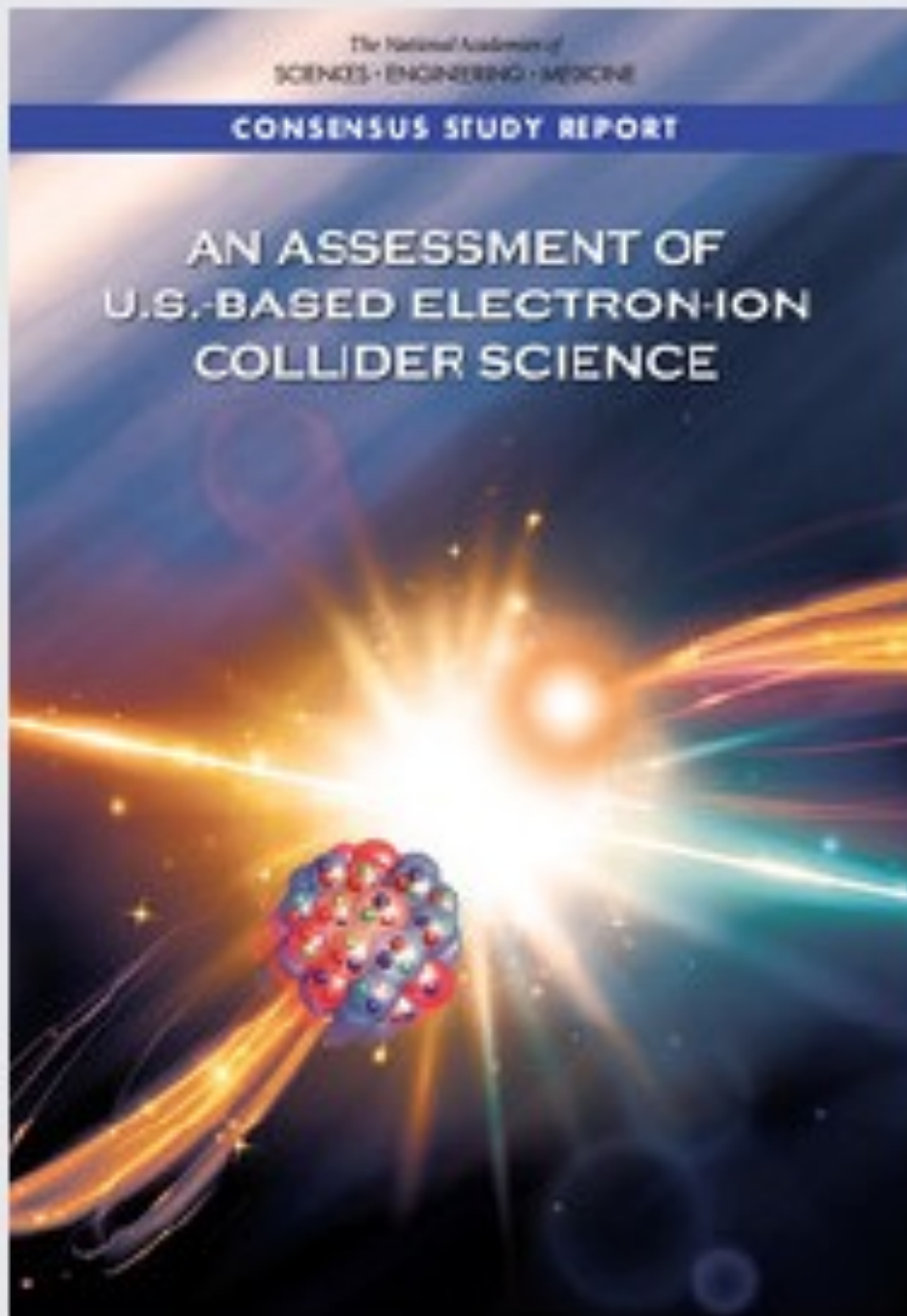


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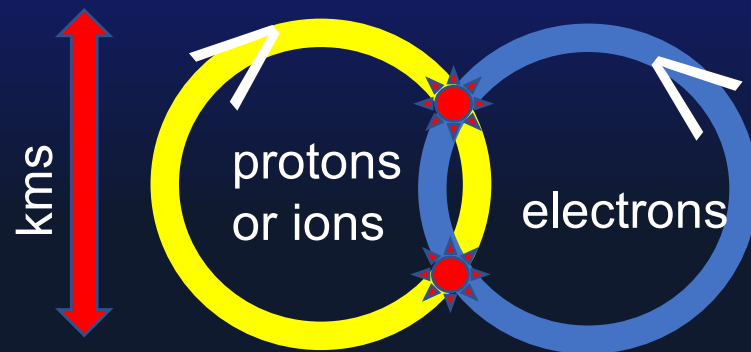
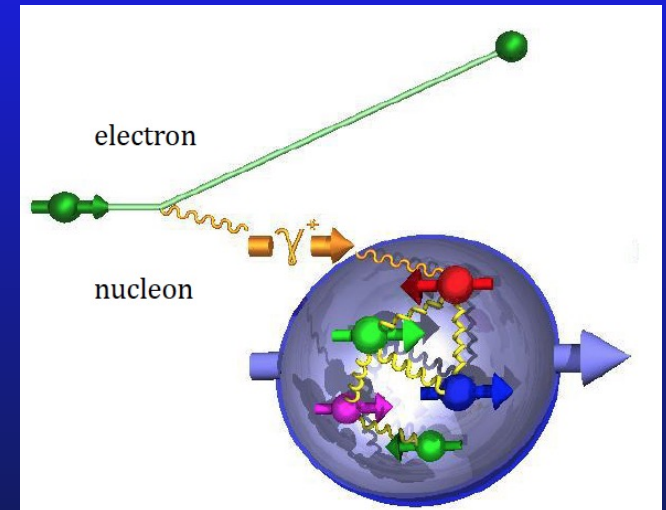
National Academy of Sciences
– National Research Council
report, July 2018, on the
science of an EIC

[Download pdf from
https://www.nap.edu/catalog/25171](https://www.nap.edu/catalog/25171)

The Electron-Ion Collider

A very big accelerator -- colliding beams of electrons with beams of protons or heavier ions (atomic nuclei).

A giant electron microscope for peering at the quarks and gluons deep inside the nucleon, as well as atomic nuclei. **QCD machine.**



Electron-ion center of mass energy:

$$\sqrt{s} \sim 28 \sim 140 \text{ GeV.}$$

High luminosity (event rate) and spin polarized beams!

*Electron
microscope
Invented 1931*



ca. 1940

Atomic nuclei: building blocks of the everyday world:



1

H

hydrogen

[1.007, 1.009]

2

He

helium

4.003

3

Li

lithium

[6.938, 6.997]

4

Be

beryllium

9.012

11

Na

sodium

22.99

12

Mg

magnesium

[24.30, 24.31]

19

K

potassium

39.10

20

Ca

calcium

40.08

21

Sc

scandium

44.96

22

Ti

titanium

47.87

23

V

vanadium

50.94

24

Cr

chromium

52.00

25

Mn

manganese

54.94

26

Fe

iron

55.85

27

Co

cobalt

58.93

28

Ni

nickel

58.69

29

Cu

copper

63.55

30

Zn

zinc

65.38(2)

31

Ga

gallium

69.72

32

Ge

germanium

72.63

33

As

arsenic

74.92

34

Se

selenium

78.96(3)

35

Br

bromine

[79.90, 79.91]

36

Kr

krypton

83.80

37

Rb

rubidium

85.47

38

Sr

strontium

87.62

39

Y

yttrium

88.91

40

Zr

zirconium

91.22

41

Nb

niobium

92.91

42

Mo

molybdenum

95.96(2)

43

Tc

technetium

44

Ru

ruthenium

101.1

45

Rh

rhodium

102.9

46

Pd

palladium

106.4

47

Ag

silver

107.9

48

Cd

cadmium

112.4

49

In

indium

114.8

50

Sn

tin

118.7

51

Sb

antimony

121.8

52

Te

tellurium

127.6

53

I

iodine

126.9

54

Xe

xenon

131.3

55

Cs

caesium

132.9

56

Ba

barium

137.3

57-71

lanthanoids

72

Hf

hafnium

178.5

73

Ta

tantalum

180.9

74

W

tungsten

183.8

75

Re

rhenium

186.2

76

Os

osmium

190.2

77

Ir

iridium

192.2

78

Pt

platinum

195.1

79

Au

gold

197.0

80

Hg

mercury

200.6

81

Tl

thallium

[204.3, 204.4]

82

Pb

lead

207.2

83

Bi

bismuth

209.0

84

Po

polonium

85

At

astatine

86

Rn

radon

87

Fr

francium

88

Ra

radium

89-103

actinoids

104

Rf

rutherfordium

105

Db

dubnium

106

Sg

seaborgium

107

Bh

bohrium

108

Hs

hassium

109

Mt

meitnerium

110

Ds

darmstadtium

111

Rg

roentgenium

112

Cn

copernicium

113

Nh

nihonium

114

Fl

flerovium

115

Mc

moscovium

116

Lv

livermorium

117

Ts

tennessine

118

Og

oganesson

57

La

lanthanum

138.9

58

Ce

cerium

140.1

59

Pr

praseodymium

140.9

60

Nd

neodymium

144.2

61

Pm

promethium

62

Sm

samarium

150.4

63

Eu

europium

152.0

64

Gd

gadolinium

157.3

65

Tb

terbium

158.9

66

Dy

dysprosium

162.5

67

Ho

holmium

164.9

68

Er

erbium

167.3

69

Tm

thulium

168.9

70

Yb

ytterbium

173.1

71

Lu

lutetium

175.0

89

Ac

actinium

227.0

90

Th

thorium

232.0

91

Pa

protactinium

231.0

92

U

uranium

238.0

93

Np

neptunium

94

Pu

plutonium

95

Am

americium

96

Cm

curium

97

Bk

berkelium

98

Cf

californium

99

Es

einsteinium

100

Fm

fermium

101

Md

mendelevium

102

No

nobelium

103

Lr

lawrencium

atomic number

Symbol

name

standard atomic weight

13

B

boron

[10.80, 10.83]

14

C

carbon

[12.00, 12.02]

15

N

nitrogen

[14.00, 14.01]

16

O

oxygen

[15.99, 16.00]

17

F

fluorine

19.00

18

Ne

neon

20.18

13

Al

aluminium

26.98

14

Si

silicon

[28.08, 28.09]

15

P

phosphorus

30.97

16

S

sulfur

[32.05, 32.08]

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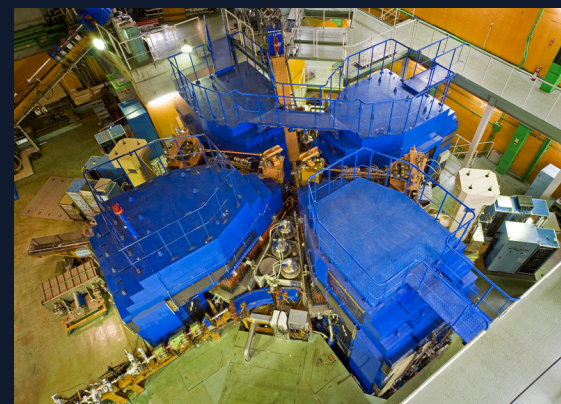
phosphorus

30.97

The periodic table of chemical elements is over 150 years old.
Are there further elements out there?
Any of them stable?

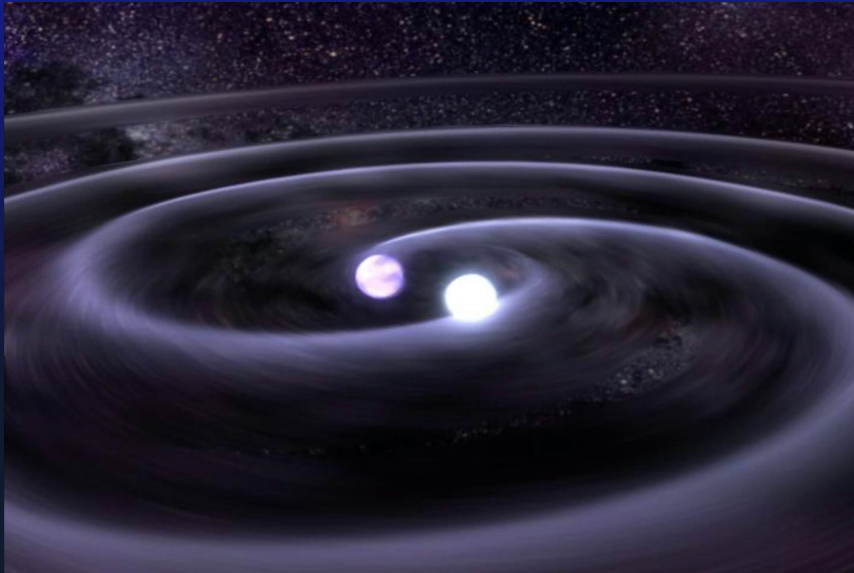
How many isotopes does each element have?
Answers from rare isotope accelerators (e.g., RIBF at RIKEN in Japan and FRIB at MSU) studying nuclei far from stability.

F = facility, RIB = rare isotope beam



RIKEN

How to make the heavier elements? Answers, remarkably, from multimessenger studies of binary neutron star mergers:
Merger GW170817 observed on 17 Aug. 2017 by LIGO and Virgo (gravitational radiation), FERMI (gamma ray telescope) + some 70 other electromagnetic observatories.



Two neutron stars merging,
emitting gravitational radiation
and post-merger forming:

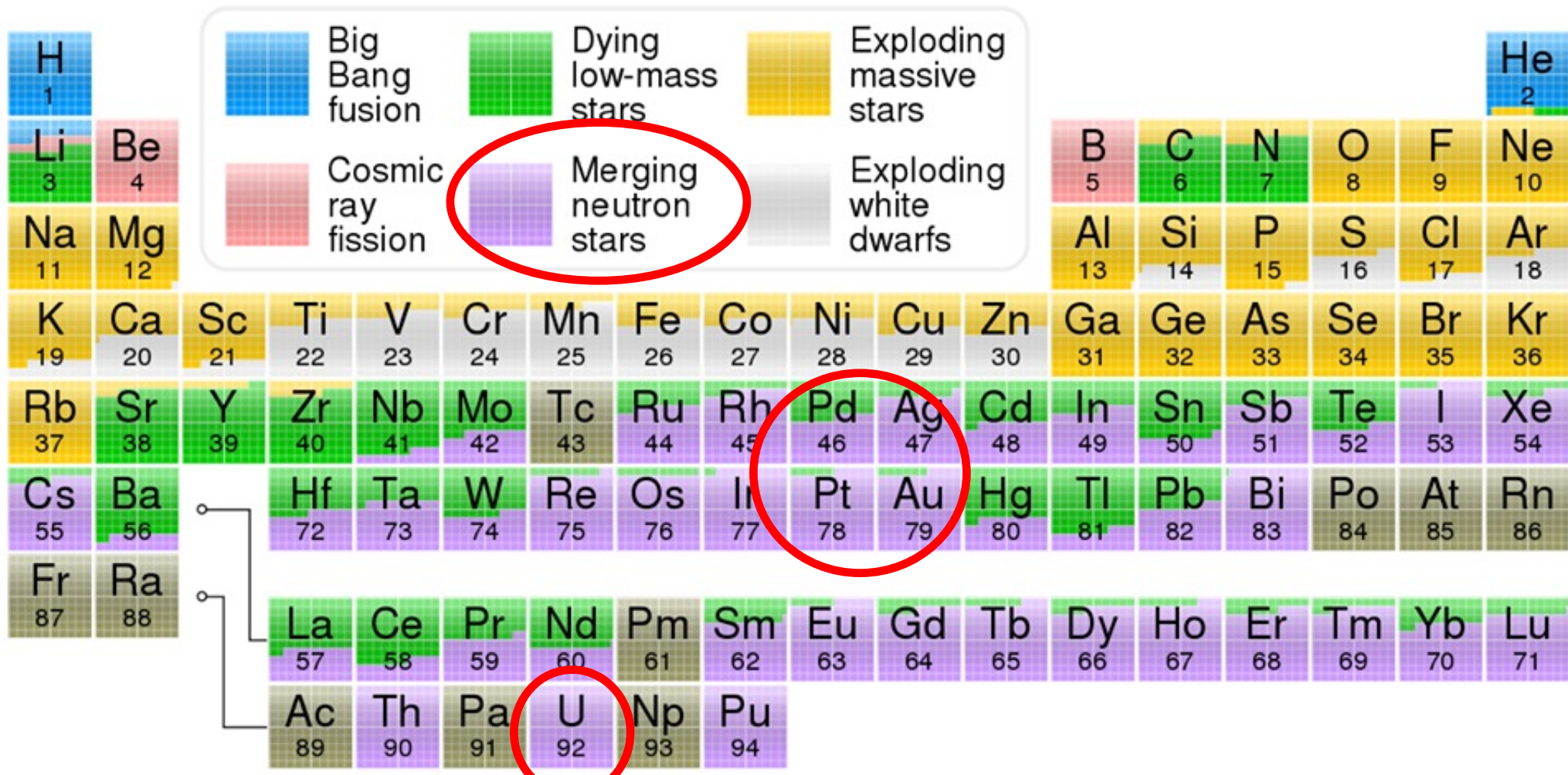
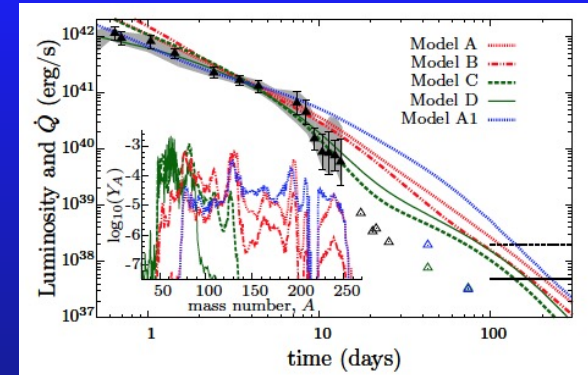


**Kilonova: neutron-rich
site of r-process**

Binary neutron star mergers likely site of heavy element production

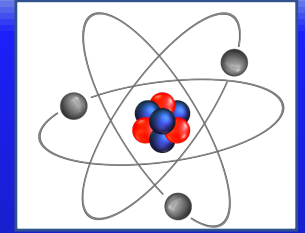
r-process in *kilonova* =>
earth-scale masses of
Au, Ag, Pt, U... and **Sr**

Kilonova
N= 50, 82, 126



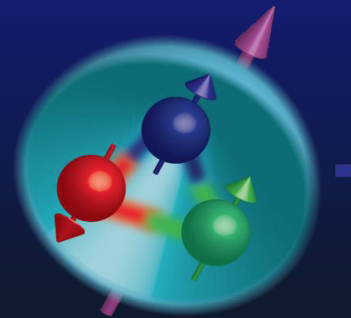
Atoms are made of electrons and nuclei.

Similarly neutrons, protons, and nuclei are made of **quarks and gluons**. But how? How to explain nucleon masses, spin, magnetic moments, etc. in terms of quarks and gluons?



Quarks = fractionally charged spin-1/2 fermions, baryon no. = 1/3, with internal SU(3) **color** degree of freedom.

Flavor	Charge/ e	Mass(MeV)
u	2/3	~2
d	-1/3	~5
s	-1/3	~ 94
c	2/3	~1280
b	-1/3	~4200
t	2/3	~175,000



proton = $u + u + d$

neutron = $u + d + d$

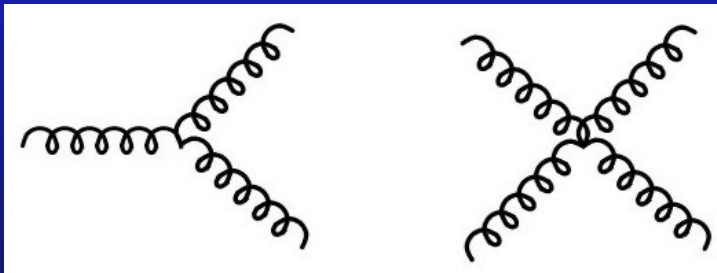
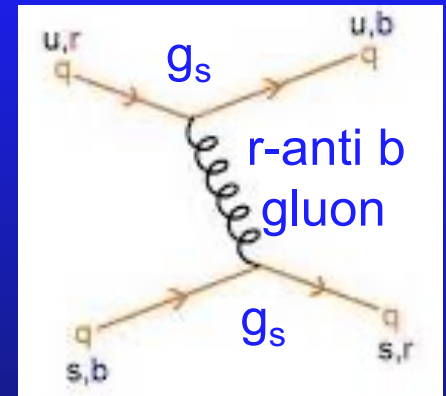
$\pi^+ = u + \bar{d}$, etc.



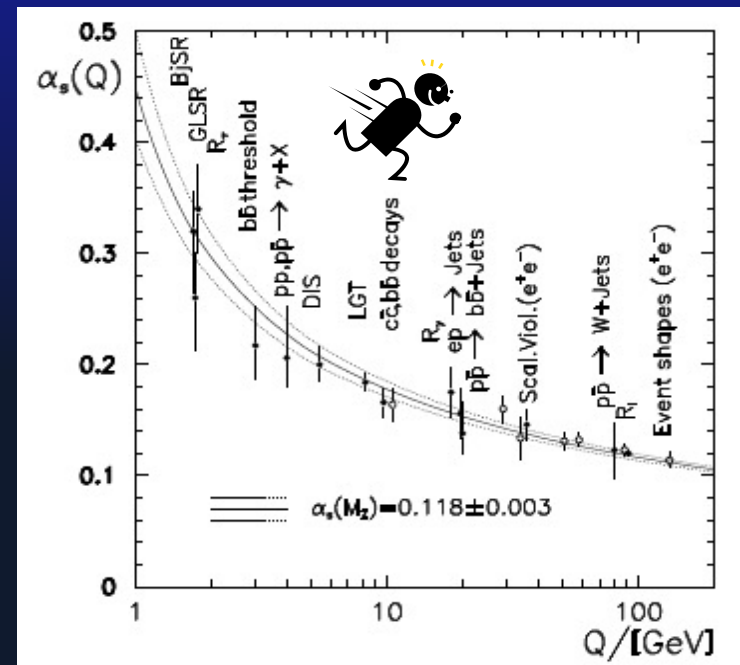
Form of baryons in the early universe at $t < 1$ microsec ($T > 100$ MeV) and in the deep interiors of heavier neutron stars.

Strong interactions – quantum chromodynamics

Quarks interact by exchanging gluons – massless vector bosons (like photon) with spin 1, and coming in 8 colors. **Gluons also interact with each other!!**



QCD



Running coupling constant

$$\alpha_s(\mu) = \frac{g_s^2}{4\pi} = \frac{6\pi}{(33 - 2N_f) \ln(\mu/\Lambda_{QCD})}$$

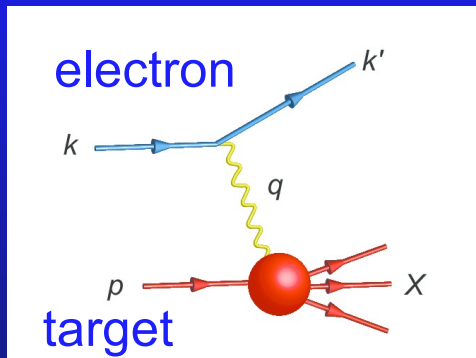
μ = energy scale $\Lambda_{QCD} \sim 340$ MeV

Asymptotic freedom as $\mu \rightarrow \infty$

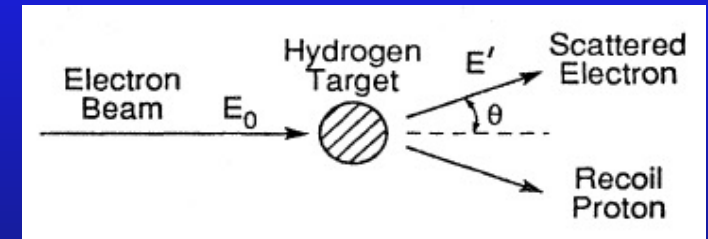
(Even at Grand Unified (GUT)
scale, 10^{15} GeV, $g_s \sim 1/2$;
cf. electrodynamics:

$$e^2/4\pi = 1/137 \Rightarrow e \sim 1/3)$$

Electron scattering on nucleons and nuclei: origins



Electrons (without internal structure) are precise probe of the complex structure of nucleons and nuclei.



First scattering of electrons on nuclei, Illinois Betatron 1951

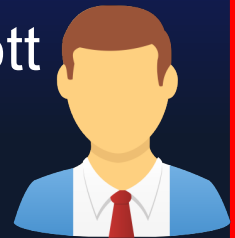
Al Hanson,



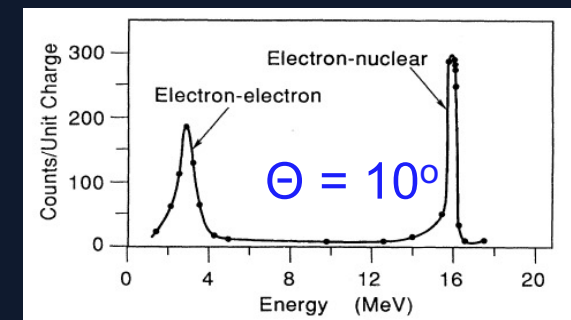
Ernie Lyman



and Merrill B. Scott

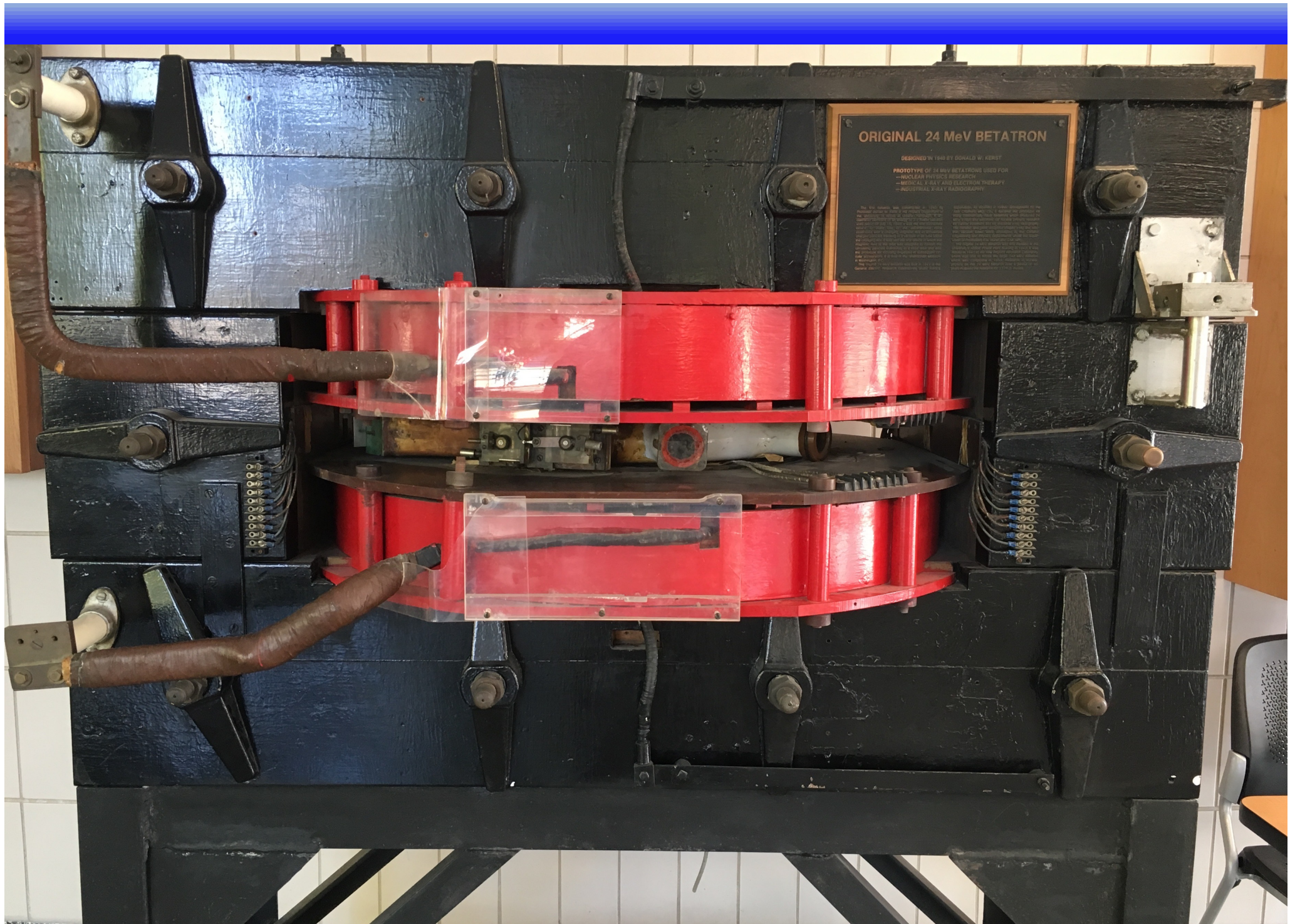


15.7 MeV electron beam on Be and Au foils





Donald Kerst, U of I, with the Betatron, ca. 1941



ORIGINAL 24 MeV BETATRON

DESIGNED IN 1940 BY DONALD W. KERST
PROTOTYPE OF 34 MeV BETATRON USED FOR
—NUCLEAR PHYSICS RESEARCH
—MEDICAL X-RAY AND ELECTRON THERAPY
—INDUSTRIAL X-RAY RADIOGRAPHY

The first machine was constructed in 1940 by Kerst and his team at the University of California, Berkeley. It was the first of a series of betatrons designed by Kerst for the production of high-energy X-rays for medical and industrial use. The machine was designed to produce X-rays with a maximum energy of 24 MeV. It was the first of a series of betatrons designed by Kerst for the production of high-energy X-rays for medical and industrial use.

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First scattering of electrons on protons and alpha particles

R. W. McAllister and Robert Hofstadter, 1956, with 187 MeV electrons at Stanford High Energy Physics Lab (HEPL) on H_2 and He gaseous targets, and then on polyacetylene (CH_2).

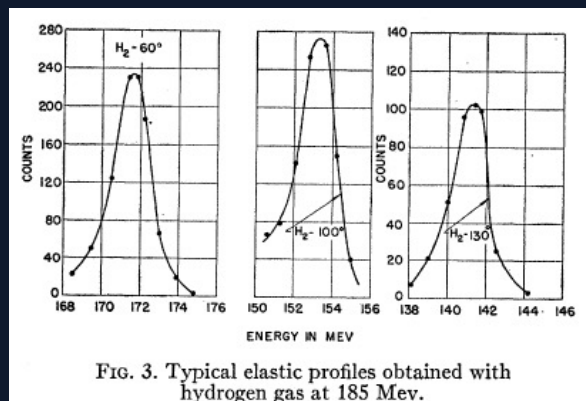


Found first hint of internal structure of the proton from the angular distribution:

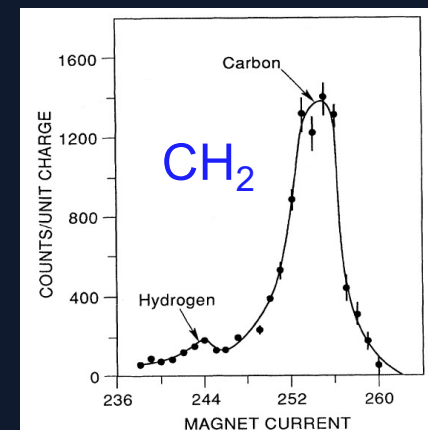
proton radius $r_{\text{proton}} \sim 0.7 \text{ fm}$

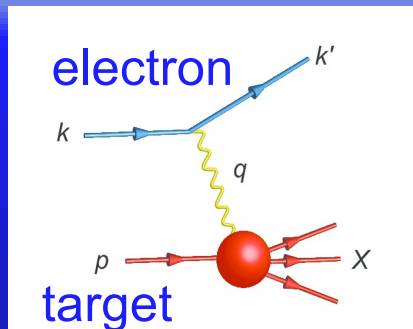
($1 \text{ fm} = 10^{-13} \text{ cm}$)

But is the inside of the proton a continuous “pudding” or are its constituents point particles?

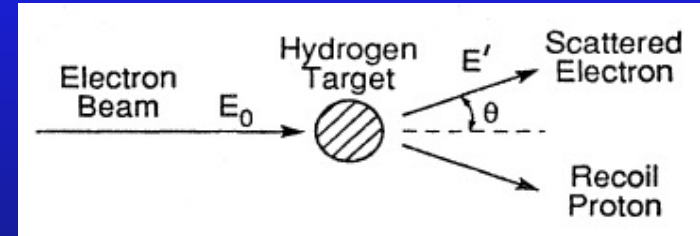


Elastic scattering on H_2 at various Θ





Kinematic variables in electron scattering from nucleons and nuclei



Electron scattering angle: θ

Electron energy transfer **in lab**: $\nu = E_e - E_e'$

3-momentum transfer from electron: \mathbf{q}

4-momentum transfer squared: $Q^2 = \mathbf{q}^2 - \nu^2$
Larger Q^2 = higher (transverse) resolution

Elastic scattering on target of mass m :

Energy conservation $(m^2 + q^2)^{1/2} = m + \nu \Rightarrow \nu = Q^2/2m$

Bjorken scaling variable: $x = Q^2/(2m_{\text{proton}} \nu)$
“shutter speed” $\sim 1/x$

The variables Q^2 and x define the landscape of electron scattering

Discovery of quarks 1967-69

Friedman, Kendall, and Taylor do **deep-inelastic scattering** (DIS) destroying proton target – at the SLAC (Stanford Linear Accelerator Center) 2 mile long, 20 GeV electron accelerator.

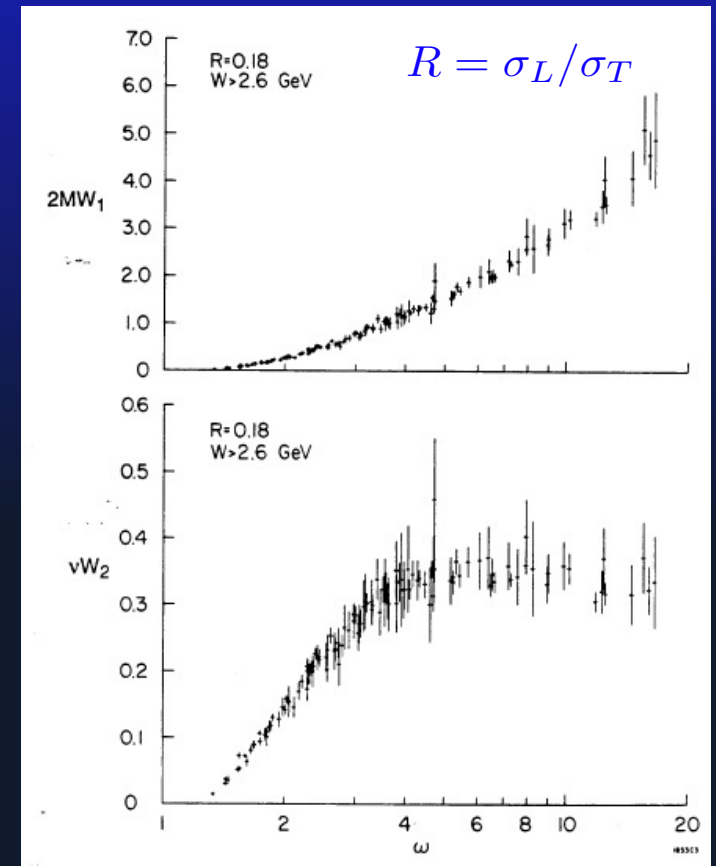
Measure electron angular cross section:

$$\sigma(E, E', \theta) = \frac{4e^4 E'^2}{Q^4} \left\{ W_2(\nu, Q^2) \cos^2 \frac{\theta}{2} + 2W_1(\nu, Q^2) \sin^2 \frac{\theta}{2} \right\}$$

For scattering from point particles inside proton, W_1 and νW_2 depend only on Bjorken scaling variable

$$x = \frac{Q^2}{2m_{\text{proton}}\nu}$$

Observe dependence on x only; shows that **proton is made of point particles**.



Bjorken scaling indicating quark structure. $\omega = 1/x$ (W = mass of recoiling target)

Proposal of quarks as mathematical model 1964

Volume 8, number 3

PHYSICS LETTERS

1 February 1964

A SCHEMATIC MODEL OF BARYONS AND MESONS *

M. GELL-MANN

California Institute of Technology, Pasadena, California

Ordinary

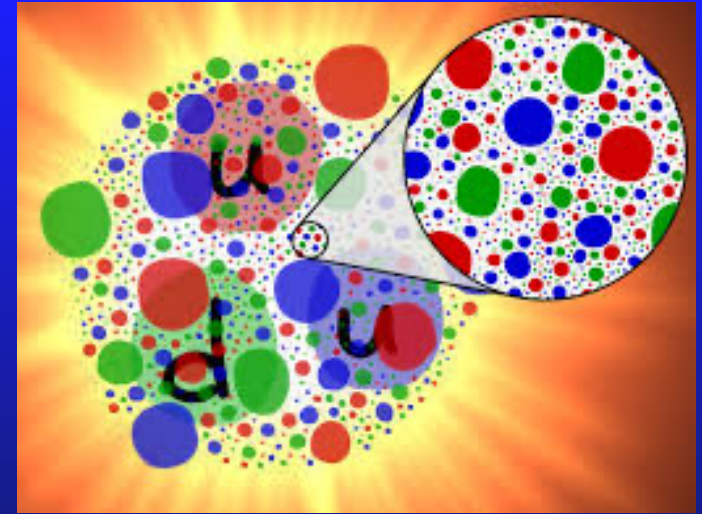
matter near the earth's surface would be contaminated by stable quarks as a result of high energy cosmic ray events throughout the earth's history, but the contamination is estimated to be so small that it would never have been detected. A search for stable quarks of charge $-\frac{1}{3}$ or $+\frac{2}{3}$ and/or stable di-quarks of charge $-\frac{2}{3}$ or $+\frac{1}{3}$ or $+\frac{4}{3}$ at the highest energy accelerators would help to reassure us of the non-existence of real quarks.

Parton model (1969)

R.P. Feynman and J.D. Bjorken:



Understand electron scattering in terms of *partons* -- **quasiparticles**



Given parton carries momentum p (in beam direction) -- a fraction **x** of the total target proton momentum p_{proton} . **Same x**

$$\frac{p_{\text{parton}}}{p_{\text{proton}}} = x = \frac{Q^2}{2m_{\text{proton}}\nu}$$

energy conservation
in scattering on parton $\nu = \frac{Q^2}{2m_{\text{parton}}}$

$$m_{\text{parton}} = x m_{\text{proton}}$$

Measure x in “infinite momentum” frame, i.e., with the proton moving at (nearly) the speed of light.

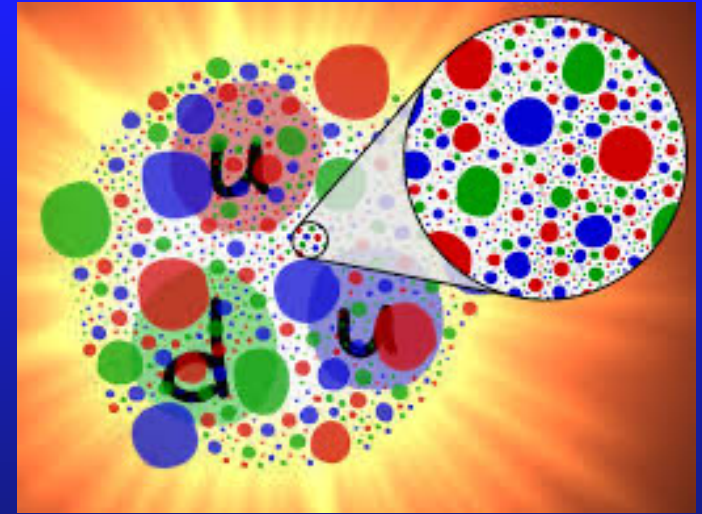
Point partons identified with quarks, antiquark, and gluons -- all governed by QCD, with asymptotic freedom. **Brodsky and Farrar**

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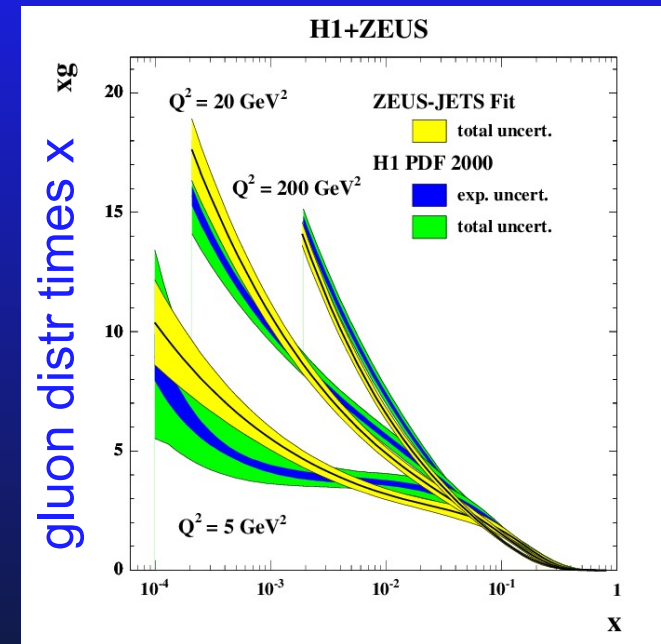
1991-2007 e scattering at HERA

(Hadron-Electron Ring Accelerator, Hamburg)

High-energy collisions of 27.5 GeV electron and positron beams (polarizable) with 920 GeV proton beams (unpolarized).

HERMES (spin) fixed-target experiment.

HERA => great abundance of low x gluons within nucleon.

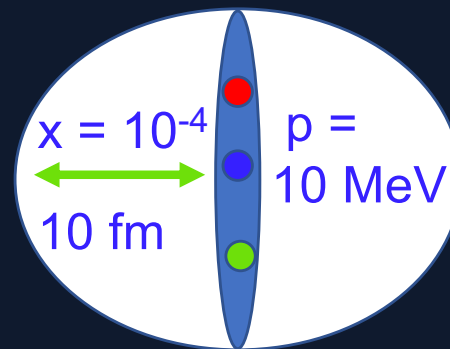


How can electrons probe electrically neutral gluons?

$g \rightarrow q + \bar{q}$ causing an effective electric dipole moment

View proton in frame in which proton is “slower” than ∞ momentum.

Heisenberg => low x “wee” partons stick out.



Lorentz contracted

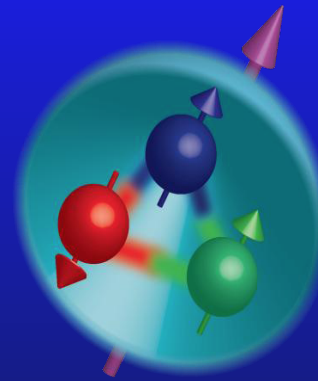


100 GeV proton

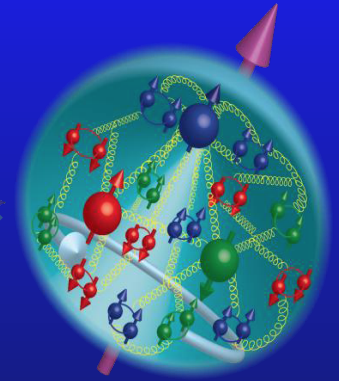
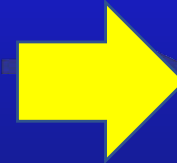
Physics goals of an EIC

Our picture of the nucleon has evolved considerably from the first simple picture as a “bag” of three valence quarks.

Have sea of quark-antiquark pairs (u, d, s) as well as cloud of gluons, all interacting!

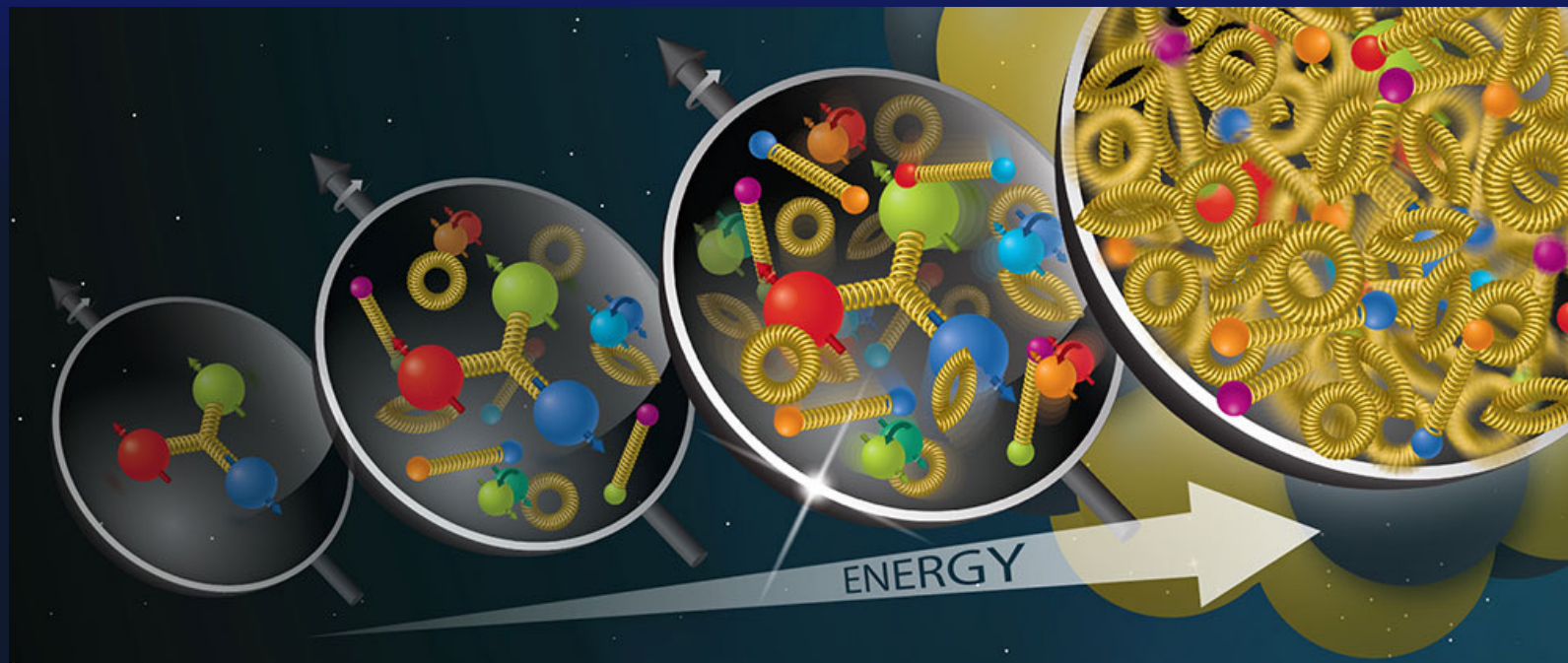


1970s cartoon



Now

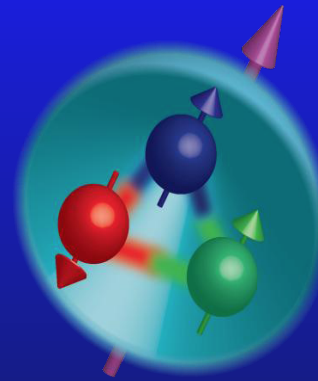
Z-E Meziani



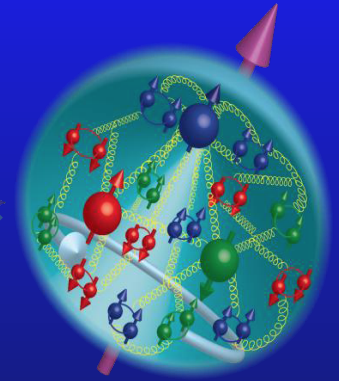
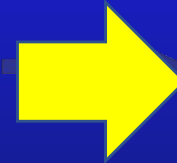
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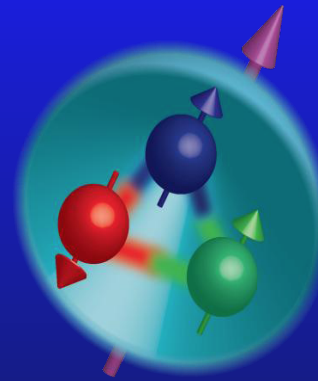
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Where are all the quarks and gluons -- in space and in momentum space? What is the many-body physics of all these interacting degrees of freedom.

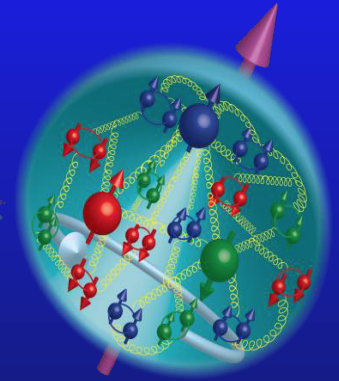
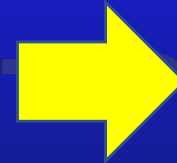
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1970s cartoon



Now

Z-E Meziani

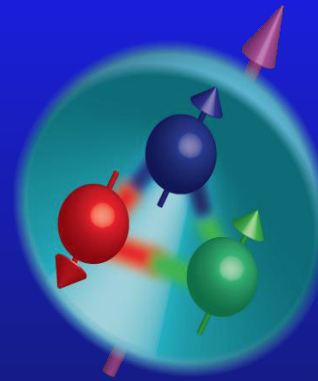
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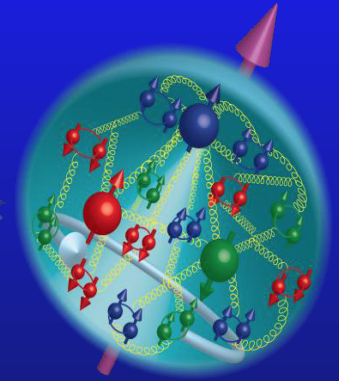
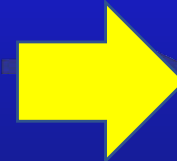
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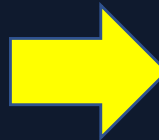
1970s cartoon



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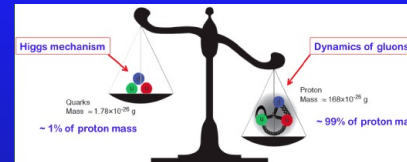


FEL.com: Philippe Crassous

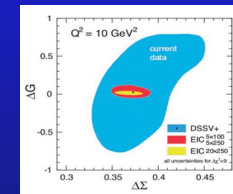
Sketch of worm vs. modern electron microscope picture (deep ocean worm)

Basic science questions for an ELC

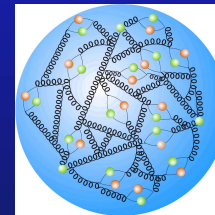
How does the nucleon get its mass?



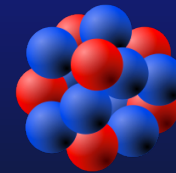
How does the spin of the nucleon arise from its elementary quark and gluon constituents?



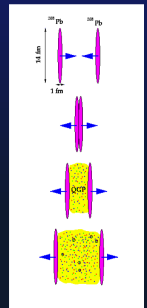
What are the emergent properties of dense systems of gluons?



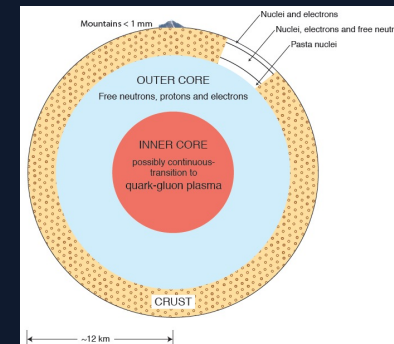
What is the internal structure of nuclei?
How do nuclei differ from being a simple collection of nucleons?



What is the initial state in ultrarelativistic heavy-ion collisions?



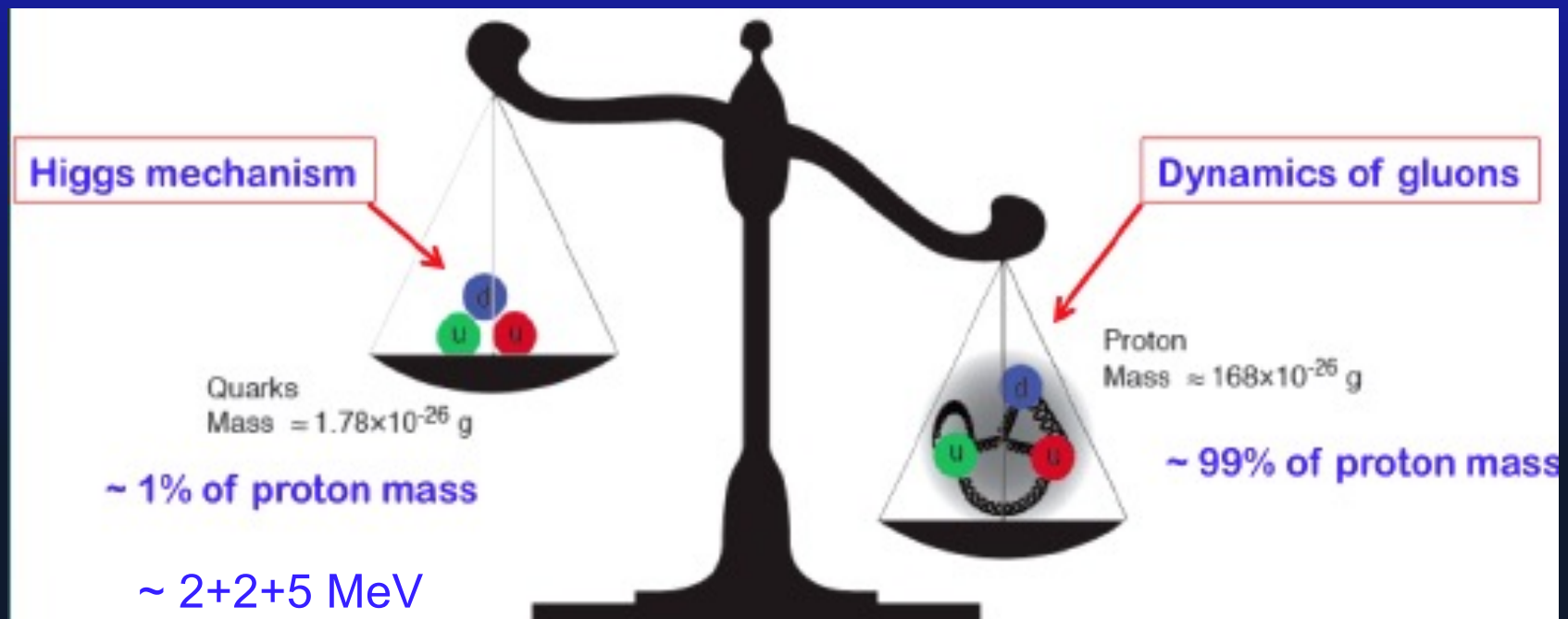
How does dense matter crossover from nucleonic degrees of freedom to quark degrees of freedom at higher density – application to neutron stars



Signals of beyond-the-standard-model physics? Effects of new particles and forces?

How does a nucleon get mass?

$m_p = 938 \text{ MeV}$ is almost 100 times greater than the masses of its valence quarks. **Cannot** be understood in terms of the Higgs mechanism!! Higgs \Rightarrow I should weigh $\sim 750 \text{ gm}$



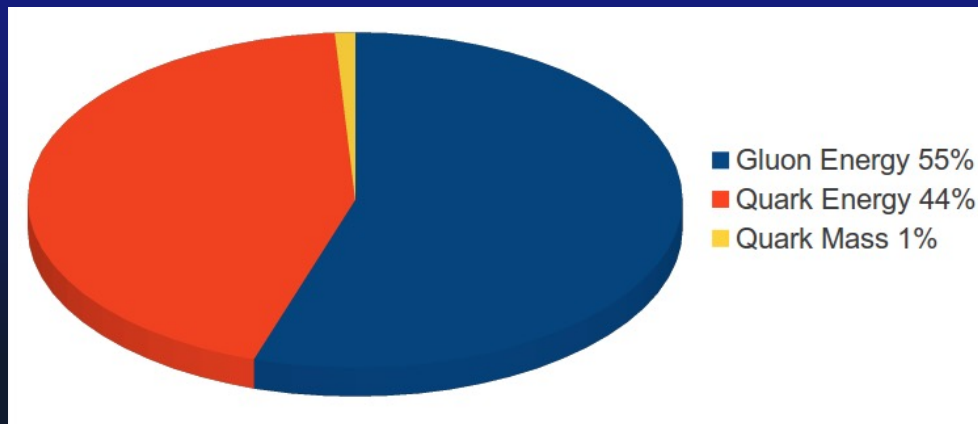
In atoms, mass = total mass of constituents minus binding energy (energy released in chemical reactions) < mass of constituents.

In nucleons, mass \gg total mass of constituents.

The mass of the nucleon

Naively, the mass of the nucleon arises from all quark, anti-quark and gluon kinetic energies – from the uncertainty principle when localizing excitations within the nucleon (radius, r_p).

$$\Delta E \sim \hbar c / r_p \text{ per quark or gluon.}$$



Energy distribution
within nucleon

How are these components distributed (in space and momentum) in the nucleon?

And how do nuclei differ from being a simple collection of nucleons – changes of quark and gluon distributions in nuclei?

Parton (quark or gluon) distributions

Start with **correlation** or Green's function $G(1, 2) = \langle N | \bar{q}(2) \cdots q(1) | N \rangle$ (here q = quarks). "... depends on process. $|N\rangle$ = nucleon.

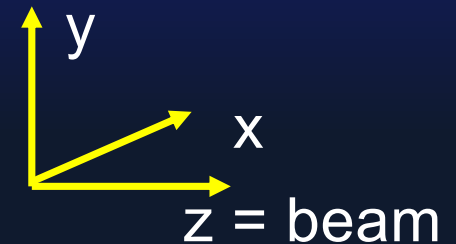
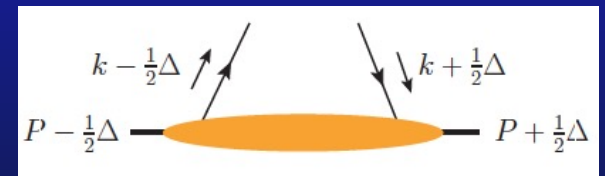
Define momentum distributions in terms of different Fourier transformations.

Fully transform in momentum space

$$f(k, \Delta) \propto \int d^4x_1 d^4x_2 e^{i\Delta(x_1+x_2)/2} e^{-ik(x_1-x_2)} G(1, 2)$$

Δ = momentum transfer to nucleon

k = relative momentum transfer



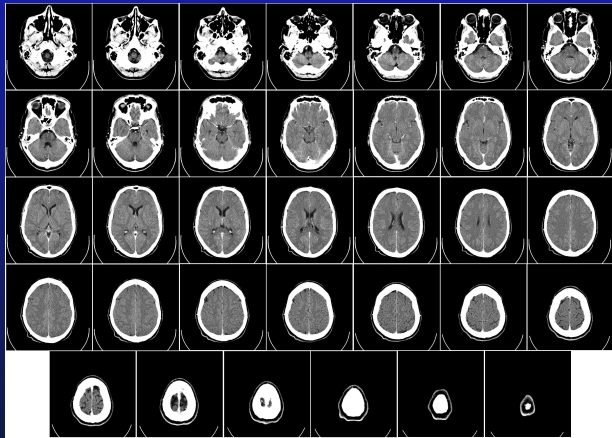
$$\vec{x}_{1\perp} - \vec{x}_{2\perp} \rightarrow \vec{k}_{\perp}$$

transverse momentum distribution alone

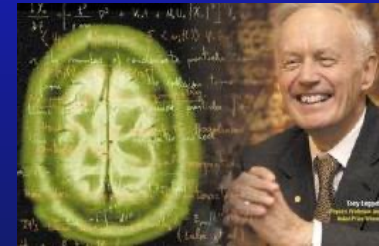
Along direction of electron momentum transfer $z_1 - z_2 \Rightarrow x_{\text{Bjorken}}$, etc.

Tomography

Build up correlation functions by measuring transverse momentum \vec{k}_{\perp} distributions at varying x . (Requires large luminosity.)

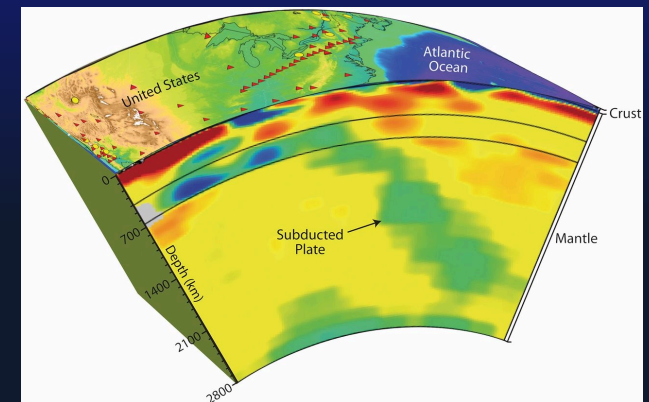


Computer assisted tomography of brain



A.J. Leggett

Seismic tomography, slices at varying depth; earthquakes as probes



Measure dependence of cross section on momentum transfer q from electron to target => information about transverse position and momentum of struck quarks and gluons.

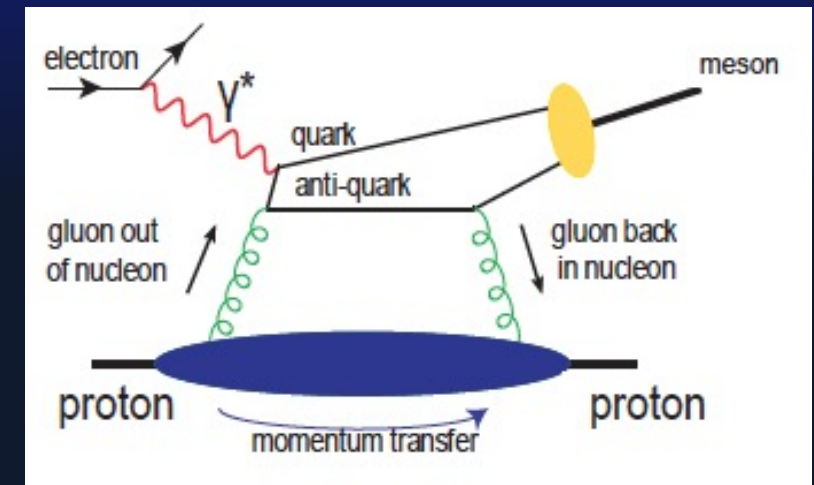
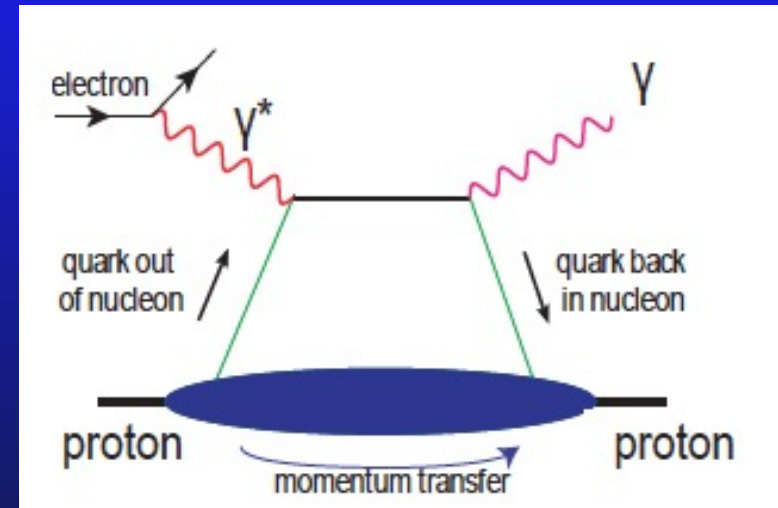
Measuring quark and gluon transverse distributions.

Real photon production:
Deeply virtual Compton scattering
sensitive to transverse
position of quarks.

Real meson production:
gives info on gluon distribution.

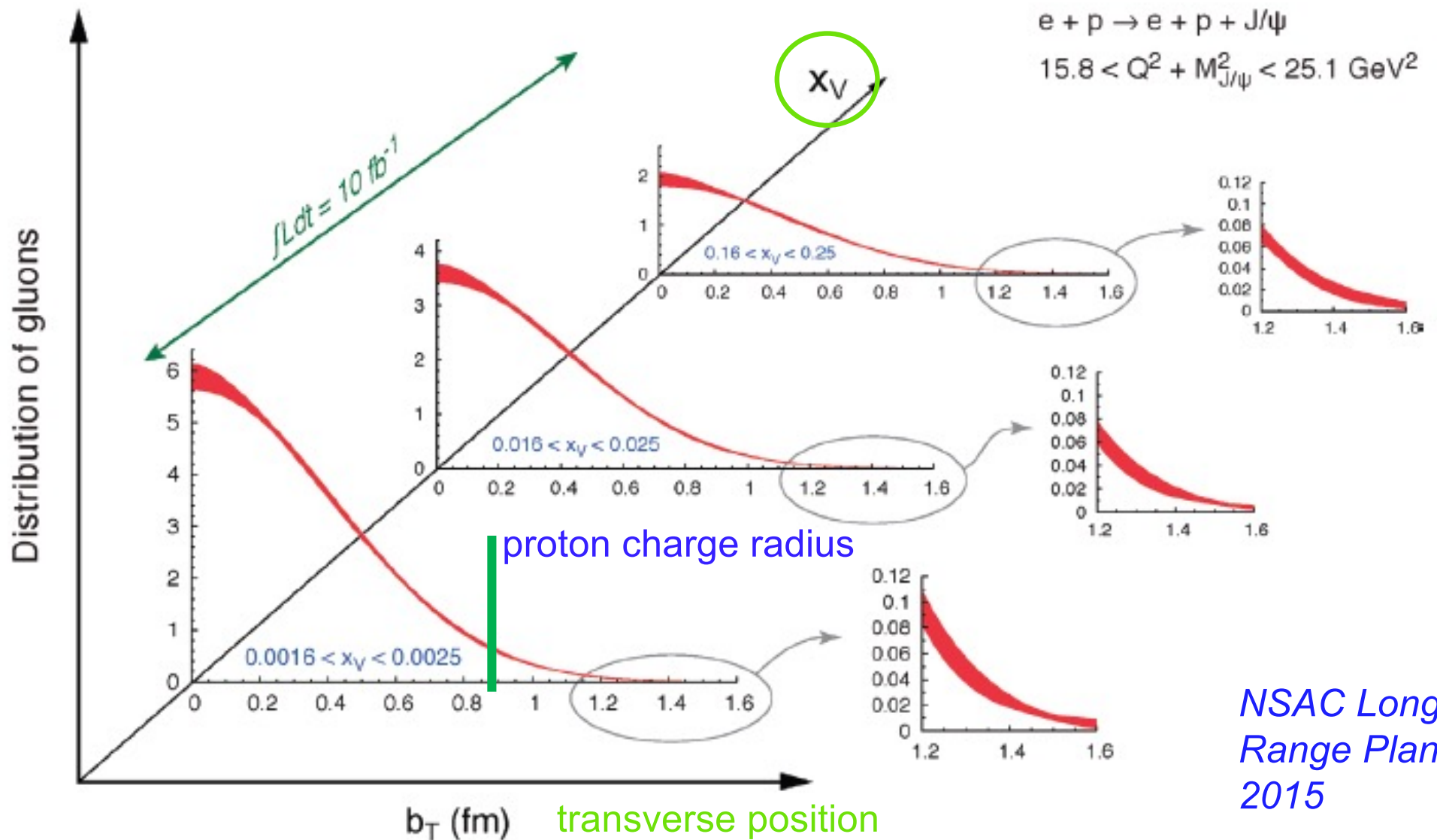
Ex., Heavy mesons J/ψ or Upsilon (Υ):

$$\gamma^* \rightarrow \bar{b}b \rightarrow \Upsilon$$



Expected transverse gluon distributions in space

Slices vs. transverse position b_T at various x

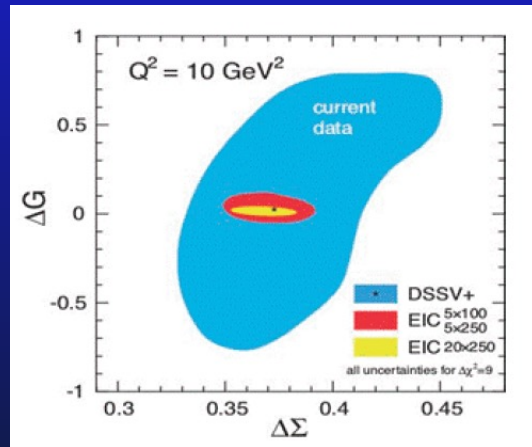


NSAC Long
Range Plan
2015

How do the quarks and gluons give the proton its spin?

Proton spin is the basis of nuclear magnetic resonance (MRI) imaging.

Spin crisis:



Current estimate of contributions to spin (in units of $\hbar/2$):

Quarks $\Delta\Sigma \sim 30\text{-}40\%$

Gluons $\Delta G \sim -70 \text{ to } +70\% \text{ ??}$

(RHIC pp)

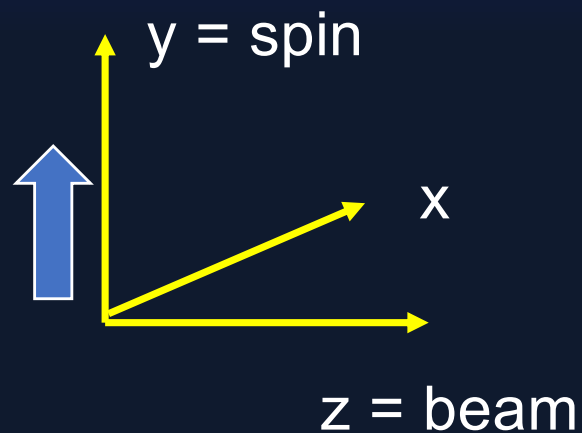
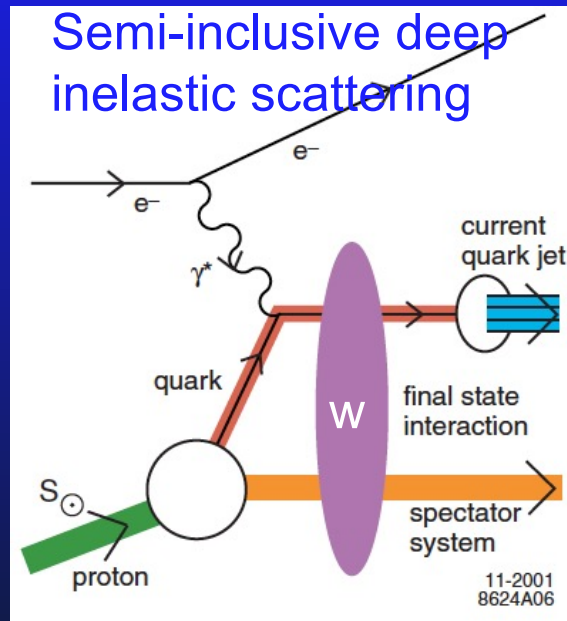
Orbital motion ?? (JLab - 12GeV)

Extract orbital contributions from transverse motion measured by spatial and momentum distributions of components in deep inelastic scattering -- **tomography**.

Extract gluon contributions from transfer of gluon polarization to $q\text{-}\bar{q}$ pair, probed by polarized electrons.

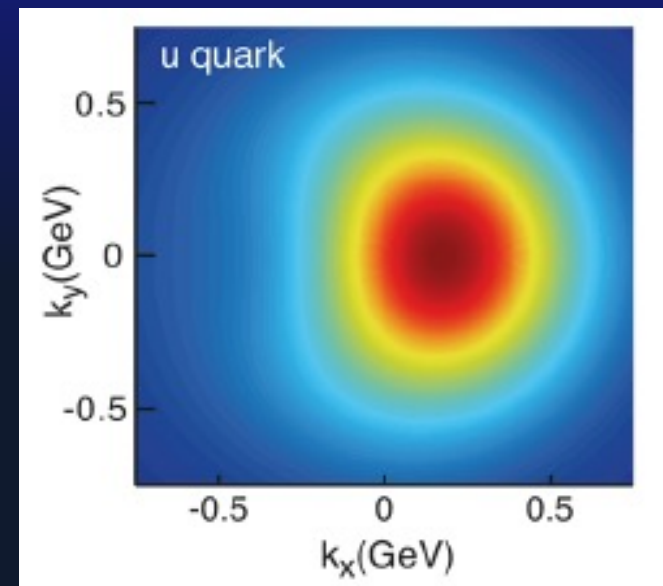
Polarized beams ($> 70\%$) critical!

Detecting quark orbital contributions to proton spin



Proton spinning (along y) with orbital motion as well leads to asymmetry of quark distributions in x direction.

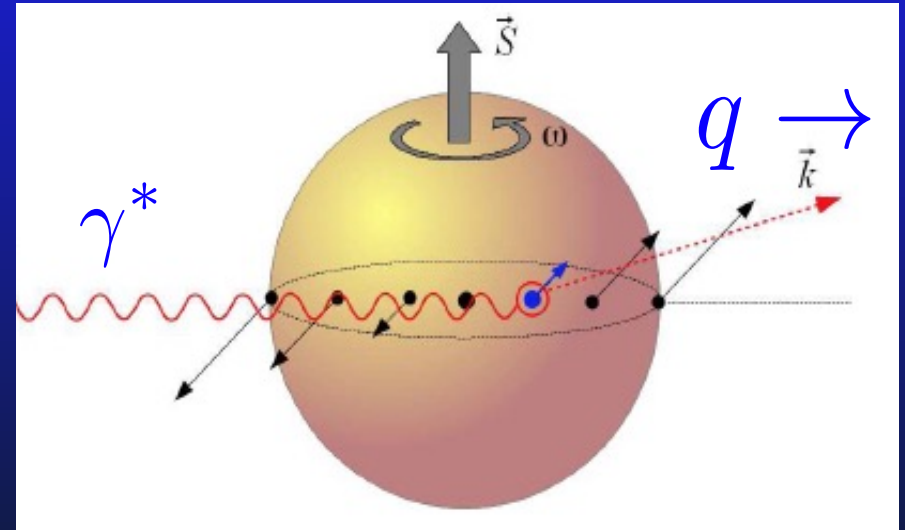
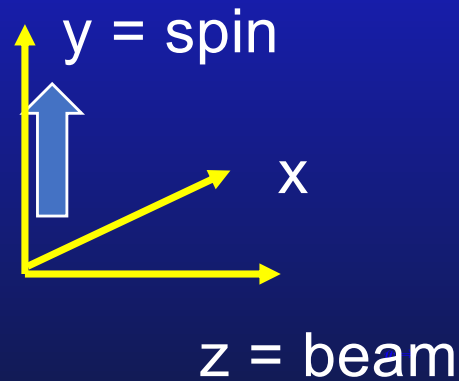
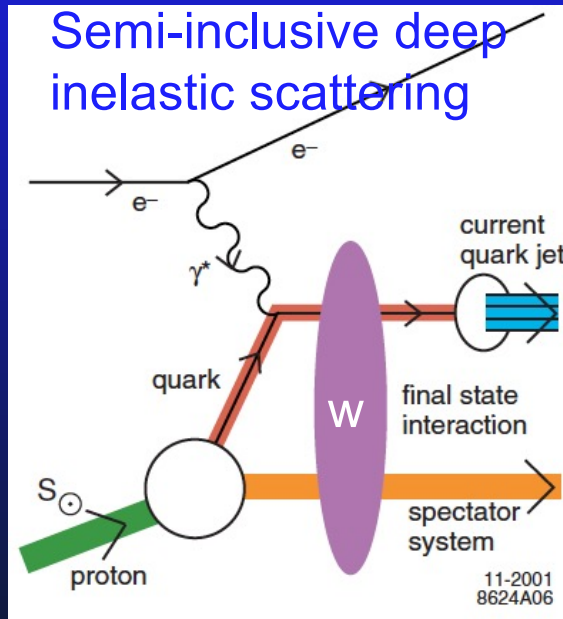
Beam along z .



Representative asymmetry of up quark distribution for $x = 0.1$.

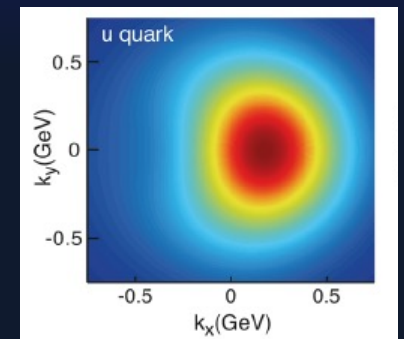
Intuitive quasi-classical scattering on rotating nucleon

Yuri Kovchegov and *Matt Sievert*, *Phys. Rev. D* 89, 054035 (2014)



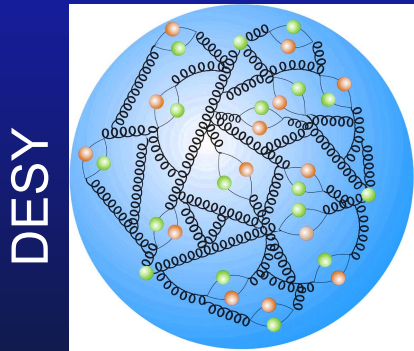
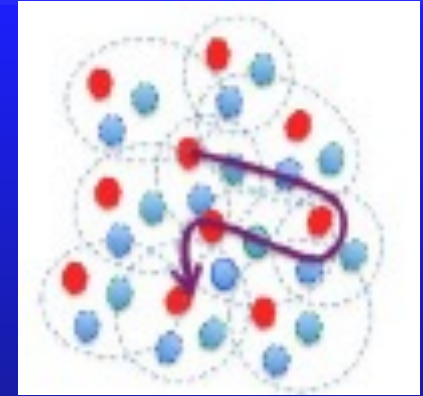
Virtual photon γ^* excites quark along z direction.
 Quark scatters strongly (W) with other partons
=> short mean free path.

Knocked-out quark more likely emitted from far (right) side,
 and thus orbital motion kicks it a bit in the $+x$ direction.



Gluon physics

Gluons in nucleons and nuclei (as well as other hadrons) are like dark matter in the universe— unseen but crucial in holding matter together.

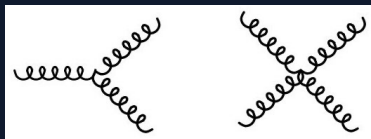


Nucleons and nuclei are in fact complex interacting many-body systems -- beyond bags of free quarks and free gluons. Ex., **nuclei exhibit composite fermions— the nucleons.** Confinement!

“The most precise picture of the proton”

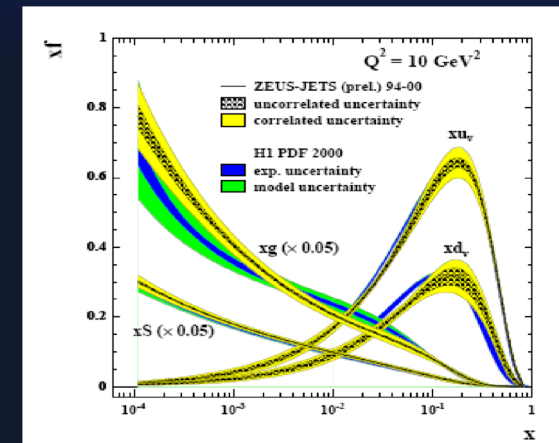
HERA => huge numbers of low momentum gluons in the nucleon -- at low x ($<10^{-4}$).

Low momentum sector (wee partons) dominated by strongly interacting gluons!. The gluon field



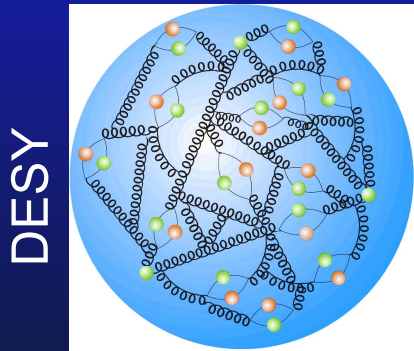
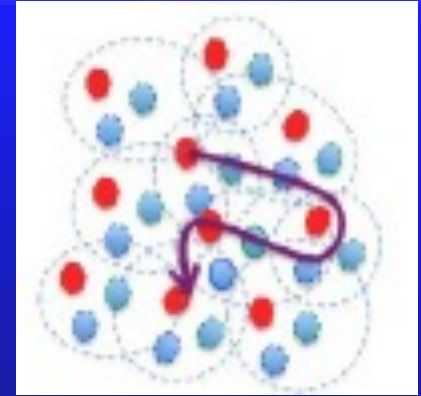
is highly non-linear!

A new many-body system! **New emergent phenomena?**



Gluon physics

Gluons in nucleons and nuclei (as well as other hadrons) are like **dark matter** in the universe— unseen but crucial in holding matter together.



Nucleons and nuclei are in fact complex interacting many-body systems -- beyond bags of free quarks and free gluons. Ex., **nuclei exhibit composite fermions (called nucleons)**. Confinement!

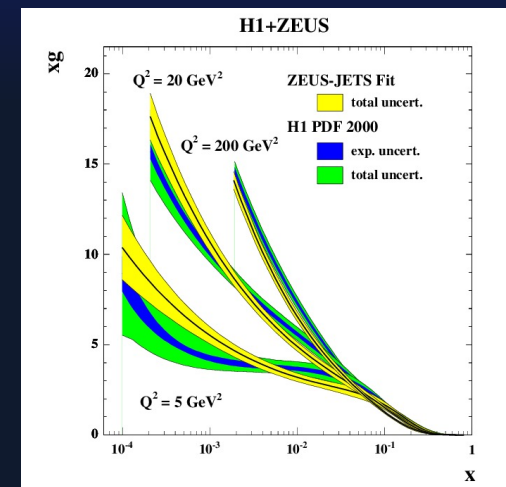
“The most precise picture of the proton”

HERA => huge numbers of low momentum gluons in the nucleon -- at low x ($<10^{-4}$). Low momentum sector (wee partons) dominated by strongly interacting gluons!. The gluon field



is highly non-linear!

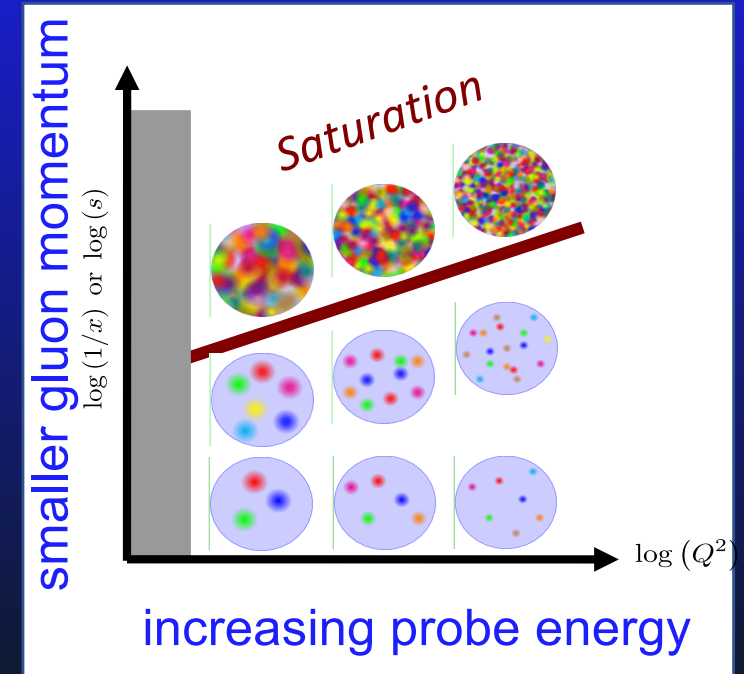
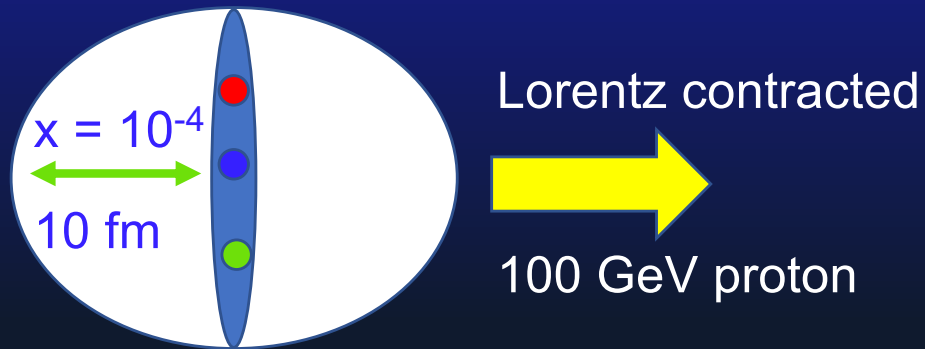
A new many-body system! **New emergent phenomena?**



Gluon self-interactions become important when there are enough: at small x

Scale of saturated gluonic matter: Q_s

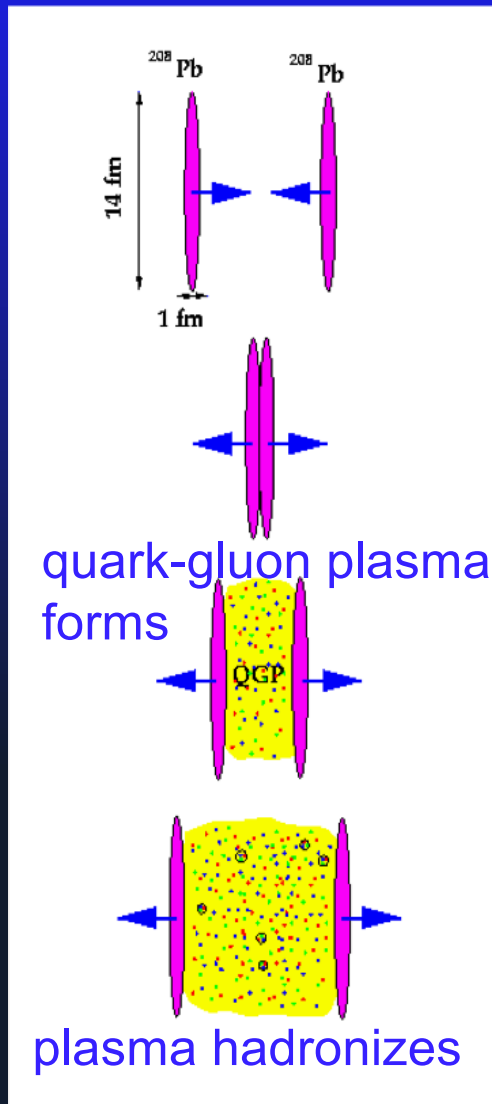
At HERA (318 GeV c.m.) $Q_s \sim 1$ GeV



First approximation, dense cloud of gluons forms a **Bose condensate** – “color glass condensate.”

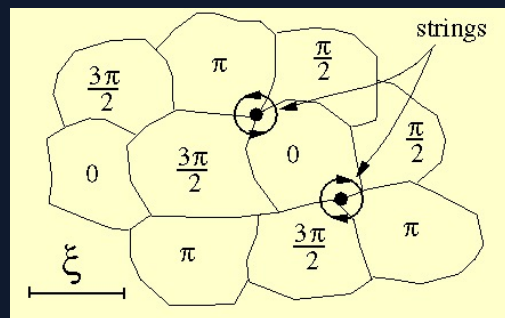
Excitations of saturated gluonic matter? Topology?

Connections to heavy ion collisions:

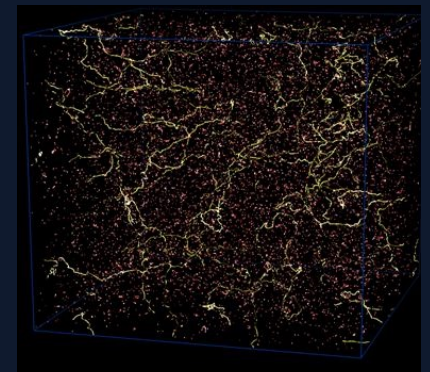


Saturated gluonic matter reachable at a sufficiently energetic EIC. Describes initial state in ultrarelativistic heavy ion collisions. **Bose-condensed gluonic matter** (color-glass condensate, ...). Condensate is metastable, decaying into quark-gluon plasma.

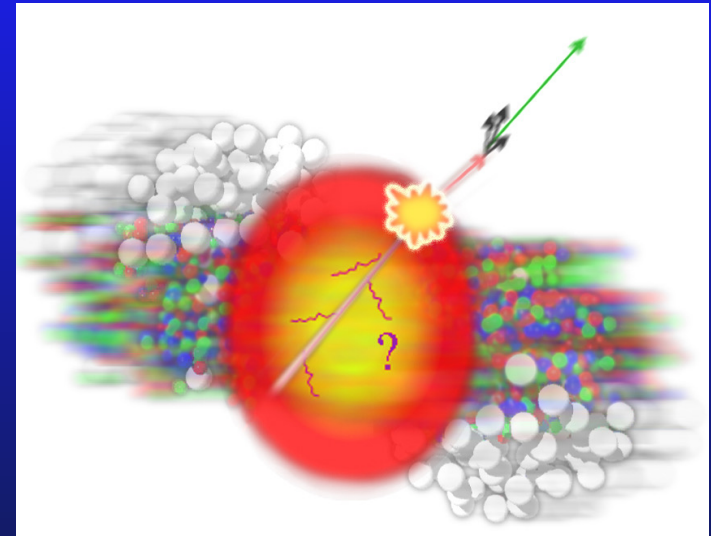
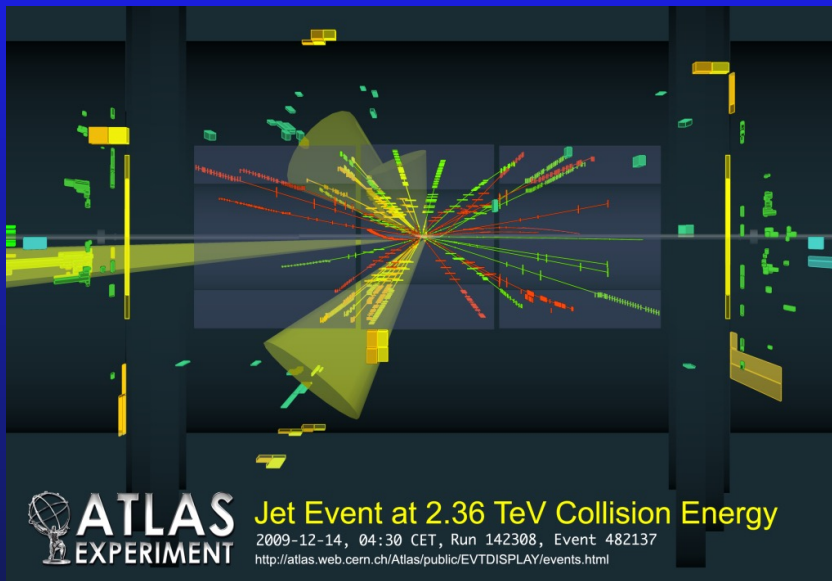
As in early universe, form topological defects, e.g., handedness asymmetry of produced $q\bar{q}$ pairs (chiral magnetic effect) related to the structure of the color field in saturated gluonic matter.



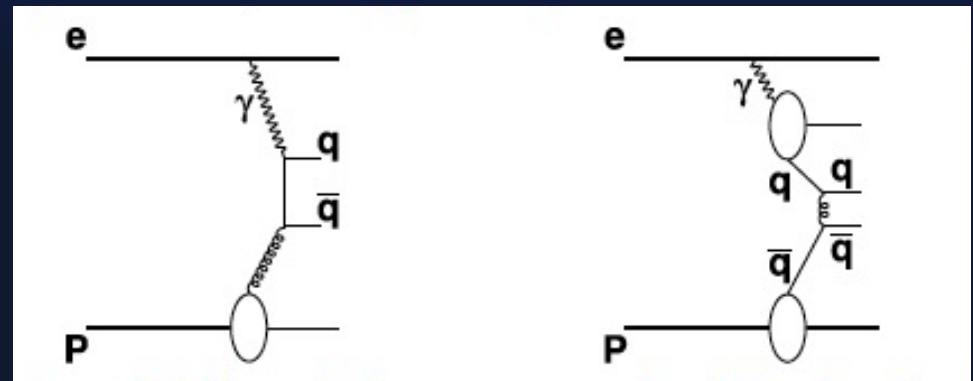
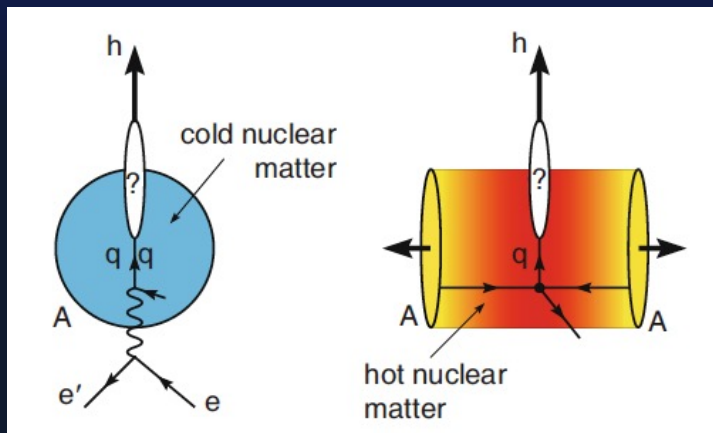
cosmic strings



Particle jets in electron-ion and ion-ion collisions



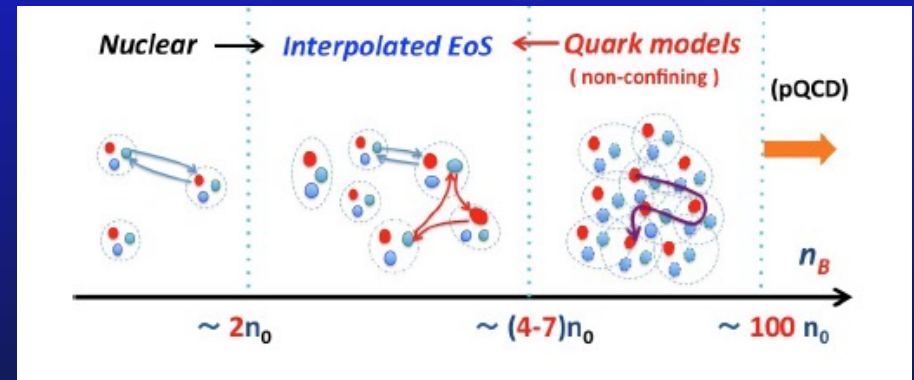
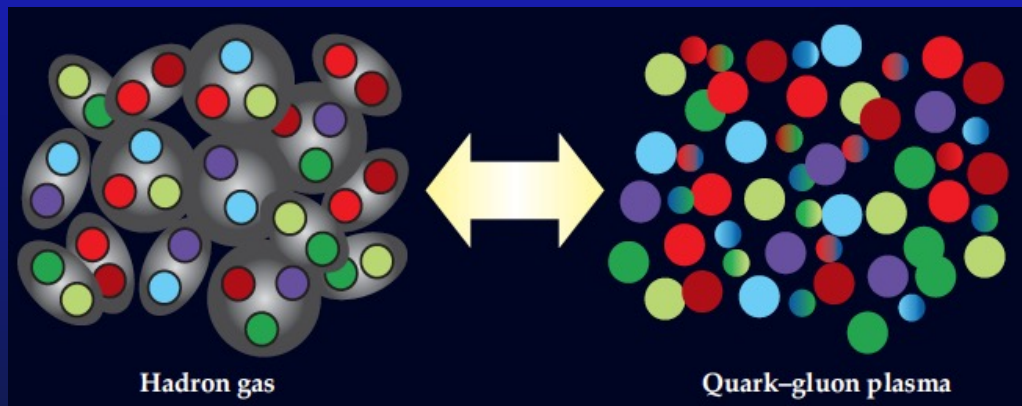
asymmetric jet in AA collision



Jets probe quark and gluon distributions

Dense matter and neutron stars:

Study transition from cold nuclear matter to quark matter – vital for neutron stars. What is energy density vs. baryon density?



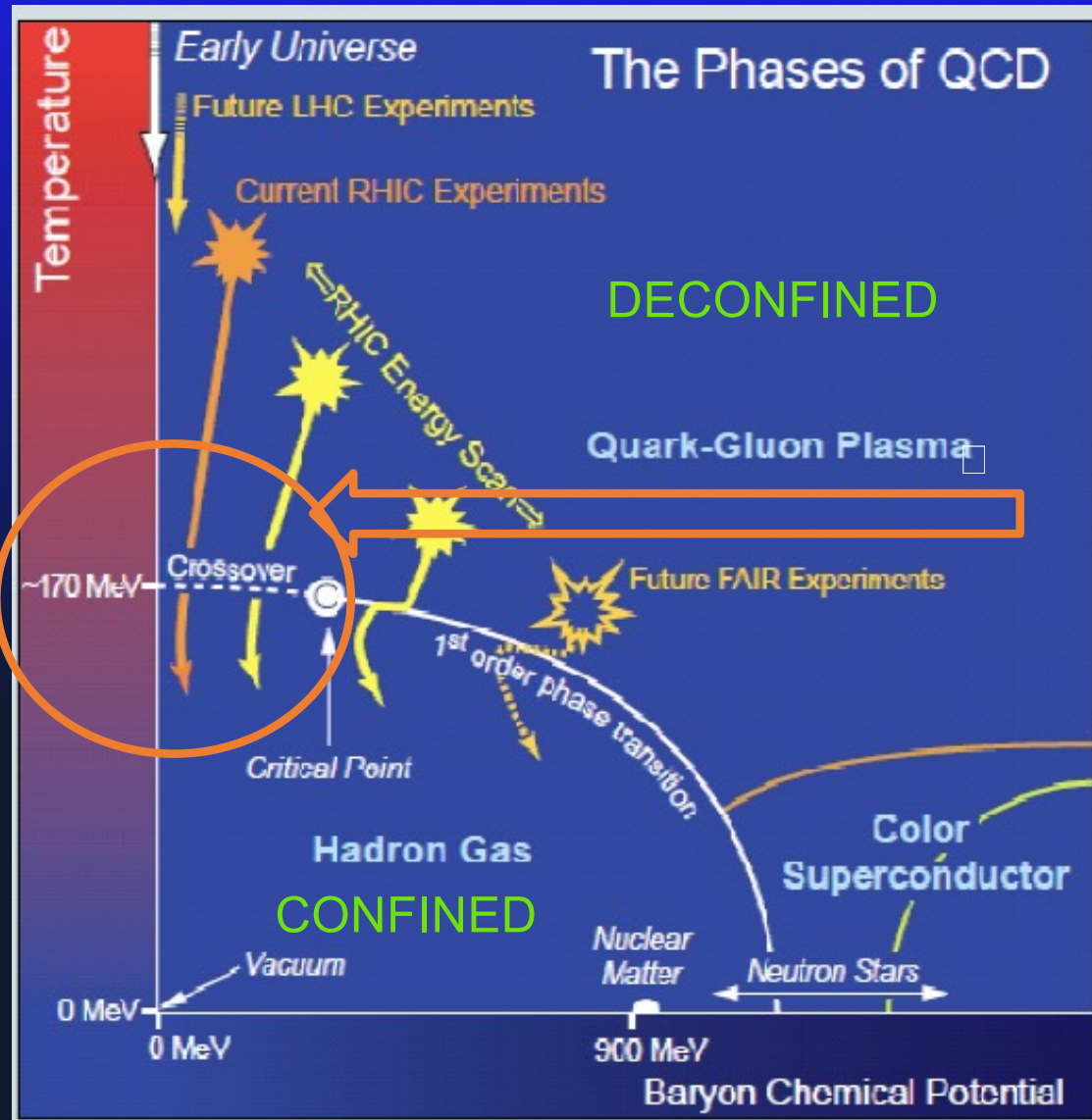
Expect “smooth” transition from nucleons to quarks

Gluon (and quark) distributions in nuclei at finer and finer scales should shed light on transition from nucleonic to quark degrees of freedom as density of matter increases.

(Can mapping of energy-momentum tensor (stress-energy tensor) in eA collisions reveal pressure vs. baryon density in dense matter?)

Modern phase diagram

Crossover at zero net baryon density



QCD lattice gauge theory – for finite light quark masses -- predicts crossover from confined phase at lower T to deconfined phase at higher T .

Wuppertal-Budapest lattice collab.
S. Borsanyi et al., PLB (2014)
A. Bazavov et al., PRD (2014)

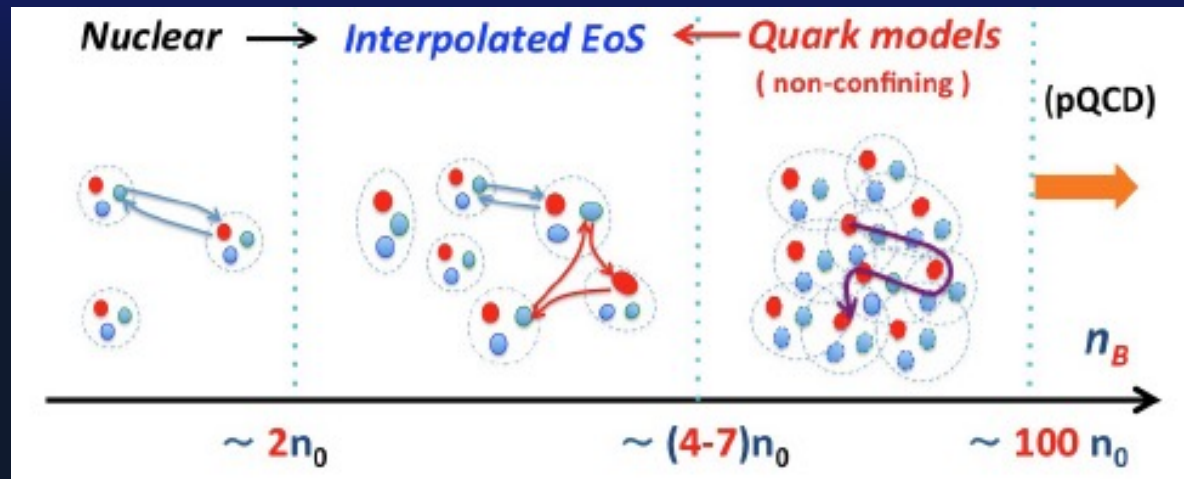
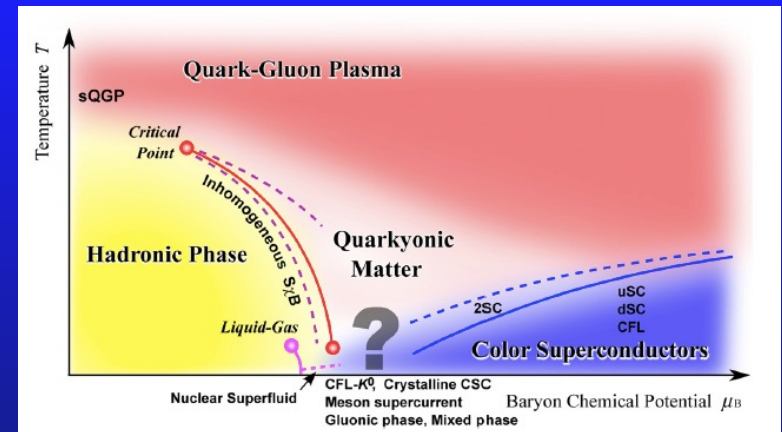
Do quarks roam freely in the deconfined phase?
If so, they must also roam freely at lower T .

Are free quarks really running about in this room?

No free quarks even above the crossover!

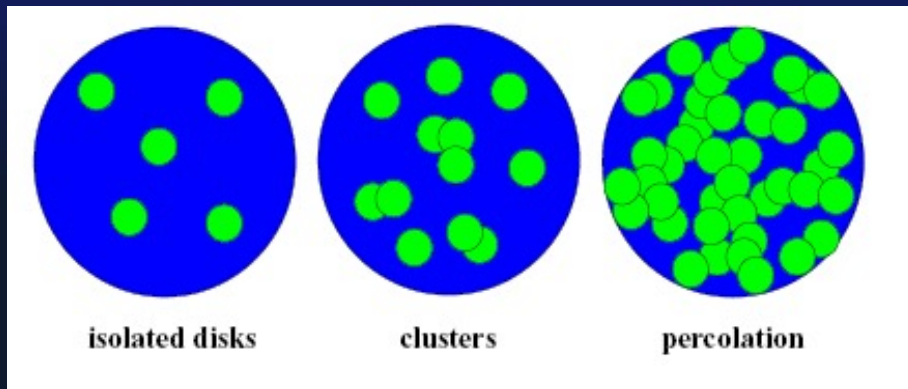
In confined region quarks are inside hadrons. Also have quarks and antiquarks in the QCD forces between hadrons. With temperature, form larger clusters, which **percolate** at the crossover. Expect similar percolation of clusters with increasing density.

In deconfined regime clusters extend across all of space.



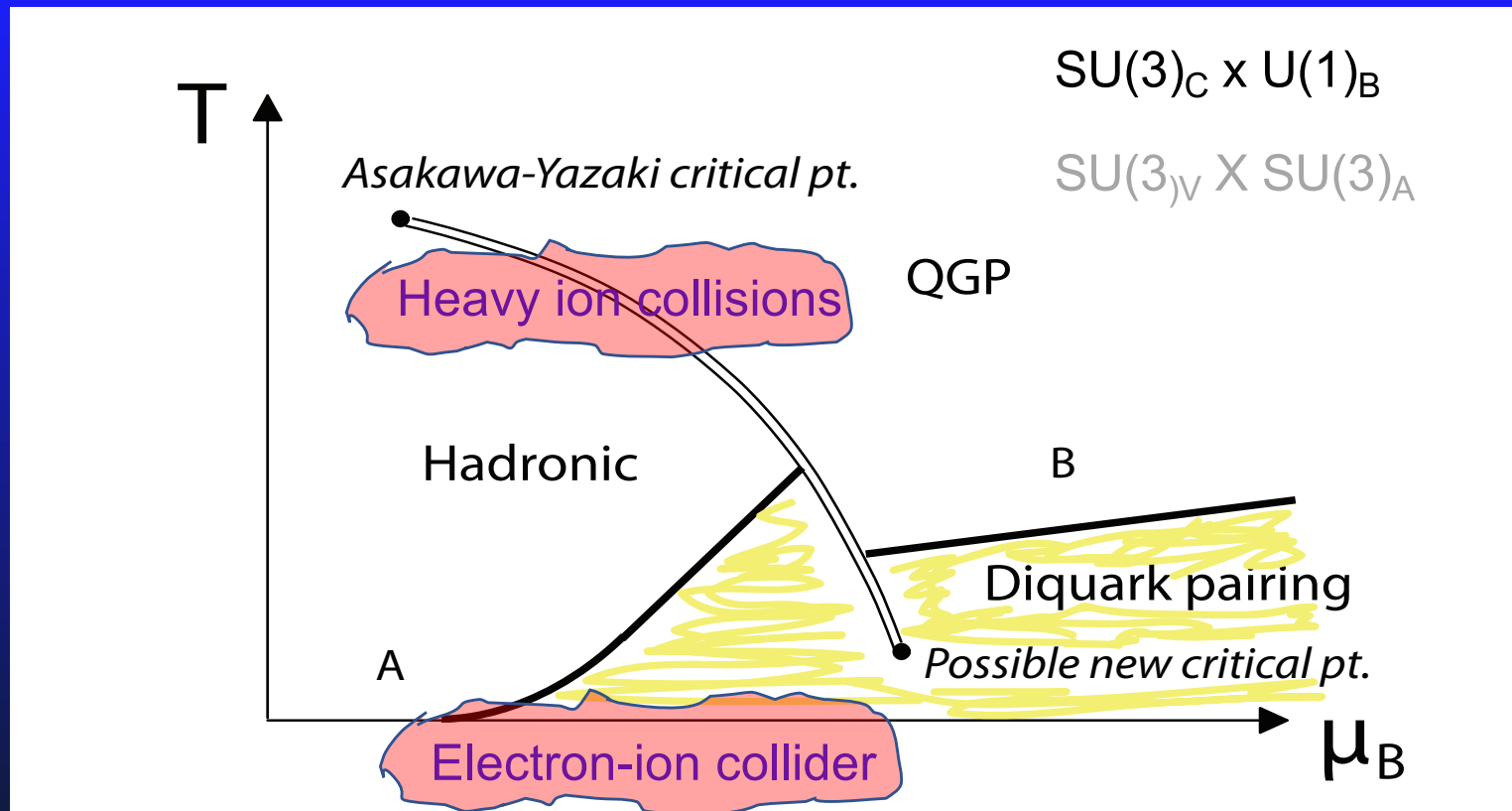
Need theoretical description of this evolution

Percolation



$$n_{\text{perc}} \sim 0.34 \left(\frac{3}{4}\pi r_n^3 \right) \text{ fm}^{-3}$$

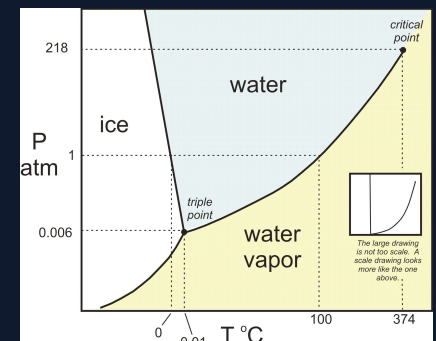
Quarks can still be bound even if deconfined.



Critical points similar to those in liquid-gas phase diagram (H_2O). **Neither critical point necessary!!**

Can go continuously from A to B around the upper critical point. Liquid-gas phase transition.

In lower shaded region have BCS pairing of nucleons, of quarks, and possibly other states (meson condensates). Different symmetry structure than at higher T .

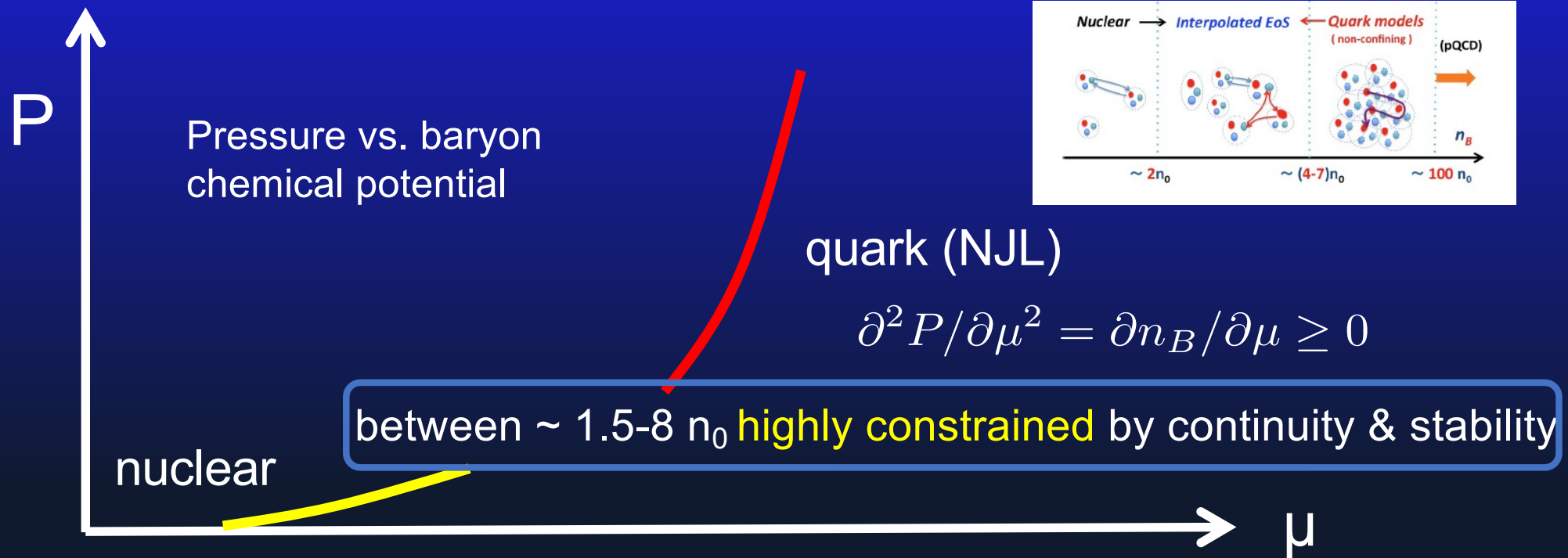


QHC21 (quark-hadron crossover) equation of state:

QHC19 --GB, S. Furusawa, T. Hatsuda, T. Kojo & H. Togashi, Ap.J. 885:42 (2019)

QHC21 --T. Kojo, GB, & T. Hatsuda, Ap. J. 934:46 (2022).

Have good idea of equation of state at nuclear and at high densities.



Quarks in Nambu-Jona-Lasinio (NJL) model with universal **repulsive short-range qq coupling** (Kunihiro) and **diquark (BCS) pairing interaction**

$$\mathcal{L}_V^{(4)} = -g_V (\bar{q} \gamma^\mu q)^2$$

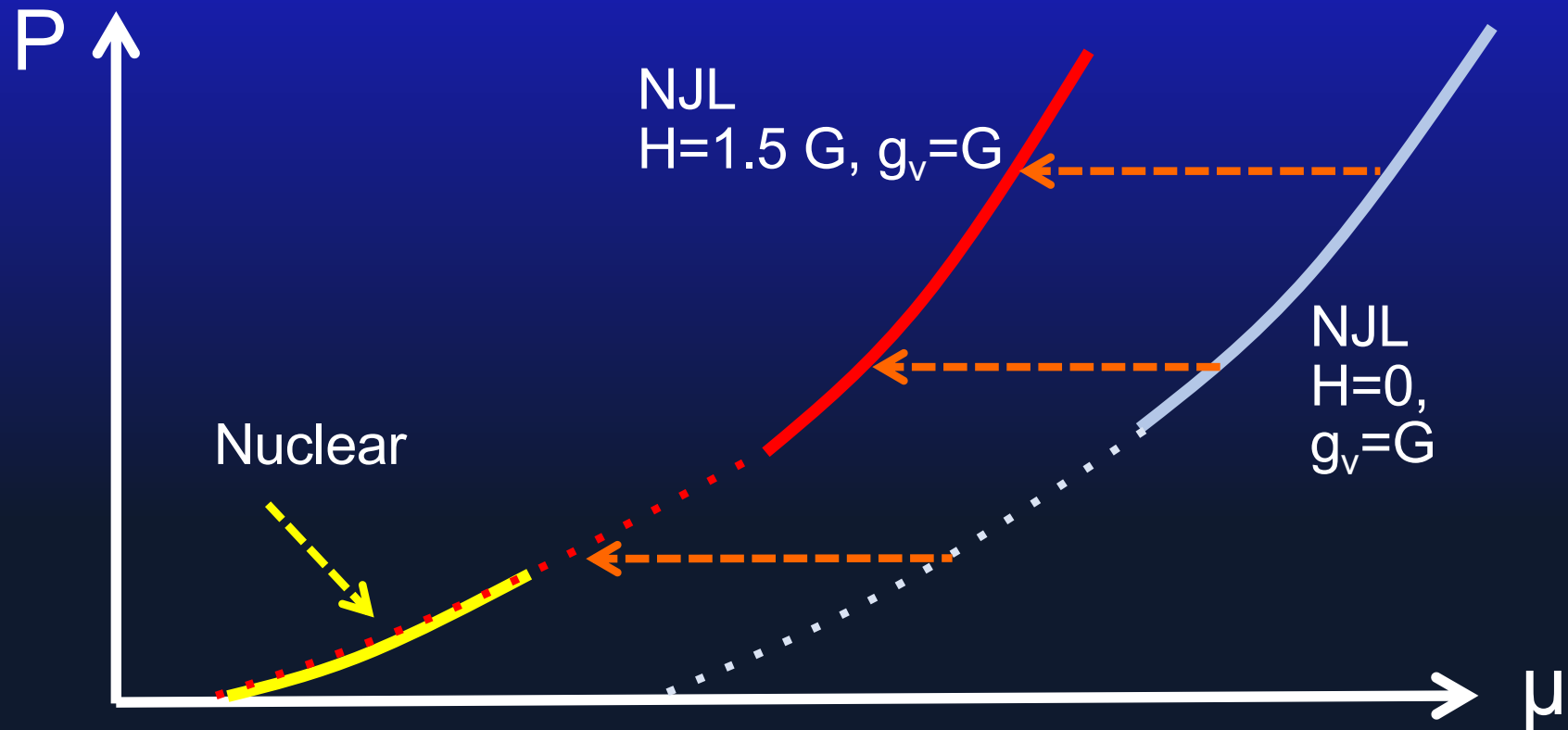
$$\mathcal{L}_d^{(4)} = H \sum_{A,A'=2,5,7} [(\bar{q} i \gamma_5 \tau_A \lambda_{A'} C \bar{q}^T)(q^T C i \gamma_5 \tau_A \lambda_{A'} q)]$$

Nuclear matter equation of state below $1.5 n_0$

For thermodynamic stability need large diquark pairing interaction, H

$$\mathcal{L}_d^{(4)} = H \sum_{A,A'=2,5,7} [(\bar{q} i \gamma_5 \tau_A \lambda_{A'} C \bar{q}^T)(q^T C i \gamma_5 \tau_A \lambda_{A'} q)]$$

$$\mathcal{L}_V^{(4)} = -g_V (\bar{q} \gamma^\mu q)^2$$

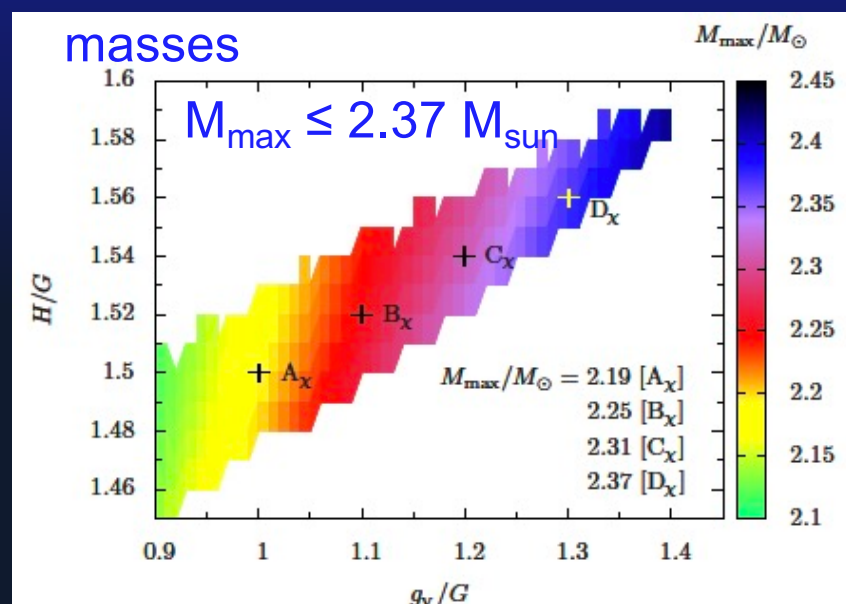
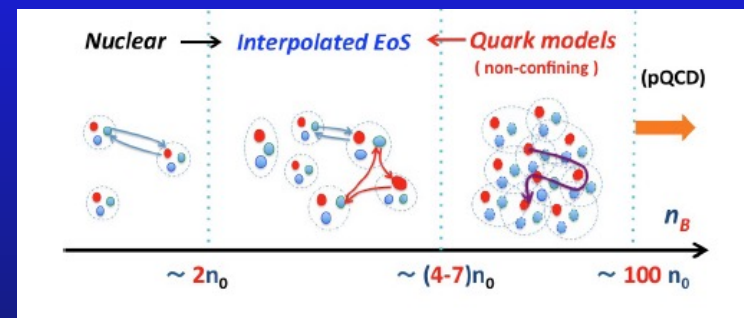


Expect onset of stronger 2-body correlations, with decreasing density, as quark matter comes nearer to becoming confined

QHC21 (quark-hadron crossover) eqn of state

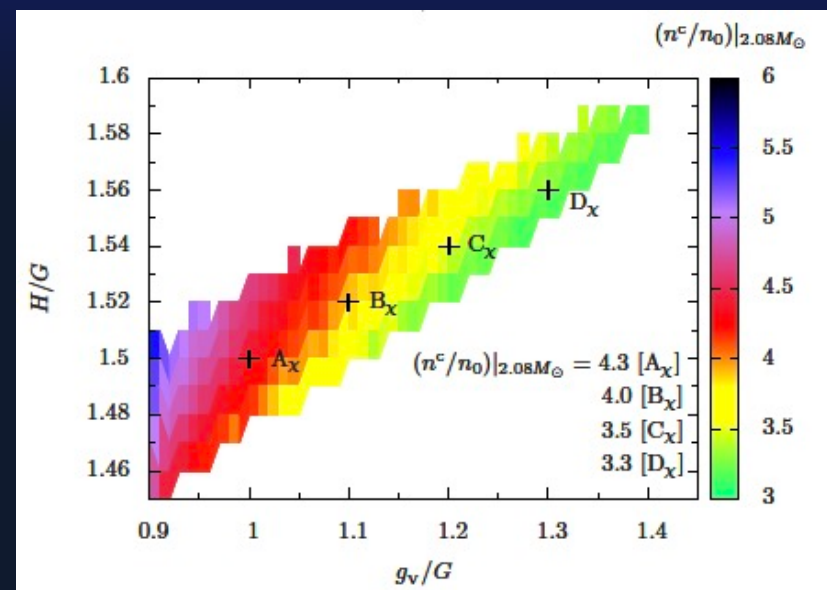
T. Kojo, GB, & T. Hatsuda, *Ap. J.* 934:46 (2022).

Parameters g_v and H must be in colored region so that
 speed of sound \leq speed of light.



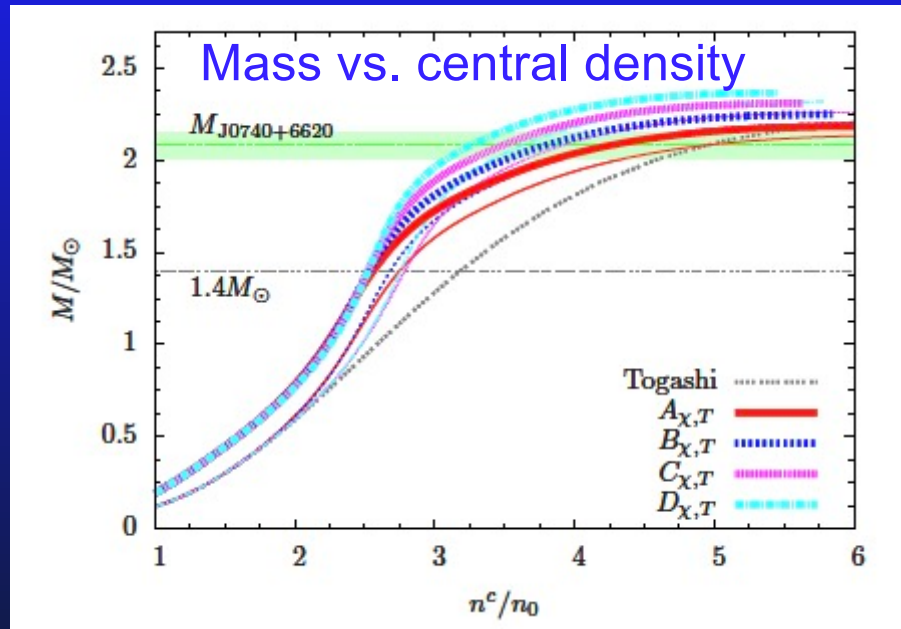
Strange quarks included

Further restricted by maximum
 neutron star mass $> 2.08 M_{\text{sun}}$



$2.08 M_{\text{sun}} \Rightarrow g_v \gtrsim 0.9 G$

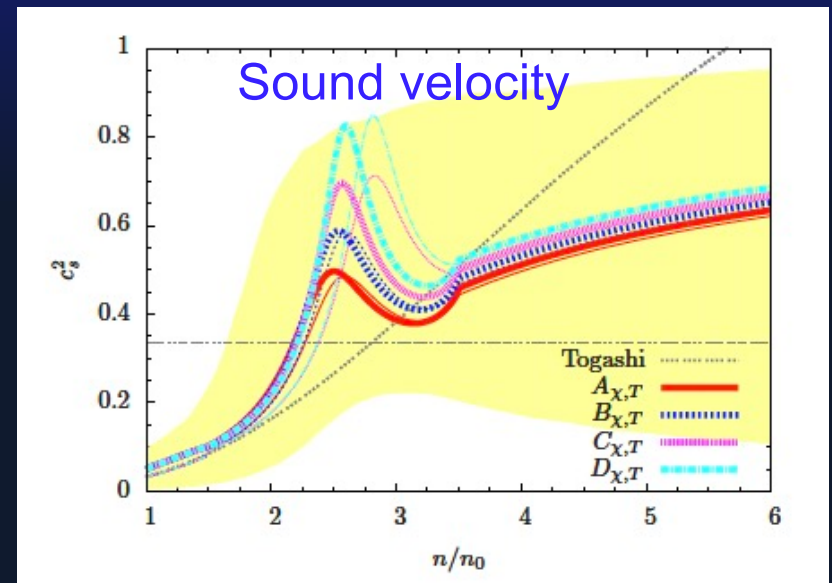
Quarks beginning to enter neutron stars



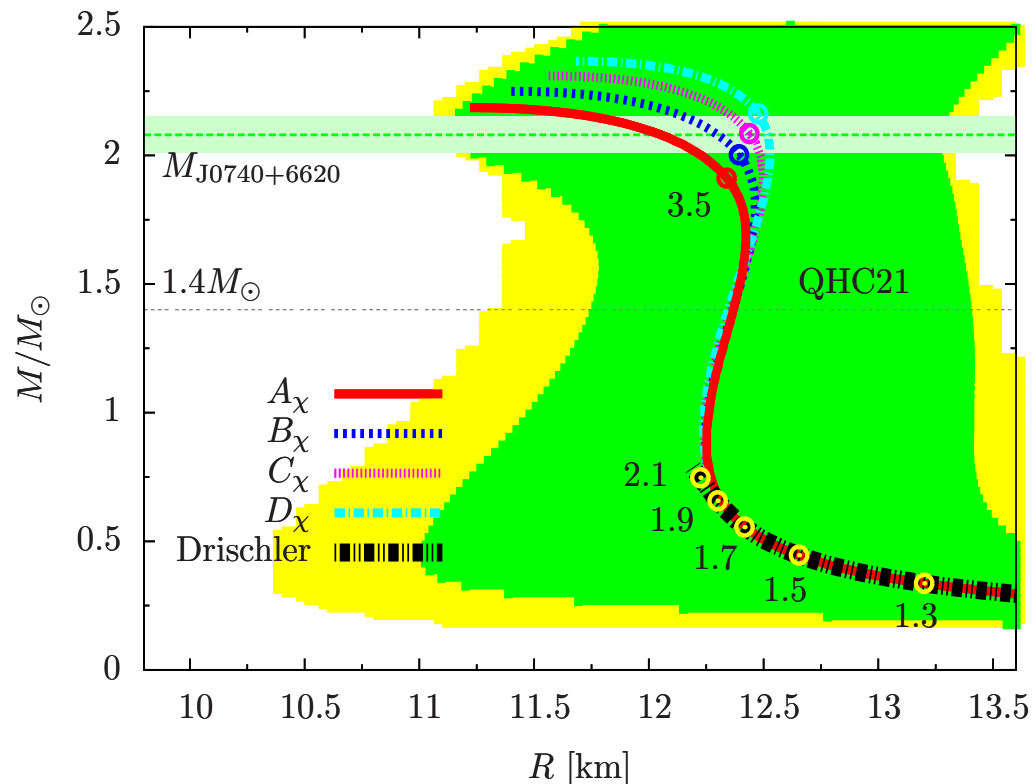
Cores of higher mass stars could reach beyond transition. Fully developed quark matter in cores? Need fully microscopic calculations of matter undergoing transition from nucleonic to quark degrees of freedom.

Central density of PSR J0740+6620 $\sim 5n_0$.

Well above densities where pure hadronic calculations are valid. Entering transition to strongly interacting quark matter.



Peak not seen in nucleonic eq. of state

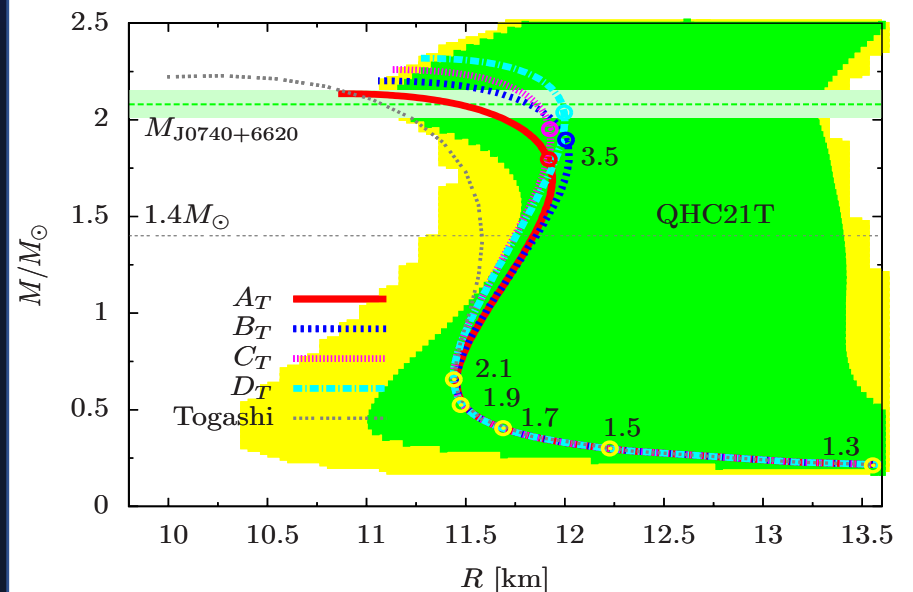


Chiral EFT +
transition to quark
matter in excellent
agreement with
NICER inferences
of radii. Rapid
pressure rise!

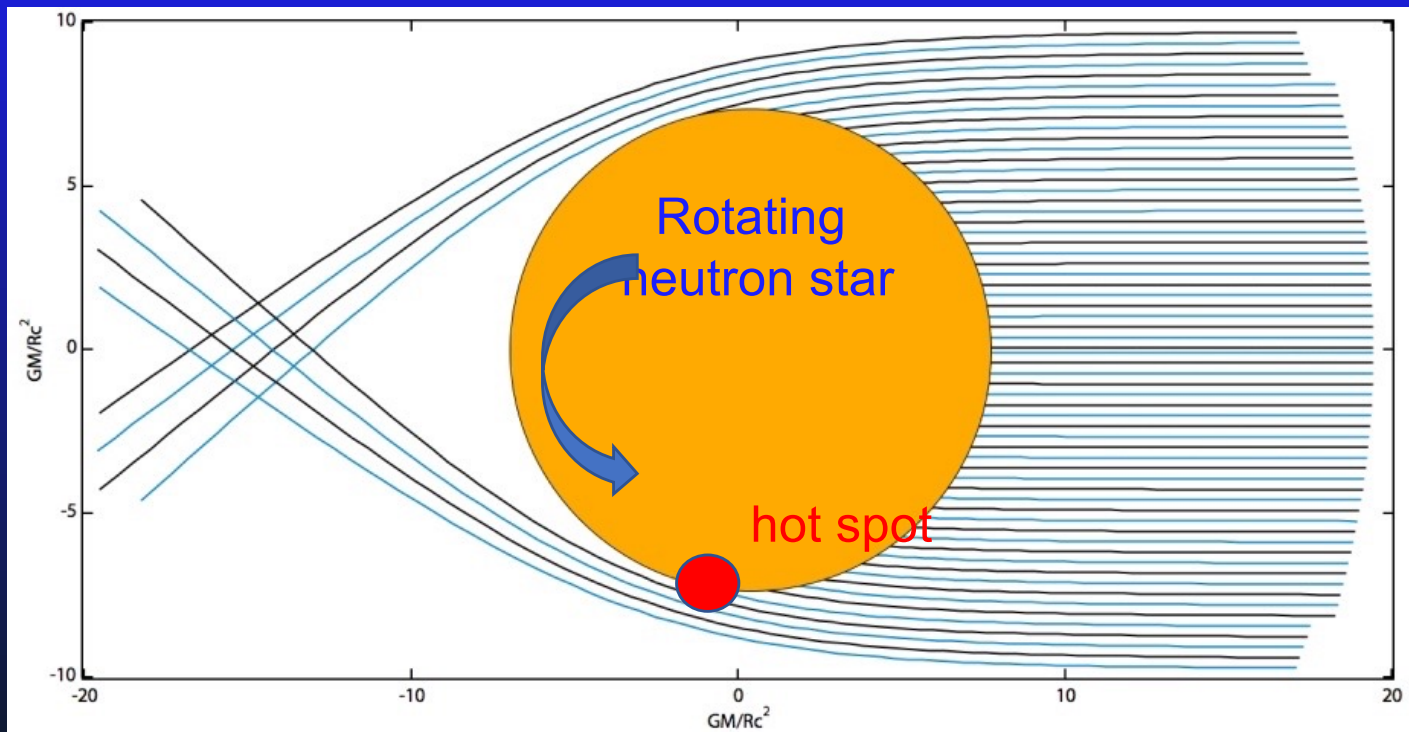
Central density for
 $2.08 M_{\text{sun}} \sim 3.6 n_0$

QHC21

APR-Togashi too soft
in nuclear regime.

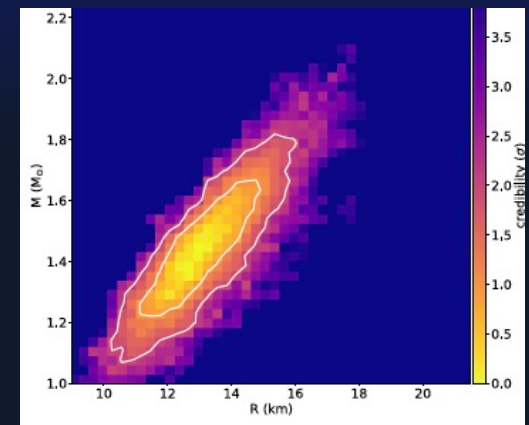


Track hot spots on neutron star. Light bending by star enables one to see spot “behind” star. Bending depends on M and R .



Observer
= NICER

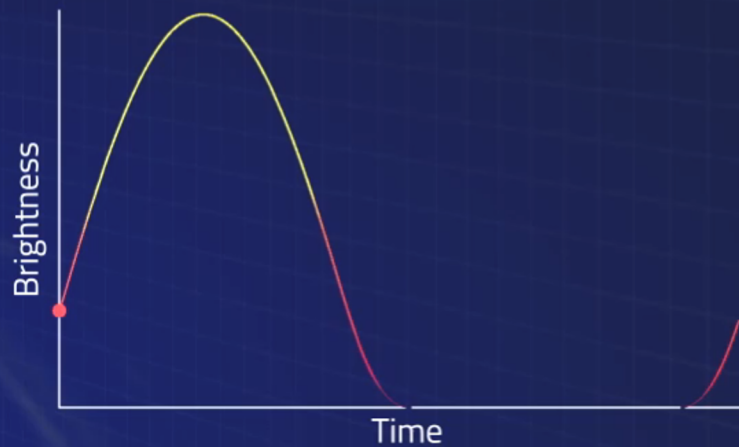
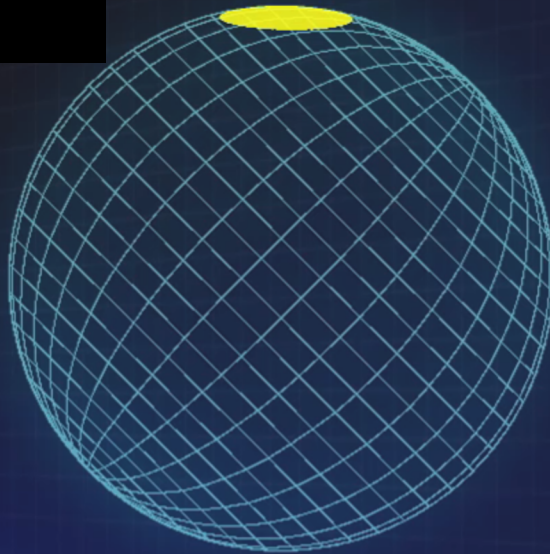
Measure amplitudes and phases
in different frequencies, construct
model of hot spots to interpret data



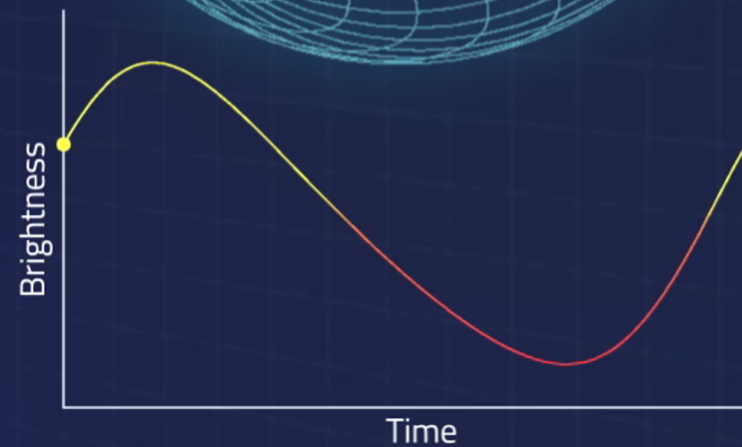
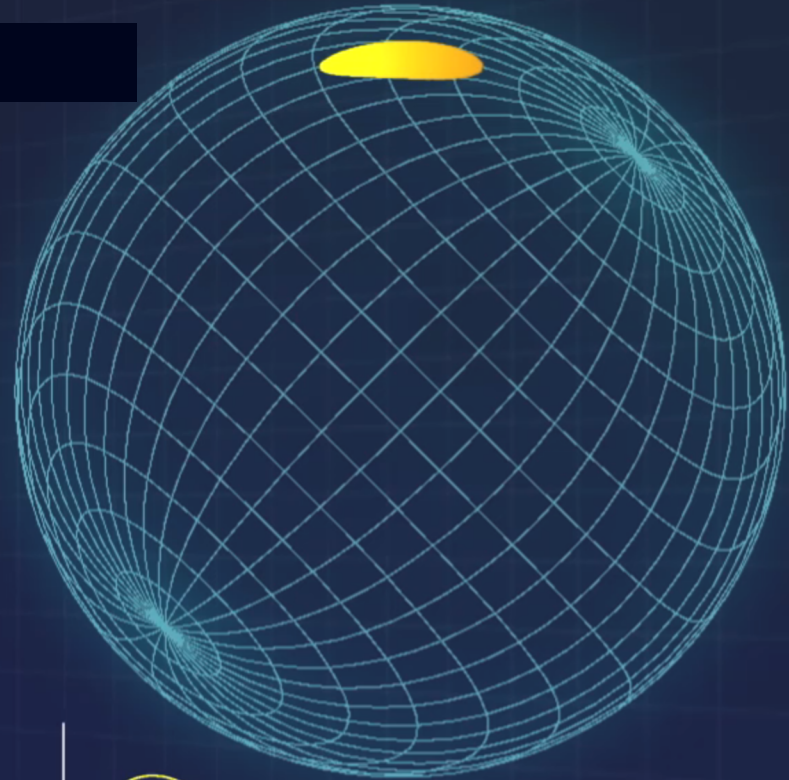
PSR J0030+0451
Mass vs Radius
Miller et al., Ap.J. (2019)

Pulse Profile Modeling (PPM)

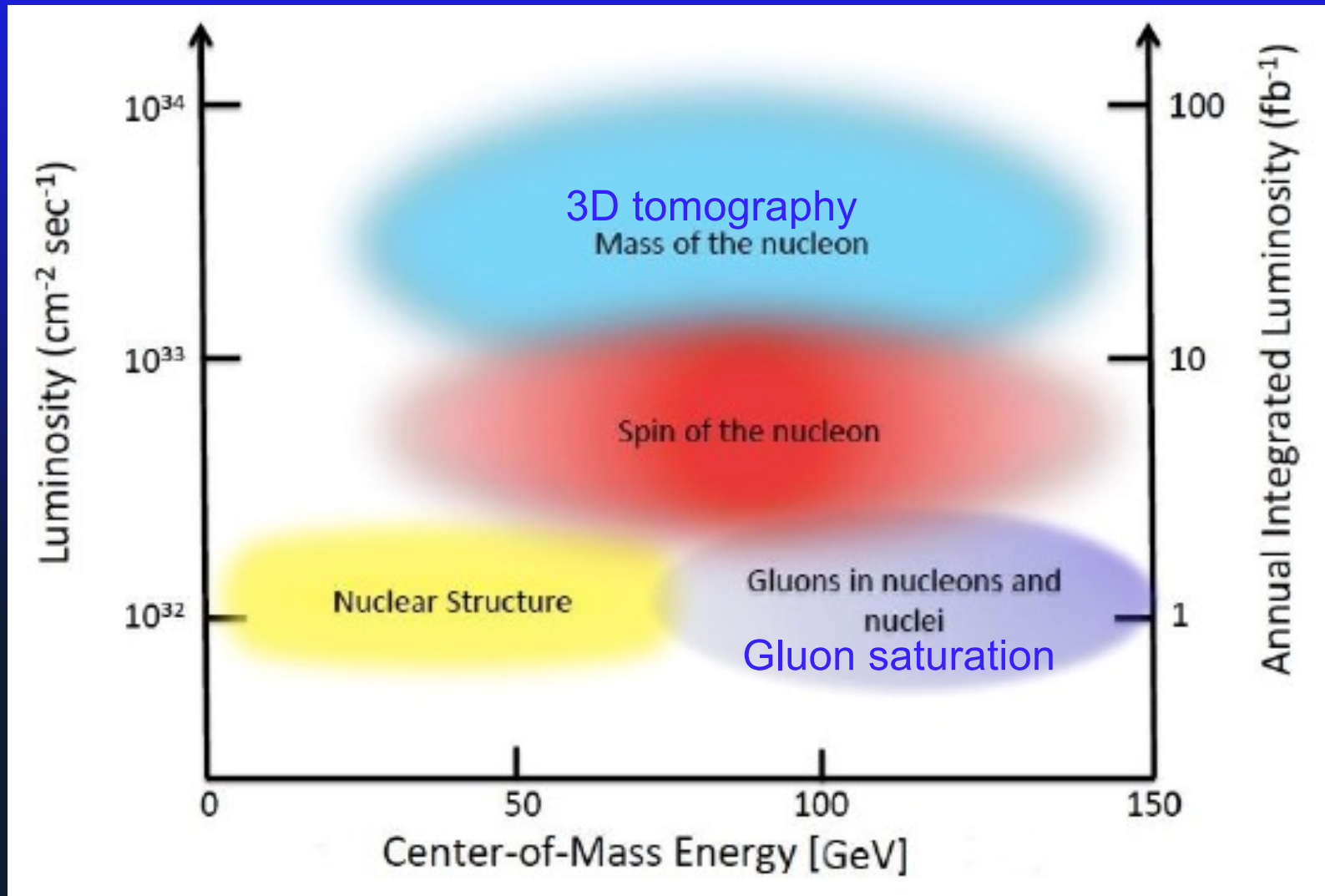
Weak
Gravity



Strong
Gravity



The accelerator



(electron-proton) c.m. energy - luminosity landscape
Luminosity measures the rate of collisions: ($L\sigma = \text{event rate}$)

Kinematic range

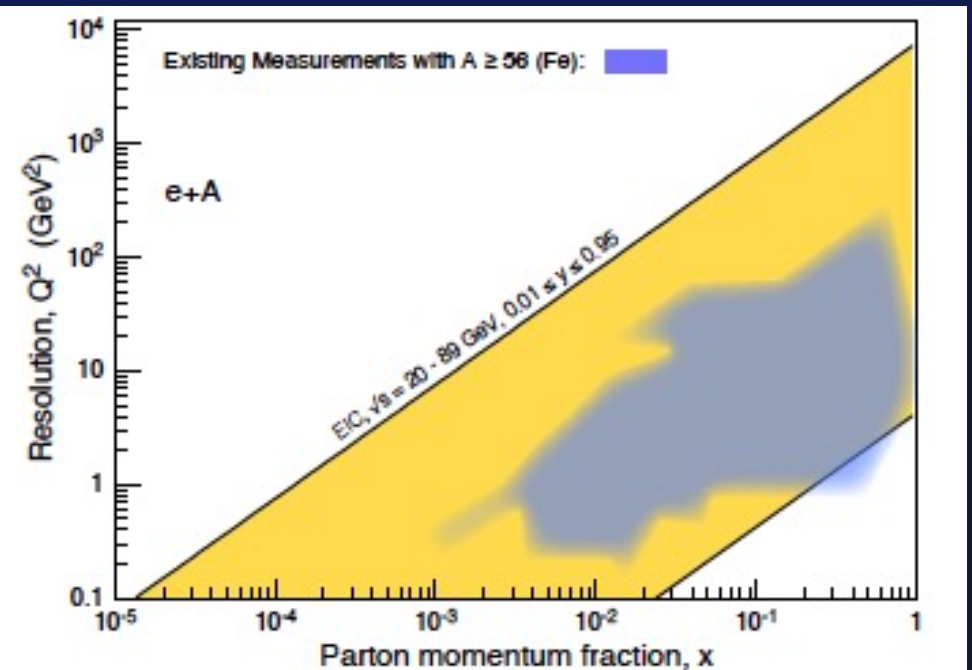
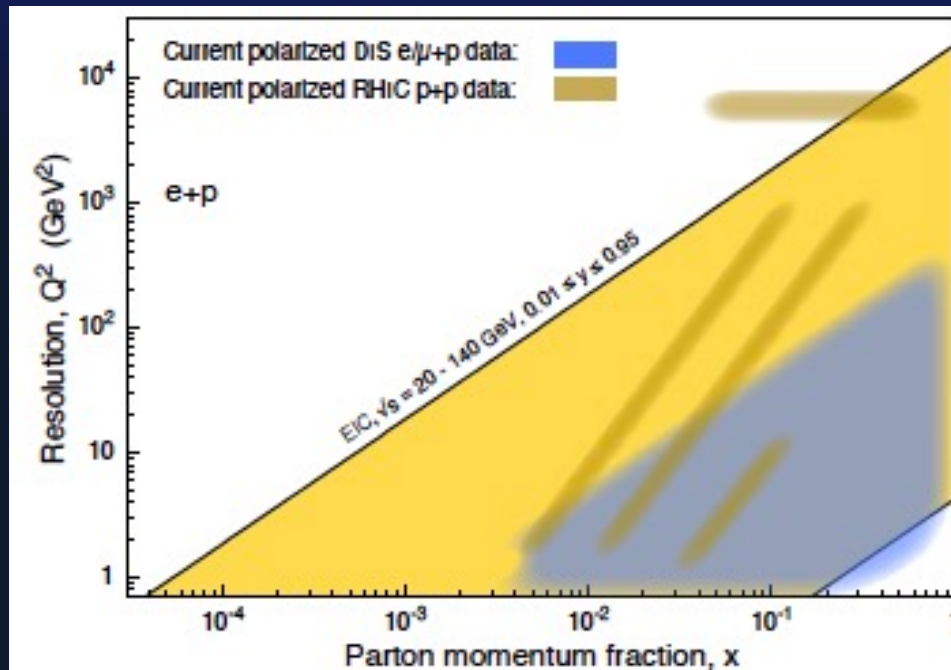
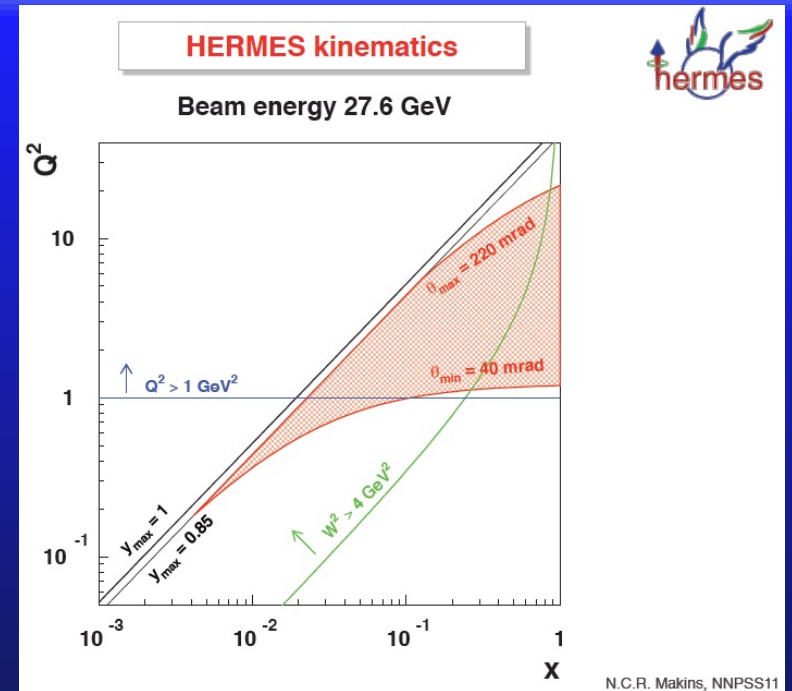
$$Q^2 = x(1 - E_e'/E_e)s \leq x s$$

\sqrt{s} = center of mass collision energy

$x \gtrsim 10^{-4}$ for $Q^2 \sim 1 \text{ GeV}^2$, $\sqrt{s} \sim 100 \text{ GeV}$

($Q^2 \sim 1 \text{ GeV}^2$ lower limit for inelastic scattering)

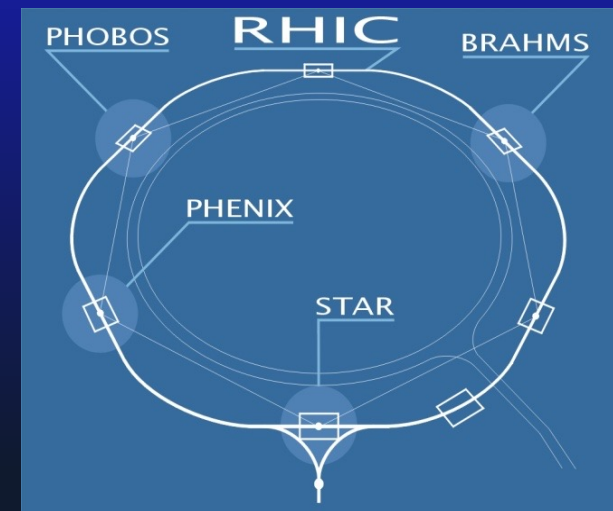
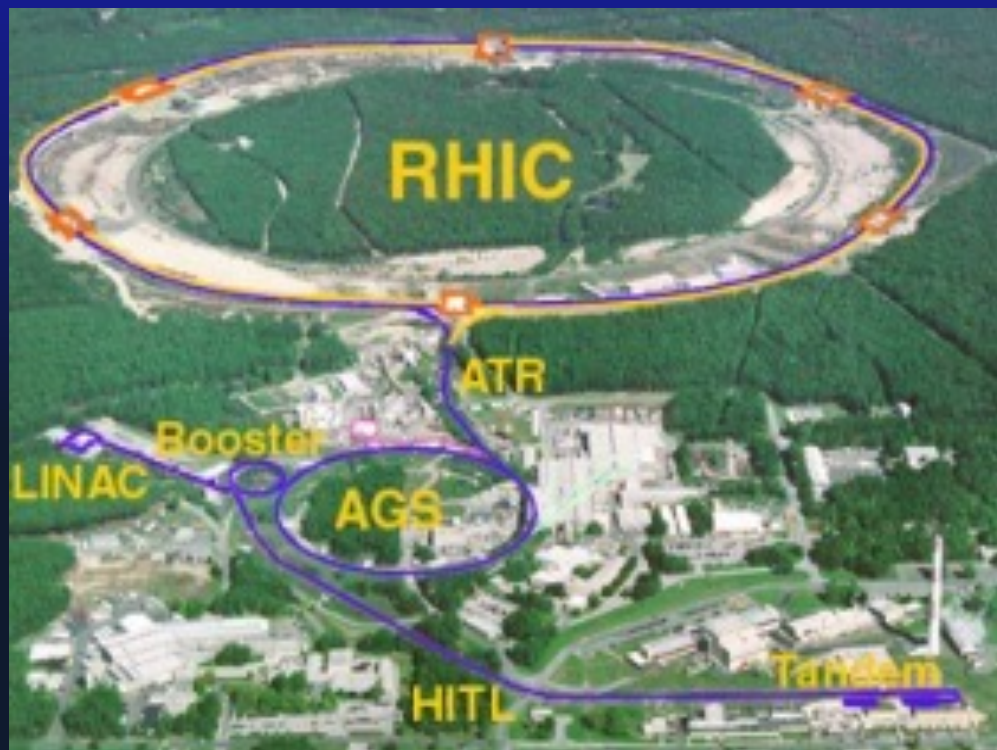
Small x region needs large beam energy



Relativistic Heavy Ion Collider (Brookhaven) since 2000 now becoming the EIC (joint project with Jefferson Lab)

RHIC: Colliding heavy ion beams 100 GeV/A

Colliding polarized proton beams $\sim 255 + 255$ GeV

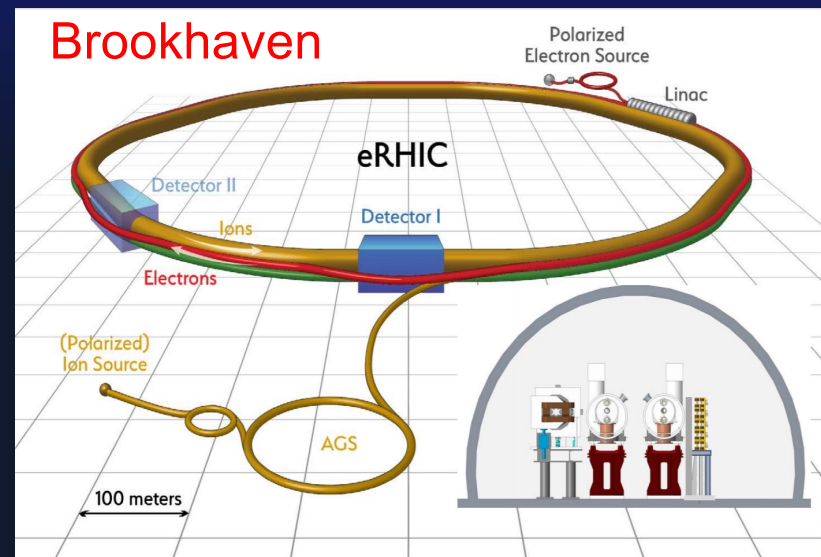
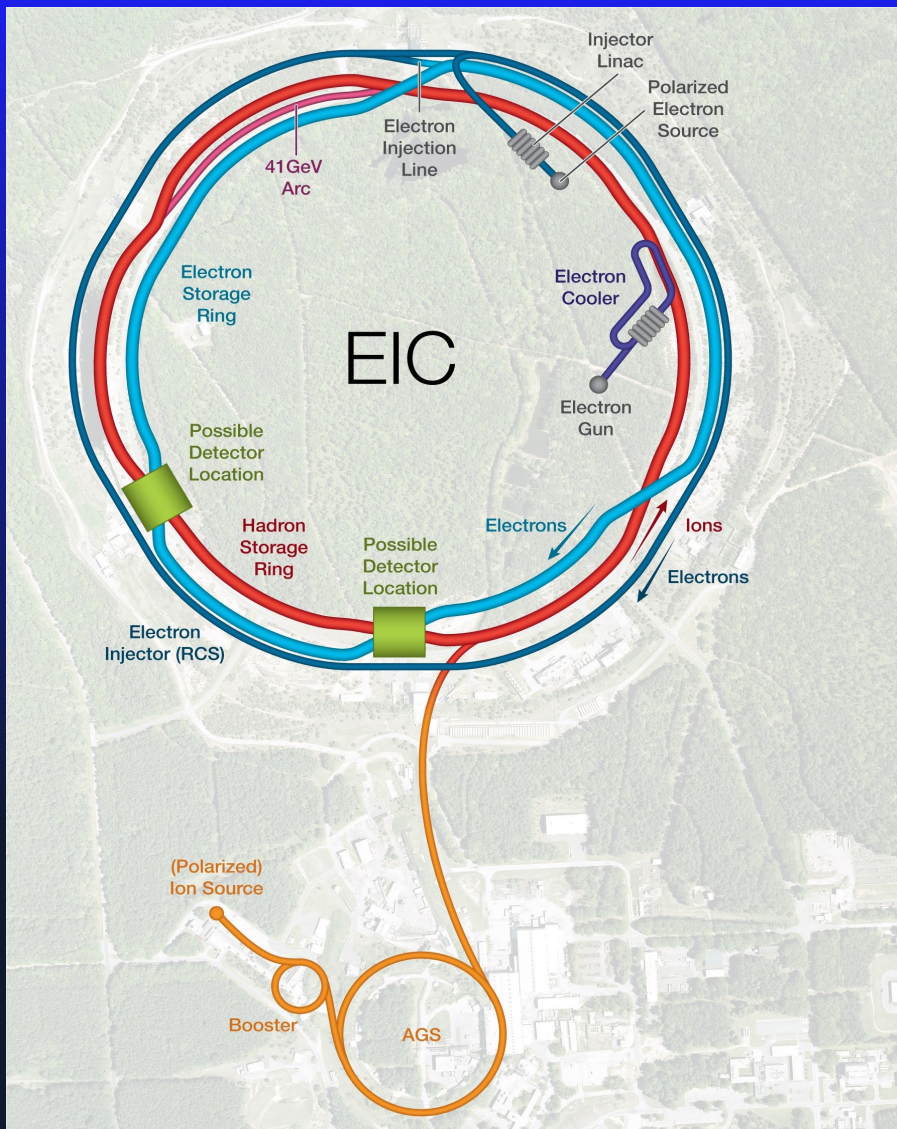


$\text{Au}(197 \times 100 \text{ GeV}) + \text{Au}(197 \times 100 \text{ GeV})$

Brookhaven eRHIC

Add 5-18 GeV electron storage ring in present RHIC tunnel.
Collide e with p to 275 GeV (vs 255 GeV now) and ions to 100 GeV/A in one RHIC ring.

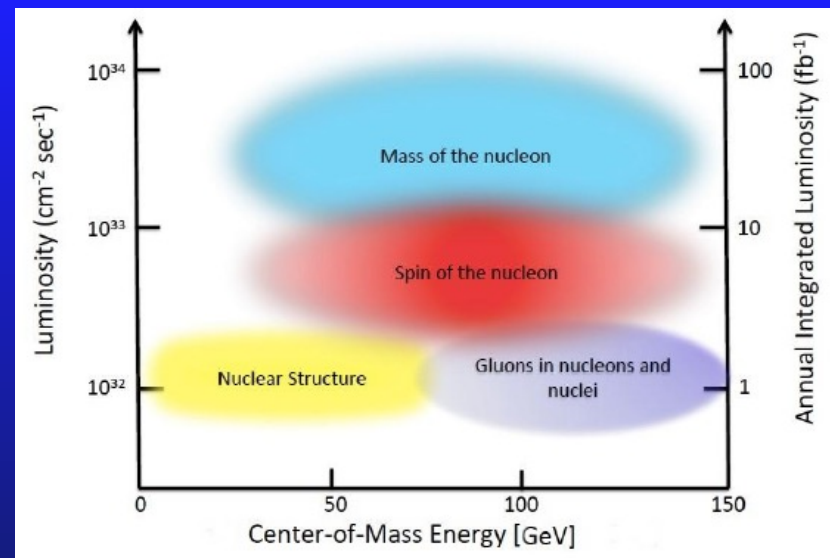
$$\sqrt{s} \simeq 2\sqrt{E_e E_p} = 75-140 \text{ GeV}$$



Accelerator requirements

To map quarks and gluons -- from nuclei (quark-gluon gas) to saturated gluonic matter -- designing for variable c.m. energy range from ~28 to ~120 GeV, upgradeable to 140 GeV.

Reach x down to 10^{-4}



Need ion beams from p to heavy stable nuclei: $A = 1 - 208$

3D imaging of gluon and sea quark distributions in nucleons and nuclei needs high luminosity:

L up to 10^{34} /cm² sec. ($\sim 10^2$ - 10^3 X HERA, and LHC at 2×10^{34})

To study correlations of gluon and sea quark distributions with spin, need polarized e-, p, (and light-ion beams) each above 70%.

Two intersection regions (IP6 and IP8): allow for two detectors

Accelerator challenges

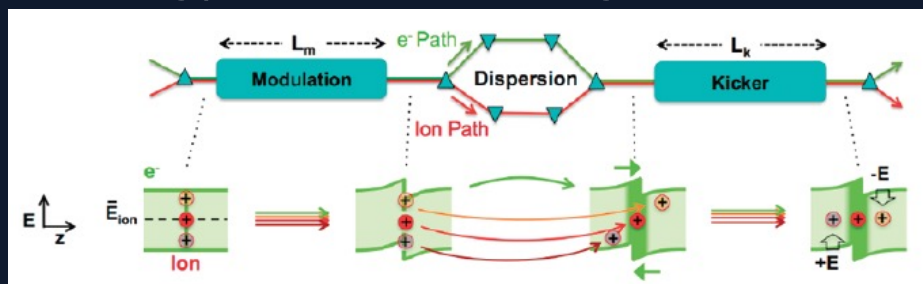
The EIC is the only large scale accelerator project in the U.S.!
Accelerator requirements beyond current technology.

High energy, spin-polarized beams colliding with high luminosity.

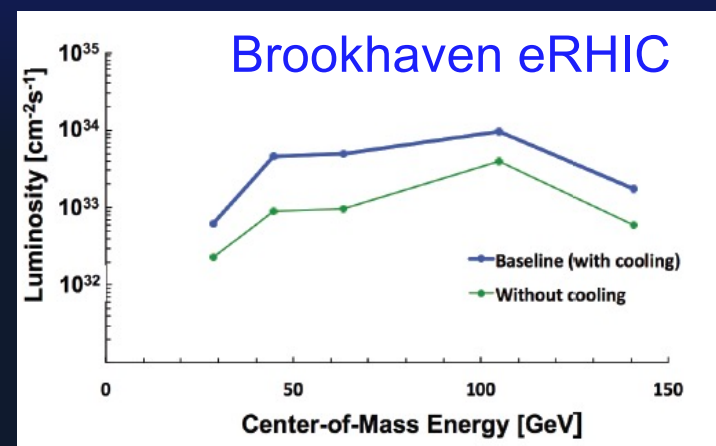
Polarized beams in a collider achieved only at

- HERA (polarized e^- or e^+ on unpolarized p) and at the
- Relativistic Heavy Ion Collider (RHIC - pp) with both proton beams polarized

Require strong **hadron beam cooling**
(focusing) to achieve high luminosity

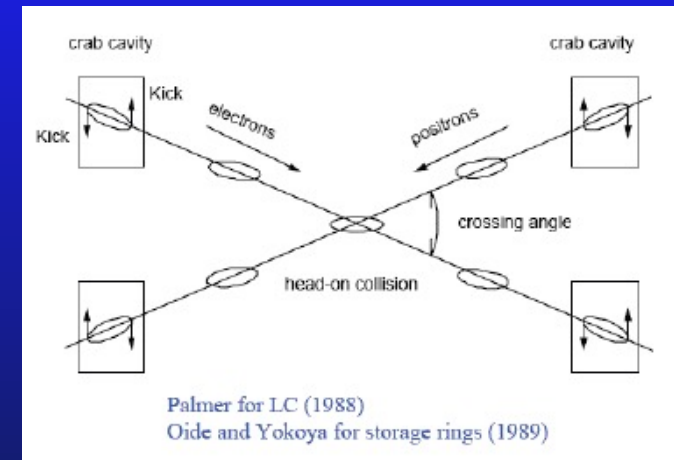
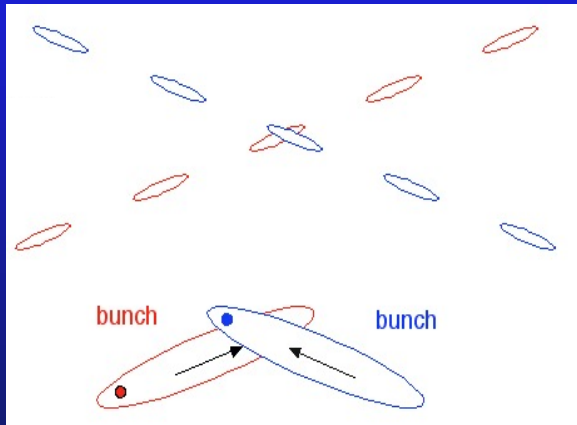


e.g., Coherent Electron Cooling: kick
slow hadrons forward, fast ones backward



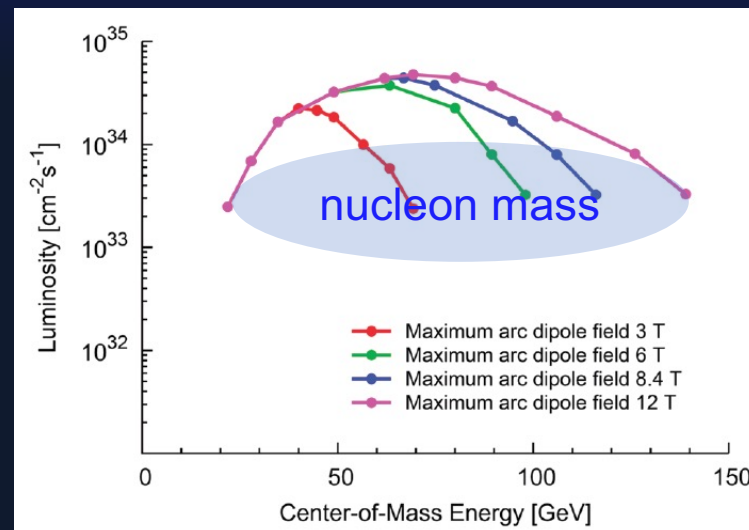
luminosity increase at
eRHIC with hadron cooling

Intersection region design: Crab crossings (beam walking sideways) to maximize collisions



Instead of bunches crossing at an angle, turn them parallel when they collide parallel. At KEK-B (e^+e^-), but never used in hadron beams.

Magnet technology

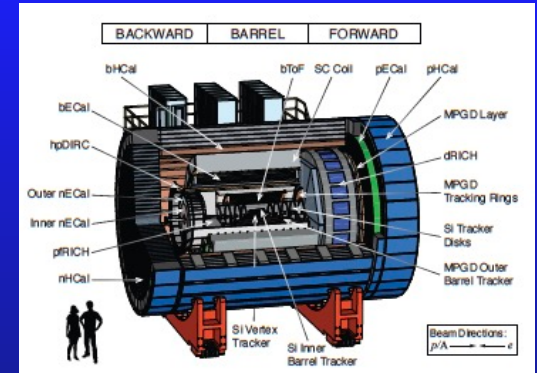


luminosity and cm ep energy vs magnetic fields (JLEIC). Red curve ~ 10 GeV e + 100 GeV p

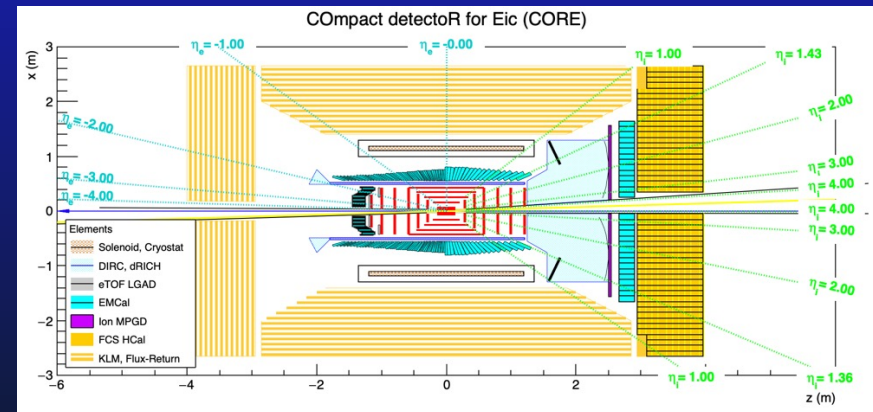
Detectors

Three proposed detectors at IP6:

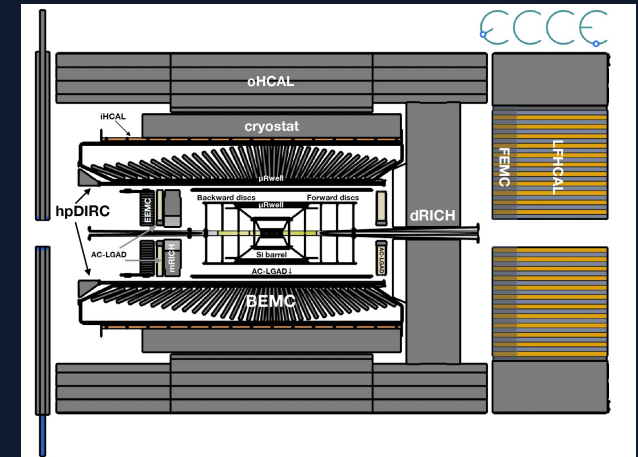
ATHENA (A Totally Hermetic Electron Nucleus Apparatus) *J. Adam et al. J. Instrument. 17, P10019 (2022)*



CORE - a COmpact detectoR for the EIC (arXiv:2209.00496)



ECCE - EIC Comprehensive Chromodynamics Experiment (arXiv:2209.02580)



EPIC (Electron Proton/Ion Collisions)
collaboration finalizing detector design.
Second detector working group formed

Timeline of EIC

Summer 2018: National Academy of Sciences report issued

Sept 2019: EIC enters U.S. budget

Dec. 19, 2019: Department of Energy first approval (CD-0)

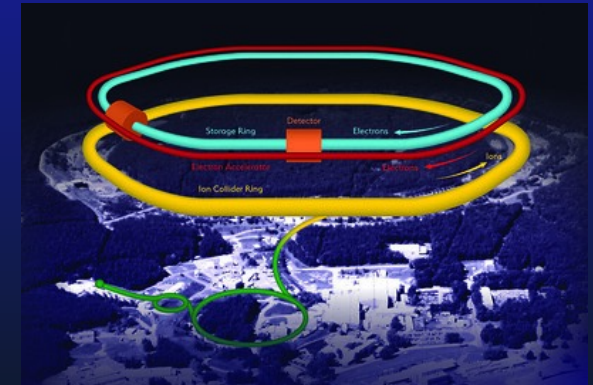
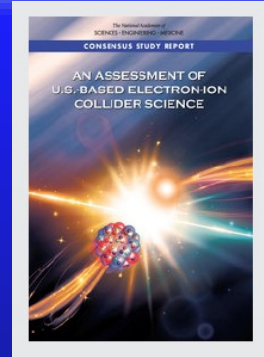
Jan. 9, 2020: Site selection -- Brookhaven

June 2021: Preliminary plan approved (CD-1)

Early 2024: Approval of final plan (CD-2)

2025: Start of construction (CD-3)

Early 2030's: Beams!! (CD-4)



Department of Energy
(DOE) approval steps

PROJECT ACQUISITION PROCESS AND CRITICAL DECISIONS					
Project Planning Phase		Project Execution Phase			Mission
Preconceptual Planning	Conceptual Design	Preliminary Design	Final Design	Construction	Operations
CD-0	CD-1	CD-2	CD-3	CD-4	
Approve Mission Need	Approve Preliminary Baseline Range	Approve Performance Baseline	Approve Start of Construction	Approve Start of Operations or Project Closeout	

EIC Users Group

EIC Users Group now has over 1200 physicists from some 250 labs worldwide.

<http://www.eicug.org/>

Link to EIC Users Group White Paper (Nov. 2022):

https://www.dropbox.com/s/yk6pvhoxd9irdz9/LRP_EIC_White_Paper_Draft.pdf?dl=0