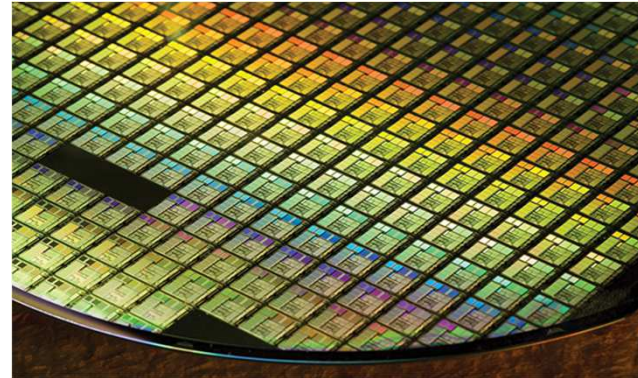
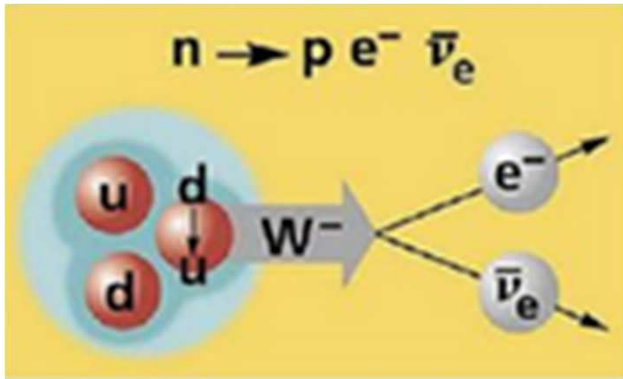


# From Neutrino to Semiconductor

A journey of Learning and Research that Started from Physics

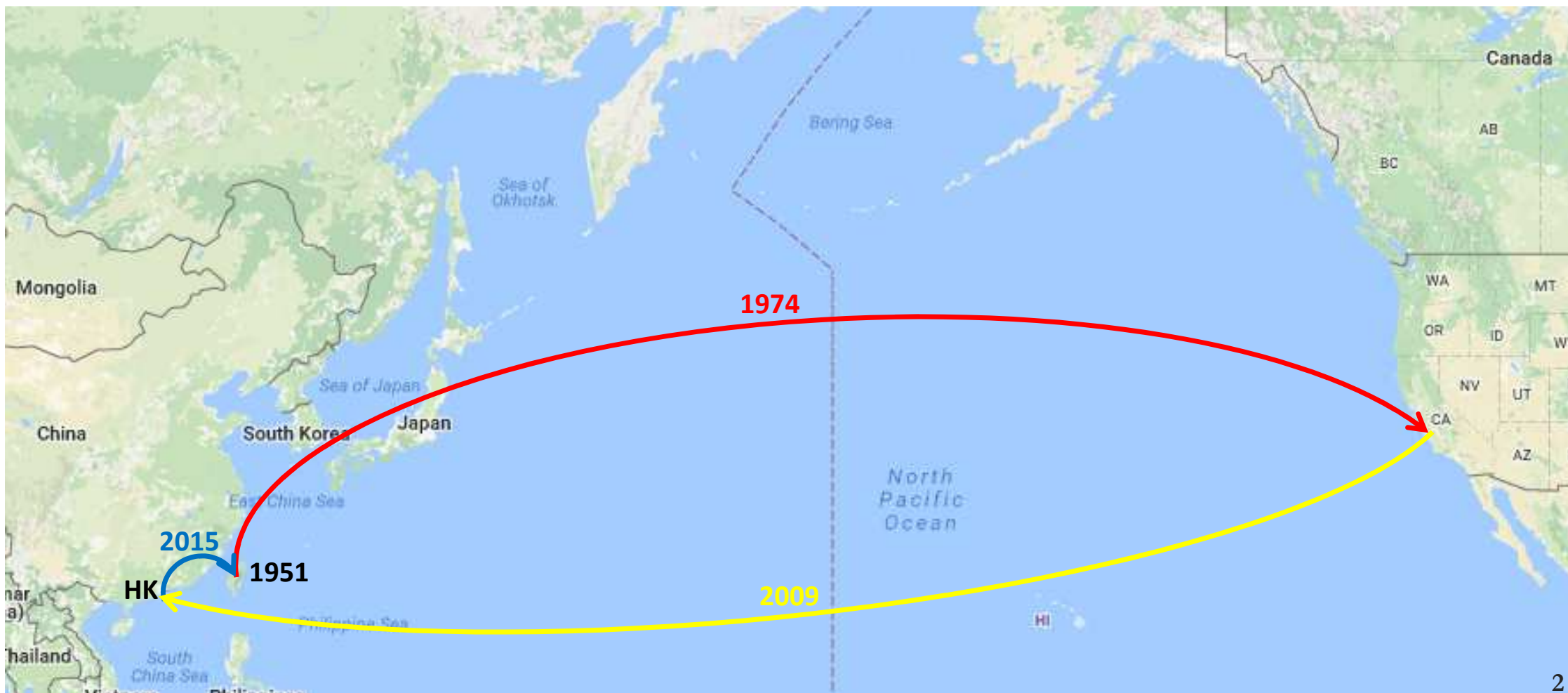


**KC Wang (王克中)**

**Macronix (旺宏電子)**

**Joint AS IOP Colloquium and PG2023 General Talk on 6/1/2023**

# A Fulfilling Journey



# Outline

- **Introduction**
- **Experience Sharing**
  - **Some Stories of Neutrino**
  - **GaAs Heterojunction Bipolar Transistor (HBT) and ICs**
  - **Semiconductor Memories**
- **Conclusion Remarks**

# Elementary Particles of Standard Model

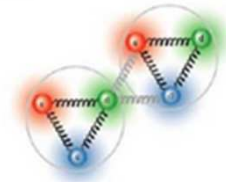
three generations of matter (fermions)			interactions / force carriers (bosons)		
	I	II	III		
mass	$\approx 2.2 \text{ MeV}/c^2$	$\approx 1.28 \text{ GeV}/c^2$	$\approx 173.1 \text{ GeV}/c^2$	0	$\approx 124.97 \text{ GeV}/c^2$
charge	$\frac{2}{3}$	$\frac{2}{3}$	$\frac{2}{3}$	0	0
spin	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	1	0
QUARKS	<b>u</b> up	<b>c</b> charm	<b>t</b> top	<b>g</b> gluon	<b>H</b> higgs
	<b>d</b> down	<b>s</b> strange	<b>b</b> bottom	<b><math>\gamma</math></b> photon	
	<b>e</b> electron	<b><math>\mu</math></b> muon	<b><math>\tau</math></b> tau	<b>Z</b> Z boson	
LEPTONS	<b><math>\nu_e</math></b> electron neutrino	<b><math>\nu_\mu</math></b> muon neutrino	<b><math>\nu_\tau</math></b> tau neutrino	<b>W</b> W boson	

There are 3 flavors of neutrinos and their anti-particles

Source: [https://en.wikipedia.org/wiki/Fundamental\\_interaction#/media/File:Standard\\_Model\\_of\\_Elementary\\_Particles.svg](https://en.wikipedia.org/wiki/Fundamental_interaction#/media/File:Standard_Model_of_Elementary_Particles.svg)

# Fundamental Interactions

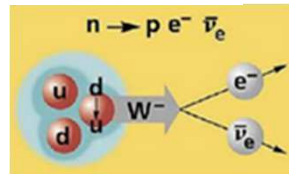
1. Strong



2. Electromagnetic



3. Weak



4. Gravitation



Interaction	Current theory	Mediators	Approximate relative strength	Long distance behavior (potential)	Range (m)
Strong	QCD	gluons	1	$\sim r$	$10^{-15}$
EM	QED	photons	$10^{-2}$	$1/r$	infinite
Weak	Electroweak theory	W and Z bosons	$10^{-6}$ to $10^{-7}$	$1/r e^{-m_{W,Z} r}$	$10^{-18}$
Gravity	General relativity	gravitons	$10^{-39}$	$1/r$	infinite

[https://en.wikipedia.org/wiki/Fundamental\\_interaction](https://en.wikipedia.org/wiki/Fundamental_interaction)

<http://hyperphysics.phy-astr.gsu.edu/hbase/Forces/couple.html>

# Outline

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# The Origin of Neutrino Concept

## Puzzle of Beta Decay



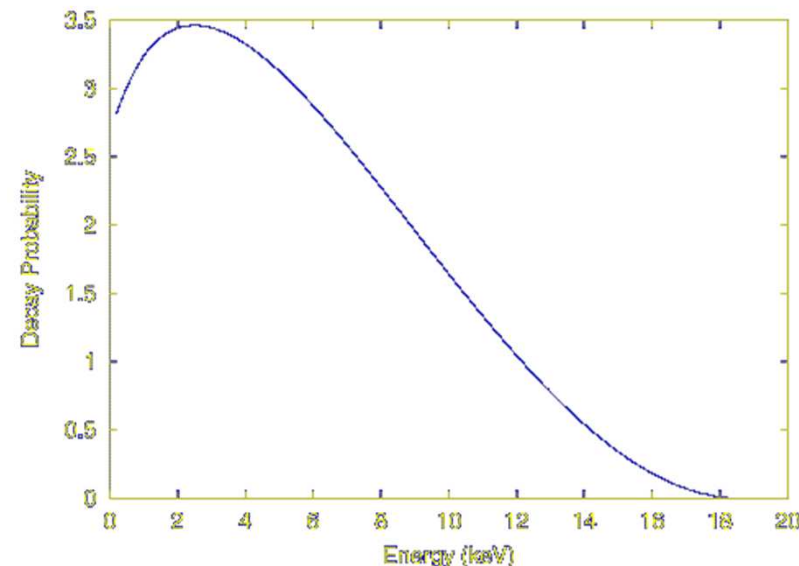
When a neutron converts to a proton and electron,  
The electron should have **single energy**



However, the observed spectrum of  
electron was **continuous**.

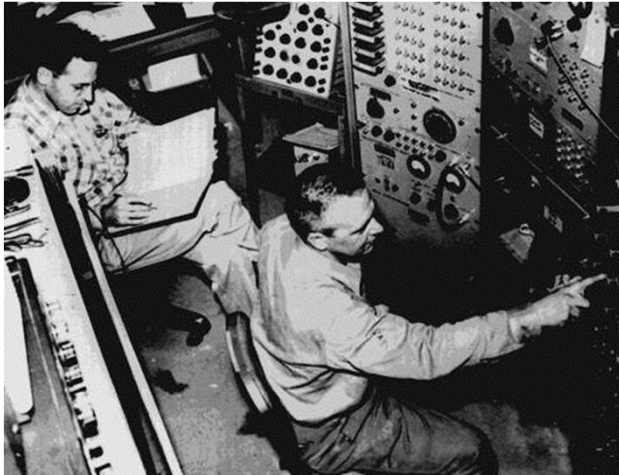
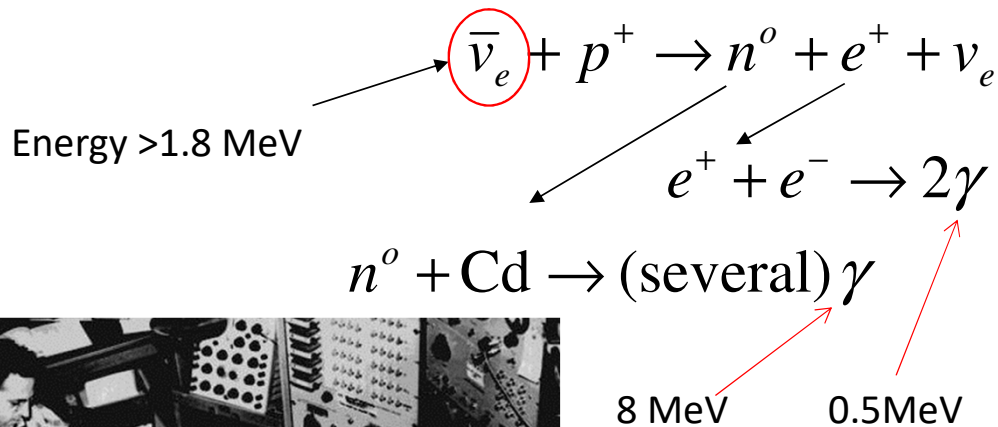
$$N1 \rightarrow N2 + e^-$$

$$E_{e^-} = E_{N1} - E_{N2}$$

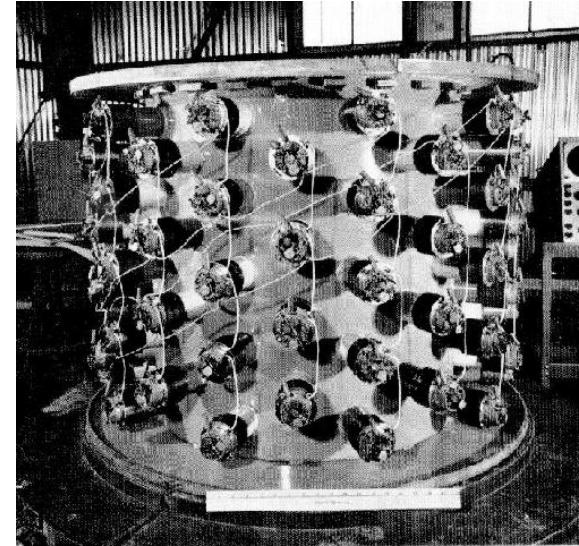


In 1930, Wolfgang Pauli proposed an idea of neutrino to save the law of conservation of energy

# Discovery of Neutrino

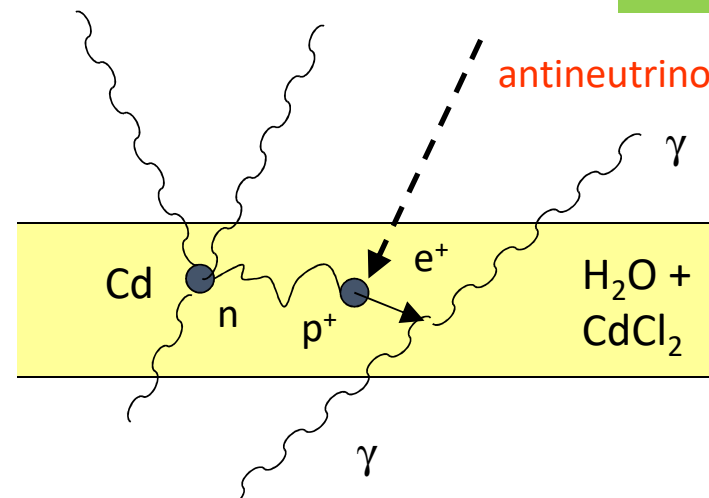


Experiment attempted at Hanford in 1953, too much background. Repeated at Savannah River in **1956**. [Flux:  $1.2 \times 10^{12}$  neutrinos/(cm<sup>2</sup> s)]



300-liter liquid scintillator with Cadmium salt using 90 PMTs

Signal  $2\gamma$ , then several  $\gamma \sim \text{few } \mu\text{s}$  later



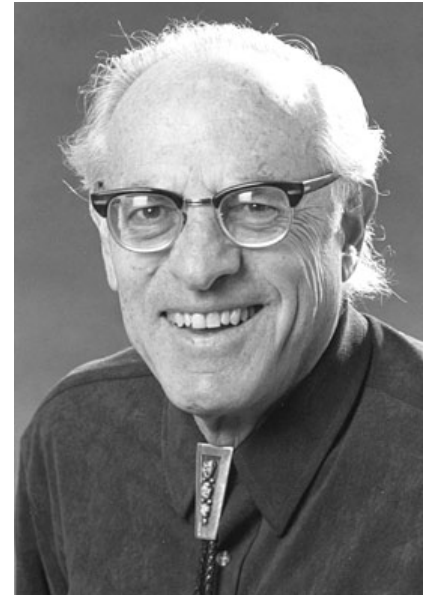


Cowan passed away in 1974

In 1995, Reines was awarded the Nobel Prize in Physics



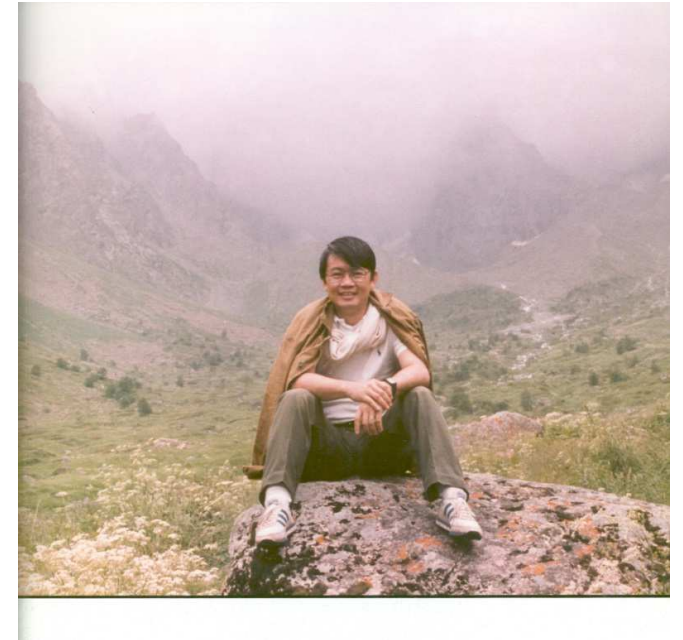
Clyde Cowan



Fred Reines

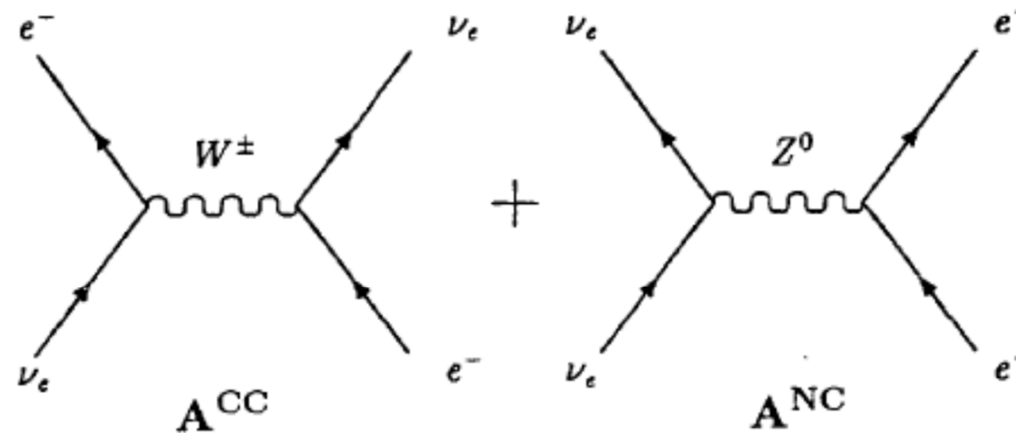
# Herbert H. Chen (陳華生) – An Exceptional Neutrino Physicist

- Born in Chunking, China in 1942 and immigrated to US in 1955
- B.S. from Caltech and Ph.D. from Princeton U.
- Joined UC Irvine in 1968, Associate Prof. in 1974, Prof. in 1980
- Led UCI-LANL  $\nu_e$  - e elastic scattering experiment
- **Designed and initiated Sudbury Neutrino Observatory (SNO) experiment**
- Pioneered development of Liquid Argon TPC
- Chaired an NSF committee on “Computing for Particle Physics by Network” which led to NSFNET that soon merged with ARPANET which eventually became Internet
- Passed away in 1988



# $\nu_e - e$ Elastic Scattering Experiment

Feynman Diagram for  $\nu_e + e^- \rightarrow \nu_e + e^-$



**FIGURE 1:** Feynman diagram for  $\nu_e + e^- \rightarrow \nu_e + e^-$  showing the weak charged and neutral current amplitudes.

## Los Alamos Meson Physics Facility (LAMPF)

- Opened in 1972
- Has an 800-m long, 800 MeV, 1mA proton linear accelerator
- Now, it is Los Alamos Neutron Science Center (LANSCE)\_



# LAMPF Beam-Stop Neutrino Source and Experimental Area

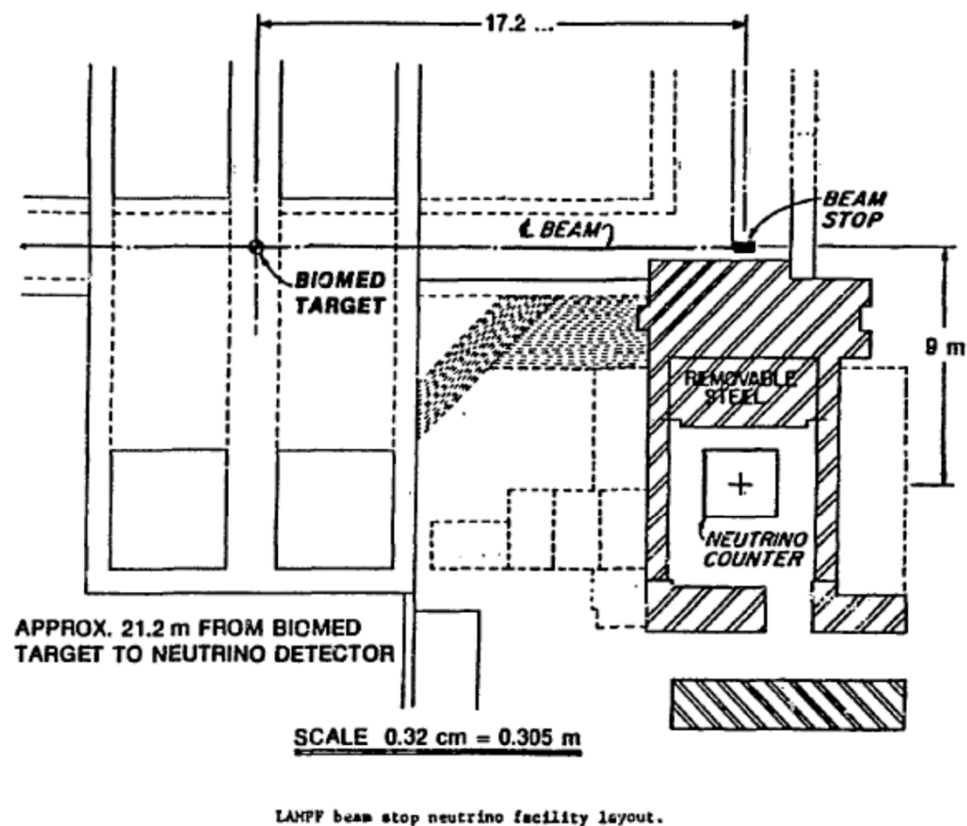


FIGURE 6: Overhead view of LAMPF beam-stop neutrino source and experimental area.

Source: LAMPF



# Energy Spectrum of Neutrinos at LAMPF

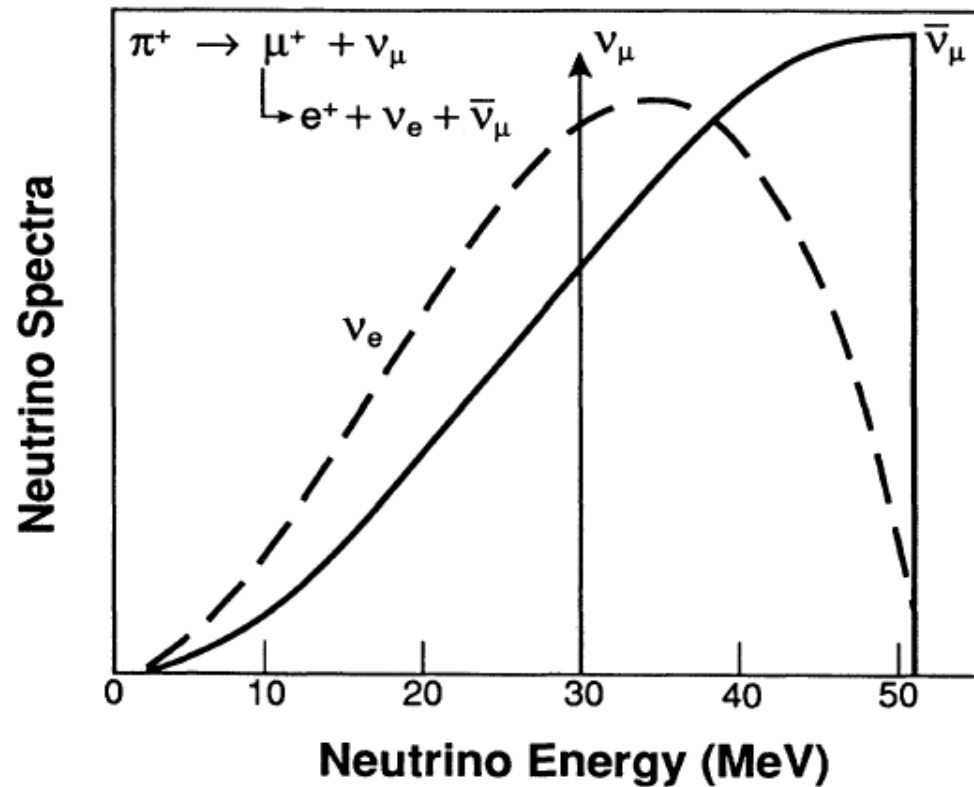


FIG. 3. Energy spectrum of neutrinos produced by  $\pi$  and  $\mu$  decay at rest in the proton beam stop. Spectra for all three neutrinos are shown.

# Neutrino-Electron Elastic Scattering Detector

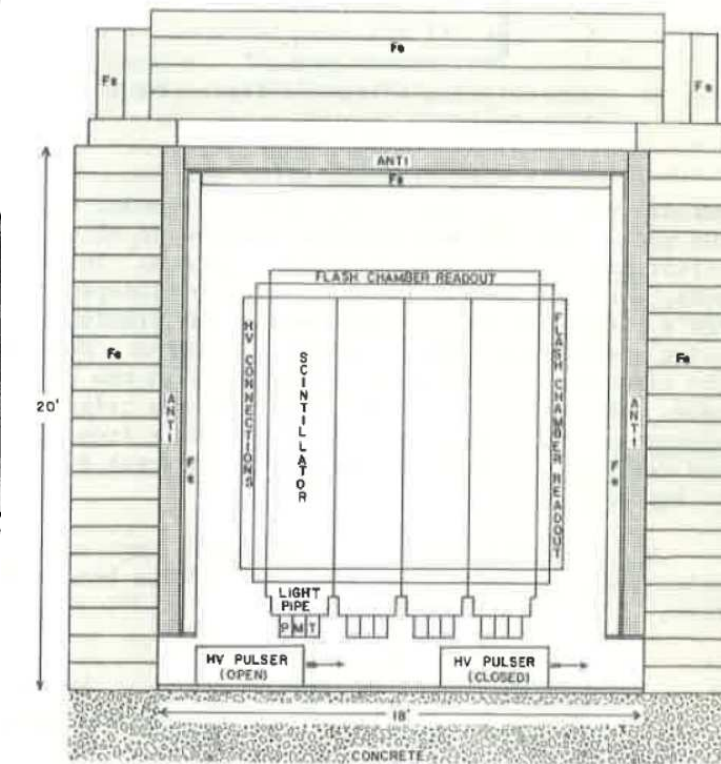
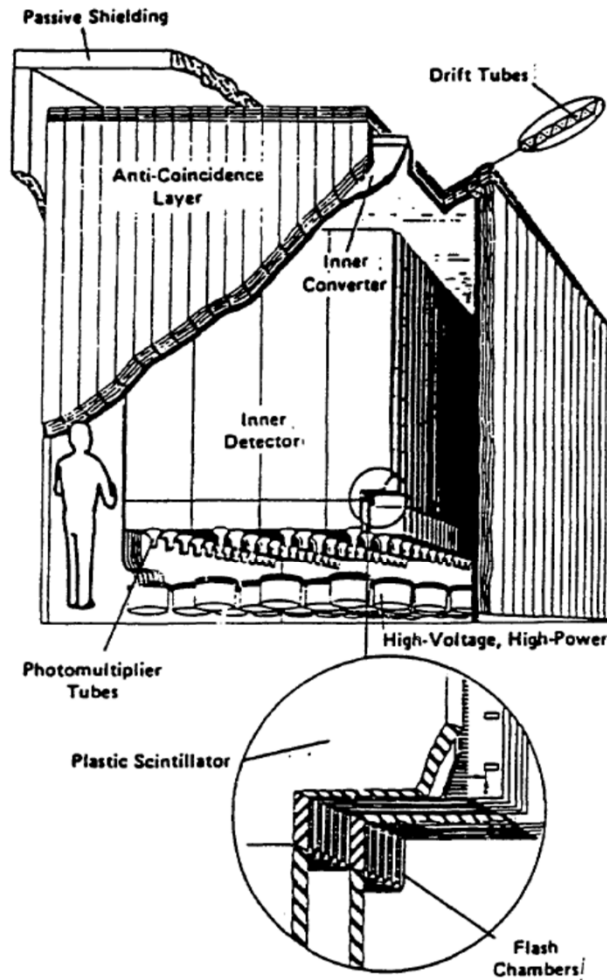


Fig. 1b. End view of the UCI/LASL neutrino detector.

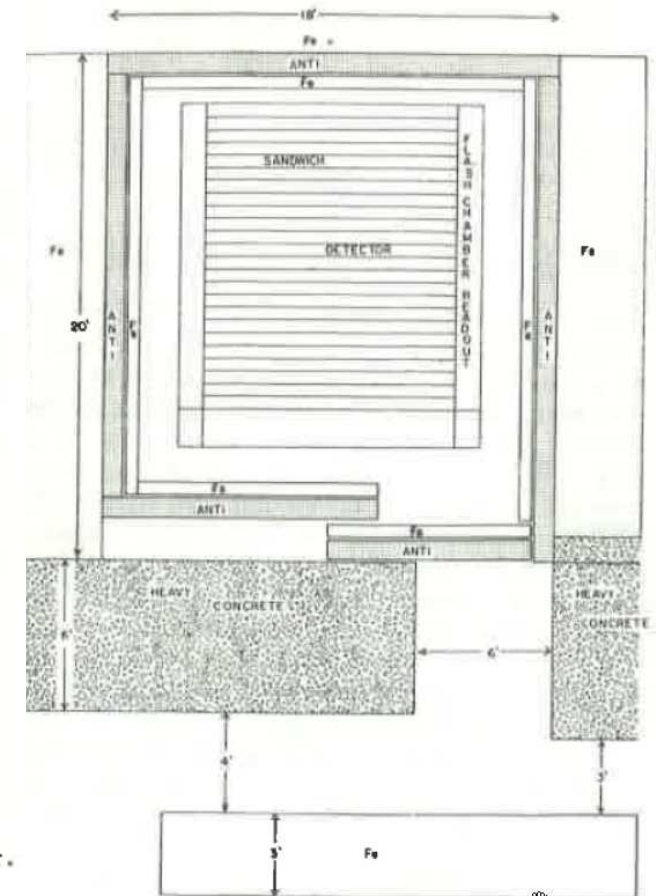


Fig. 1a. Top view of the UCI/LASL neutrino detector.

**End View**

**Top View**

FIGURE 7: Cutaway view of the neutrino detector showing how the central detector was placed within the anticoincidence shield and cosmic-ray shielding. Insets show details of the central detector's FCM and scintillation counter stacking and details of the MWPC wire arrangement.

Source: UCI-LANL Neutrino Collaboration

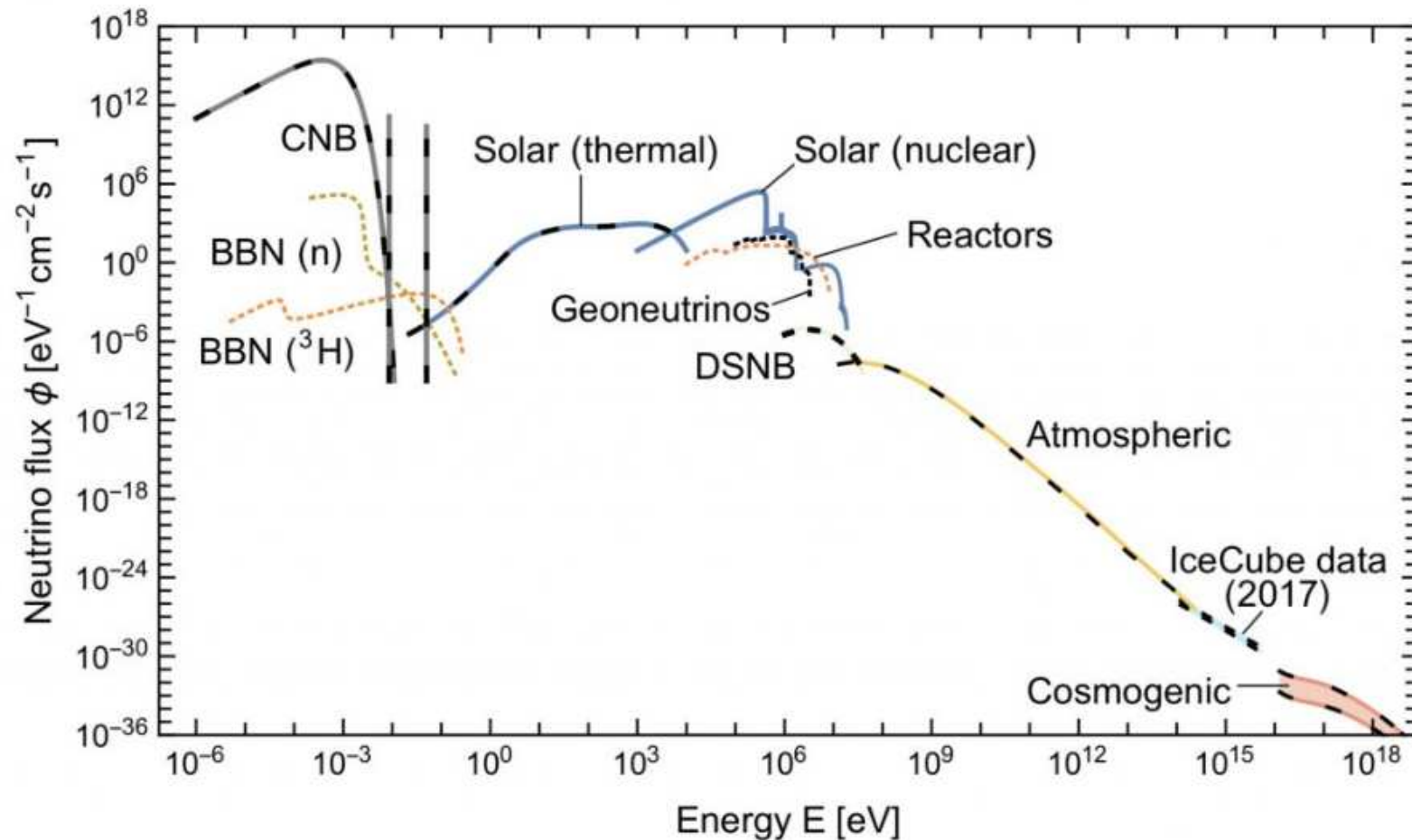
K. C. Wang and H. H. Chen, IEEE Tran. on Nuclear Science, Vol. NS-28, No. 1, 1981

## Key $\nu_e e \rightarrow \nu_e e$ Results

- Observed  $236 \pm 35$   $\nu_e e \rightarrow \nu_e e$  events
- Determined total elastic scattering cross-section to be  $10.0 \pm 1.5(\text{stat}) \pm 0.9(\text{syst}) \times 10^{-45} \text{ cm}^2 \times [E_\nu (\text{MeV})]$
- The results showed that the interference of weak charged and neutral currents was destructive with a strength of  $I = -1.07 \pm 0.21$
- It agreed well with the standard model (SM) prediction of  $I = -1.08$

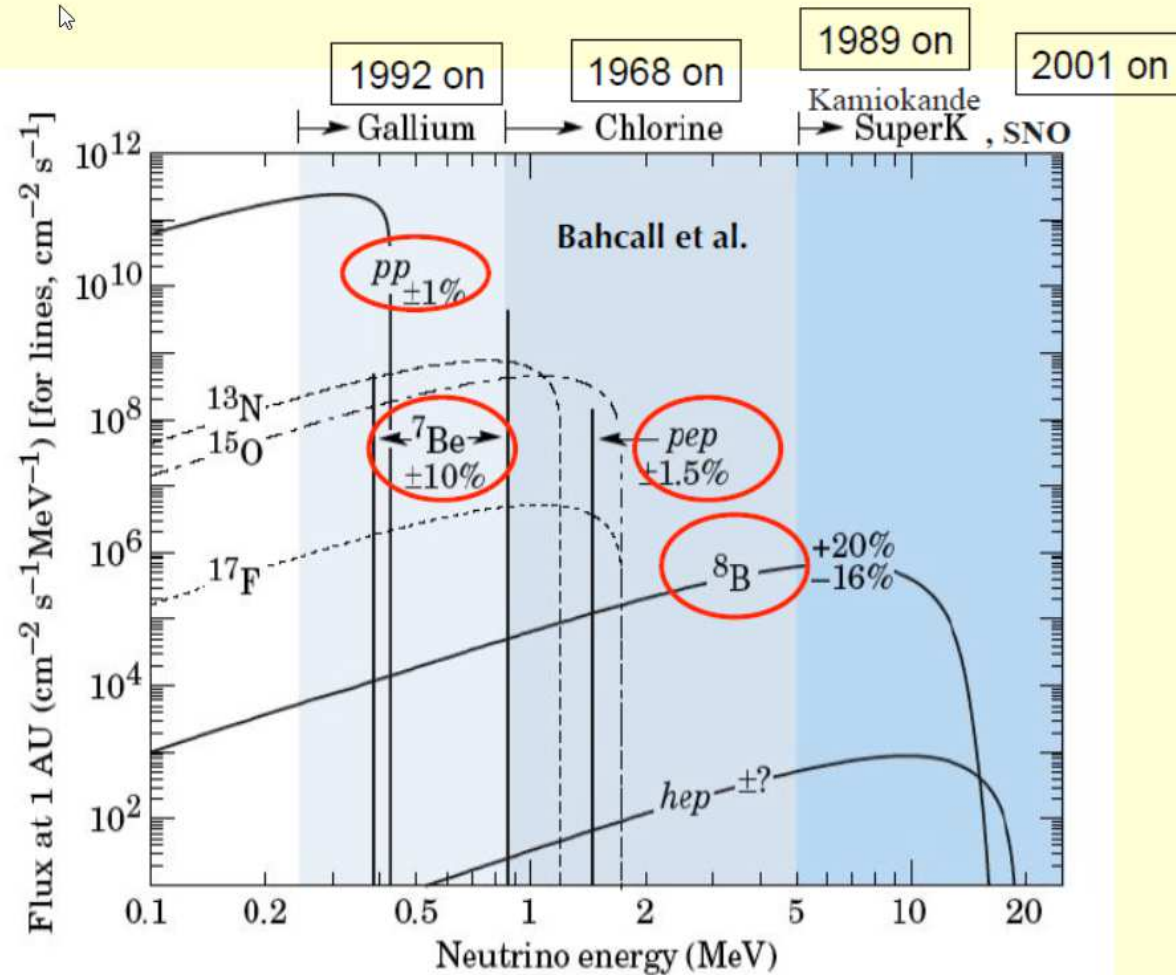
# Neutrino Sources and Flux on Earth

## - Grand Unified Neutrino Spectrum



E. Vitagliano, I. Tamborra, G. Raffelt, Rev. Mod. Phys. 92 (2020) 45006; arXiv:1910.11878  
<https://neutrino-history.in2p3.fr/introduction-to-neutrino-sources/>

## Including other solar neutrino measurements



Source: Art McDonald, Queen's University, for SNO Collaboration



## Pioneers of Solar Neutrino Physics: Davis, Bahcall, Pontecorvo & Gribov



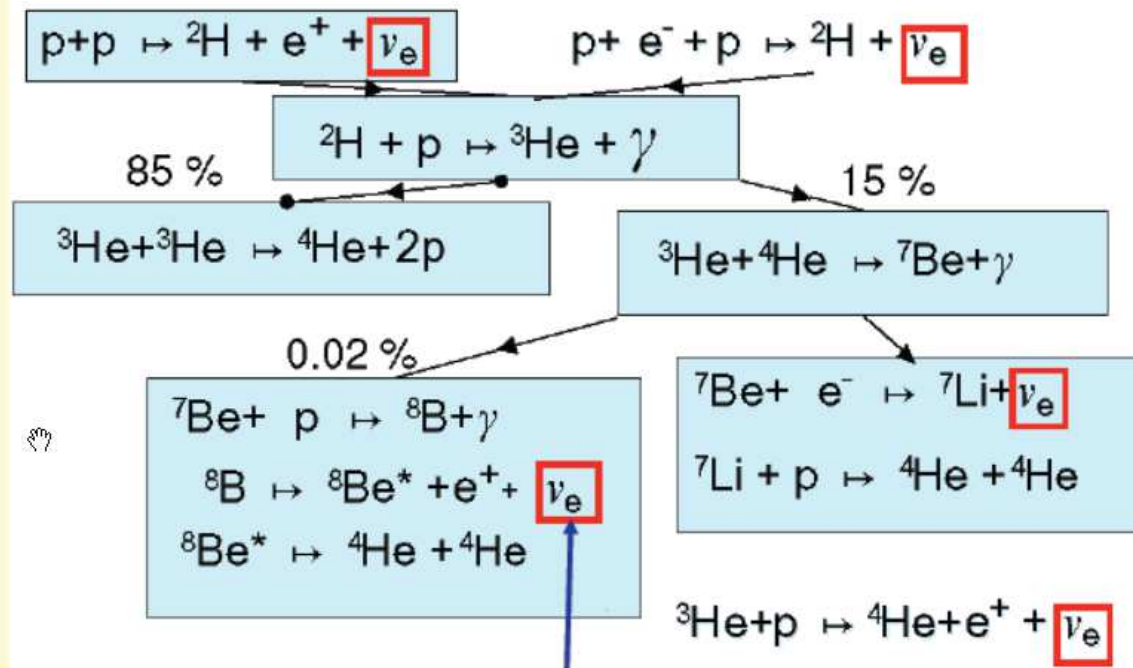
1968: Davis' Measurements of electron neutrinos with Chlorine-based detector show 3 times fewer than Bahcall's calculations.

Ray Davis: Nobel Laureate 2002



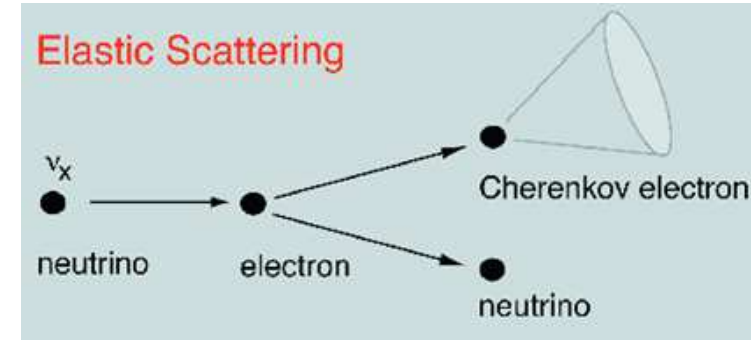
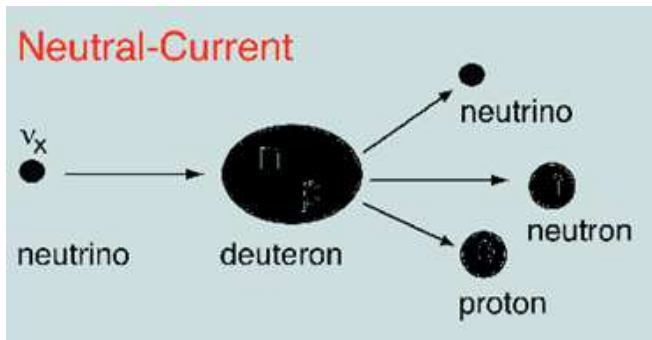
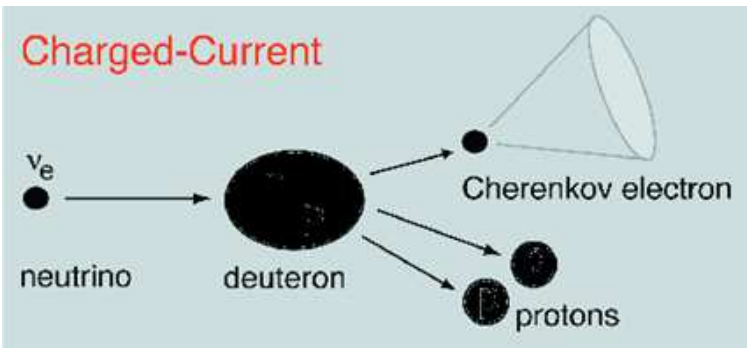
1969: Gribov and Pontecorvo suggest flavor change (oscillation) of electron neutrinos to muon neutrinos as a possible reason.

## SOLAR FUSION CHAIN



**1984: Herb Chen proposes heavy water to search for direct evidence of flavor transformation for neutrinos from  ${}^8\text{B}$  decay in the Sun. Electron neutrinos and all active neutrinos are measured separately to show flavor change independent of solar model calculations. SNO collaboration is created with Chen and George Ewan as Spokesmen.**

# Herb Chen's Idea (1984): Using Heavy Water



VOLUME 55, NUMBER 14

PHYSICAL REVIEW LETTERS

30 SEPTEMBER 1985

## Direct Approach to Resolve the Solar-Neutrino Problem

Herbert H. Chen

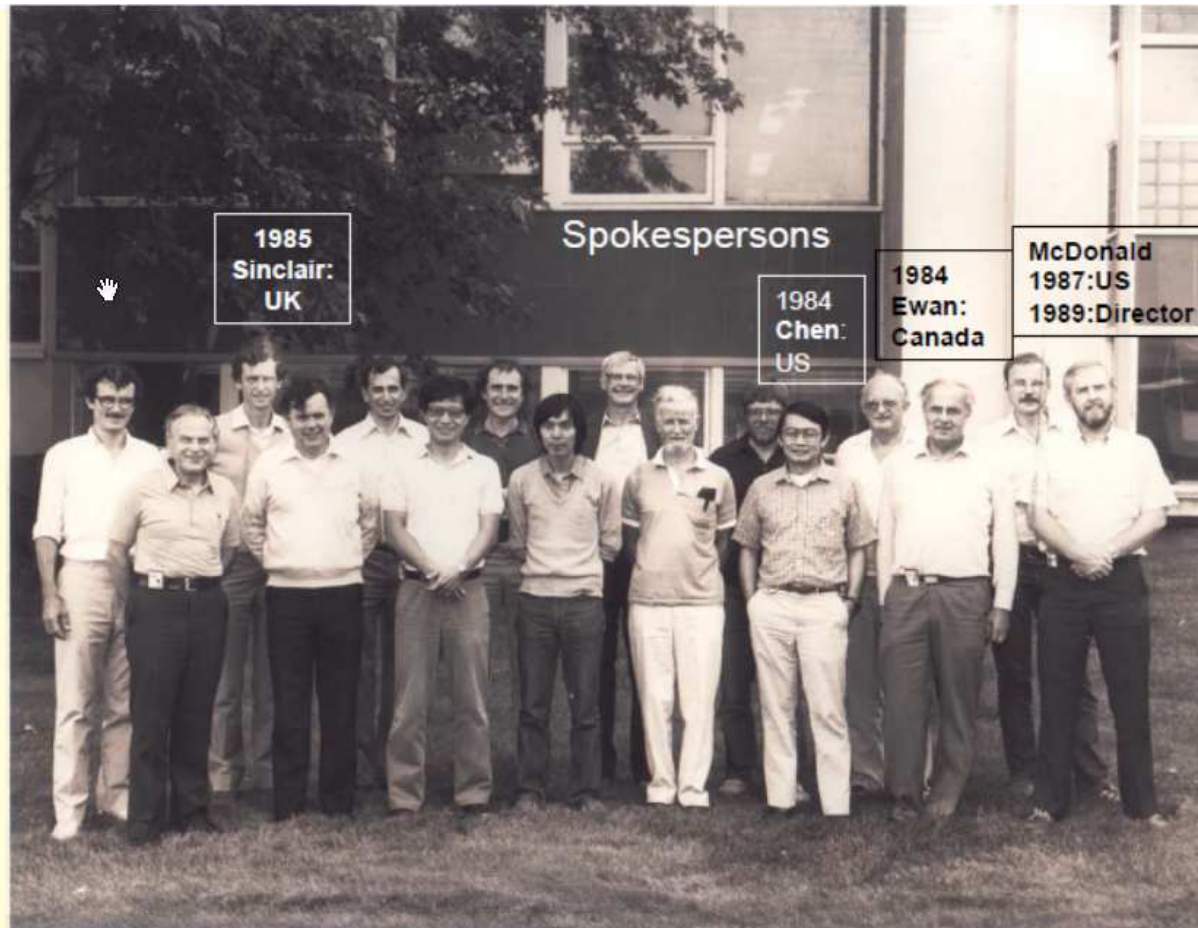
*Department of Physics, University of California, Irvine, California 92717*

(Received 27 June 1985)

A direct approach to resolve the solar-neutrino problem would be to observe neutrinos by use of both neutral-current and charged-current reactions. Then, the total neutrino flux and the electron-neutrino flux would be separately determined to provide independent tests of the neutrino-oscillation hypothesis and the standard solar model. A large heavy-water Cherenkov detector, sensitive to neutrinos from  $^8\text{B}$  decay via the neutral-current reaction  $\nu + d \rightarrow \nu + p + n$  and the charged-current reaction  $\nu_e + d \rightarrow e^- + p + p$ , is suggested for this purpose.

PACS numbers: 96.60.Kx, 14.60.Gh





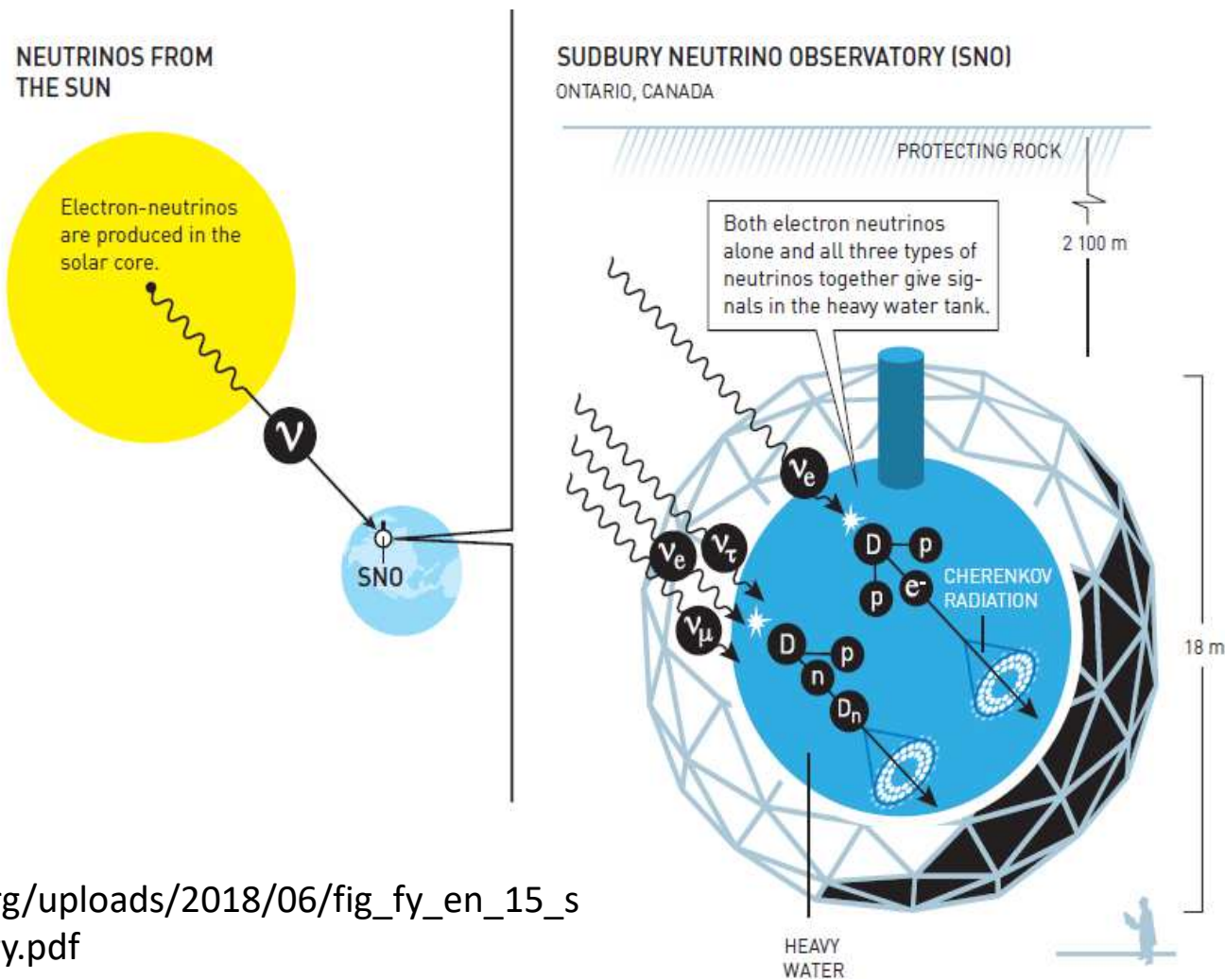
SNO Collaboration Meeting, Chalk River, 1986

**PROPOSAL TO BUILD A NEUTRINO OBSERVATORY IN SUDBURY, CANADA**

D. Sinclair, A.L. Carter, D. Kessler, E.D. Earle, P. Jagam, J.J. Simpson, R.C. Allen, H.H. Chen, P.J. Doe, E.D. Hallman, W.F. Davidson, A.B. McDonald, R.S. Storey, G.T. Ewan, H.-B. Mak, B.C. Robertson Il Nuovo Cimento C9, 308 (1986)

Source: Art McDonald, Queen's University, for SNO Collaboration

# Sudbury Neutrino Observatory



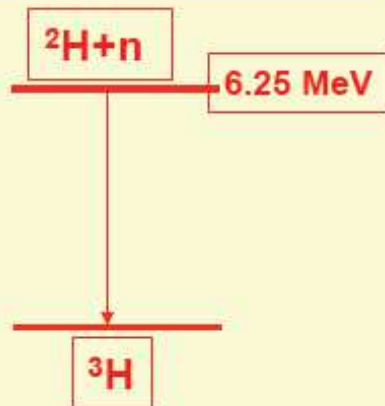
[https://www.nobelprize.org/uploads/2018/06/fig\\_fy\\_en\\_15\\_sudburyneutrinoobservatory.pdf](https://www.nobelprize.org/uploads/2018/06/fig_fy_en_15_sudburyneutrinoobservatory.pdf)



# 3 neutron (NC) detection methods (systematically different)

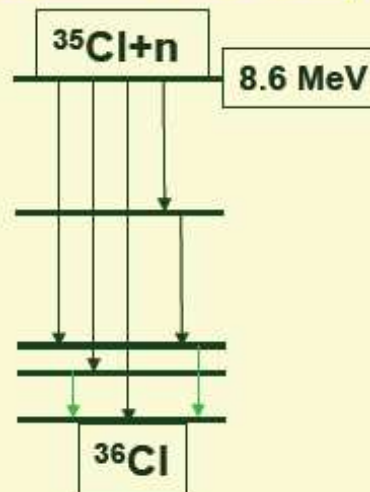
## Phase I ( $D_2O$ ) Nov. 99 - May 01

n captures on  $^2H(n, \gamma)^3H$   
Effic.  $\sim 14.4\%$   
NC and CC separation by energy, radial, and directional distributions



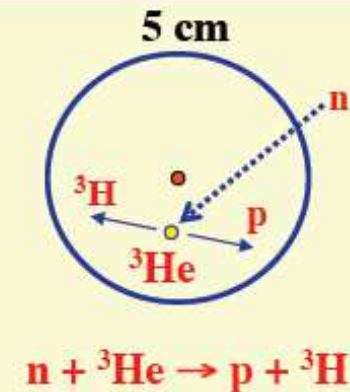
## Phase II (salt) July 01 - Sep. 03

2 tonnes of NaCl  
n captures on  $^{35}Cl(n, \gamma)^{36}Cl$   
Effic.  $\sim 40\%$   
NC and CC separation by event isotropy



## Phase III ( $^3He$ ) Nov. 04-Dec. 06

400 m of proportional counters  
 $^3He(n, p)^3H$   
Effic.  $\sim 30\%$  capture  
Measure NC rate with entirely separate detection system.



# Sudbury Neutrino Observatory (SNO)

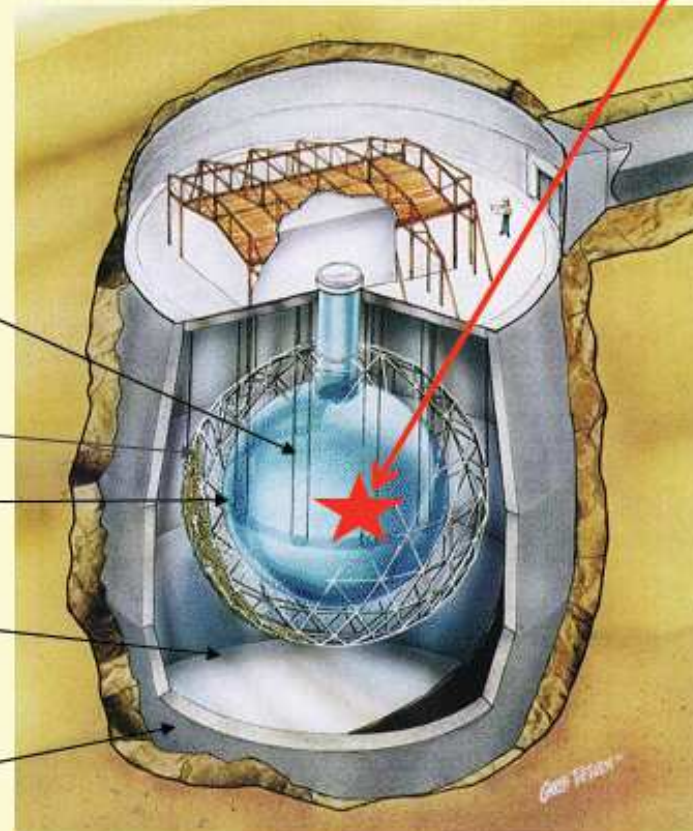
Neutrinos are very difficult to detect so our detector had to be very big with low radioactivity, deep underground.

1000 tonnes of heavy water:  $D_2O$   
\$ 300 million on Loan for \$1.00

9500 light sensors  
12 m Diameter Acrylic Container

Ultra-pure Water:  $H_2O$ .

Urylon Liner and Radon Seal



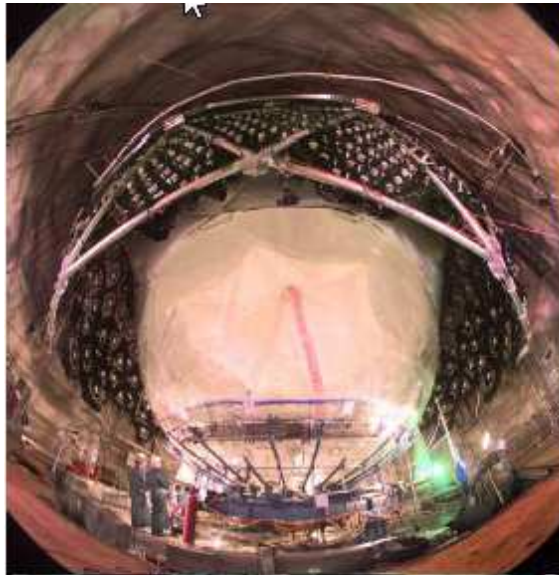
NEUTRINO

34 m  
or  
~ Ten  
Stories  
High!

2 km  
below  
the  
ground

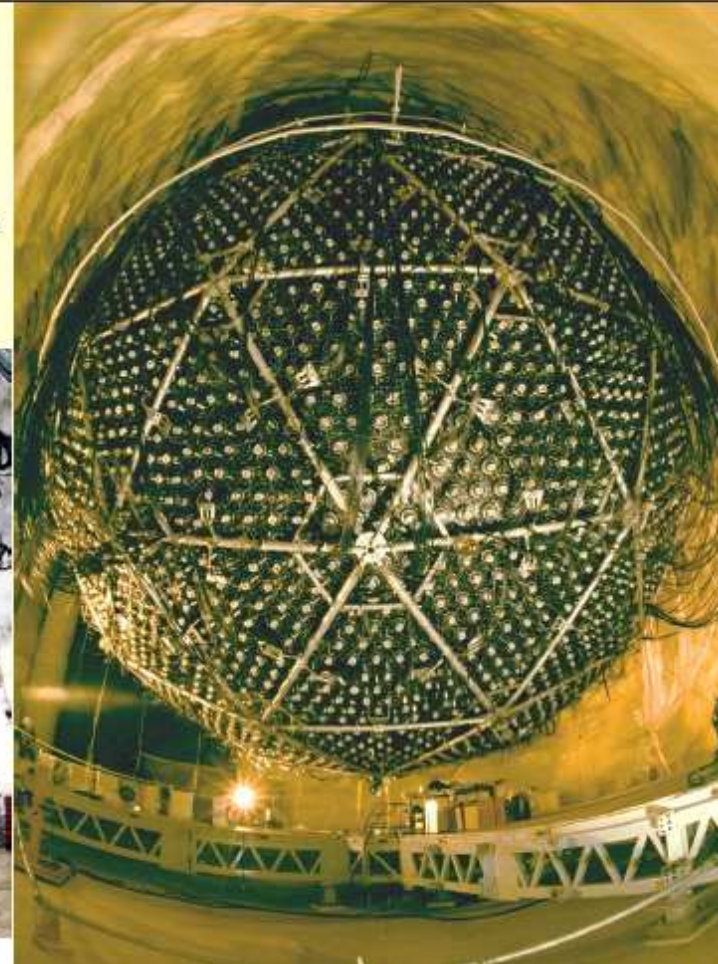
Source: Art McDonald, Queen's University, for SNO Collaboration





**SNO: One million pieces transported down in the 3 m x 3 m x 4 m mine cage and re-assembled under ultra-clean conditions. Every worker takes a shower and wears clean, lint-free clothing.**

70,000  
showers  
during the  
course of the  
SNO project

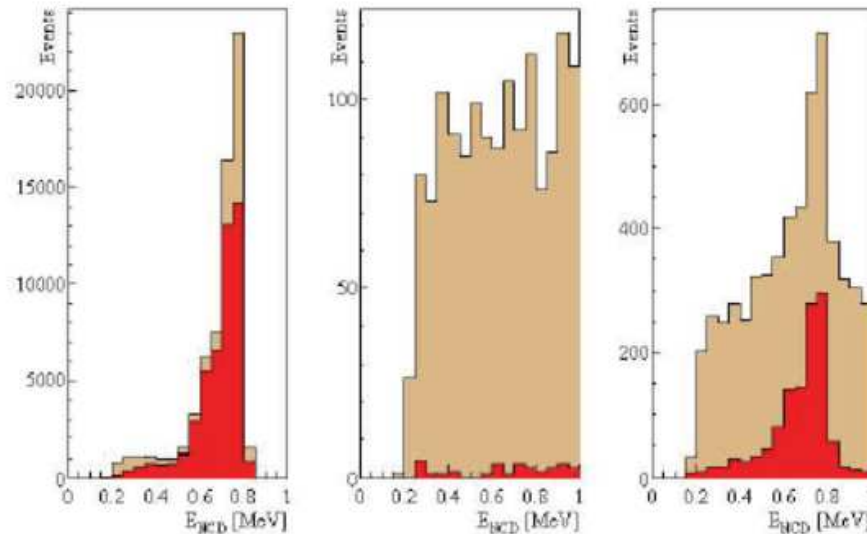


Source: Art McDonald, Queen's University, for SNO Collaboration <sup>26</sup>

# Final Complete Analysis of SNO solar data

The SNO Collaboration (B. Aharmim et al) Phys. Rev. C 88, 025501 (2013)

**NCD pulse shape analysis to identify neutron events**



**<sup>3</sup>He detectors:**  
neutron source

**<sup>4</sup>He detectors:**  
alpha backgnd

**<sup>3</sup>He detectors:**  
neutrino data

$$CC/NC = 0.317 \pm 0.016(stat) \pm 0.009(syst)$$

implies flavor change at far more than  $7\sigma$  and shows that  $\theta_{12}$  is non-maximal by more than  $5\sigma$ .

Full joint analysis of solar data from all three phases provides best sensitivity with all correlations, backgrounds, systematic uncertainties included.

Individual results from all three phases are very consistent within uncertainties

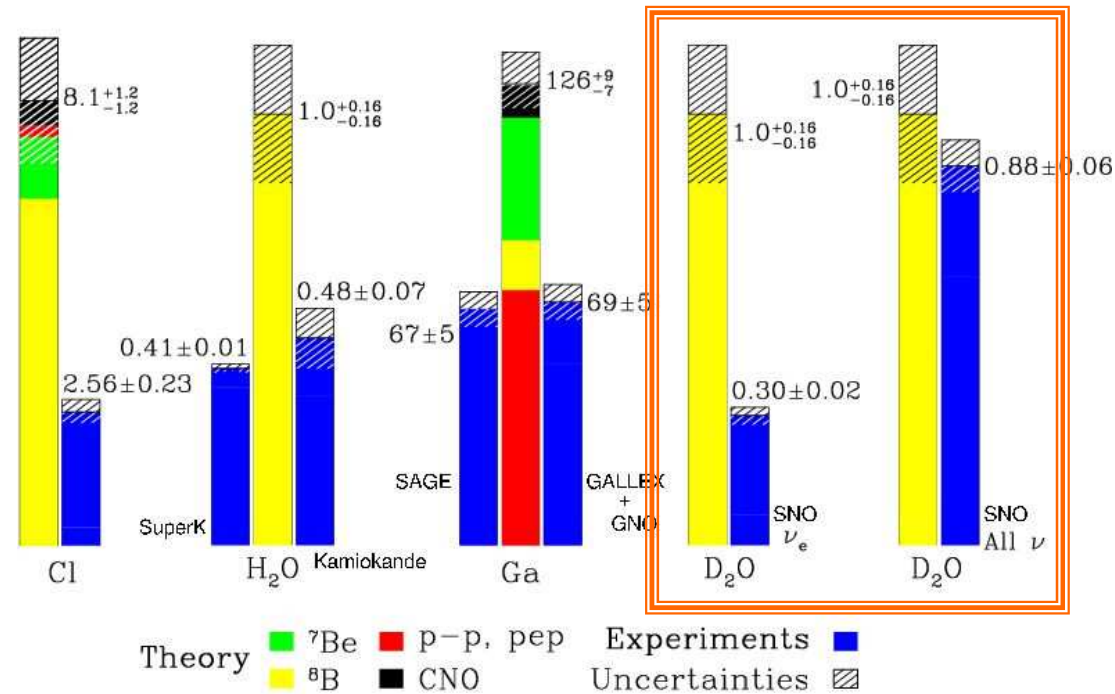
$$\Phi_{8B} = 5.25 \pm 0.16^{+0.11}_{-0.13}$$

More accurate than current solar models and lying between the fluxes predicted for two values of metallicity in the sun



# Experimental Results of Solar $\nu$ Experiments

Total Rates: Standard Model vs. Experiment  
Bahcall-Serenelli 2005 [BS05(OP)]



Final SNO results (2013): Total  $\nu$  flux =  $5.25 \pm 0.16$  (stat)  $^{+0.11}_{-0.12}$  (sys)  $\times 10^6 \text{cm}^{-2} \text{s}^{-1}$   
In very good agreement with theory



**Arthur B. McDonald and Takaaki Kajita were awarded Nobel Prize in Physics in 2015 for "for the discovery of neutrino oscillations, which shows that neutrinos have mass"**



Arthur B. McDonald



Takaaki Kajita

## 6 Neutrino Experiments Won Nobel Prizes

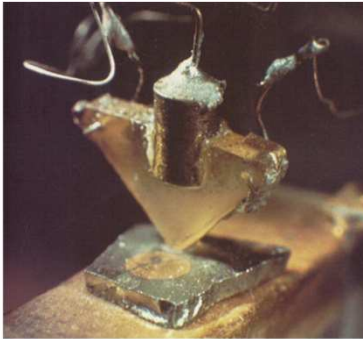
1. 1988 Leon Lederman, Melvin Schwartz, and Jack Steinberger (at Brookhaven National Lab., USA)
2. 1995 Frederick Reines (at Savannah River Nuclear Reactor, USA)
3. 2002 Raymond Davis, Jr. (at Homestake Gold Mine, USA)
4. 2002 Masatoshi Koshiro (at Kamioka Observatory, Japan)
5. 2015 Takaaki Kajita (at Super Kamiokande, Japan)
6. 2015 Arthur B. McDonald (at Sudbury Neutrino Observatory, Canada)

**And more are coming...**

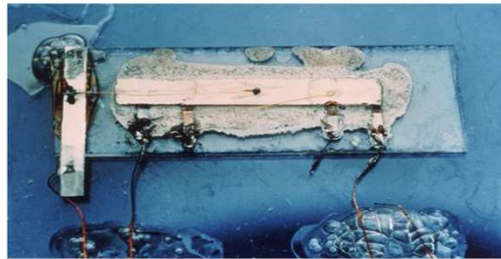
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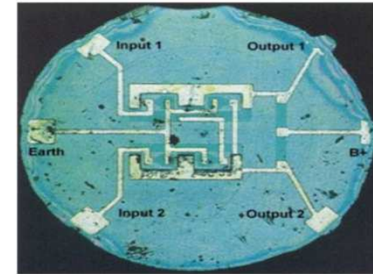
# First Transistor and ICs



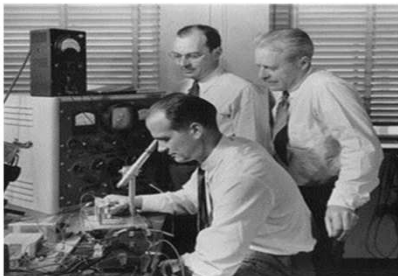
The 1<sup>st</sup> Transistor -1947



The 1<sup>st</sup> IC -1958



The 1<sup>st</sup> planar IC - 1960



William Shockley, John Bardeen, Walter Brattain  
Nobel Prize Winner, 1956



Jack Kilby  
Nobel Prize Winner, 2000



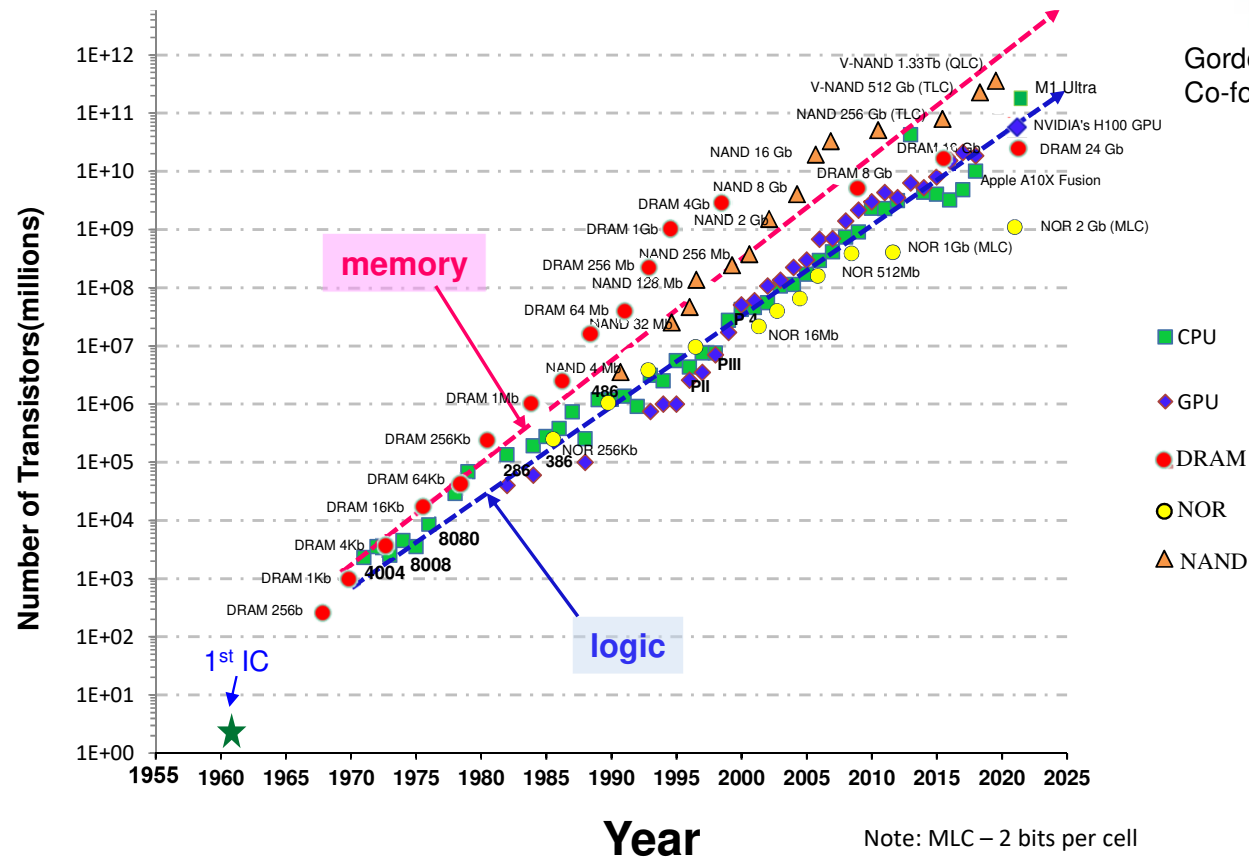
Robert Noyce

# Moore's Law

The number of transistors per chip doubles each 18~24 months \*



Gordon Moore  
Co-founder of Intel 1968

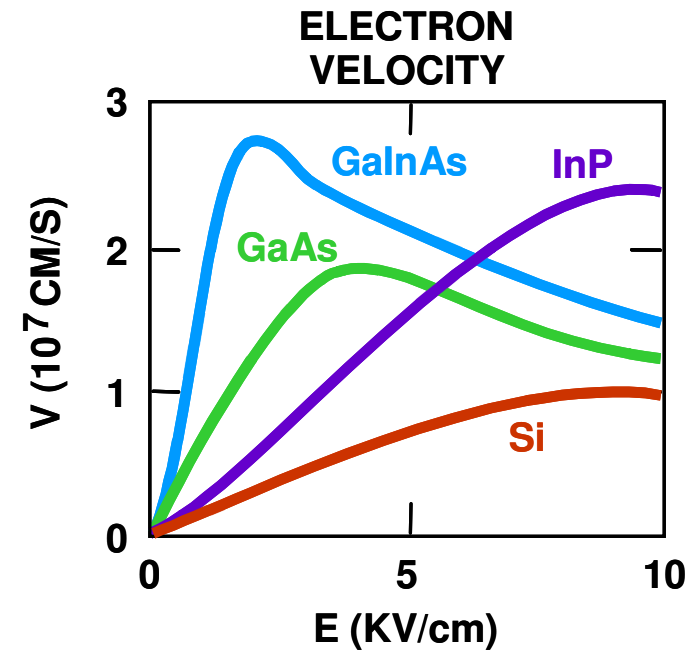
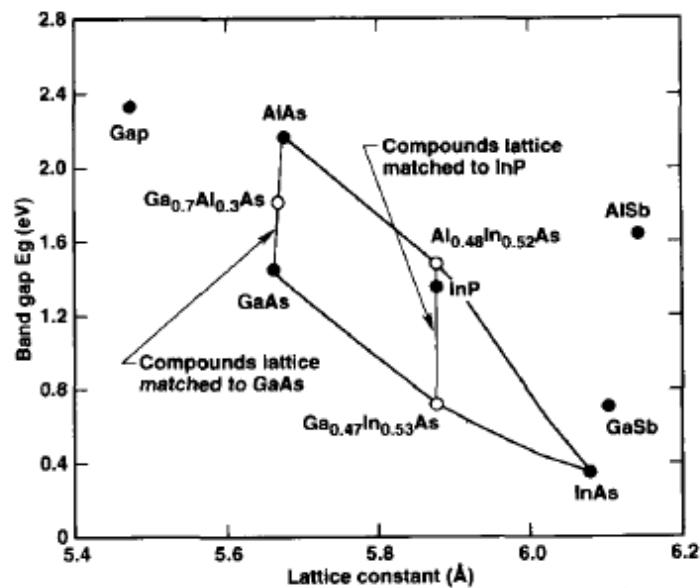


Data source: [https://en.wikipedia.org/wiki/Transistor\\_count](https://en.wikipedia.org/wiki/Transistor_count)  
[https://en.wikipedia.org/wiki/Random-access\\_memory](https://en.wikipedia.org/wiki/Random-access_memory)  
[https://en.wikipedia.org/wiki/Flash\\_memory#NAND\\_flash](https://en.wikipedia.org/wiki/Flash_memory#NAND_flash)  
 \* [https://en.wikipedia.org/wiki/Moore%27s\\_Law#cite\\_note-26](https://en.wikipedia.org/wiki/Moore%27s_Law#cite_note-26)

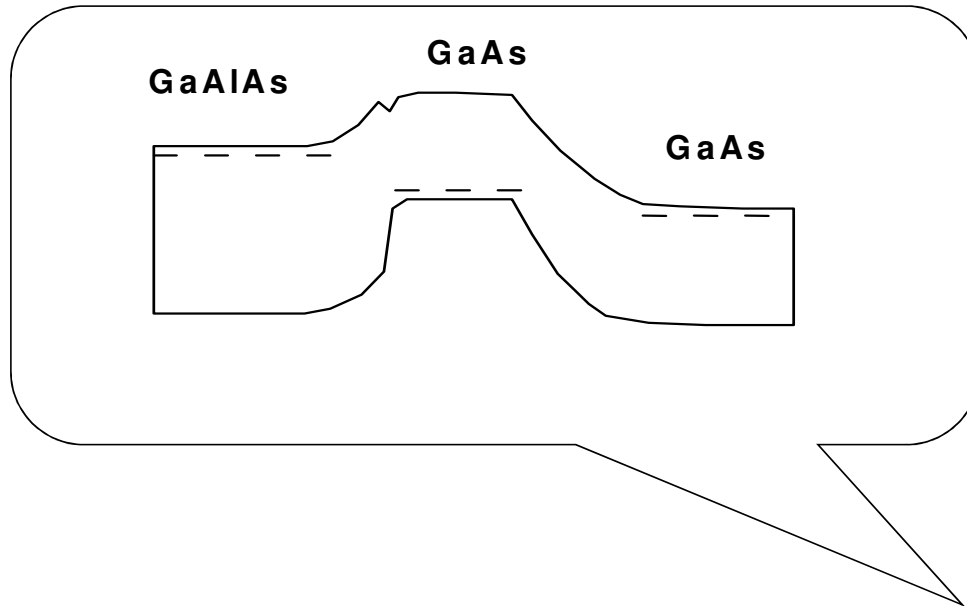


# III-V Compound Semiconductors for High-speed Electronics

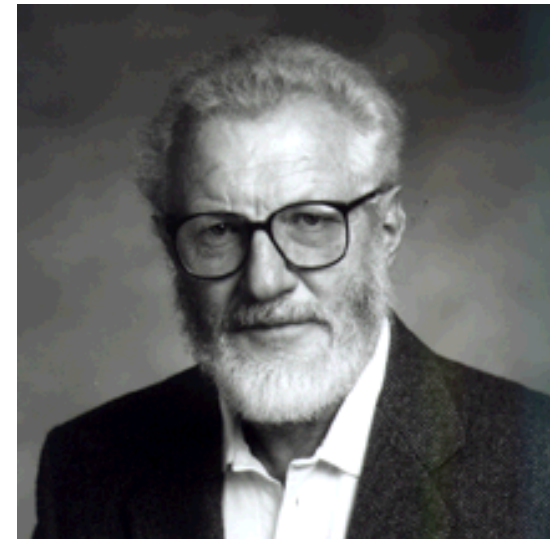
<i>I</i>	<i>II</i>	<i>III</i>	<i>IV</i>	<i>V</i>	<i>VI</i>	<i>VII</i>	<i>VIII</i>
Li	Be	B	C	N	O	F	Ne
Na	Mg	Al	Si	P	S	Cl	Ar
K	Ca	Ga	Ge	As	Se	Br	Kr
Rb	Sr	In	Sn	Sb	Te	I	Xe



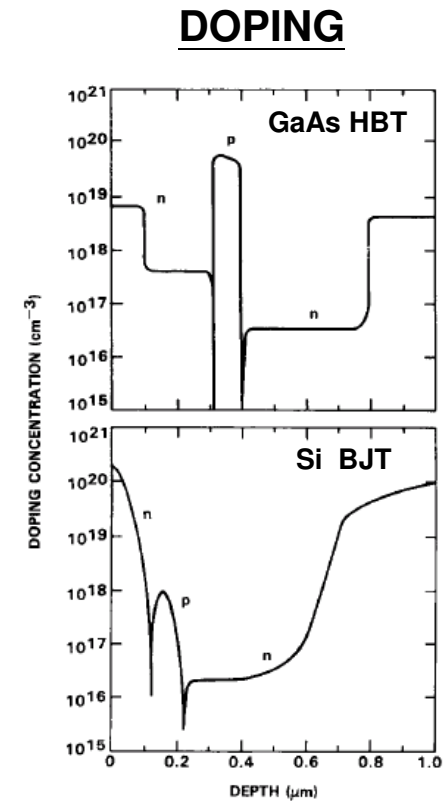
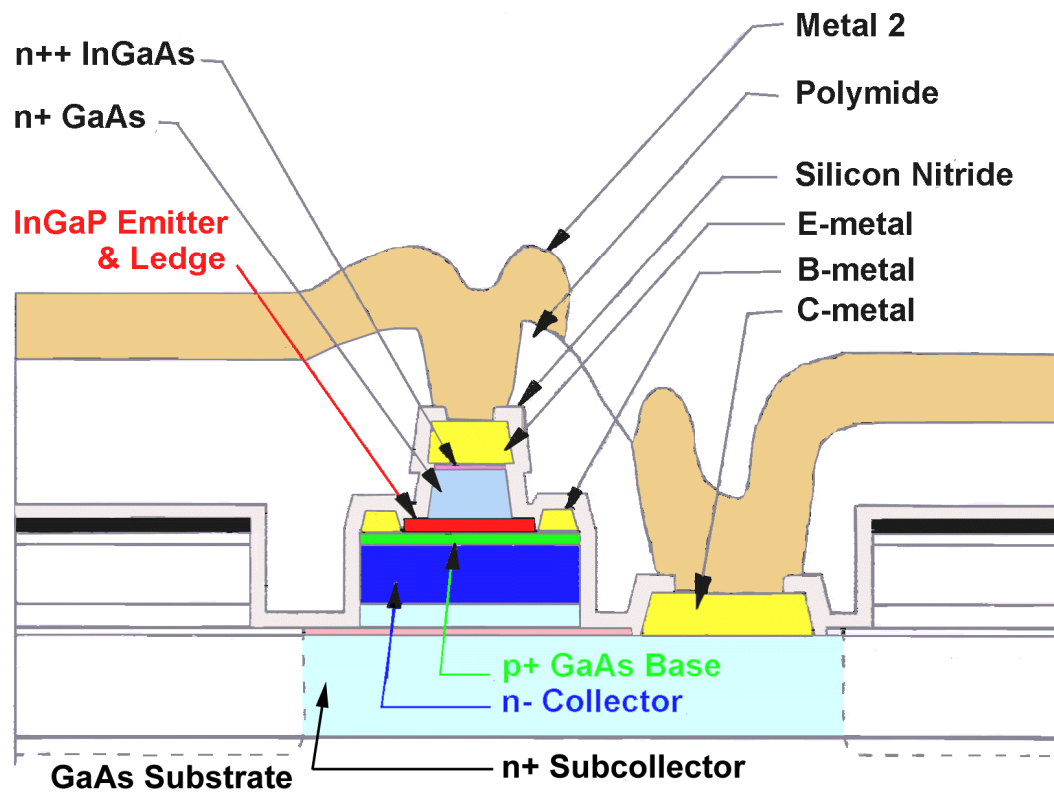
# An Enthusiastic HBT Promoter



**Herb Kroemer**  
**2000 Nobel Laureate in Physics**



# InGaP/GaAs HBT



Source: Peter Asbeck, UCSD

# HBT Power Amplifiers Enhanced Mobile Communications

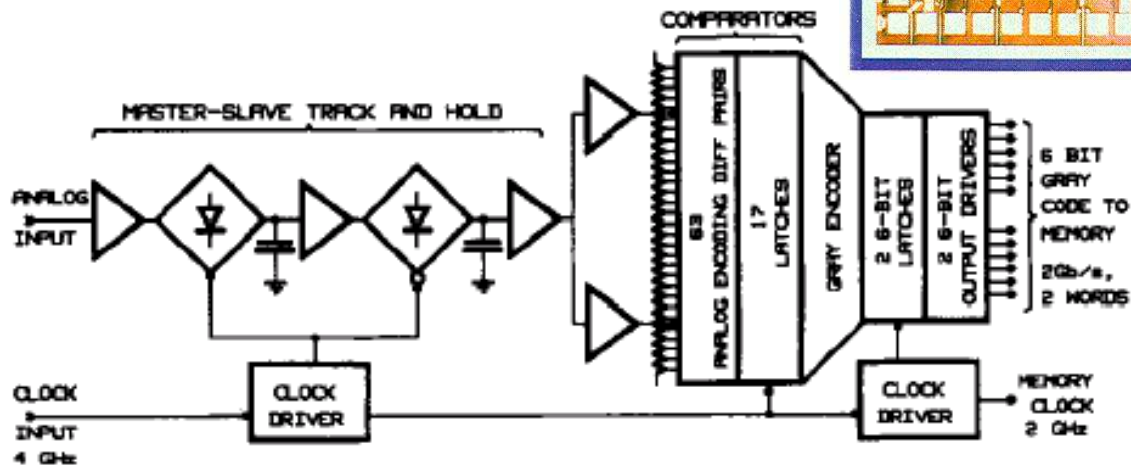
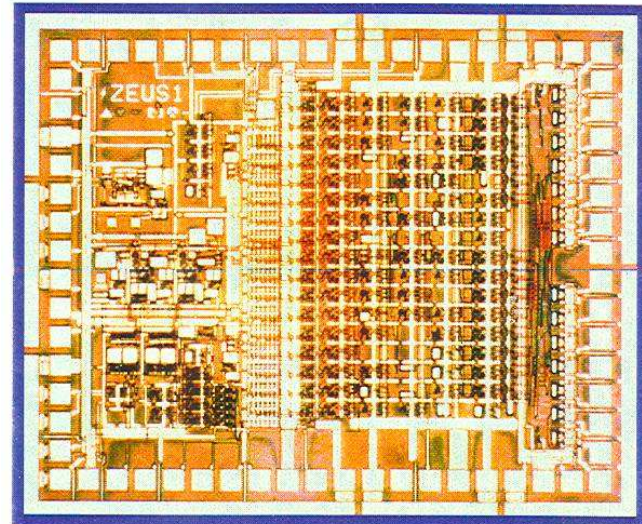
## Skyworks at a Glance



[https://www.skyworksinc.com/About/-/media/SkyWorks/Documents/Downloads/skyworks\\_overview.pdf](https://www.skyworksinc.com/About/-/media/SkyWorks/Documents/Downloads/skyworks_overview.pdf)

# 6-Bit 4 GS/s ADC with AlGaAs/GaAs HBTs

- Supported by AFWL
- 6-Bit ADC with on-chip 1:2 demux
- 1.8 GHz analog BW at 4GS/s
- 6 W power

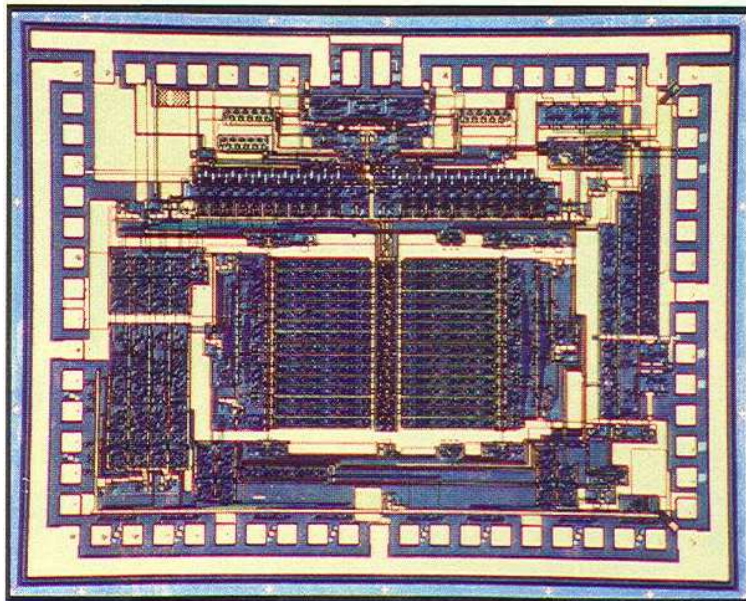


Source: Rockwell & Agilent

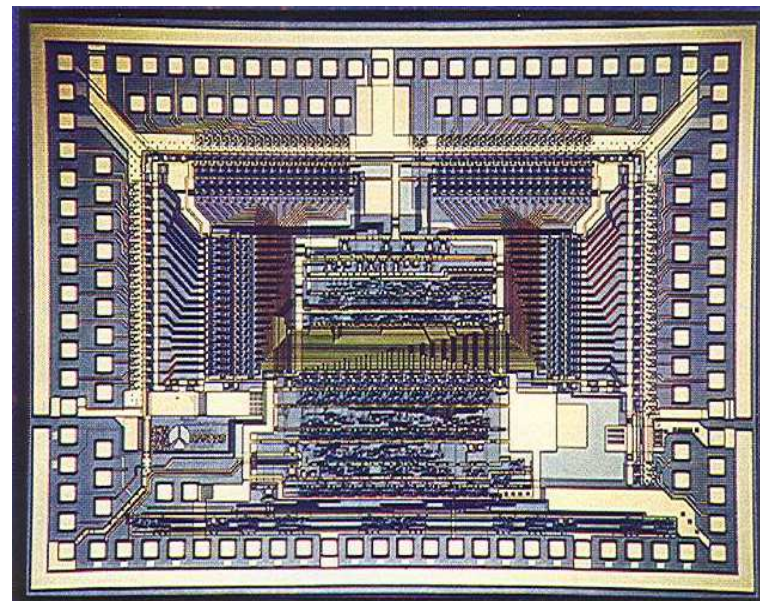


# 8-Bit 2-Gs/s ADC and 8:64 DEMUX

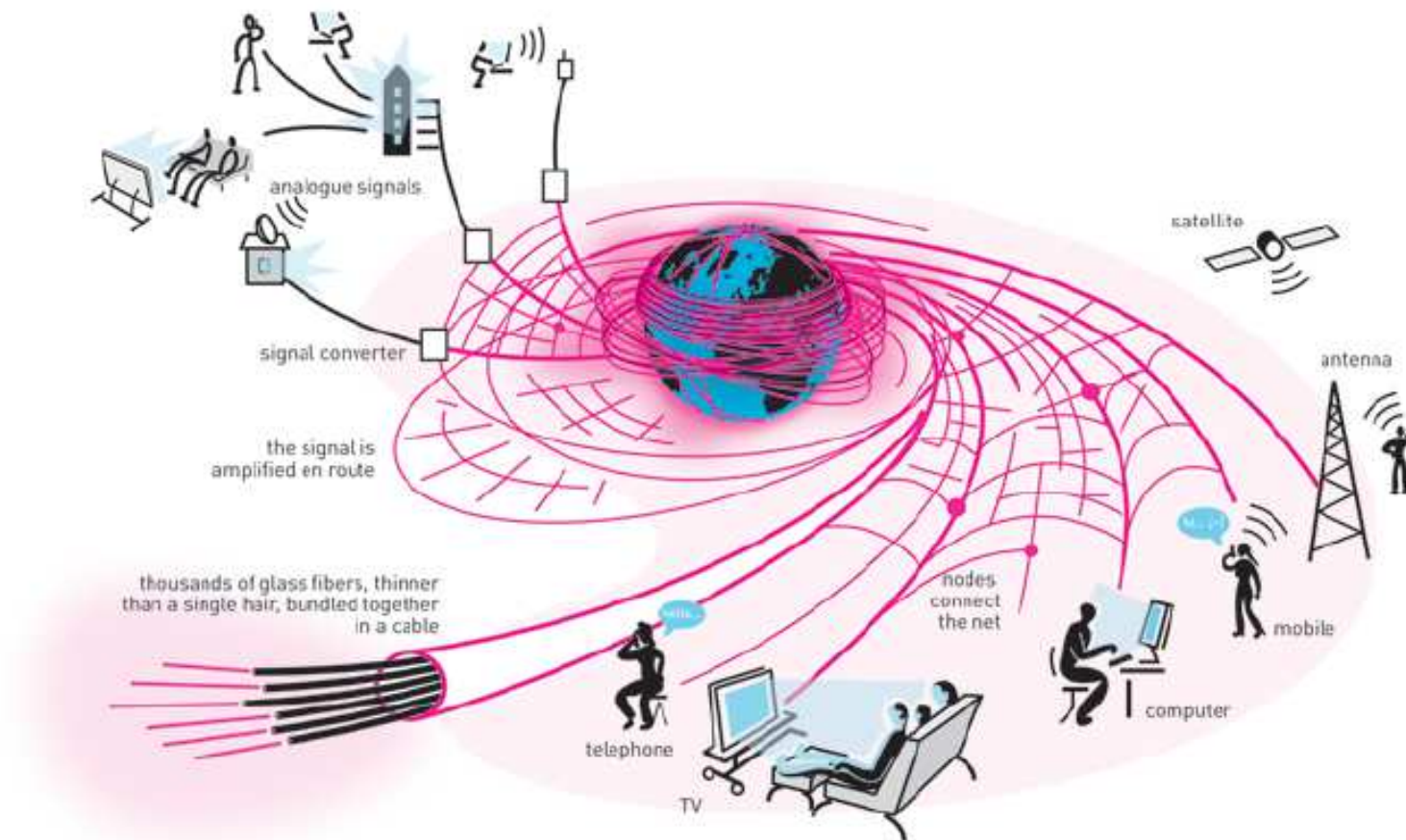
- 2.5GS/s sampling rate
- 3 GHz analog input BW
- IQ Folding and Interpolation architecture
- 5 W power



- Gray to Binary converter
- 3 GB/s
- BIST PRN generator
- 3 W power



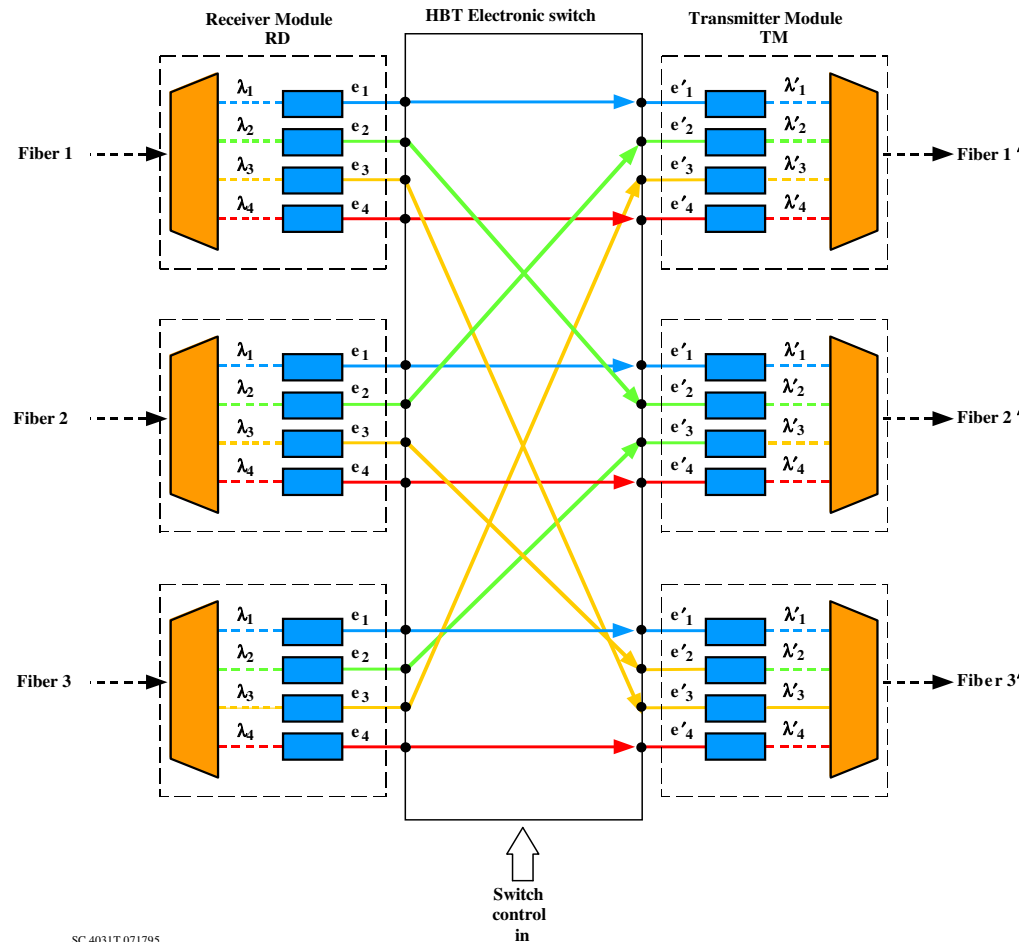
# Artistic View of Global Communication



Sources: The Royal Swedish Academy of Sciences.

# WDM with Electronic Switch Technology (WEST)

## WEST Optoelectronic 3 x 3 WDM Switch



- Low risk
- Fast reconfiguration
- Clock and data regeneration
- Minimum crosstalk
- Flexible switching (Wavelength translation)
- Low cost

WDM: Wavelength Division Multiplexing

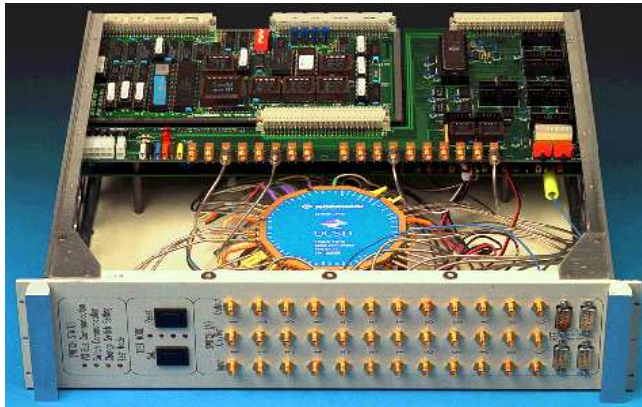
WEST Consortium: DAARPA, Rockwell, Ortel, Caltech/JPL, UCSB, UCSD

Source: Rockwell Science Center



# 12x12 VXI Switch Module

## Feature Highlights

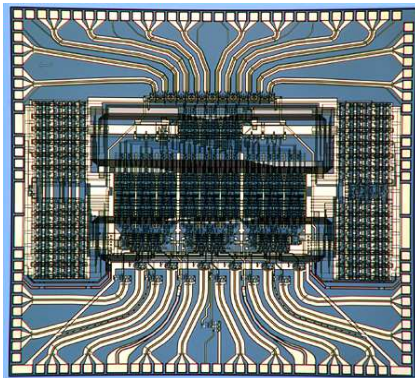


### Switch Module Features:

- 120 Gb/s data throughput
- Twelve 10 Gb/s channels
- VXI Management and Control

### High Speed Package Features:

- Clean high speed interface
- High Isolation
- Thermal management



### Switch Chip Highlights:

- Suitable for larger switching fabrics
- GaAsAlGaAs HBT
- Low crosstalk & jitter generation
- Die size: 4.8 x 5.1 mm<sup>2</sup>
- 4600 transistors
- $P_{diss}$ : 7.4 W



10Gb/s

Source: Rockwell Science Center

# WEST Demonstration



Source: Rockwell Science Center



# WEST Team



Source: Rockwell Science Center

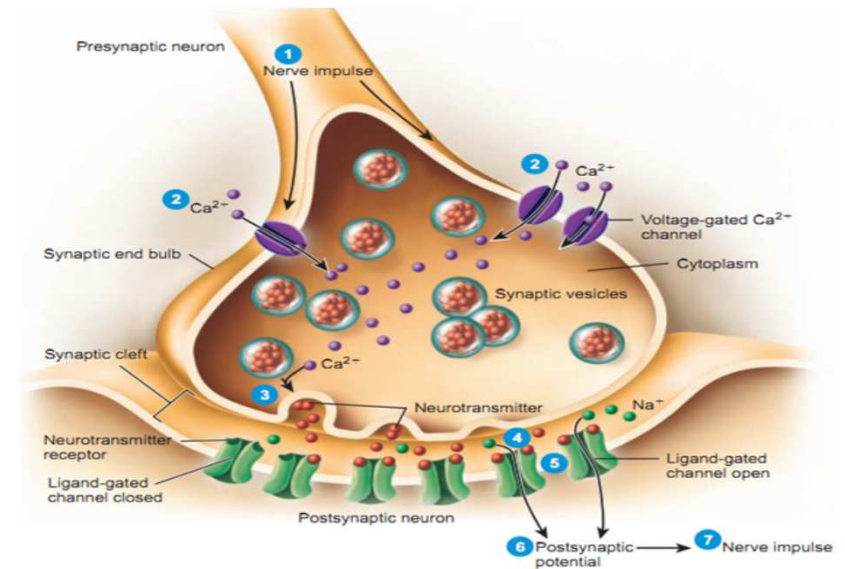
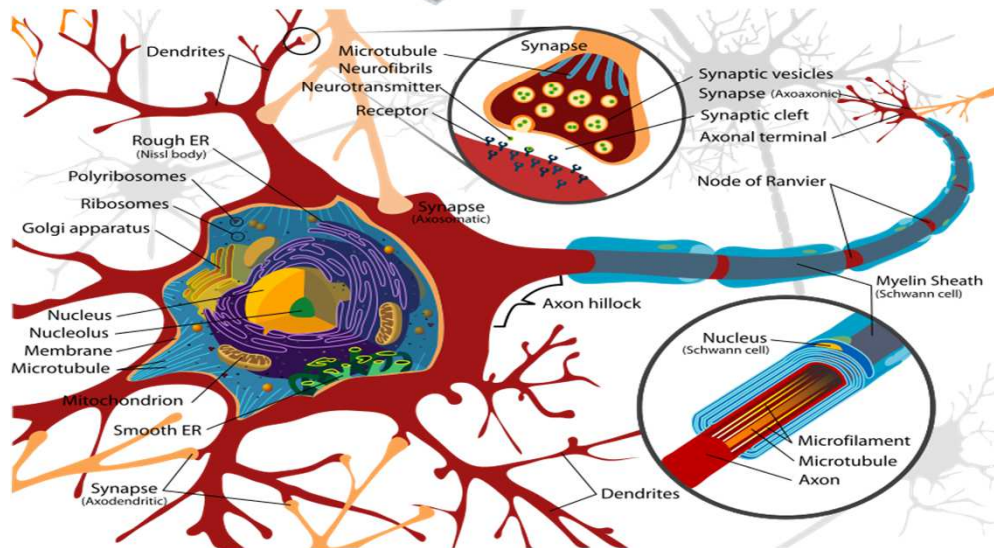
# Outline

- Introduction
- **Experience Sharing**
  - Some Stories of Neutrino
  - GaAs Heterojunction Bipolar Transistor (HBT) and ICs
  - **Semiconductor Memories**
- Conclusion Remarks

# An Amazing Nature Memory

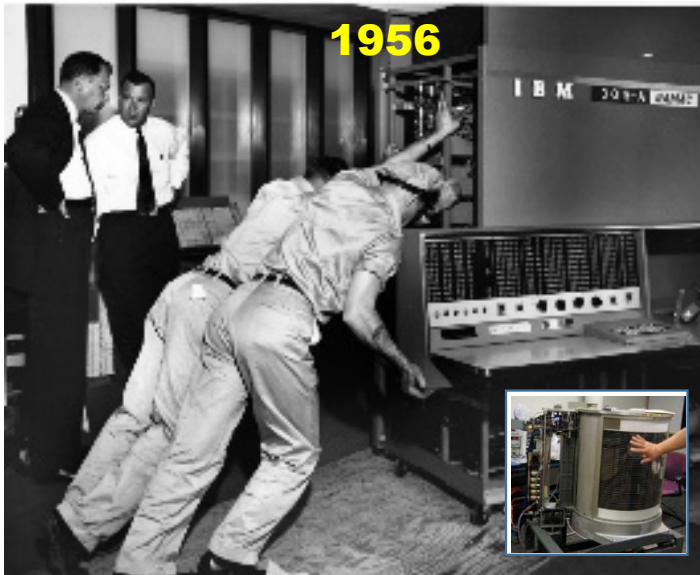


- Fits in a shoe box
- Runs on a tuna sandwich and a cup of water
- Ultra-low power and always on
- 86B neurons, 16B in cerebral cortex, Most of the 16 B neurons have ~10K synapses each
- 1.5Kg (1.2Kg in cerebral cortex)



**A brain keeps information and processes it**

# The Power of Semiconductor Technology



## IBM 350 Disk Storage Unit

Storage Capacity: 5 million characters (3.75 MB)

Weight: over 1 ton

Leased: \$750 per month

One GB capacity needs more than 308 sets IBM 350.

The total rent for one GB storage memory per month is \$ 231,000.



## 2 TB Solid State Drive

## Solid State Drive

Storage Capacity: 2 TB

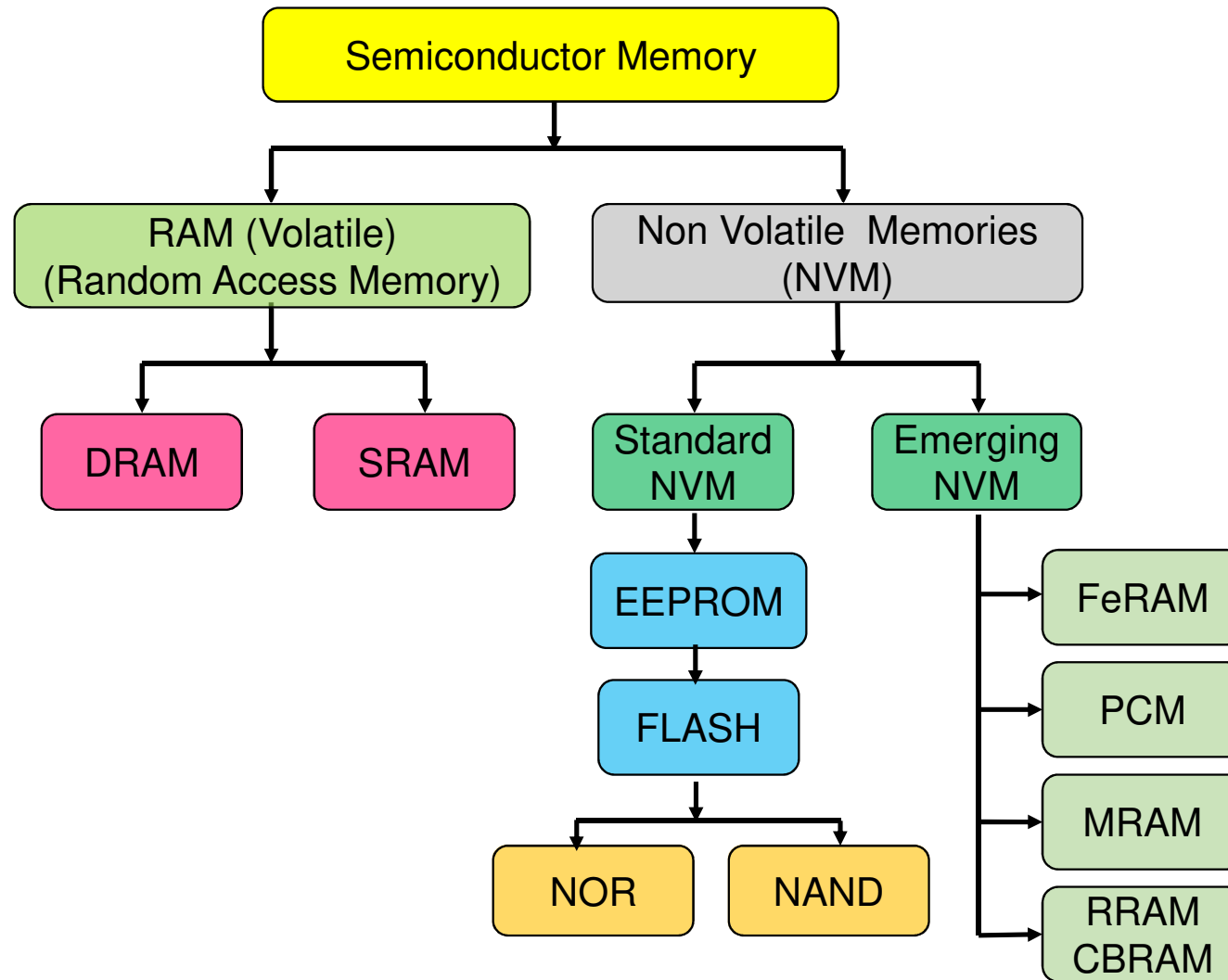
Weight: ~ 80 g

Price: \$250 ( ~ per GB costs \$ 0.125 )

The price drops to 6 order magnitudes



# Semiconductor Memory Family

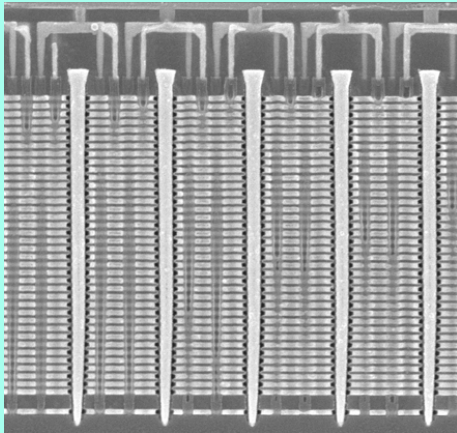
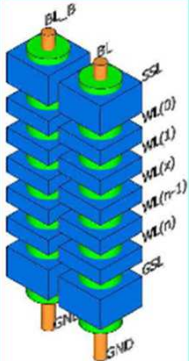


# Mature Semiconductor Memory Devices

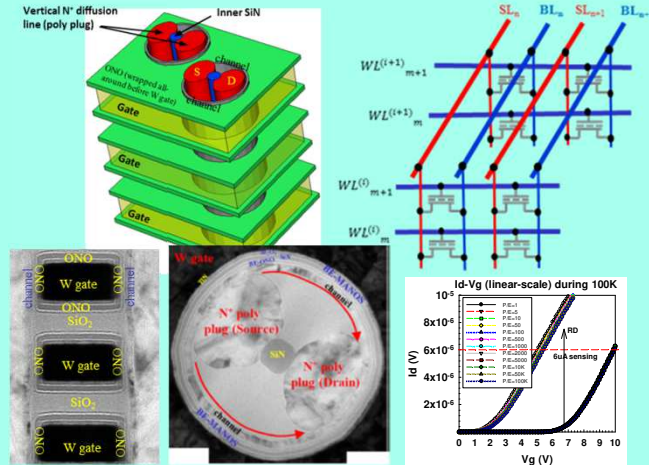
Features	SRAM	DRAM	FLASH
Density	Low	High	High
Speed	Very Fast	Fast	Slow
Leak Power	Low	Medium	Low
Non-volatile	No	No	Yes
Device	6T	1T1C	1T
Cell Structure			

# Landscape of Non-Volatile Memories

## 3D NAND Flash

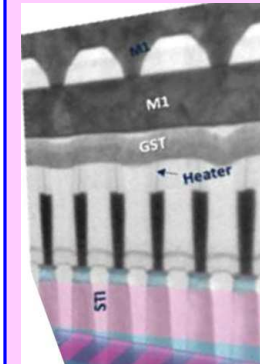


## 3D NOR Flash (in R&D)

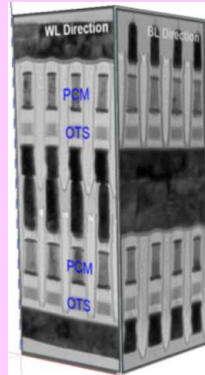
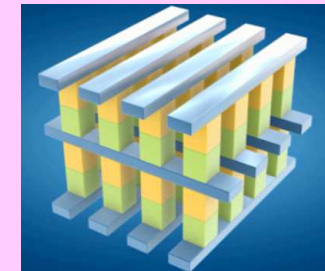


## Phase-Change Memory and 3DXP

1T1R

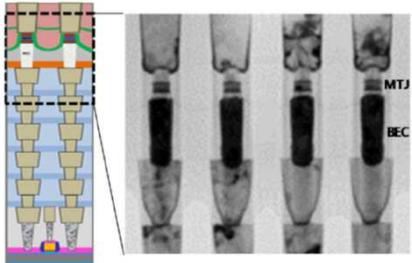


### 3D Cross-Point with selector

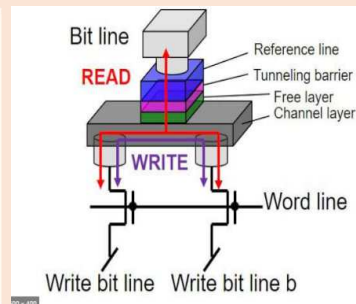


## MRAM

## STT

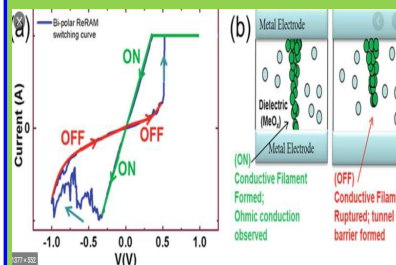


## SOT

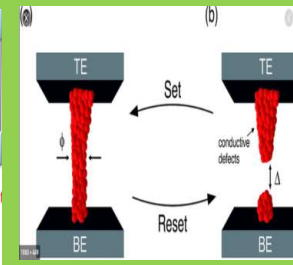


## ReRAM

## Metal Oxide

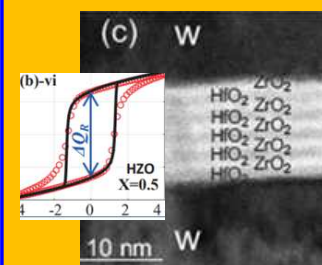


## Conductive-Bridge

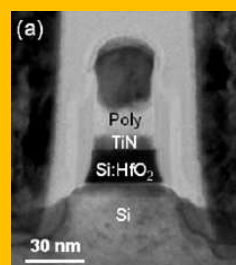


## Ferroelectric RAM

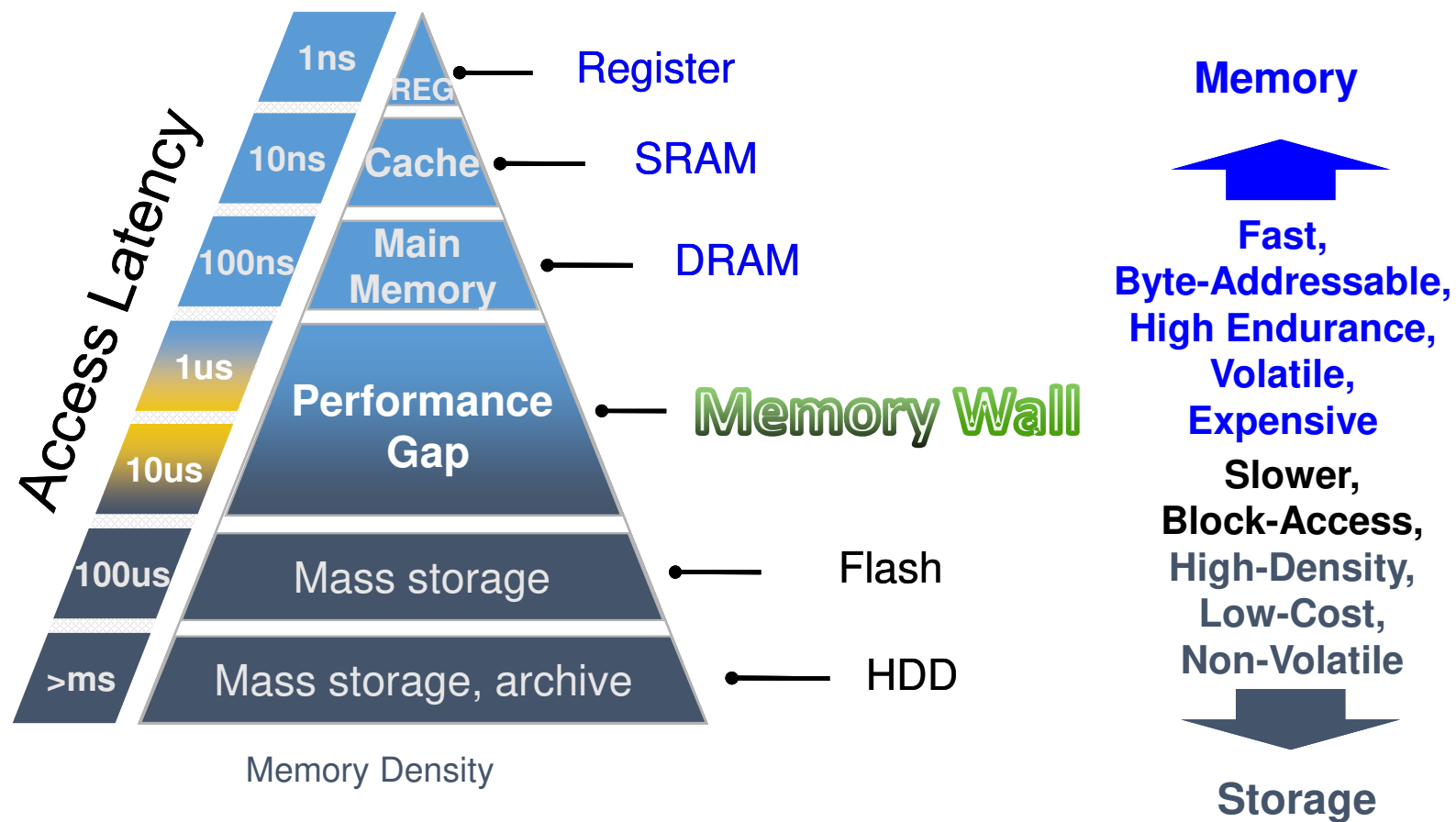
1T1C or 2T2C



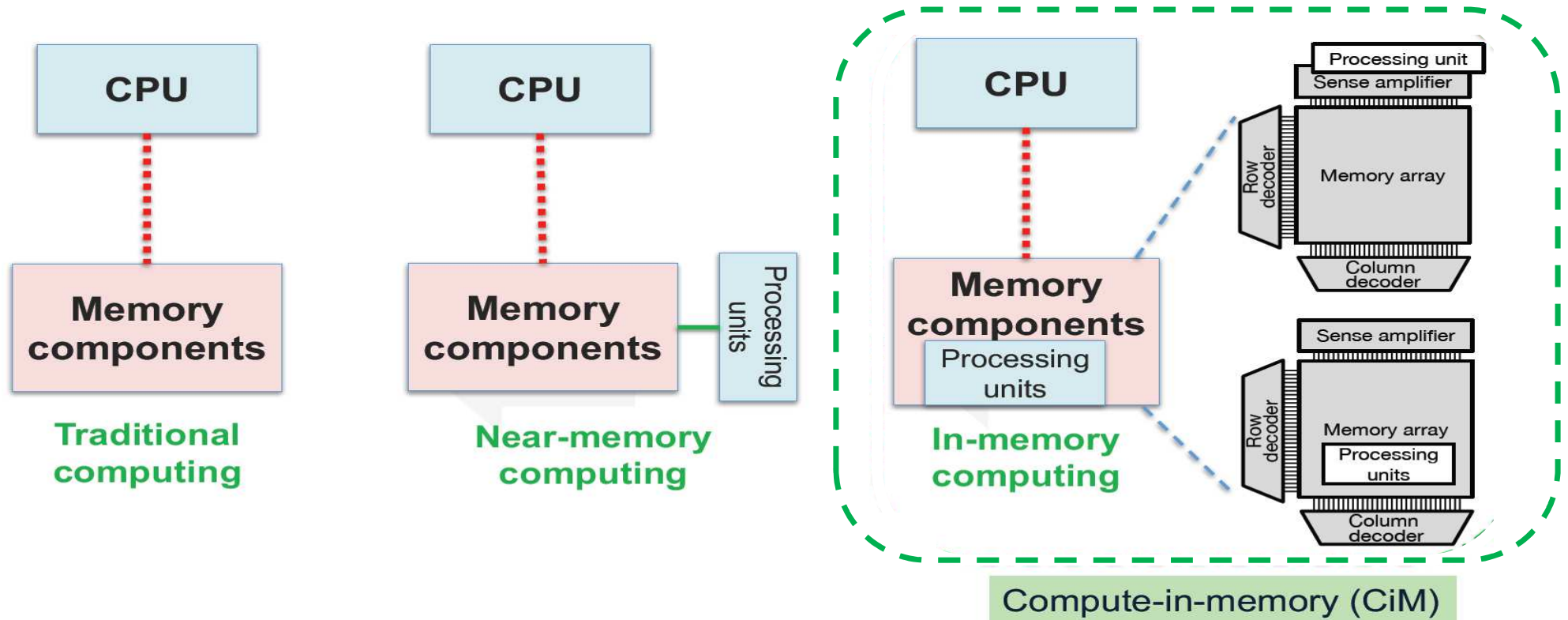
## 1T FeFET



# Memory Hierarchy



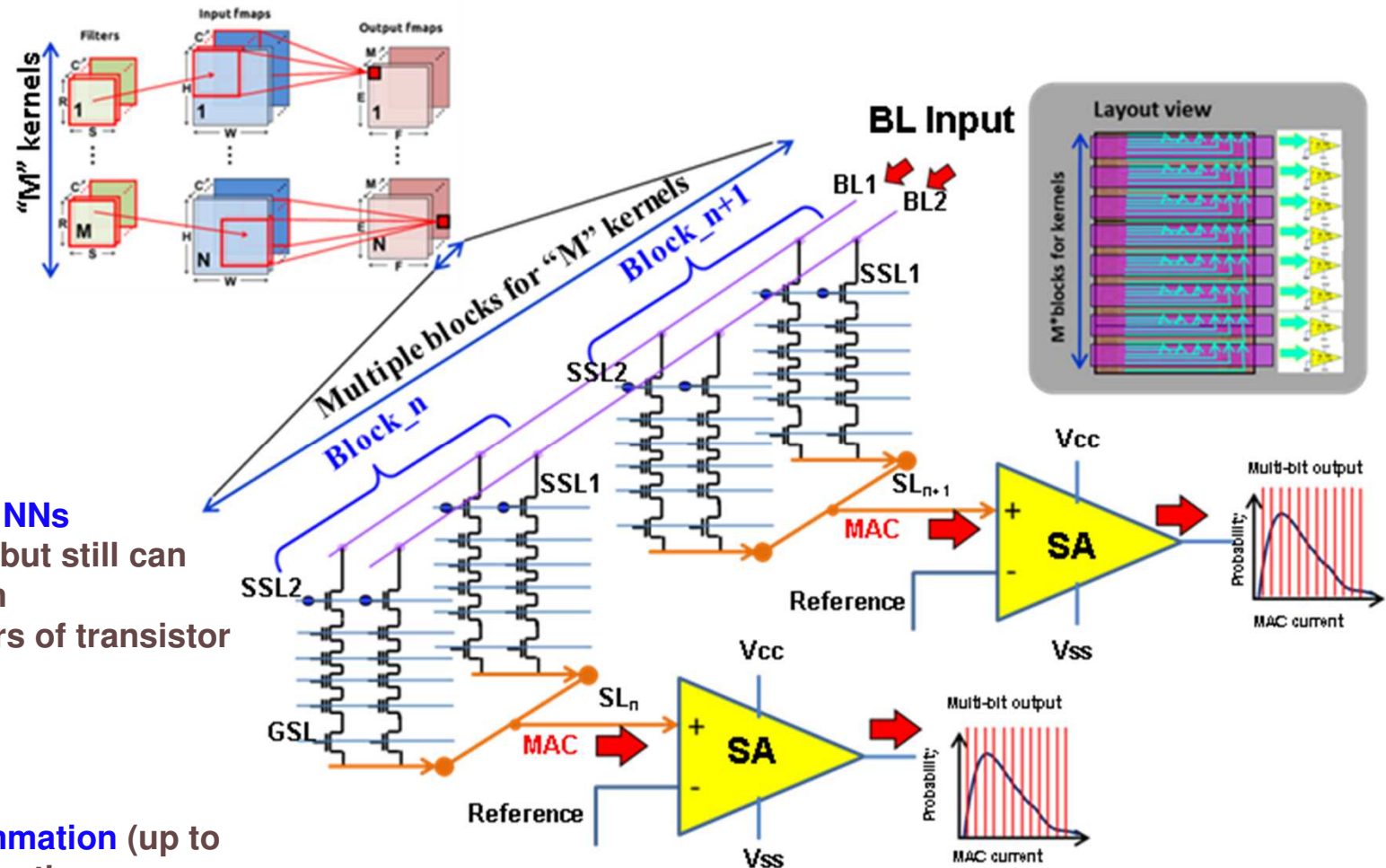
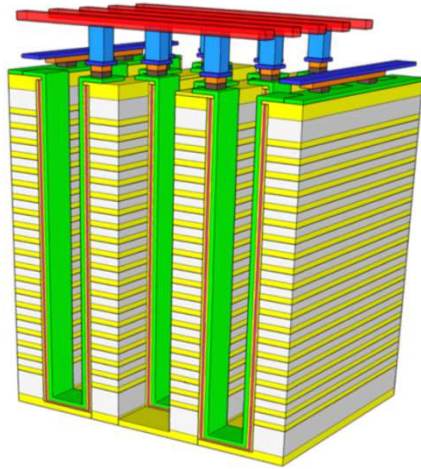
# Memory-Centric Computing is Emerging to Complement Process-Centric Computing



1. Processing units can be simple logic/arithmetic operations or more complex operation sequences
2. “Memory components” can be SRAM, DRAM, NAND, NOR, PCM, ReRAM, MRAM, FeRAM, HDD, etc. depending on applications



# In-Memory Computing Using 3D NAND

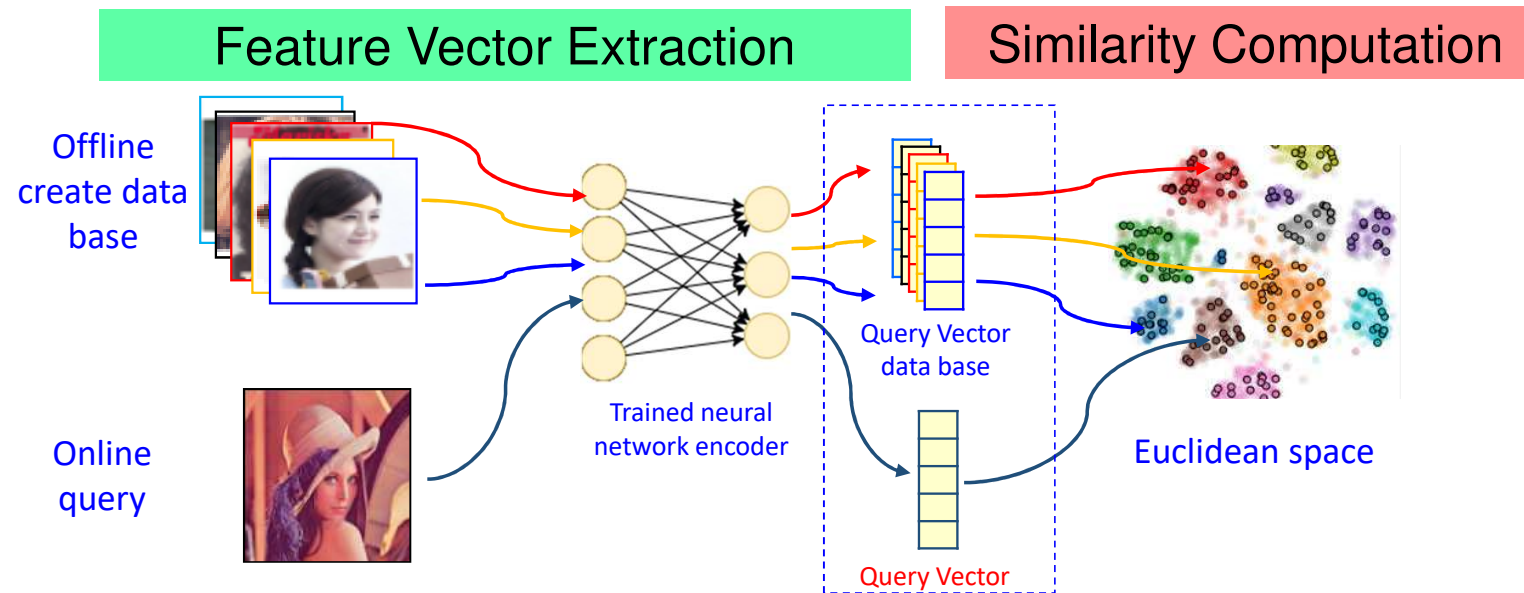


## Key features:

1. High capacity (>64Gb) for **large NNs**
2. SLC storage for best reliability, but still can provide multi-bit MAC operation
3. The large ON/OFF ratio >4 orders of transistor  $I_{cell}$
4. Low  $I_{cell} \sim nA$
5. Tight  $V_t$  distributions

➡ Enable **large and parallel summation** (up to ~10,000) to enhance the MAC operation bandwidth

# In-Memory Search Technology



**An in-memory search engine can be embedded in a data search system as a memory module (device)**

- ✓ nvTCAM: Non-volatile Ternary Content-Addressable Memory
- ✓ Similarity search with Cosine Similarity or Hamming Distance sorter

# In-Memory Search with Flash Memories

- Key Advantages
  - High capacity 512Gb per NAND, 4Gb per NOR
  - Non-volatility, Low cost, Low power (a few 100mW),
  - low read disturbance ( $> 1G$ ) + refresh
- Key Dis-advantages
  - Slow read for NAND ( $\sim 50\mu s$ ), NOR ( $\sim 120ns$ )
  - Limited endurance for NAND ( $\sim 1e4$ ), NOR ( $\sim 1e5$ )
- Some Example Applications



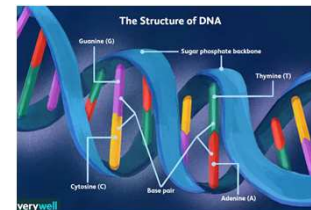
Face Recognition



Large Language Model



Car Localization



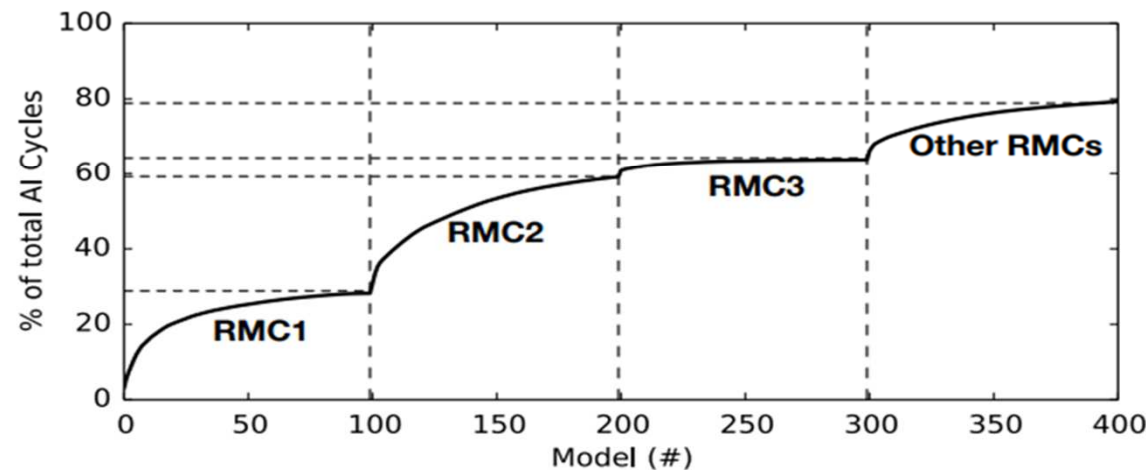
DNA Mapping



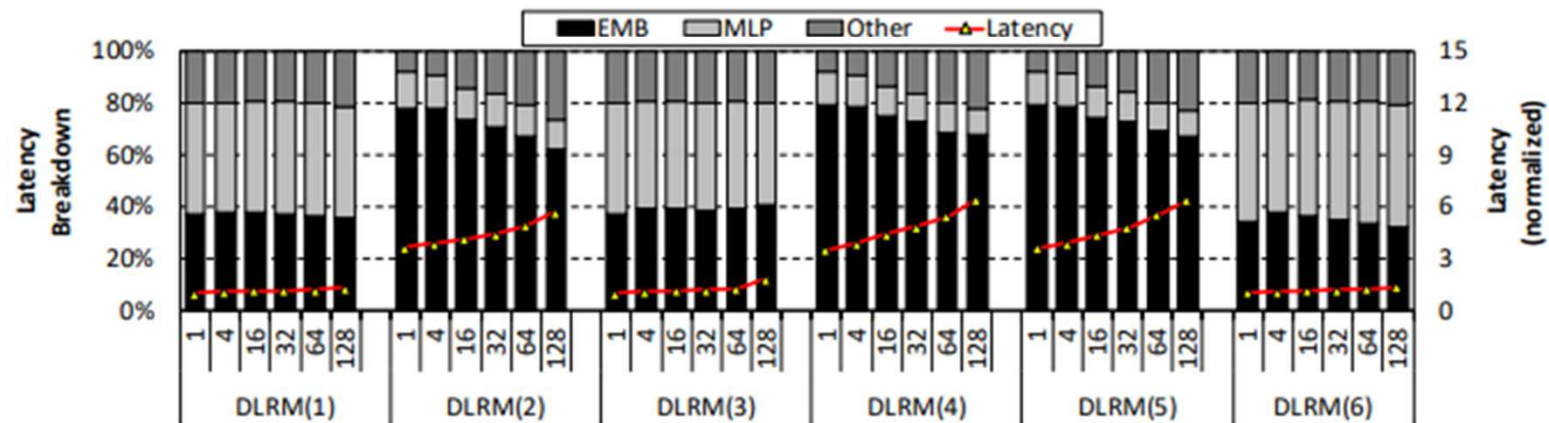
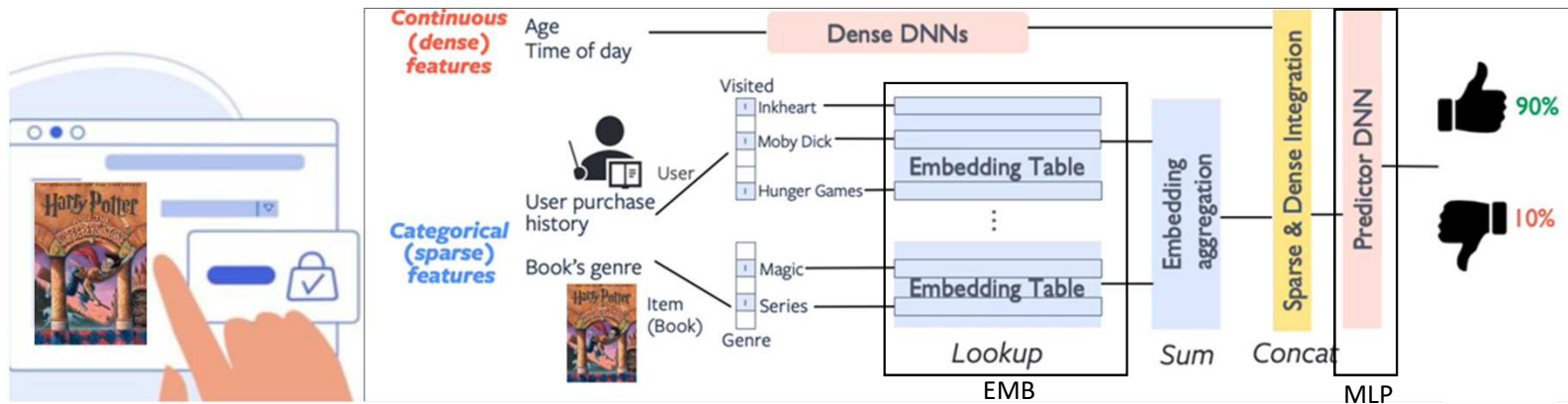
Recommendation

# Recommendation (1)

- DNN-based personalized recommendation models comprise up to 79% of AI inference cycles in a production-scale data center



# Recommendation (2)



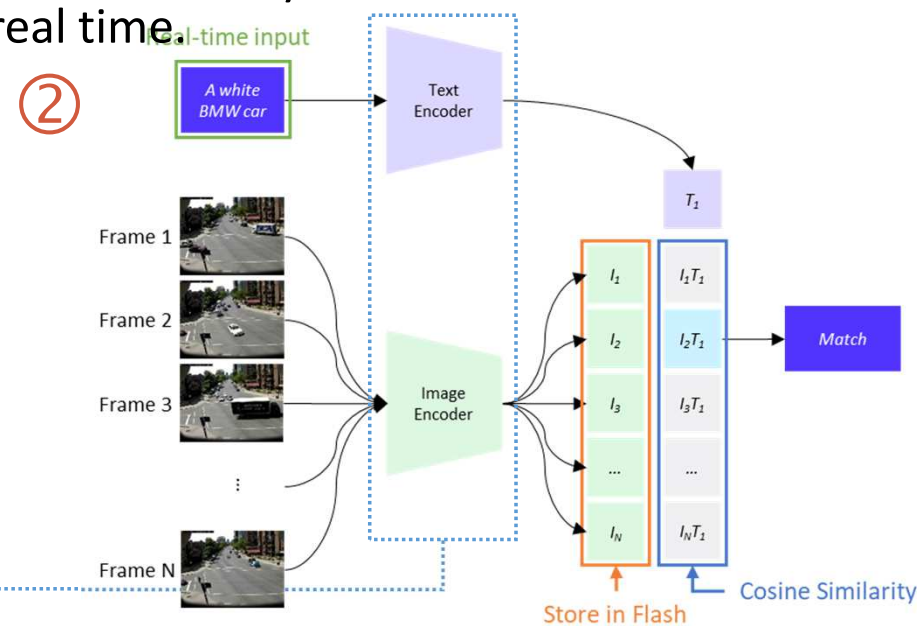
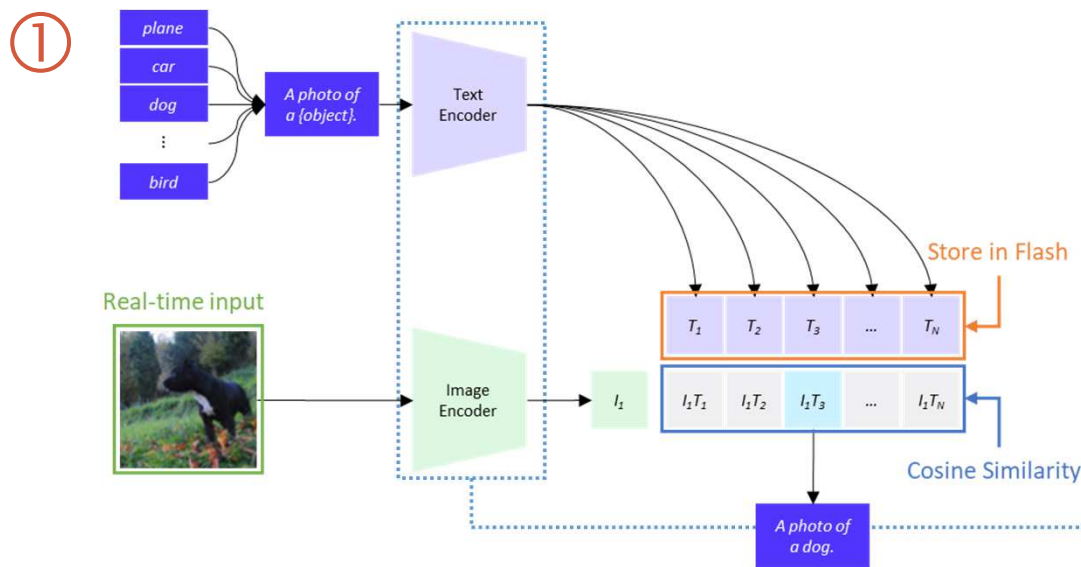
R. Hwang et al. "Centaur: A Chiplet-based, Hybrid Sparse-Dense Accelerator for Personalized Recommendations," (ISCA2020)



# Surveillance (1)

## CLIP (Contrastive Language-Image Pre-Training)

- CLIP is trained to match images with the corresponding captions and vice versa. Therefore, there are two ways to use CLIP
  - A bunch of categories are given and encoded as features. The input image is encoded and searched against the features of the categories in real time.
  - Given a bunch of images, encode them into features. The text you want to find is encoded and searched against the features of the category in real time.



CLIP is a really powerful neural network proposed by OpenAI. It is trained with 400 million (image, text) pairs collected from the internet.

<https://openai.com/blog/clip/>  
<https://arxiv.org/abs/2103.00020>

## Surveillance (2)

### CLIFS: Contrastive Language-Image Forensic Search

- CLIFS is a proof-of-concept for free text searching through videos for video frames with matching contents.
- The searching is done by first extracting features from video frames using the CLIP image encoder and then getting the features for the search query through the CLIP text encoder.
- The features are then matched by cosine similarity and the top results are returned, if above a set threshold.



A truck with the text "odwalla"



A white BMW car



A truck with the text "JCN"



A bicyclist with a blue shirt

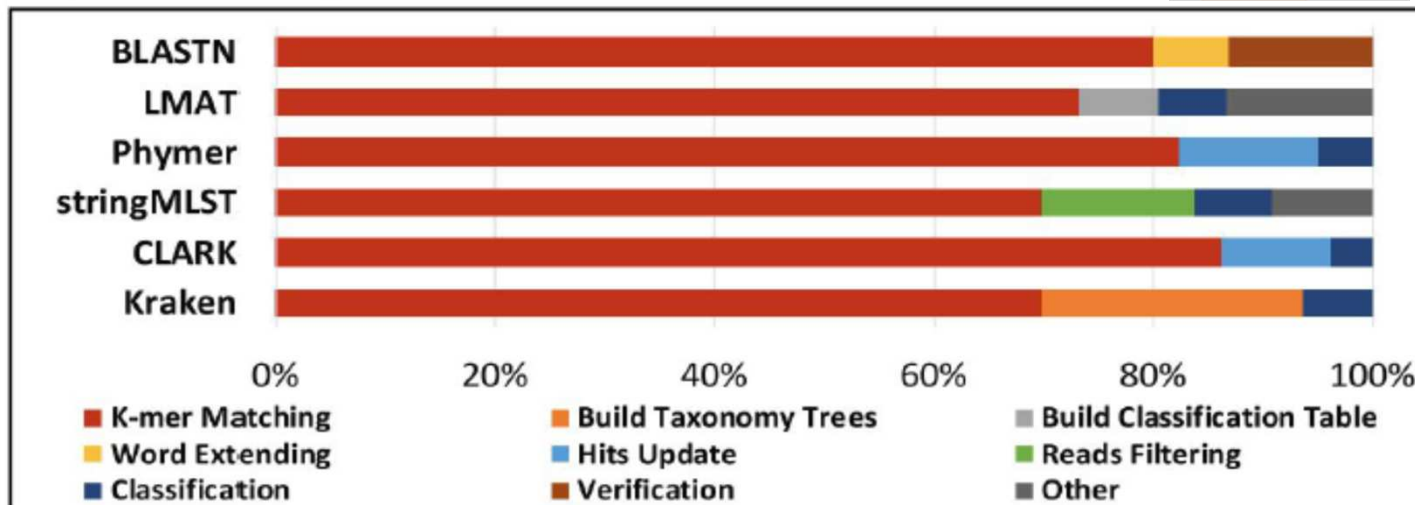
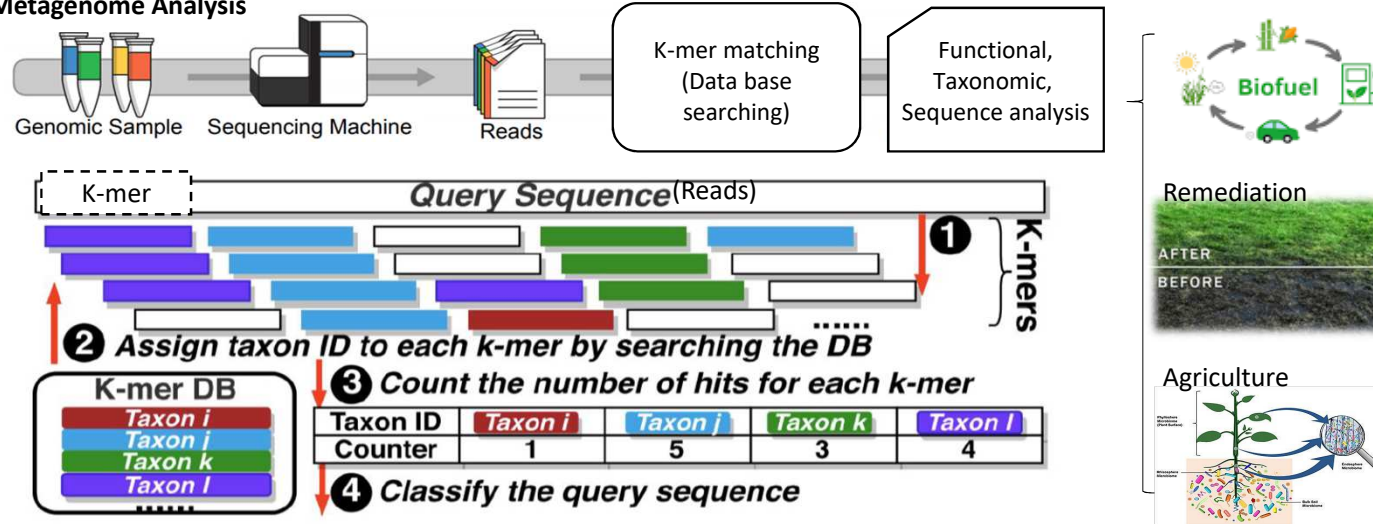


A blue SMART car

- Scenario: There are about 15,000 surveillance camera in Taipei City. Each video will be saved for 1 month.
  - Each frame is embedded to a 512 dimension FP32 feature: 2KB/feature
  - For each second, suppose that one frame is encoded and stored. Then 2,592,000 features will be stored for one month of video.
  - Total storage requirement:  $2\text{KB/feature} * 2,592,000 \text{ features/month} * 15,000 \text{ surveillance camera} = 72.4 \text{ TB}$

# Genomic Analysis

## Metagenome Analysis

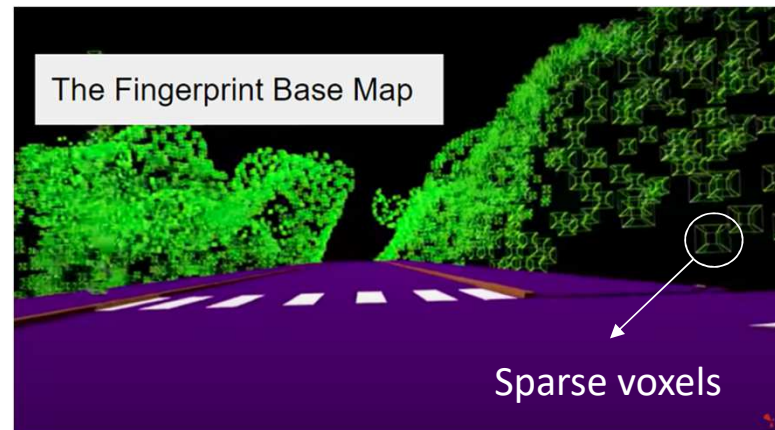
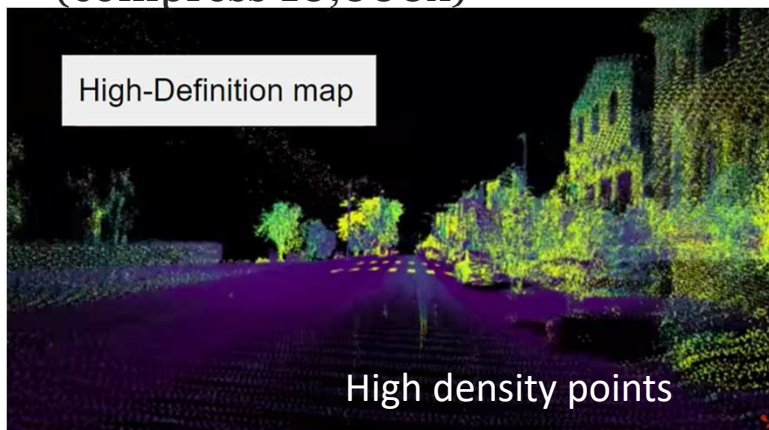


Wu, Lingxi, et al. "Sieve: Scalable in-situ dram-based accelerator designs for massively parallel k-mer matching." (ISCA 2021)

# Car Localization

## Civil Maps' The Fingerprint Base Map™

- A high-definition map (HD map) is a highly accurate map used in **localization** for autonomous driving.
- Civil Maps' The Fingerprint Base Map™ is an edge-based map for real-time localization
  - Each voxel contains its unique fingerprint.
  - It allows a **dictionary lookup** of the location in **0.1 second** and achieves **15 cm – 20 cm** accuracy at high speed.
  - A HD map of San Francisco, which is about 4,000 kilometers, can take up to **4 terabytes**, and Base Map of only **~400 megabytes** for that same region (compress 10,000x)





# Outline

- Introduction
- Experience Sharing
  - Some Stories of Neutrino
  - GaAs Heterojunction Bipolar Transistor (HBT) and ICs
  - Semiconductor Memories
- **Conclusion Remarks**

# It is a Privilege Being a Physicist

(Victor F. Weisskopf)

- **Studying physics is a lucky, interesting, and practical journey. It trains us to learn new things fast and be leaders of many fields of research and work**
- **Semiconductor industry has been advancing exponentially following Moore's Law since 1960's**
  - It enabled many electronic products that enriched our lives
  - It accelerates developments of AI, IoT, communication, sensing, biomedicine, etc.
  - Memory-centric computing is emerging to complement process-centric computing
- **Neutrino physics research continues to be exciting and important for exploring frontier knowledge of Nature**
- **Competitions in both fields are severe. The keys to success are passion, innovation, and persistence**

# Acknowledgement

- 夏期岳先生夫人
- Felix Boehm
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- Jackson Hu
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- Rich Liu
- Teachers
- Classmates
- Colleagues
- Sponsors
- Customers
- Collaborators
- Friends
- Father
- Wife, Children, Relatives
- ...



# Thank You!

[kcwang@mxic.com.tw](mailto:kcwang@mxic.com.tw)