

# Chiral magnetic waves in quark matter inside neutron stars and resulting gravitational waves

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# Outline

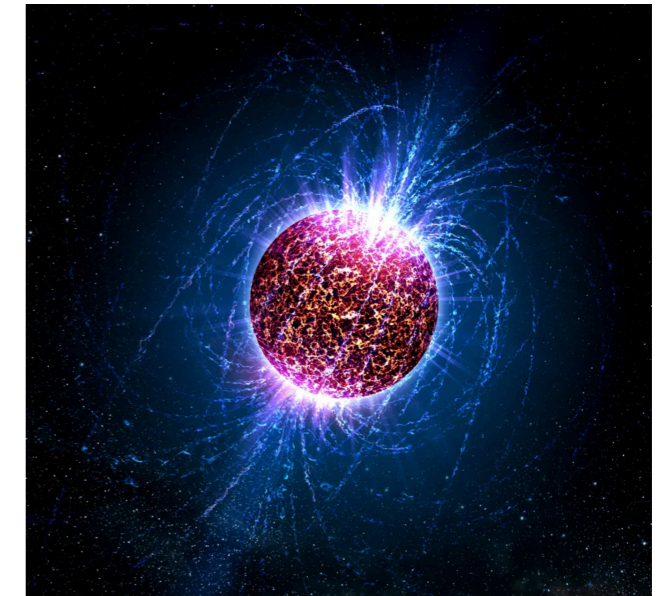
- Introduction
- Chiral transport phenomena and chiral waves
- Chiral magnetic waves in quark matter
- Gravitational waves of chiral magnetic mode
- Summary & outlook

# Introduction

# Neutron stars

- Supernova explosion occurs when nuclear fuel runs out
- Star left in the core of supernovae : Neutron star
- Typical quantities
  - mass  $\sim$  solar mass
  - radius  $\sim$  10 km
  - magnetic field (surface)  $\sim 10^{12}$ - $10^{15}$  Gauss
  - temperature  $\sim 10^6$ - $10^9$  K

density  $\sim 10^{15}$  g/cm<sup>3</sup>

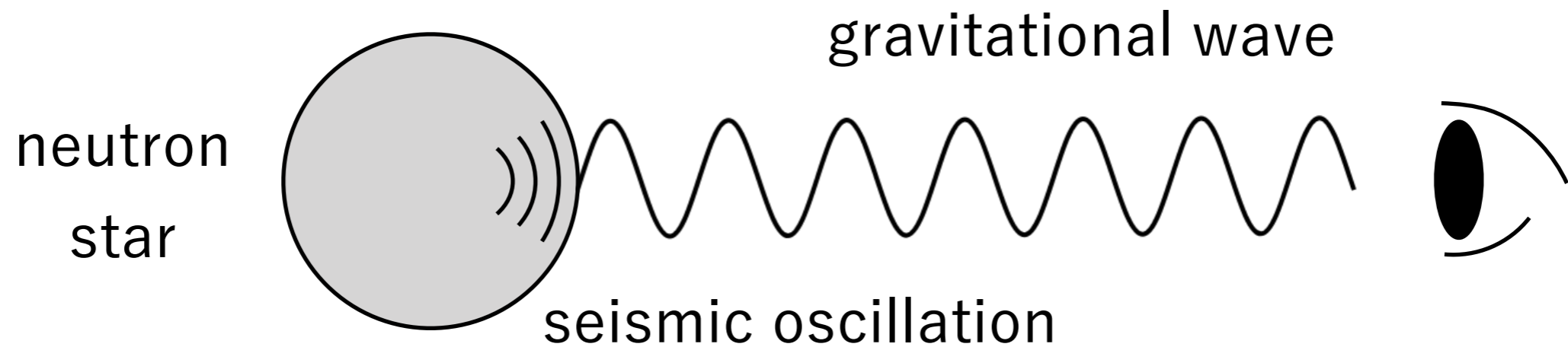


Credit: Casey Reed/  
Penn State University

Internal structure of neutron stars ?  $\rightarrow$  **Asteroseismology**

# Asteroseismology of neutron stars

- Seismic oscillations inform us about the internal structure

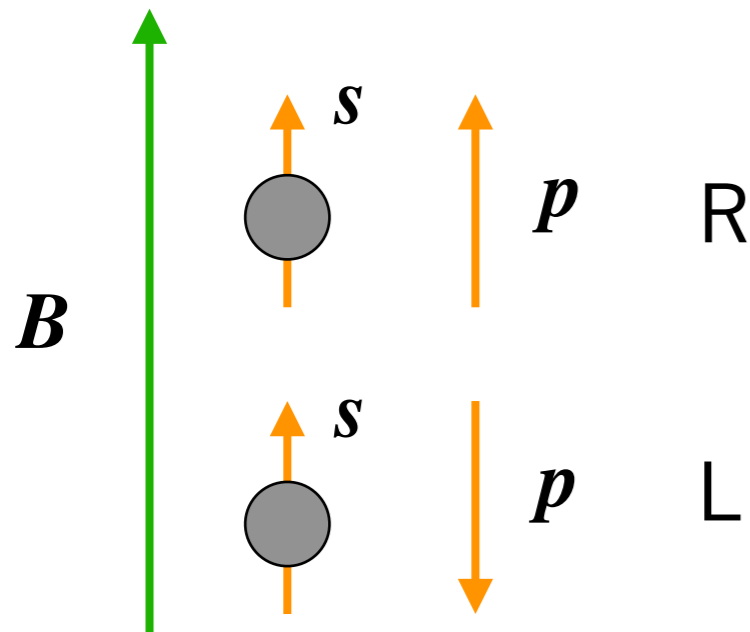


- Oscillation modes depending on physical origins
  - p-mode, g-mode, r-mode, ...

The **chirality** of quarks leads to a new type of seismic oscillation and gravitational wave

# Chiral transport phenomena and chiral waves

# Chiral magnetic/separation effect



- right-handed/left-handed current

$$\mathbf{j}_R = \frac{e\mu_R}{4\pi^2}\mathbf{B}, \quad \mathbf{j}_L = -\frac{e\mu_L}{4\pi^2}\mathbf{B}$$

- vector/chiral chemical potential

$$\mu = \frac{\mu_R + \mu_L}{2}, \quad \mu_5 = \frac{\mu_R - \mu_L}{2}$$

$$\mathbf{j} = \mathbf{j}_R + \mathbf{j}_L = \frac{e\mu_5}{2\pi^2}\mathbf{B}, \quad \mathbf{j}_5 = \mathbf{j}_R - \mathbf{j}_L = \frac{e\mu}{2\pi^2}\mathbf{B}$$

CME

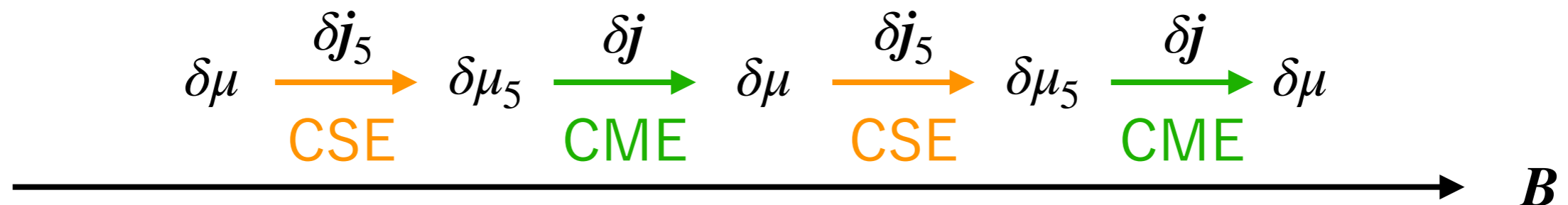
CSE

Vilenkin (1980); Nilsen, Ninomiya (1983);  
Fukushima, Kharzeev, Warringa (2008); ...

Son, Zhitnitsky (2004);  
Metlitsky, Zhitnitsky (2005); ...

# Chiral magnetic wave(CMW)

- CMW is a density wave by CME & CSE



- Wave equation & dispersion relation

$$\left[ \partial_t^2 - \left( \frac{eB}{2\pi^2\chi} \cdot \nabla \right)^2 \right] \delta n = 0 \quad \longrightarrow \quad \omega_{\text{CMW}} = \frac{eB}{2\pi^2\chi} \cdot k$$

- CMW can propagate **without** chirality imbalance at the equilibrium
- Discussed in the context of quark-gluon plasma



# Chiral magnetic waves in quark matter

Hanai, Yamamoto (2022)

# Chiral magnetic mode

- Quarks in neutron stars are relativistic
- CMW can propagate without chirality imbalance at the equilibrium
- Quark number density : no damping by electric conductivity  
 $\sim e^{-\sigma t}$
- Possible strong magnetic field in neutron stars ( $\sim 10^{18}$  Gauss)  
[Lai, Shapiro \(1991\)](#); [Cardall, Prakash, Lattimer \(2001\)](#); [Ferrer, et al. \(2010\)](#); ...

New oscillation mode (**CM-mode**) can appear  
in the quark matter inside neutron stars

# Chiral magnetic wave (CMW)

- Quark mass term

$$\mathcal{L}_{\text{mass}} = - m_q \bar{\psi}_L \psi_R + \text{h.c.} \quad \longrightarrow \quad \text{chirality flipping}$$

mixing

- Chirality flipping rate

$$\Gamma_{\text{flip}} \sim \frac{\alpha_s^2 m_q^2}{\bar{\mu}_q^2 q_D} T^2$$

(The similar expression is valid for QED process)

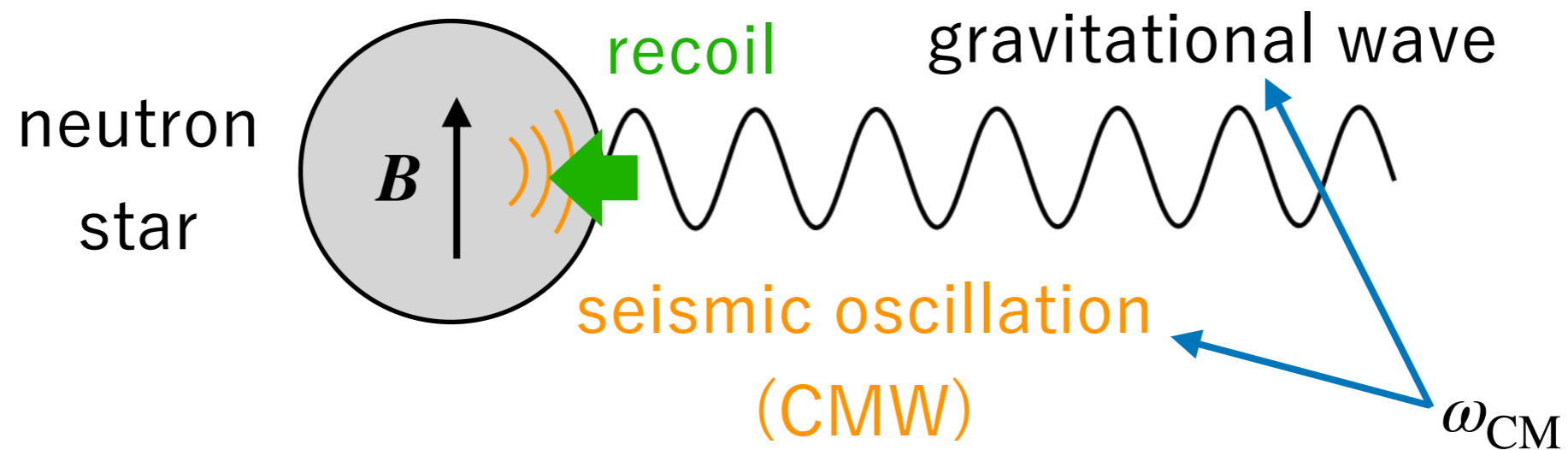
CMW is damped by the chirality flipping

# Gravitational waves of chiral magnetic mode

Hanai, Yamamoto (2022)

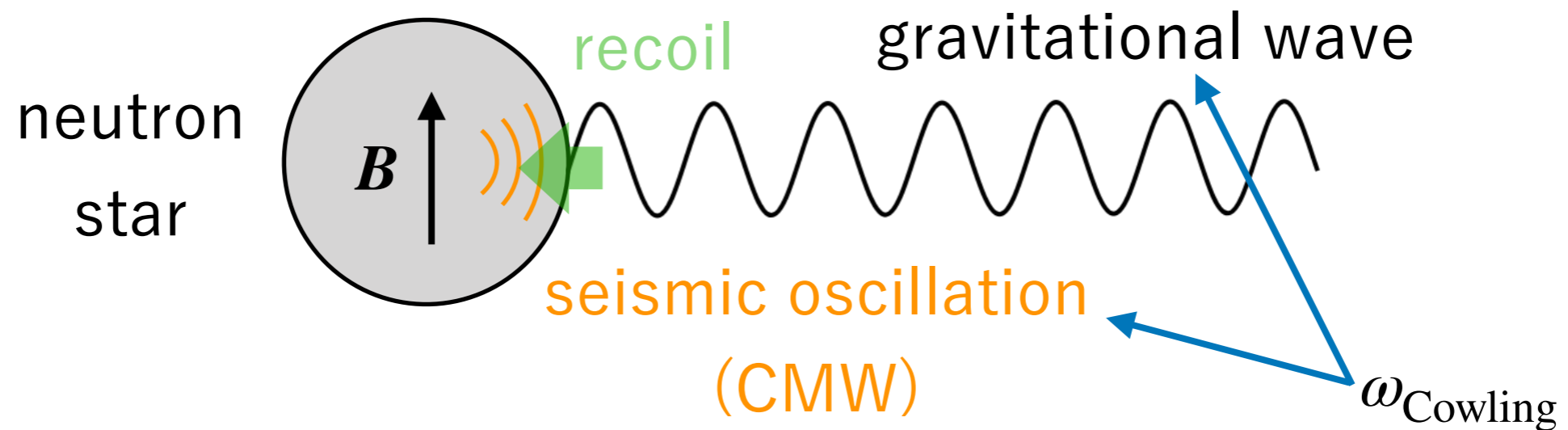
# Cowling approximation

- Newton gravity : ignore the recoil by gravitational potential  
Cowling (1941)
- General relativity : ignore the recoil by gravitational waves



# Cowling approximation

- Newton gravity : ignore the recoil by gravitational potential  
Cowling (1941)
- General relativity : ignore the recoil by gravitational waves



## Cowling approximation

$$\omega_{\text{Cowling}} \simeq \omega_{\text{CM}}$$

# Frequency of CM-mode

- Dispersion relation of the CM-mode

$$\omega_{\text{CM}} \simeq \underbrace{V_{\text{CM}} |k_z|}_{\text{chirality flipping}} - i \underbrace{\frac{\Gamma_{\text{flip}}}{2}}_{\text{diffusion}} - i e^{\lambda-\rho} D k_z^2 \quad V_{\text{CM}} \equiv e^{\lambda-\rho} \frac{N_c e B}{6\pi^2 \chi}$$

- The range of the frequency

$$\text{chirality flipping } \underbrace{\frac{\Gamma_{\text{flip}}}{4\pi}}_{\text{chirality flipping}} \ll f_{\text{CM}} \ll \underbrace{\frac{V_{\text{CM}}^2}{2\pi D}}_{\text{diffusion}} = \frac{3V_{\text{CM}}^2}{2\pi\tau} \quad \text{diffusion}$$

- Relaxation time

$$\frac{1}{\tau} \sim \# \alpha_s^2 \underbrace{\frac{T^2}{q_D}}_{\text{usual Fermi liquid}} + \# \alpha_s^2 \underbrace{\frac{T^{5/3}}{q_D^{2/3}}}_{\text{Landau damping}}$$

Heiselberg, Pethick (1993)

usual Fermi liquid

Landau damping

# Estimate of the frequency

- Settings :  $\alpha_s \simeq 0.5$ ,  $\bar{\mu} \simeq 500$  MeV,  $m_{u,d} \sim 3-5$  MeV, 2SC

- The range of the frequency

$$10 \text{ Hz} \left( \frac{T}{10^6 \text{ K}} \right)^2 \ll f_{\text{CM}} \ll 10^3 \text{ Hz} \left( \frac{B}{10^{18} \text{ Gauss}} \right)^2 \left( \frac{T}{10^6 \text{ K}} \right)^{5/3}$$

- Dependence of the physical parameters

$$f_{\text{CM}} \sim 10^2 \text{ Hz} \left( \frac{B}{10^{18} \text{ Gauss}} \right) \left( \frac{\bar{\mu}}{500 \text{ MeV}} \right)^{-2} \left( \frac{k}{10^{-5} \text{ /cm}} \right)$$

CM-mode can be a new probe of the magnetic field  
& quark matter in neutron stars



# Amplitude of the gravitational wave

- Effective amplitude of gravitational waves

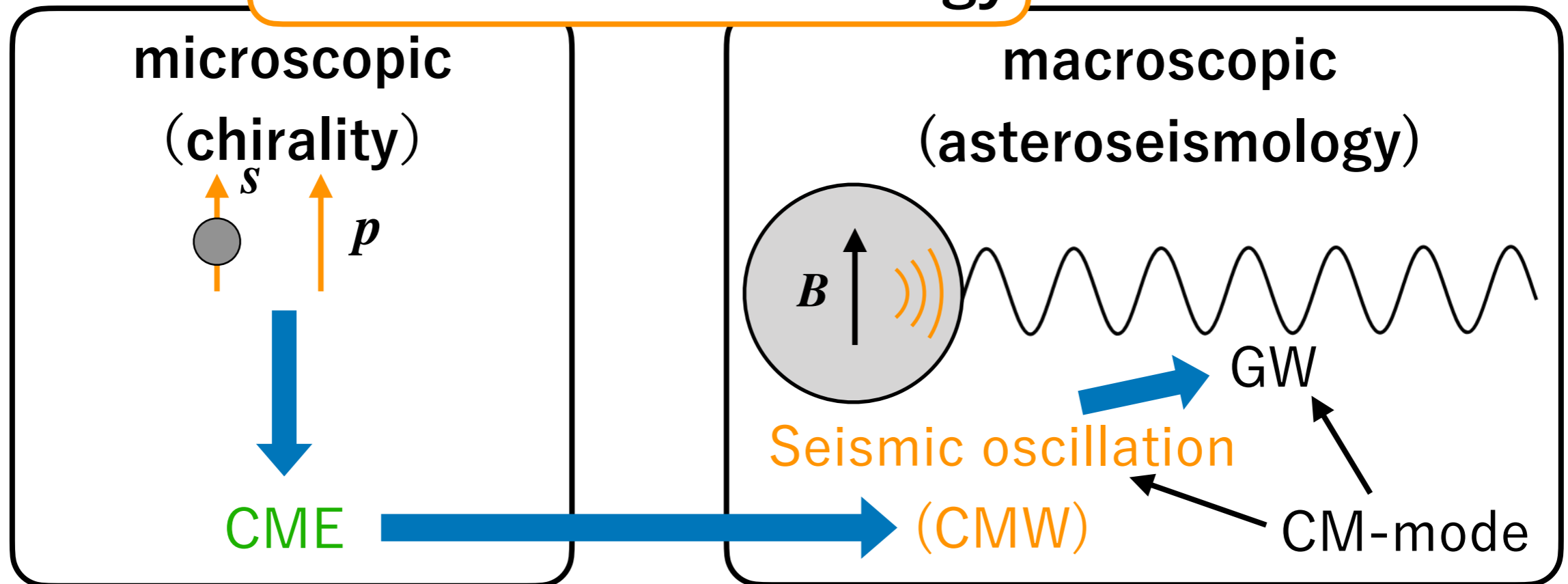
$$h \sim \frac{1}{d} \sqrt{\frac{GE_{\text{GW}}}{f}}$$

- The released energy  $E_{\text{GW}}$  depends on events
- The amplitude of the gravitational waves
  - cf. giant flare  $\sim 10^{46}$  erg, radius of our galaxy  $\sim 10$  kpc

$$h_{\text{CM}} \sim 10^{-21} \left( \frac{E_{\text{GW}}}{10^{44} \text{ erg}} \right)^{1/2} \left( \frac{f_{\text{CM}}}{10^2 \text{ Hz}} \right)^{-1/2} \left( \frac{d}{10 \text{ kpc}} \right)^{-1}$$

# Summary & outlook

## Chiral asteroseismology



- CM-mode : a new probe of magnetic fields & quark matter
- CV-mode : another mode caused by CVW in supernovae
- Contribution of magnetic fields to  $\Gamma_{\text{flip}}$  ,  $D$  ?

# Appendix

# Parity of transport

- Electric fields cause electric currents (Ohm's law)

$$ej = \sigma E \quad \longrightarrow \quad e(-j) = \sigma(-E)$$

parity transformation

- Do magnetic fields cause electric currents ?

$$ej = \sigma_m B \quad \longrightarrow \quad e(-j) = \sigma_m B$$

parity transformation

If  $\sigma_m \rightarrow -\sigma_m$  is satisfied, the current can be generated.

# Chiral magnetic wave (CMW)

- Density fluctuations:

$$\mu = \bar{\mu} + \delta\mu, \quad \delta\mu = \frac{1}{\chi}\delta n$$

( $\chi$ : susceptibility)

$$\mu_5 = \bar{\mu}_5 + \delta\mu_5, \quad \delta\mu_5 = \frac{1}{\chi}\delta n_5$$

- Current fluctuations caused by CME & CSE

$$\begin{aligned} \mathbf{j} &= \frac{e\mu_5}{2\pi^2}\mathbf{B} \\ \mathbf{j}_5 &= \frac{e\mu}{2\pi^2}\mathbf{B} \end{aligned} \quad \longrightarrow \quad \begin{aligned} \delta\mathbf{j} &= \frac{e\mathbf{B}}{2\pi^2\chi}\delta n_5 \\ \delta\mathbf{j}_5 &= \frac{e\mathbf{B}}{2\pi^2\chi}\delta n \end{aligned}$$

# Chiral magnetic wave (CMW)

- Linearized continuity equations

$$\partial_t \delta n + \nabla \cdot \delta \mathbf{j} = 0, \quad \partial_t \delta n_5 + \nabla \cdot \delta \mathbf{j}_5 = 0$$

- Wave equation of the CMW

$$\left[ \partial_t^2 - \left( \frac{e\mathbf{B}}{2\pi^2\chi} \cdot \nabla \right)^2 \right] \delta n = 0 \quad \longrightarrow \quad \omega_{\text{CMW}} = \frac{e\mathbf{B}}{2\pi^2\chi} \cdot \mathbf{k}$$

- Dispersion relation is independent of  $\bar{\mu}_5$
- Discussed in the context of quark-gluon plasma

# CMW of electron matter

- Electrons in neutron stars and supernovae are relativistic
- For electrons, charge and number density are not independent
- CMW has a gap by the chiral anomaly
- Charge density : damped by conductivity [Shovkovy, Rybalka, Dorbar \(2018\)](#)  
 $\sim e^{-\sigma t}$

$$\Gamma_{\text{flip}} \ll \frac{eB}{2\pi^2\chi_e} \frac{1}{l_{\text{mfp}}} \ll \sigma$$



CMW can not appear

# Scattering in electron matter

- Candidates of scattering processes :
  - electron-electron scattering
  - electron-proton scattering
  - electron-photon scattering
- e: degenerate, p: degenerate (NS) or non-degenerate (SN)
- photon density  $\sim T^3 \ll$  electron density  $\sim \bar{\mu}_e^3$
- Typical energy of electron:  $\bar{\mu}_e \ll m_p$

The e-p scattering (Rutherford scattering) is dominant.



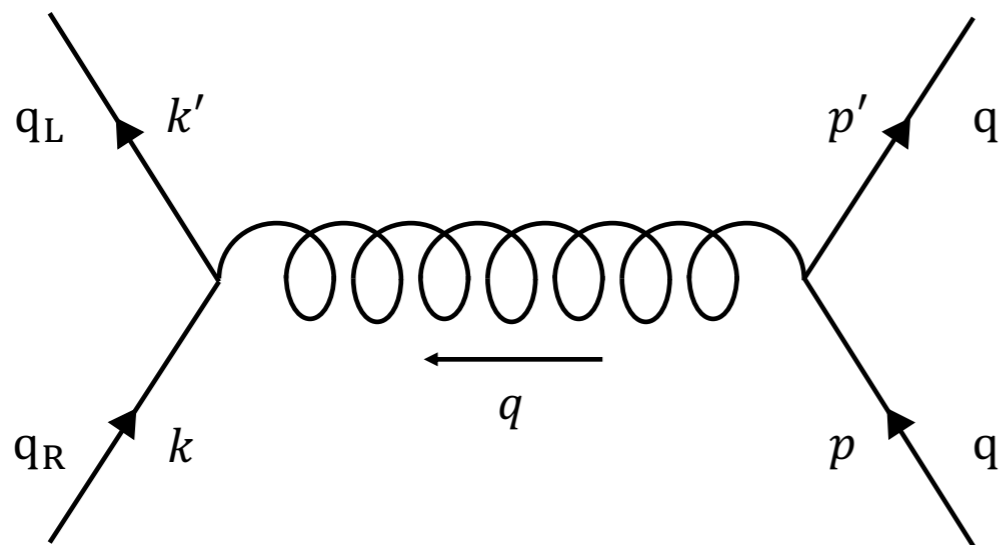
# Chirality flipping

- Quark mass term

$$\mathcal{L}_{\text{mass}} = - m_q \bar{\psi}_L \psi_R + \text{h.c.} \quad \longrightarrow \quad \text{chirality flipping}$$

mixing

- Quark-quark scattering



$$|M_1|^2 \sim \left| \frac{j_{-+}^0 j_h^0}{q^2 + \Pi_1} \right|^2$$

longitudinal

$$|M_t|^2 \sim \left| \frac{\mathbf{j}_{-+,t} \cdot \mathbf{j}_{h,t}}{-(q^0)^2 + q^2 + \Pi_t} \right|^2$$

transverse

# Chirality flipping

- Chirality flipping rate

$$\Gamma_{\text{flip}} \equiv -\frac{\dot{n}_5}{n_5}$$

Boltzmann equation  
 $\dot{f}_{R/L} = \mp C(k, t)$

$$\sim \# \frac{\alpha_s^2 m_q^2}{\bar{\mu}_q^2 q_D} T^2 + \# \frac{\alpha_s^2 m_q^2}{\bar{\mu}_q^4 q_D^{2/3}} T^{11/3} - \# \frac{\alpha_s^2 m_q^2}{\bar{\mu}_q^3 q_D^2} T^4 + \# \frac{\alpha_s^2 m_q^2}{\bar{\mu}_q^2 q_D^{10/3}} T^{13/3}$$

**longitudinal**

(usual Fermi liquid)

**transverse**

(Landau damping)

CMW is damped by the chirality flipping

# 2-flavor color superconductivity(2SC)

- Quark condensation

$$\langle q_{Li}^a q_{Lj}^b \rangle = \epsilon^{ab3} \epsilon_{ij} \Delta$$

One of colors is  
unpaired

$(a, b : \text{color}, i, j : \text{flavor})$

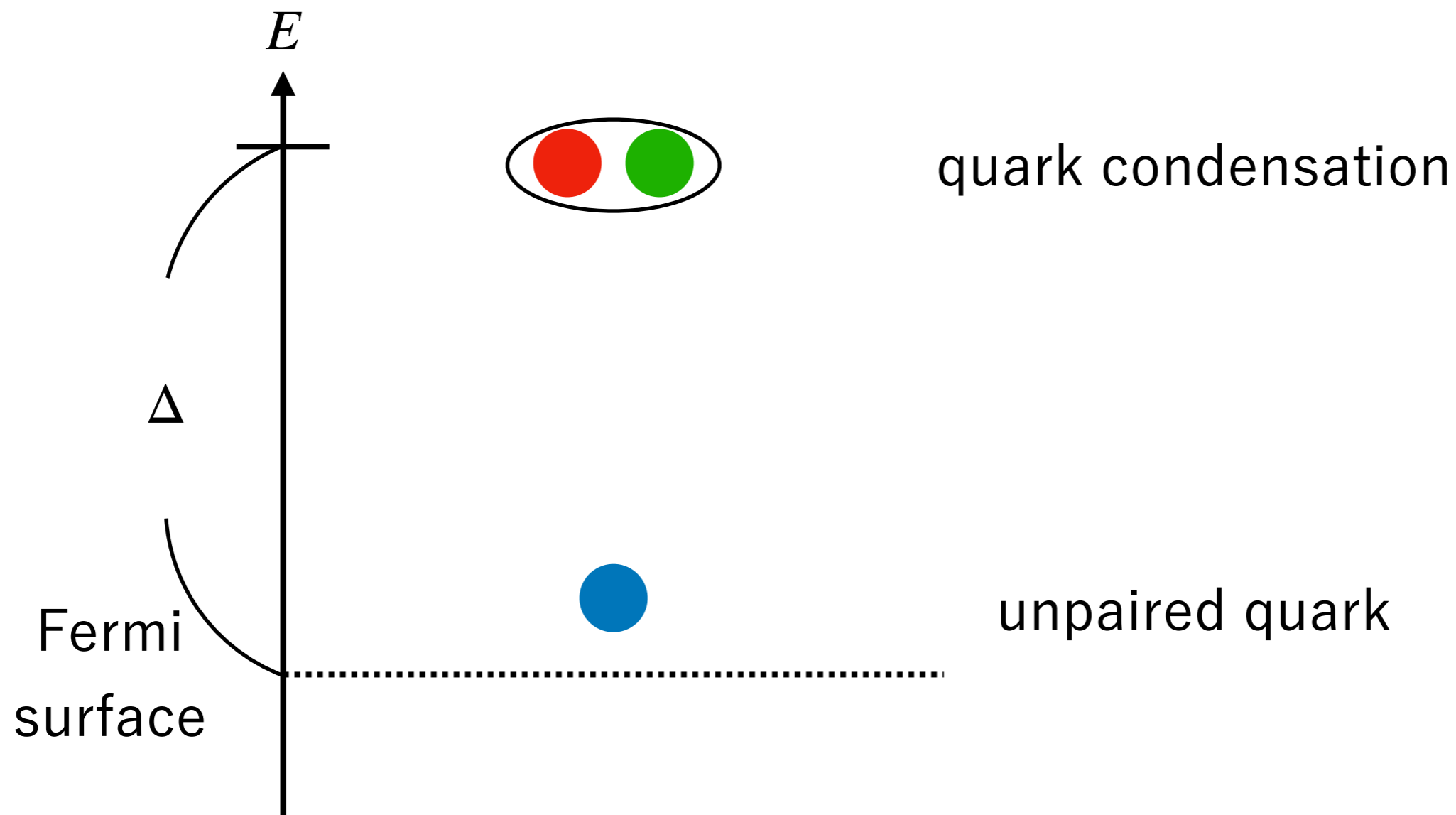
- Pattern of symmetry breaking

$$\begin{aligned} & \underline{SU(3)_C} \times SU(2)_L \times SU(2)_R \times U(1)_B \\ & \rightarrow \underline{SU(2)_C} \times \underline{SU(2)_L} \times \underline{SU(2)_R} \times U(1)_Q \end{aligned}$$

chiral symmetry  
is unbroken

# 2-flavor color superconductivity(2SC)

- Focusing on the energy smaller than the gap  $\Delta$



# Analysis of gravitational waves

- Gravitational waves are the fluctuations of spacetime
- Metric in cylindrical coordinates

$$\mathbf{B} = B\mathbf{e}_z$$

$$g_{\alpha\beta} = \bar{g}_{\alpha\beta} + \underline{h_{\alpha\beta}}$$

gravitational waves

$$\bar{g}_{\alpha\beta}dx^\alpha dx^\beta = -e^{2\lambda}dt^2 + e^{2\nu}dr^2 + r^2d\phi^2 + e^{2\rho}dz^2$$

- It is hard to solve the linearized Einstein equations



Cowling approximation

# Fundamental equations

- Metric in cylindrical coordinates

$$g_{\alpha\beta} = \bar{g}_{\alpha\beta} + \cancel{h_{\alpha\beta}} \quad \text{Cowling approximation}$$

$$\bar{g}_{\alpha\beta} dx^\alpha dx^\beta = -e^{2\lambda} dt^2 + e^{2\nu} dr^2 + r^2 d\phi^2 + e^{2\rho} dz^2$$

- Current fluctuations by CME & CSE with diffusion

$$\delta \mathbf{j} = \frac{e\mathbf{B}}{2\pi^2\chi} \delta n_5 - D \nabla \delta n, \quad \delta \mathbf{j}_5 = \frac{e\mathbf{B}}{2\pi^2\chi} \delta n - D \nabla \delta n_5$$

- Linearized continuity equations

$$\nabla_\alpha \delta j^\alpha = 0, \quad \nabla_\alpha \delta j_5^\alpha = \frac{-\Gamma_{\text{flip}} e^{-\lambda} \delta n_5}{\text{chirality flipping}}$$

# Relaxation time of quark matter

- Medium effects
  - longitudinal: Debye screening (usual Fermi liquid like)
  - transverse: Landau damping

$$\frac{1}{\tau} \sim \# \alpha_s^2 \frac{T^2}{q_D} + \# \alpha_s^2 \frac{T^{5/3}}{q_D^{2/3}}$$

Heiselberg, Pethick (1993)

usual Fermi liquid

$$(q_{\text{IR}} \sim q_D)$$

Landau damping

$$(q_{\text{IR}} \sim (q_D^2 T)^{1/3})$$

# CME & chiral anomaly

Nielsen, Ninomiya (1983)

- Energy necessary for the chirality imbalance

$$\mu_5 \frac{dQ_5}{dt} = \mu_5 \int d^3x \frac{e^2}{2\pi^2} \mathbf{E} \cdot \mathbf{B} \quad \text{chiral anomaly}$$

- Energy is supplied by electric currents

$$\mu_5 \frac{dQ_5}{dt} = \int d^3x \mathbf{j} \cdot \mathbf{E}$$

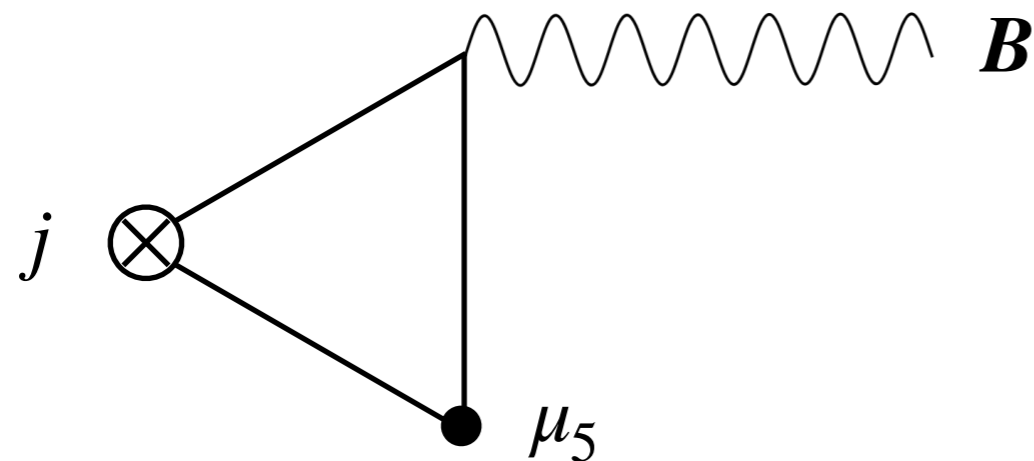
- For arbitrary electric fields,

$$\longrightarrow \mathbf{j} = \frac{e\mu_5}{2\pi^2} \mathbf{B}$$



# 2 flavor VS 3 flavor

- Coefficient of CME



$$j \propto \text{tr}(VAQ)\mu_5 B$$

$$V = \mathbf{1}_{N_f},$$

$$A = \mathbf{1}_{N_f},$$

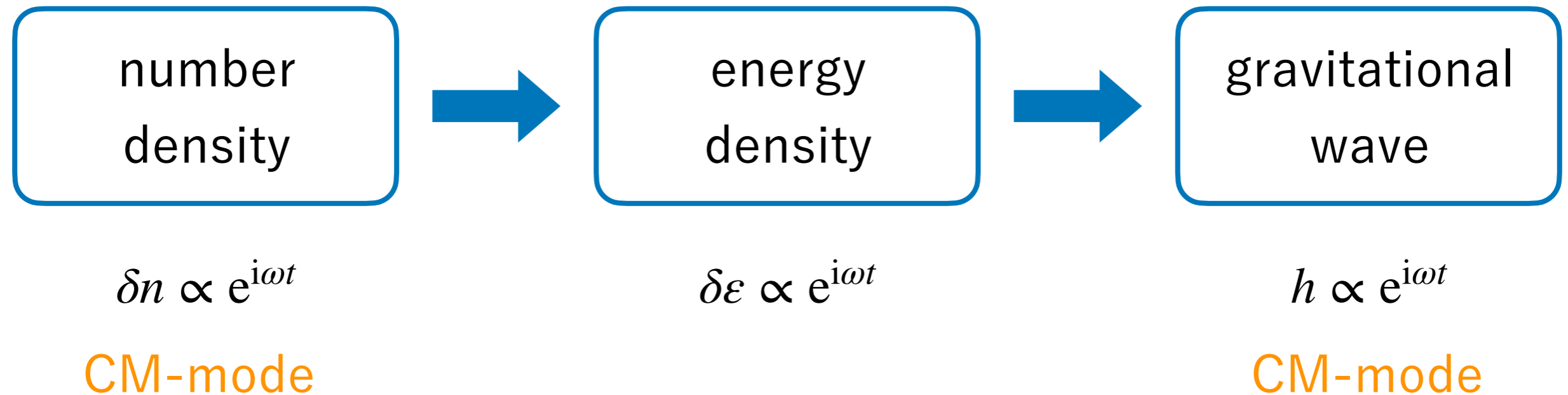
$$Q = \begin{cases} \text{diag} \left( +\frac{2}{3}, -\frac{1}{3} \right) & (N_f = 2) \\ \text{diag} \left( +\frac{2}{3}, -\frac{1}{3}, -\frac{1}{3} \right) & (N_f = 3) \end{cases}$$

Since  $j \propto \text{tr}(Q) = 0$  ( $N_f = 3$ ),  
the CMW exists only for two-flavor case.

Khazzev, Son (2011)

# Generation of gravitational waves

- The CMW is the fluctuation of quark number density



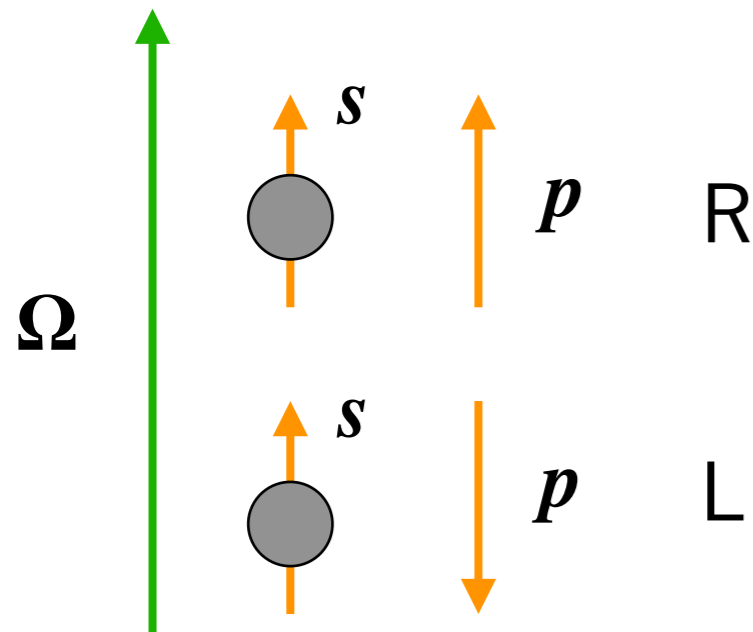
# Contribution of metric functions

- The metric functions include the correction of general relativity
- Typical metric functions in the case of neutron stars

$$\lambda, \rho \sim \frac{GM}{R} \sim 10^{-1}$$

- It is reasonable to set the metric functions to be 0

# Chiral vortical effect



- Correspondence between  $\mathbf{B}$  and  $\boldsymbol{\Omega}$

$$\underbrace{e\mathbf{v} \times \mathbf{B}}_{\text{Lorentz Force}} \leftrightarrow \underbrace{2m\mathbf{v} \times \boldsymbol{\Omega}}_{\text{Coriolis force}}$$

- Relativistic case

$$m \rightarrow E = p$$

CVE

$$\mathbf{j}_R = \left( \frac{\mu_R^2}{4\pi^2} + \frac{T^2}{12} \right) \boldsymbol{\Omega}$$

$$\mathbf{j}_L = \left( \frac{\mu_R^2}{4\pi^2} + \frac{T^2}{12} \right) \boldsymbol{\Omega}$$

Vilenkin (1979); Son, Surowka (2000); Landsteiner (2011); ...

# Chiral vortical wave

- Wave equation

$$\left( \partial_t \pm \frac{\bar{\mu}_{R/L} \mathbf{\Omega}}{2\pi^2 \chi} \cdot \nabla \right) \delta n_{R/L} = 0$$

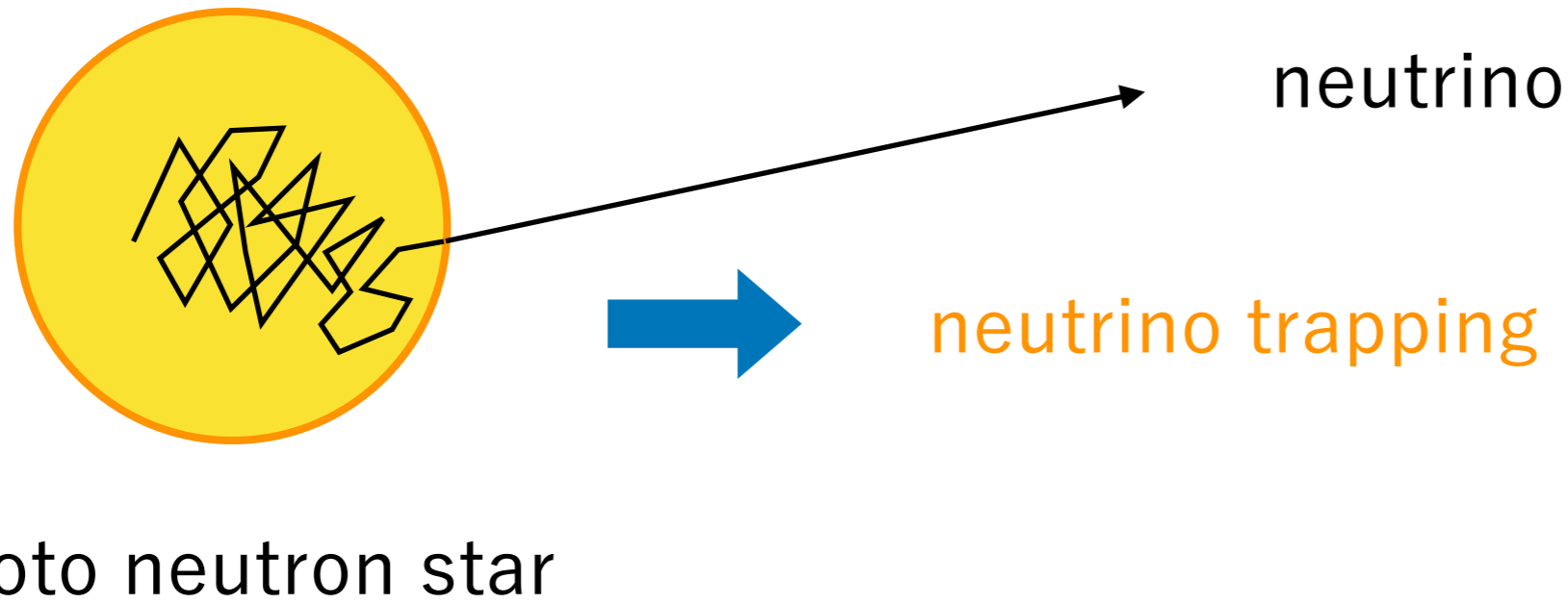
- Dispersion relation

$$\omega_{R/L} = \pm \frac{\bar{\mu}_{R/L} \mathbf{\Omega}}{2\pi^2 \chi} \cdot \mathbf{k}$$

- CVW needs the chirality imbalance to propagate

# Neutrino matter

- Electron capture:  $p + e_L^- \rightarrow n + \nu_{e,L}$
- Timescale: gravitational collapse  $<$  neutrino diffusion



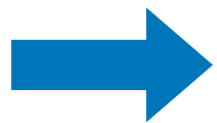
- Matter with **only** left-handed neutrinos : chiral matter

# Frequency of CV-mode

- Dispersion relation

$$\omega_{\text{CM}} \simeq -V_{\text{CV}}k_z - ie^{\lambda-\rho}Dk_z^2 \quad V_{\text{CV}} \equiv e^{\lambda-\rho} \frac{\bar{\mu}_\nu \Omega}{2\pi^2 \chi_\nu}$$

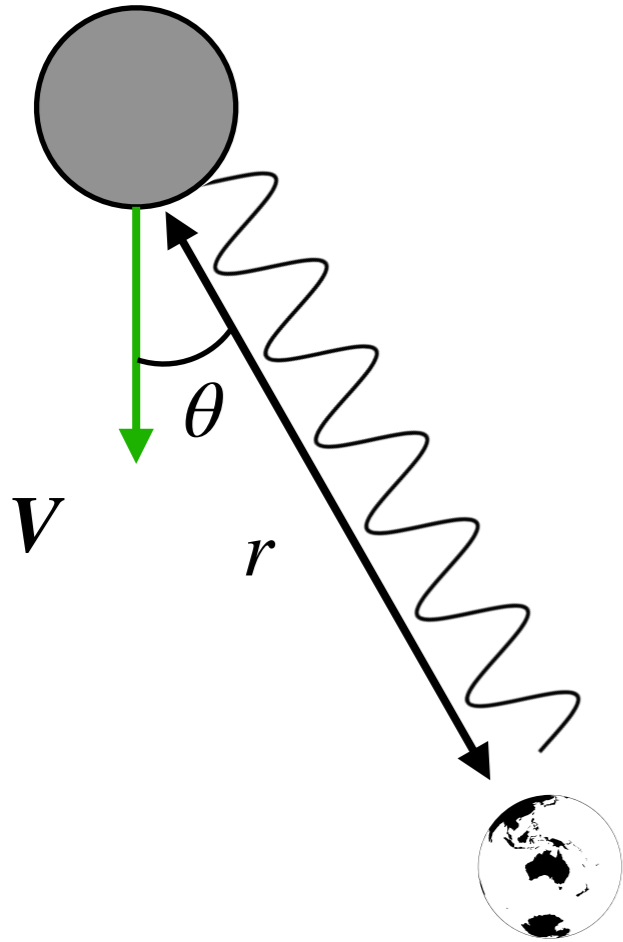
- Energy scale of rotation is small



speed of CV-mode < diffusion effect

$$\frac{V_{\text{CV}}}{D} \simeq \frac{3\Omega}{\bar{\mu}_\nu l_{\text{mfp}}} \sim 10^{-33} \text{ MeV} \left( \frac{\Omega/2\pi}{1 \text{ Hz}} \right)$$

# Angular dependence of GW



- Can assume that the source runs in constant velocity

- Quadrupole moment formula

$$\tilde{h}_{ab} = \frac{2G}{r} \frac{\partial^2}{\partial t^2} I_{ab}(t - r), \quad I_{ab} = \int d^3\mathbf{x}' T^{00}(t, \mathbf{x}) x'_a x'_b$$

- Angular dependence

$$\tilde{h}_{\theta\theta} \propto \sin^2 \theta$$

Gravitational waves are emitted strongly at  $\theta = \pi/2$