

# Rapid Parameter Estimation Using Reduced Order Quadrature

10<sup>th</sup> KAGRA International Workshop (KIW-10)

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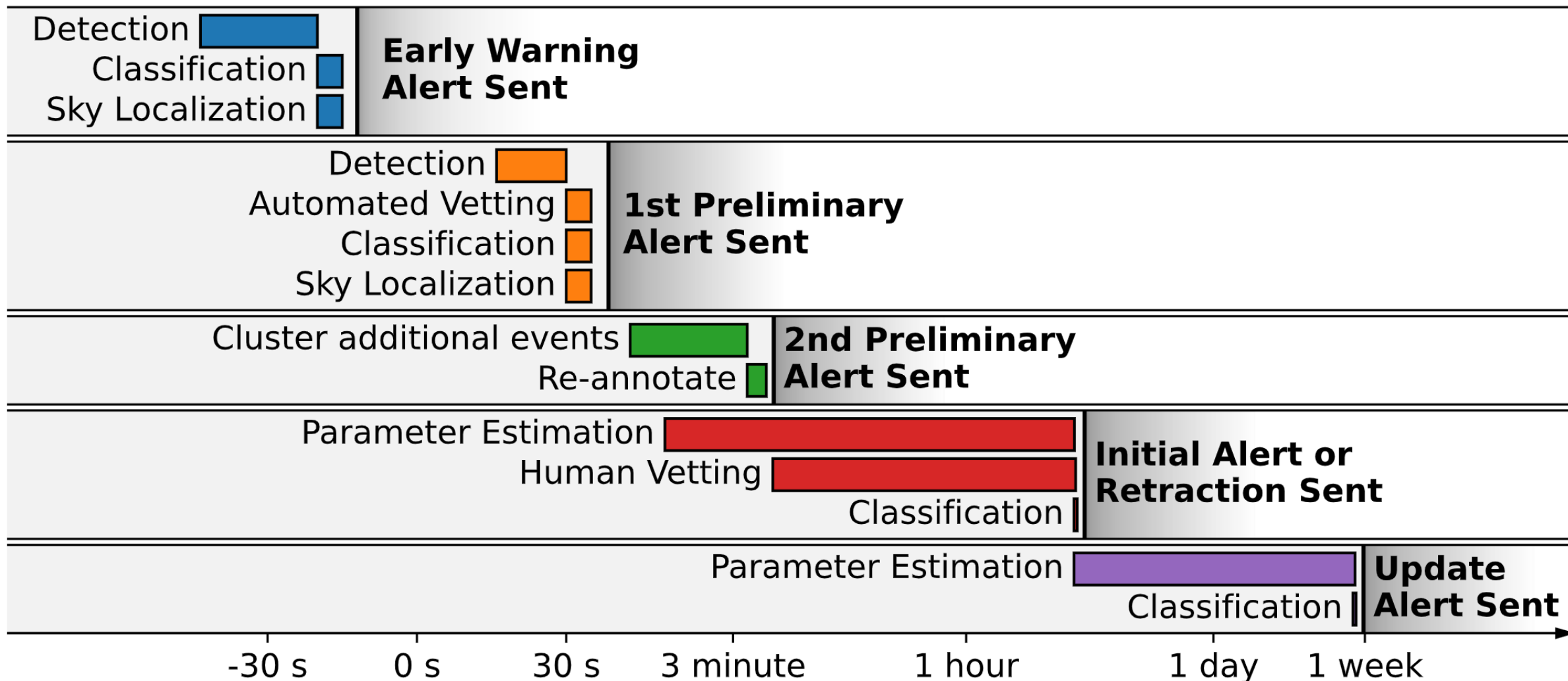
THE UNIVERSITY OF TOKYO



# Low-latency alert in O4

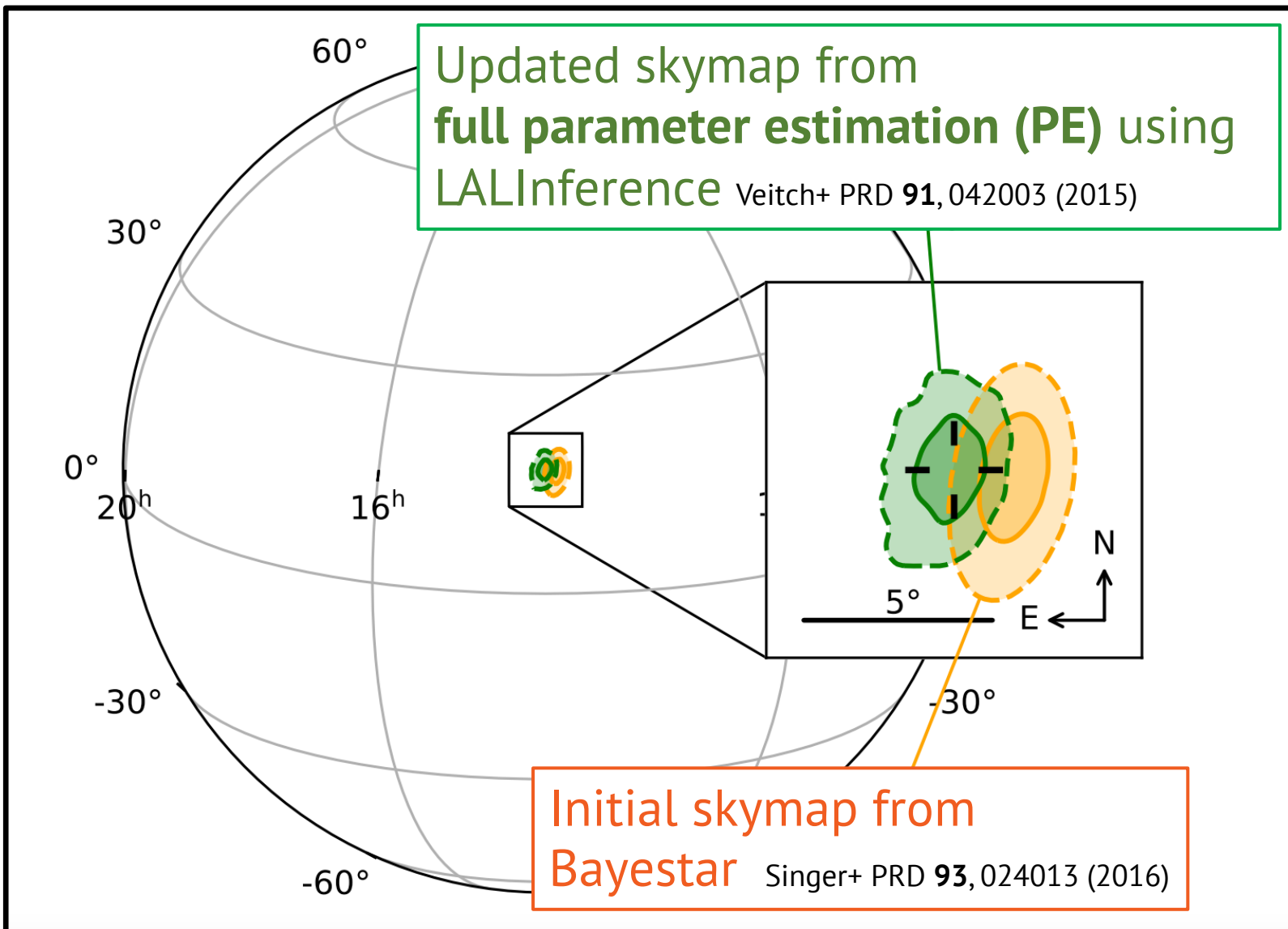
Credit: Public Alerts User Guide  
(<https://emfollow.docs.ligo.org/userguide/index.html>)

Time relative to gravitational-wave merger



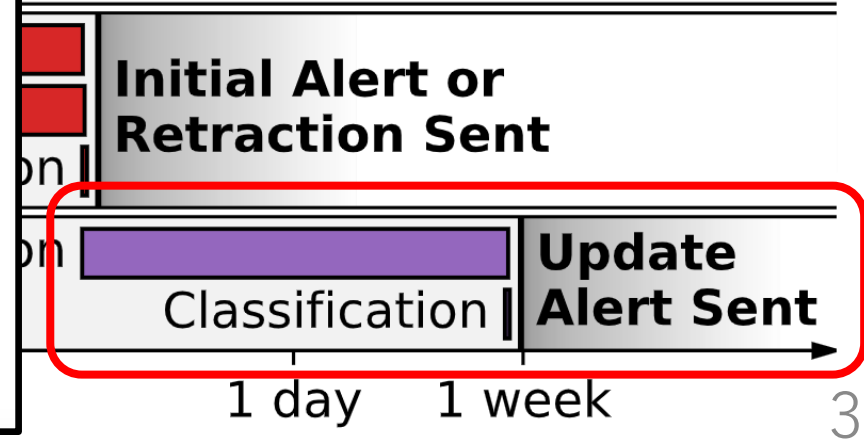
# Low-latency alert in O4

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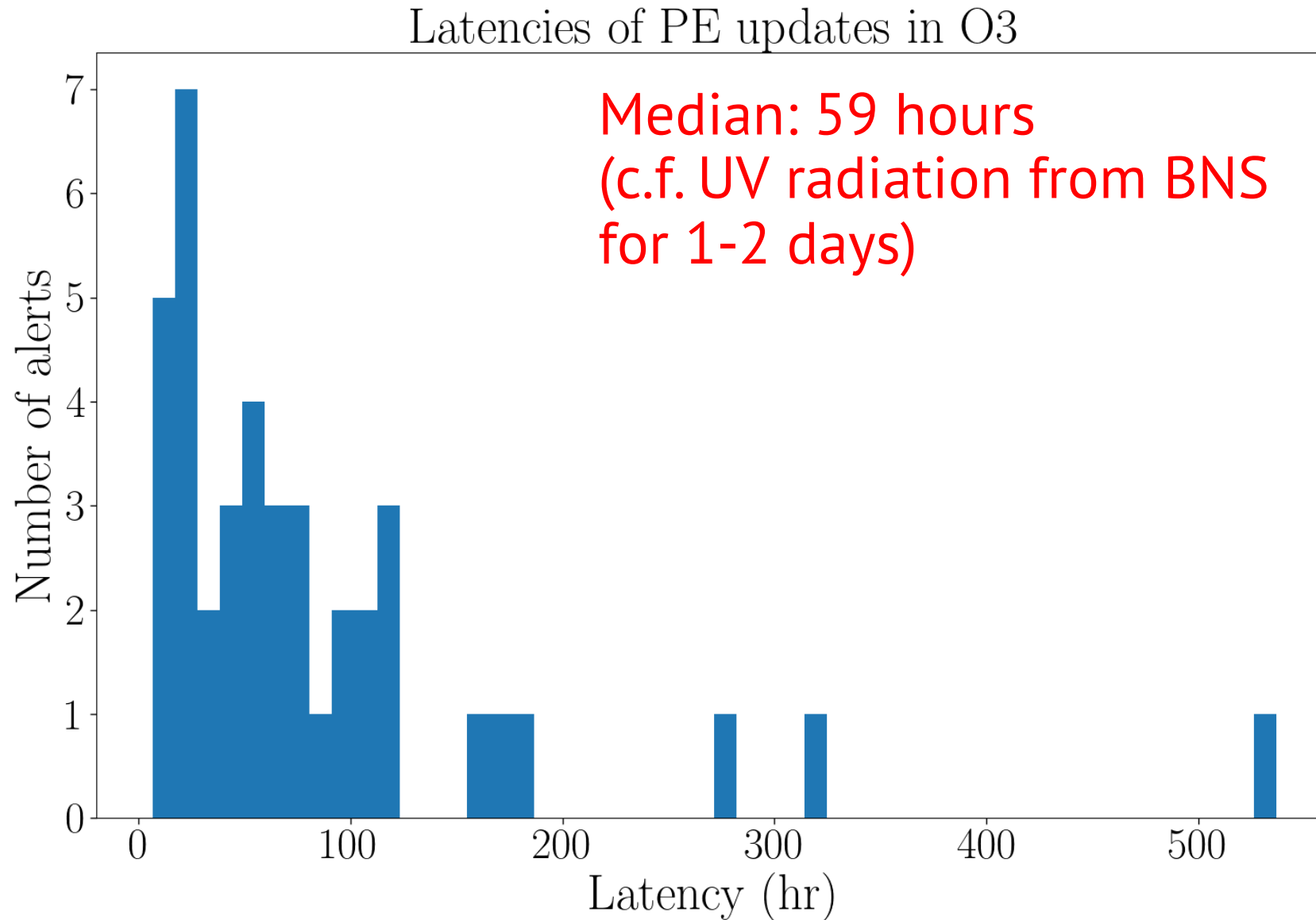


merger

ry



# PE latency in O3



# Reduced Order Quadrature (ROQ)

P. Canizares et al., Phys. Rev. Lett. **114**, 071104 (2015).  
R. Smith et al., Phys. Rev. D **94**, 044031 (2016).

- PE requires more than millions of waveform generation.
- Its cost is proportional to  $L$ , the number of frequency points  $\{f_l\}_{l=1}^L$ .
- Interpolate waveform values at  $K \ll L$  interpolation nodes  $\{F_k\}_{k=1}^K$ ,

$$\tilde{h}(f_l) \simeq \sum_{k=1}^K \tilde{h}(F_k) \underbrace{B_k(f_l)}_{\text{ROQ basis}}. \quad (l = 1, 2, \dots, L)$$

→ Speed-up by  $L/K$

# Optimizing ROQ for O4

- Significant speed-up if bases are constructed for targeted mass-spin space.  
See **SM** and Vivien Raymond, Phys. Rev. D **102**, 104020 (2020).  
→ Use **multiple ROQ bases**, each constructed for targeted chirp-mass range.
- **Adaptive frequency resolution** to speed up pre-computation of **ROQ weights**.

$$\langle d, h \rangle = \frac{4}{T} \sum_{l=1}^L \frac{\tilde{d}^*(f_l) \tilde{h}(f_l)}{S(f_l)} = \sum_{k=1}^K \tilde{h}(F_k) \omega_k, \quad \omega_k \equiv \frac{4}{T} \sum_{l=1}^L \frac{\tilde{d}^*(f_l) B_k(f_l)}{S(f_l)}.$$

See **SM**, Phys. Rev. D **104**, 044062 (2021).

Those improvements were implemented in **Bilby**, one of the LVK PE libraries.

Ashton+ ApJS **241** 27 (2019),  
Romero-Shaw+ MNRAS **499** 3 (2020).

# ROQ bases for low-latency PE of BNS

- Waveform model: **IMRPhenomD** Husa+ PRD **93**,044006 (2016), Khan+ PRD **93**,044007 (2016).
- Frequency range:  $20\text{Hz} < f < 1024\text{Hz}$
- Parameter range
  - Chirp mass:  $0.6M_{\odot} < \mathcal{M} < 4M_{\odot}$
  - Mass ratio:  $Q \equiv m_1/m_2 < 8$
  - Spins:  $\mathbf{a_1, a_2} < \mathbf{0.05}$  (based on BNS observations)
- Trained against calibration errors of  $< 20\%$  in amplitude and  $< 15\text{deg}$  in phase.
- Speed-up gain: **400 - 3500**
- Run time: **~7min** for SNR~25

# ROQ bases for NSBH and BBH

Both cover **mass ratio up to 20** and **spin magnitude up to 0.99**.

## NSBH

- Waveform model: **IMRPhenomPv2** Hannam+ PRL **113**, 151101 (2014)
- Chirp-mass range:  $1.4M_{\odot} < \mathcal{M} < 21M_{\odot}$
- Run time: **~3 hours** for SNR~20

## BBH

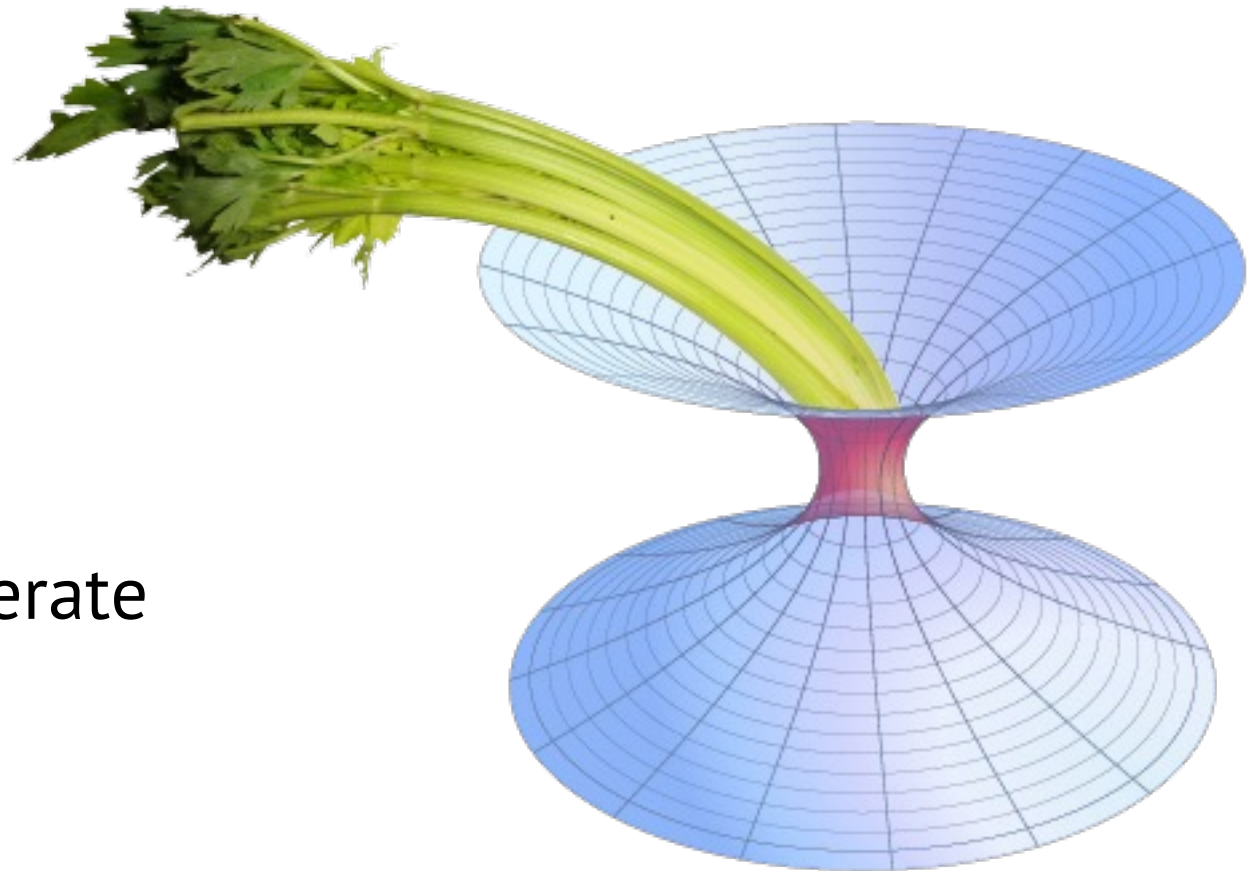
- Waveform model: **IMRPhenomXPHM** Pratten+ PRD **103**, 104056 (2021)
- Chirp-mass range:  $10M_{\odot} < \mathcal{M} < 200M_{\odot}$
- Run time: **~4 hours** for SNR~30



# Automation with **GWCelery** <https://git.ligo.org/emfollow/gwcelery>

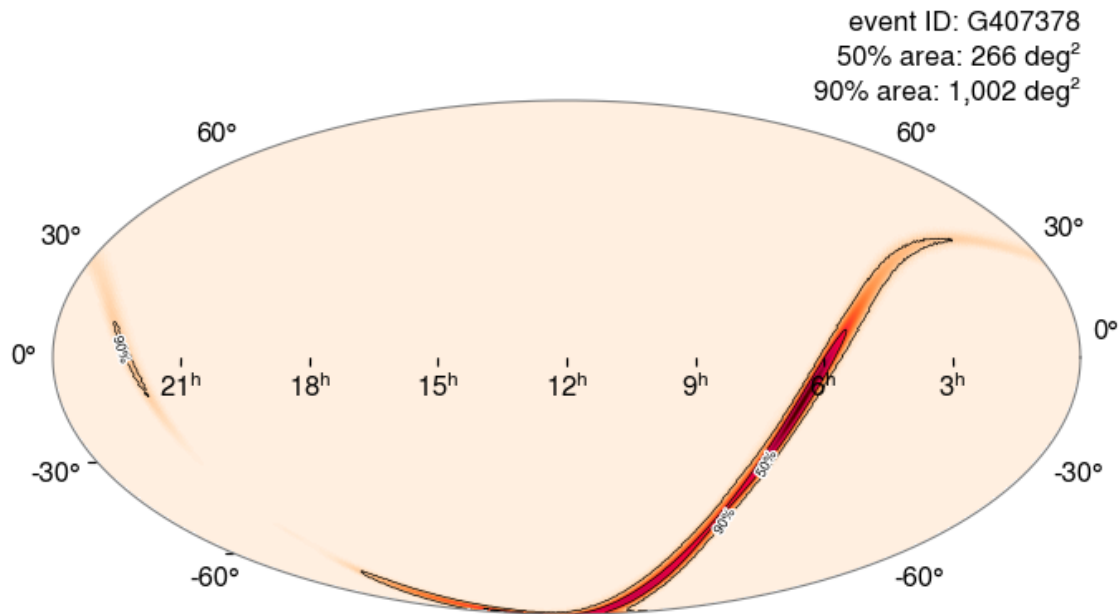
Parameter estimation is automated by **GWCelery**, the LIGO-Virgo-KAGRA alert pipeline.

- Submit and monitor PE jobs automatically.
- Postprocess PE results to generate skymap and EM-bright probs.

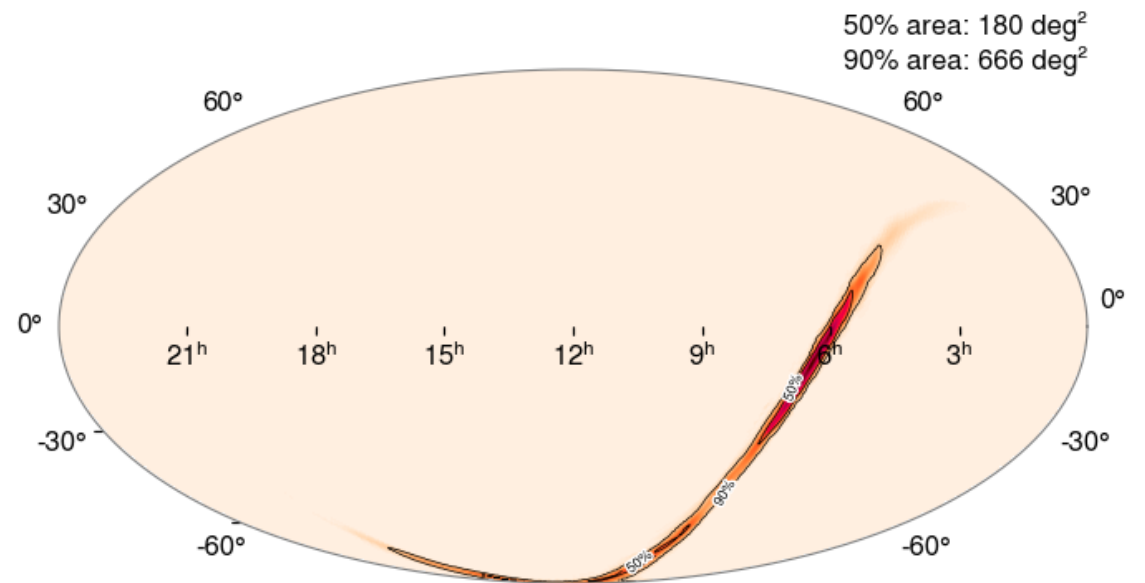


# S230518h

The updated skymap was sent in **~4 hours** after the detection, reducing 90% area **from 1002deg<sup>2</sup> to 666deg<sup>2</sup>** (See GCN circular 33816).



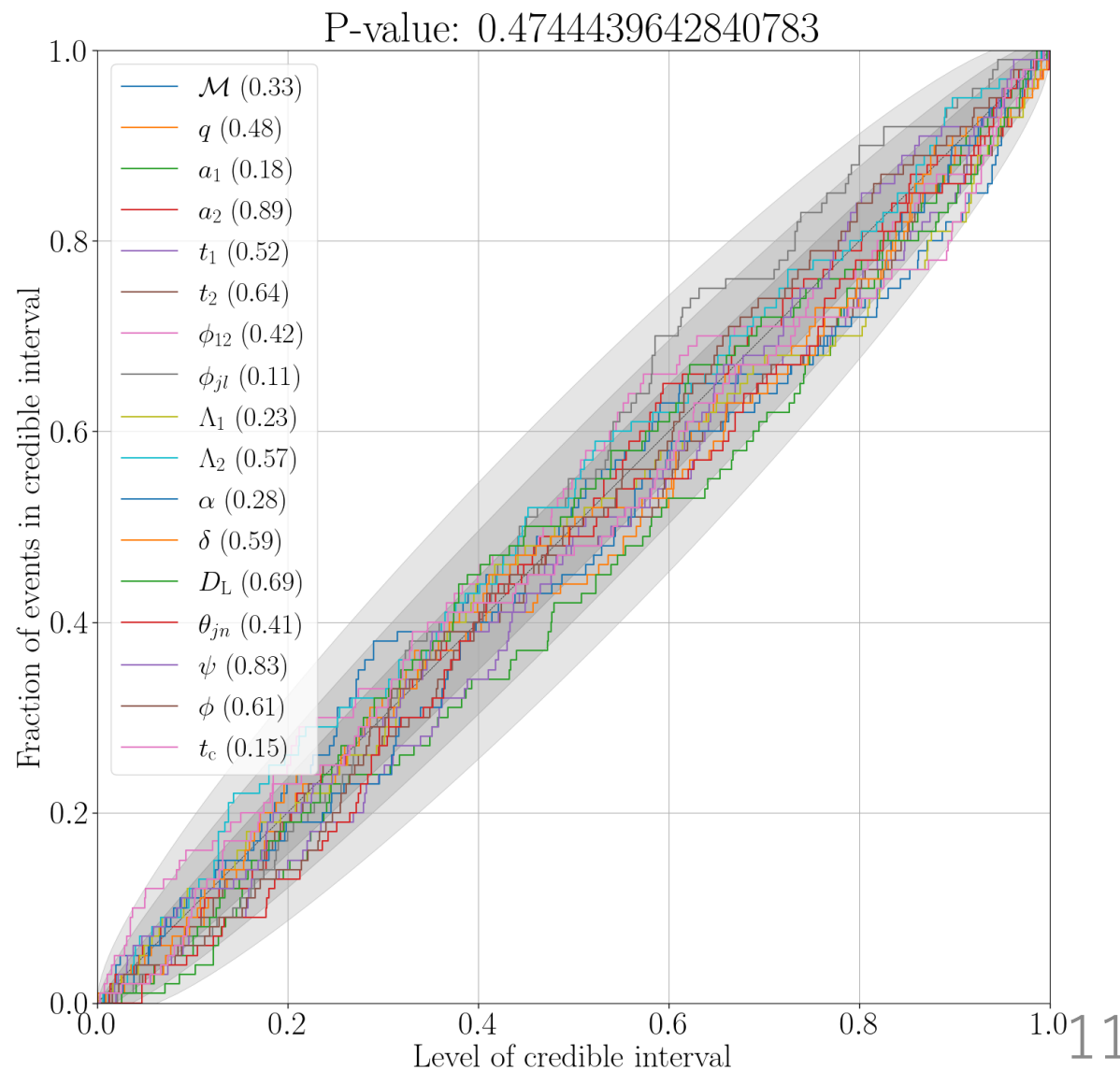
**Bayestar skymap**



**Bilby skymap**

# ROQ bases for BNS incorporating tides

- Waveform model:  
**IMRPhenomPv2\_NRTidalv2**  
Dietrich+ PRD 100,044003 (2019)
- Parameter range
  - Chirp mass:  $0.6M_{\odot} < \mathcal{M} < 4M_{\odot}$
  - Mass ratio:  $Q \equiv m_1/m_2 < 8$
  - **Spins:  $a_1, a_2 < 0.99$**
  - **Tides:  $\Lambda_1, \Lambda_2 < 5000$**
- Trained against calibration uncertainties.
- **Enable sampler tests within a few hours.**



# Summary

- Optimization to reduced order quadrature (ROQ) for O4
  - Multiple ROQ bases in a single parameter estimation (PE) run
  - Adaptive frequency resolution to speed up pre-computation
- Improved ROQ enables rapid (minutes-hours) updates of skymap in O4.
- Can be applied to test sampler for more detailed PE.