

Exploring new frontiers in laser-driven ion acceleration with the aid of artificial intelligence

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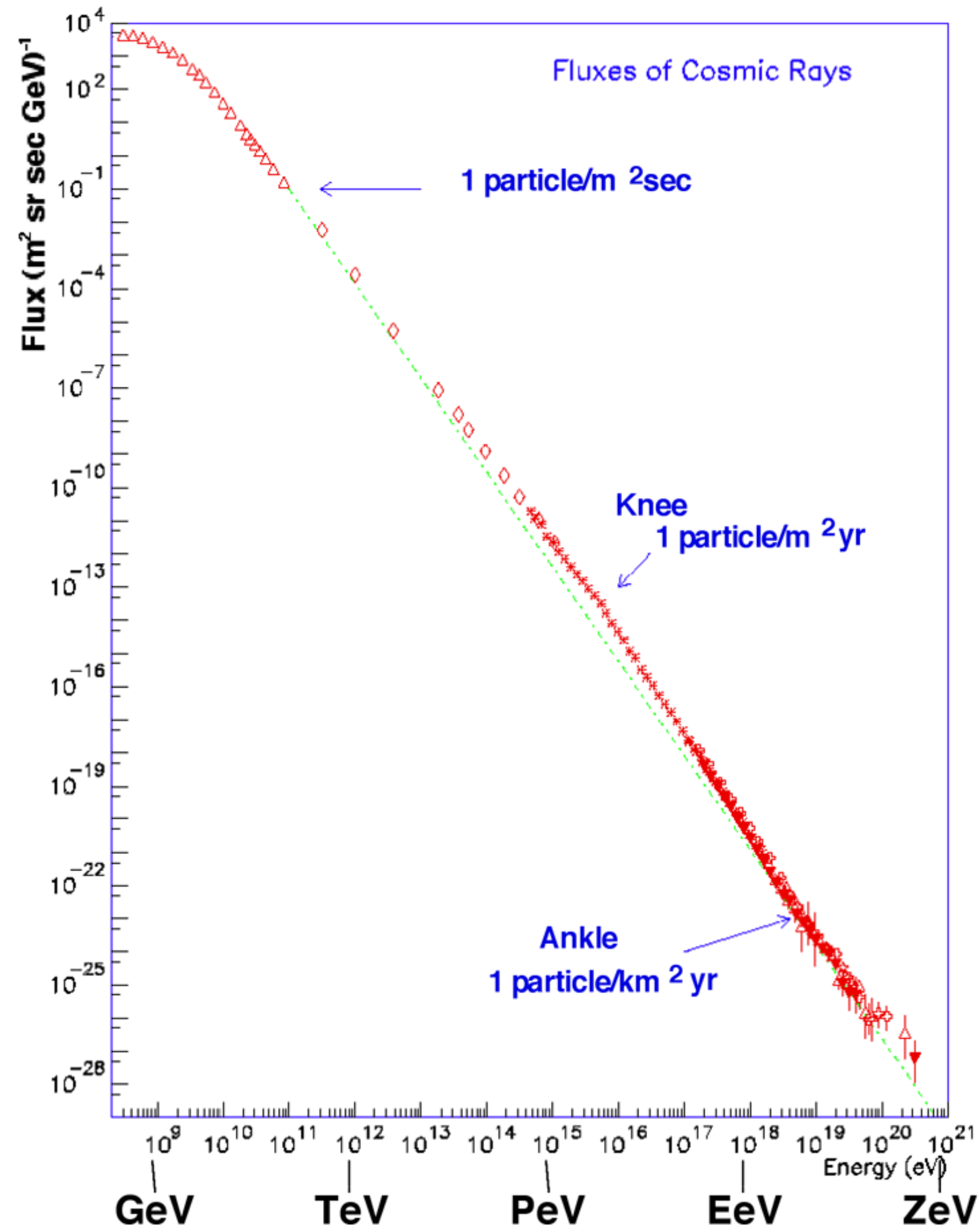
Graduate School of Engineering, Osaka University

CHiP Annual Meeting

23–24 Nov 2023 Evergreen Palace Hotel (Chiayi)

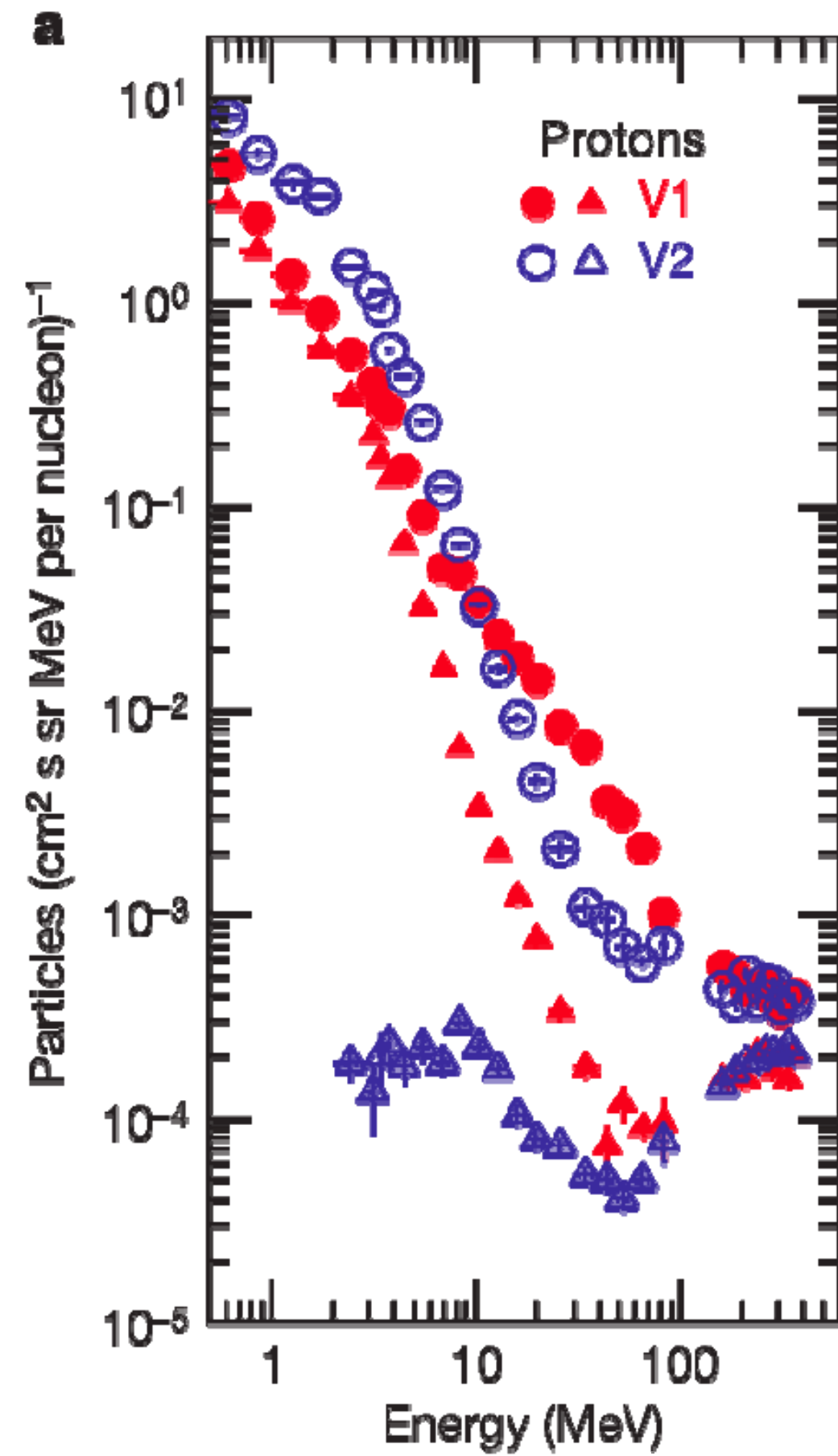
Nonthermal Universe

Cosmic Rays



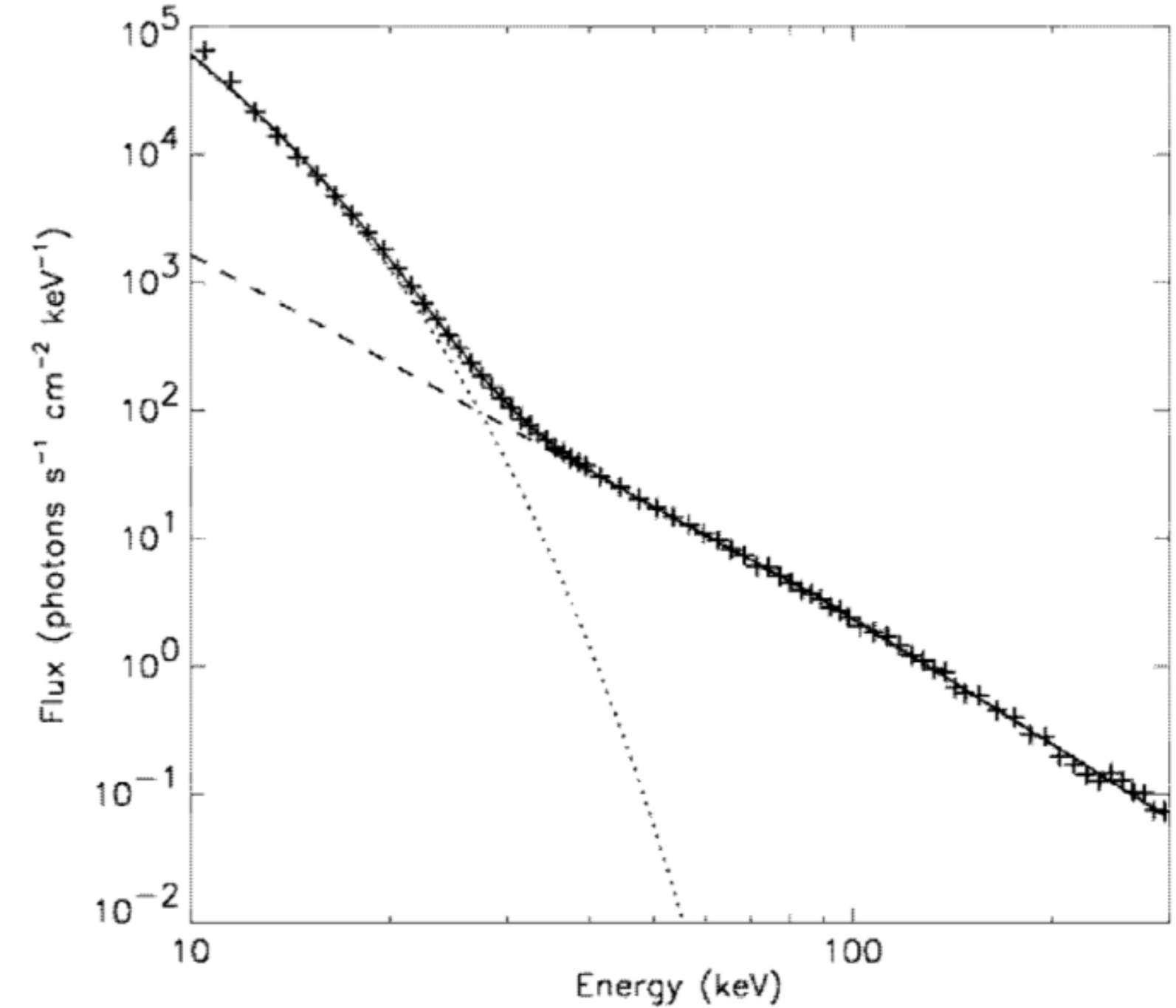
[Nagano & Watson, 2000]

ACRs



[Stone et al., 2008]

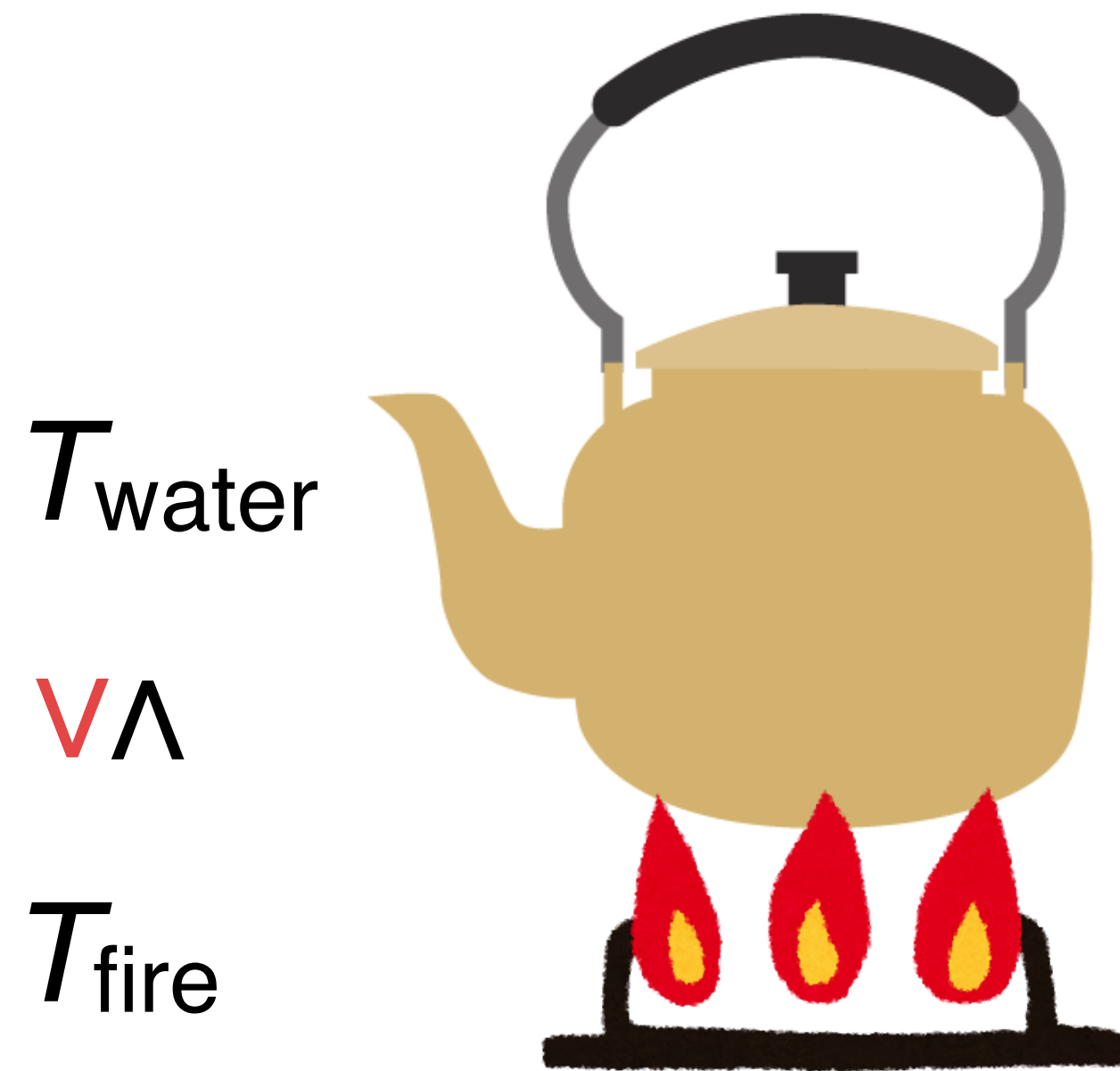
Solar Flares



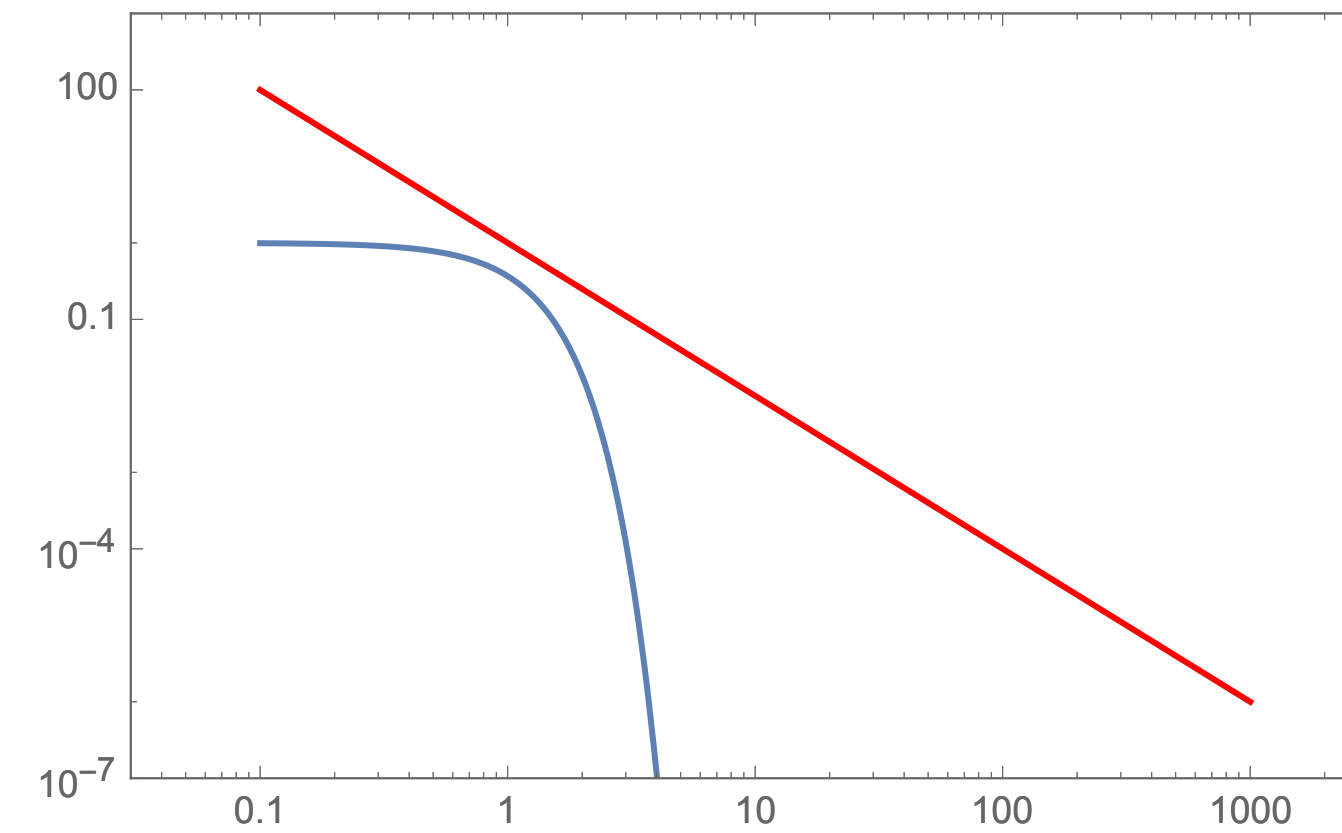
[Lin et al., 2003]

from Hoshino

Non-Thermal process



Water never be hotter than fire.

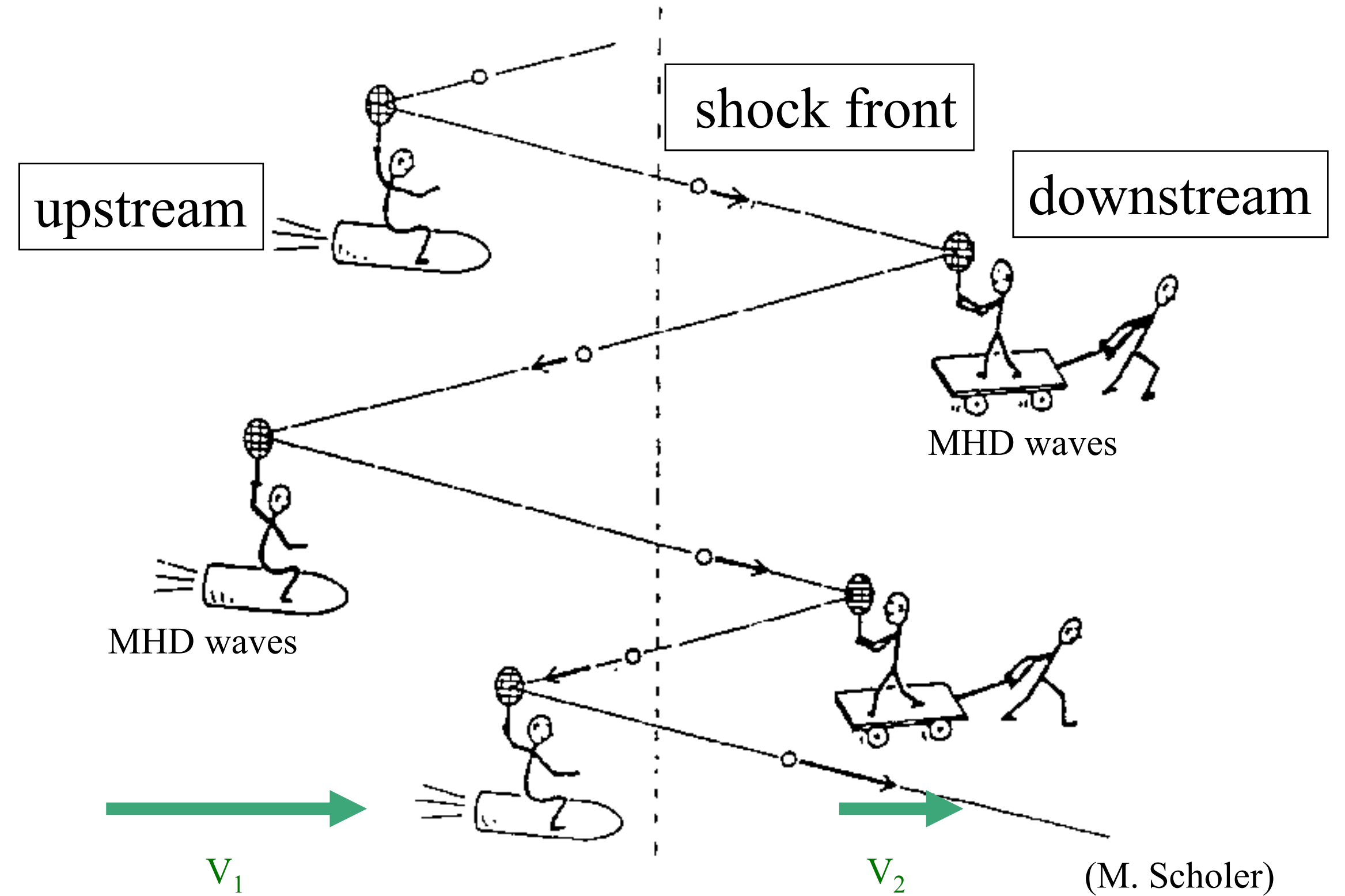


Thermal distribution

Nonthermal distribution

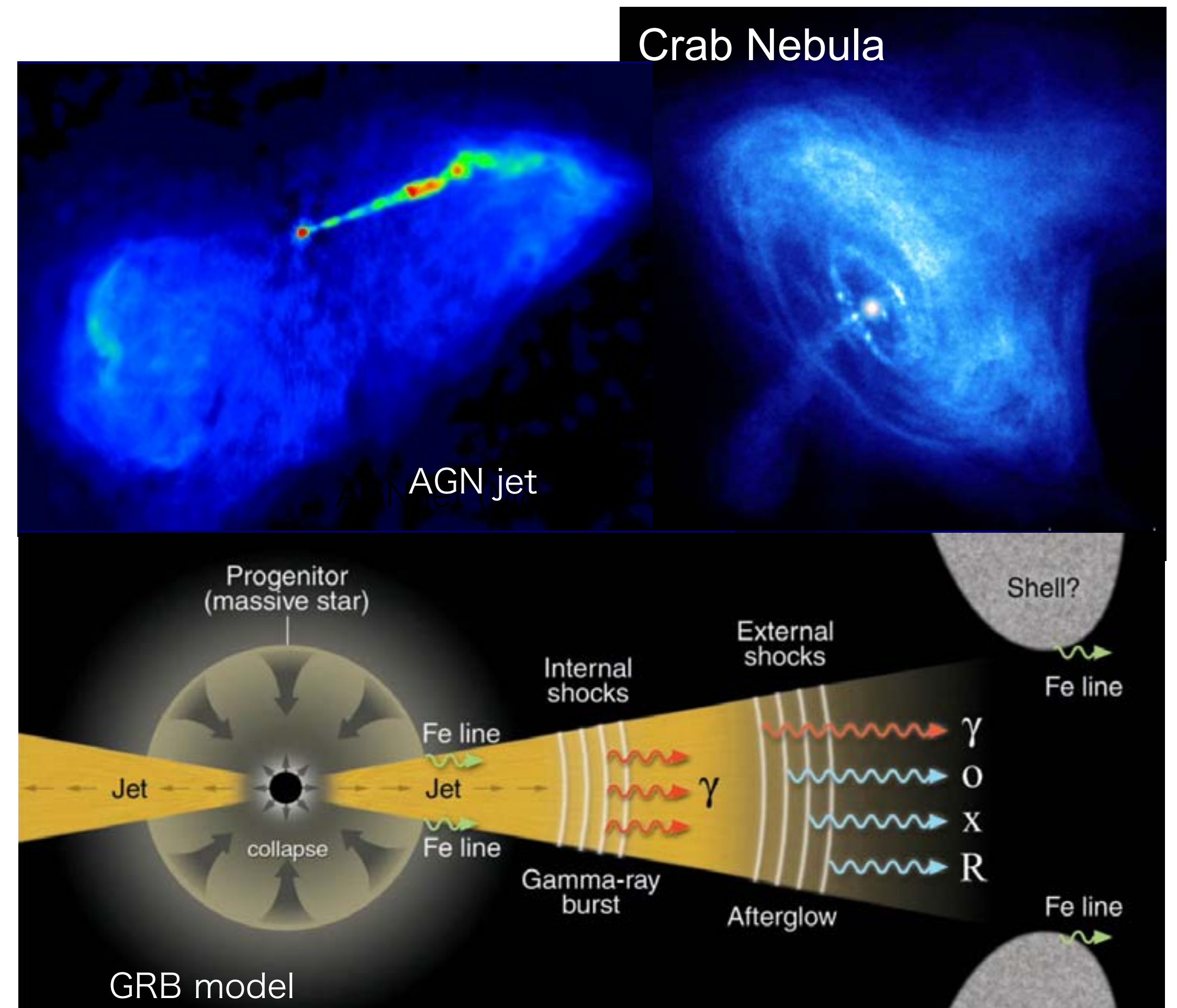
Diffusive shock acceleration

- Fermi acceleration (1949)
 - cosmic ray acceleration by moving magnetic fields
- Diffusive shock acceleration (DSA), Axford et al. (1977), Bell (1978), Blandford & Ostriker (1978)
 - Fermi acceleration at collisionless shocks
 - Always acceleration due to $U_1 > U_2$ for one cycle flight
 - High acceleration efficiency $\sim (U_1 - U_2)/c$
 - naturally and universally explains cosmic ray spectra,
 $f(\gamma) \propto \gamma^{-2}$



Extragalactic cosmic rays

- Possible sources: Relativistic collisionless shocks
 - Active galactic nucleus (AGN) jets ($\gamma \sim 10$)
 - Gamma-ray bursts ($\gamma > 100-1000$)
 - Pulsar wind ($\gamma \sim 10^{6-7}$)
- A possible mechanism
 - wakefield acceleration
Chen+ 2002 PRL
Lyubarsky 2006 ApJ
Hoshino 2008 ApJ
Kuramitsu+ 2008 ApJL
...
Iwamoto+ ...



from Hoshino

Wakefield Acceleration By Radiation Pressure In Relativistic Shock Waves

1. Shock formation
2. Excitation of electromagnetic (light) waves
3. Electrostatic field (wakefield) excitation by the light
4. Acceleration of particles by the wakefield

Two governing parameters

a_0 : normalized wave amplitude

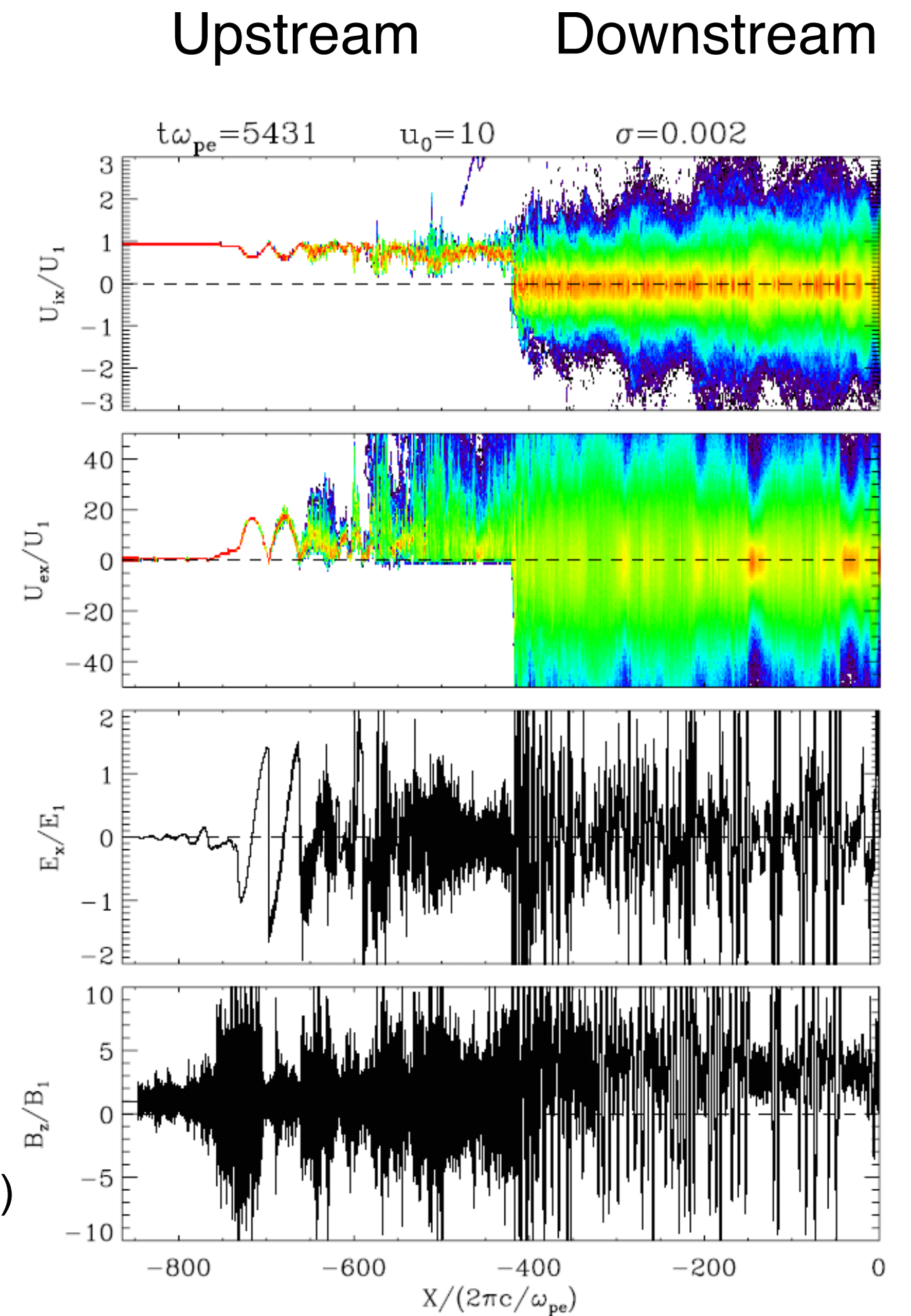
ω_p/ω_L : frequency ratio between plasma and light

ION

ELECTRON

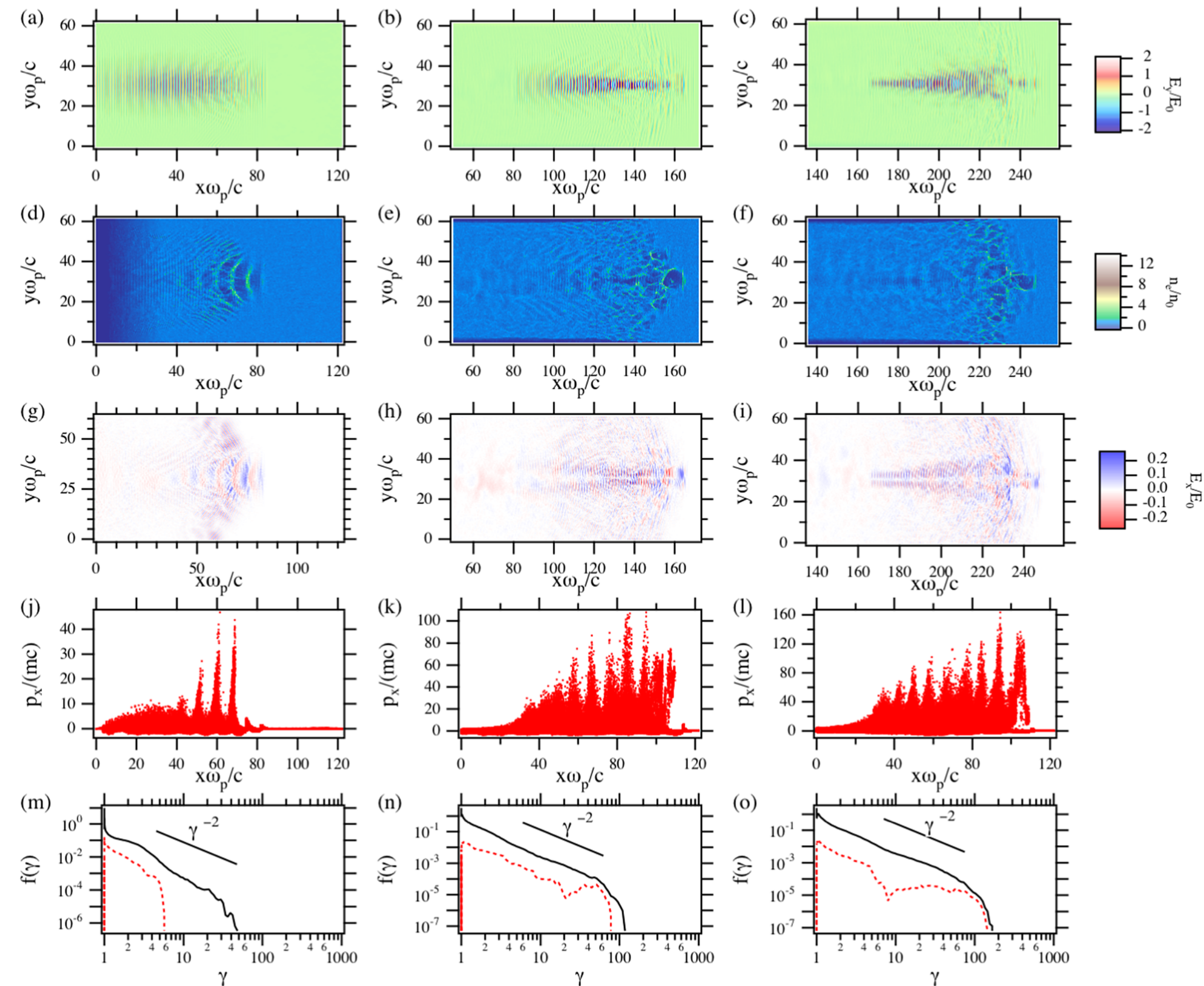
Electrostatic Waves (E_x),
(wakefield)

Electromagnetic Waves (B_z),
(precursor wave \rightarrow laser beam)



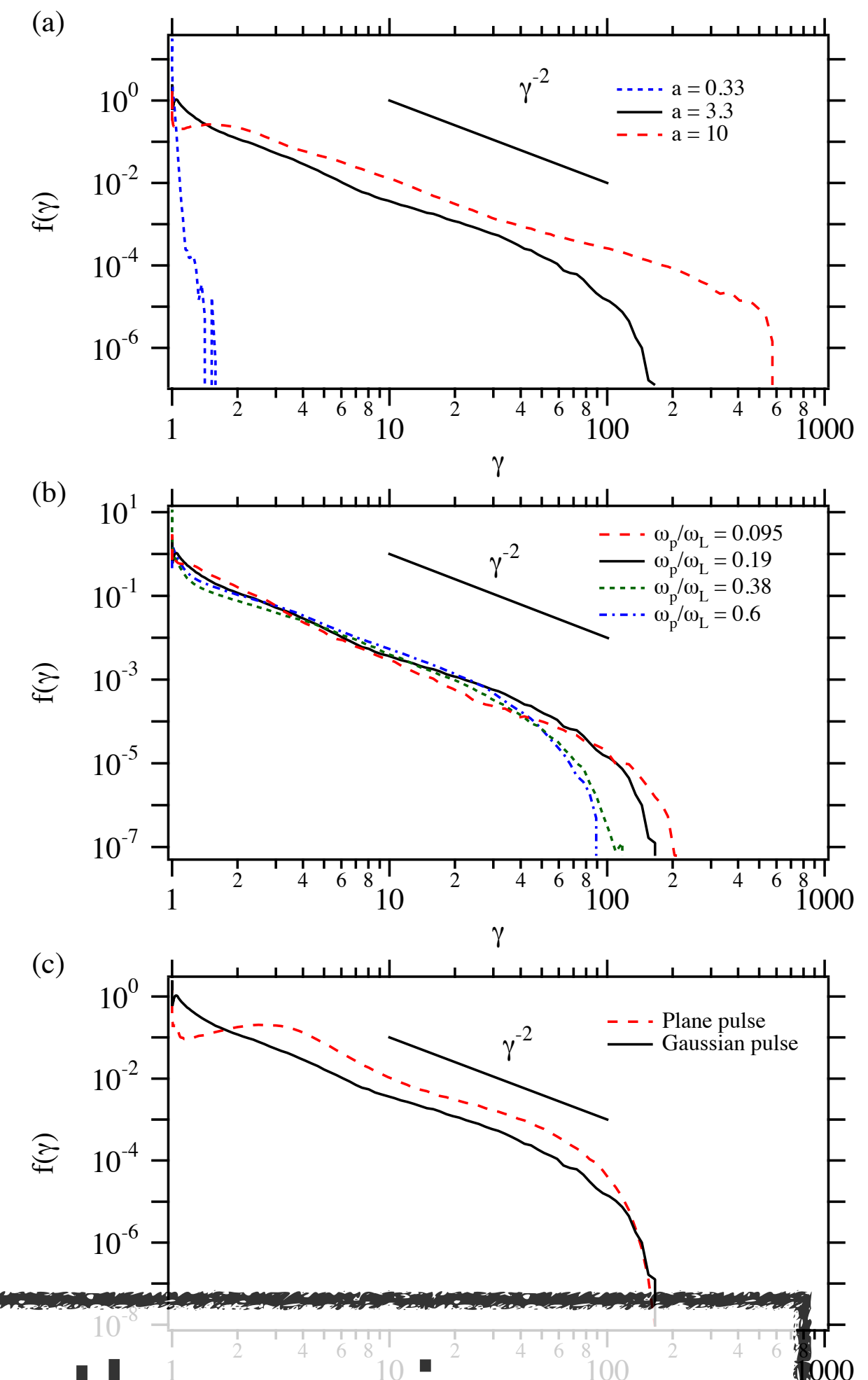
Nonthermal electron acceleration by turbulent wakefield

- Assuming large amplitude light waves propagating in a plasma,
- Independent of light amplitude $\sim a$
- Independent of plasma density $\sim \omega_p/\omega_L$
- Independent of pulse shape
- Universal production of power law spectra with an index of ~ -2
- Cyclotron and synchrotron emission free.



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It is impossible to observe this in the universe.

Model experiments of cosmic ray acceleration in laboratories (1)

- Astrophysical situation to be modeled is

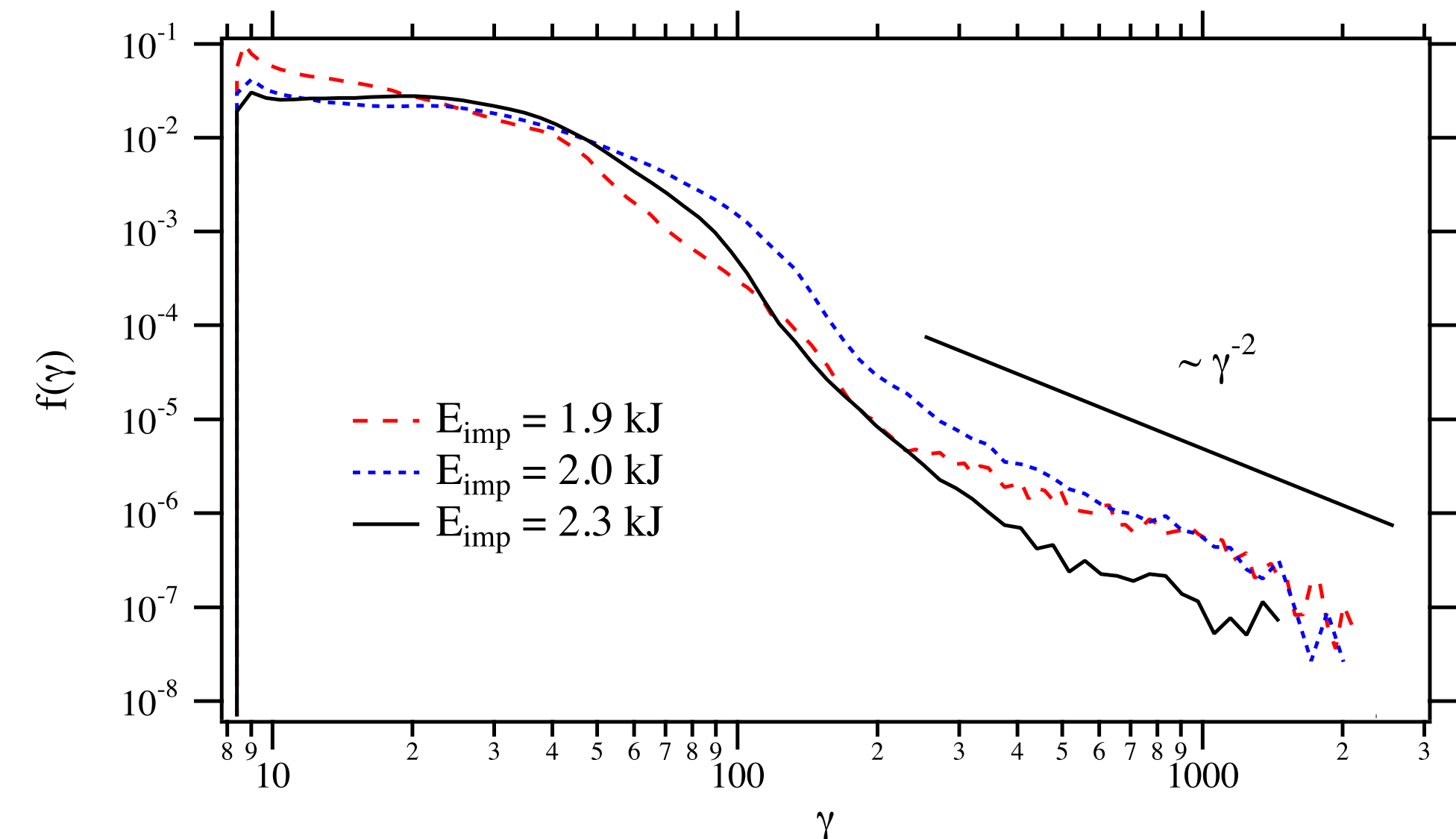
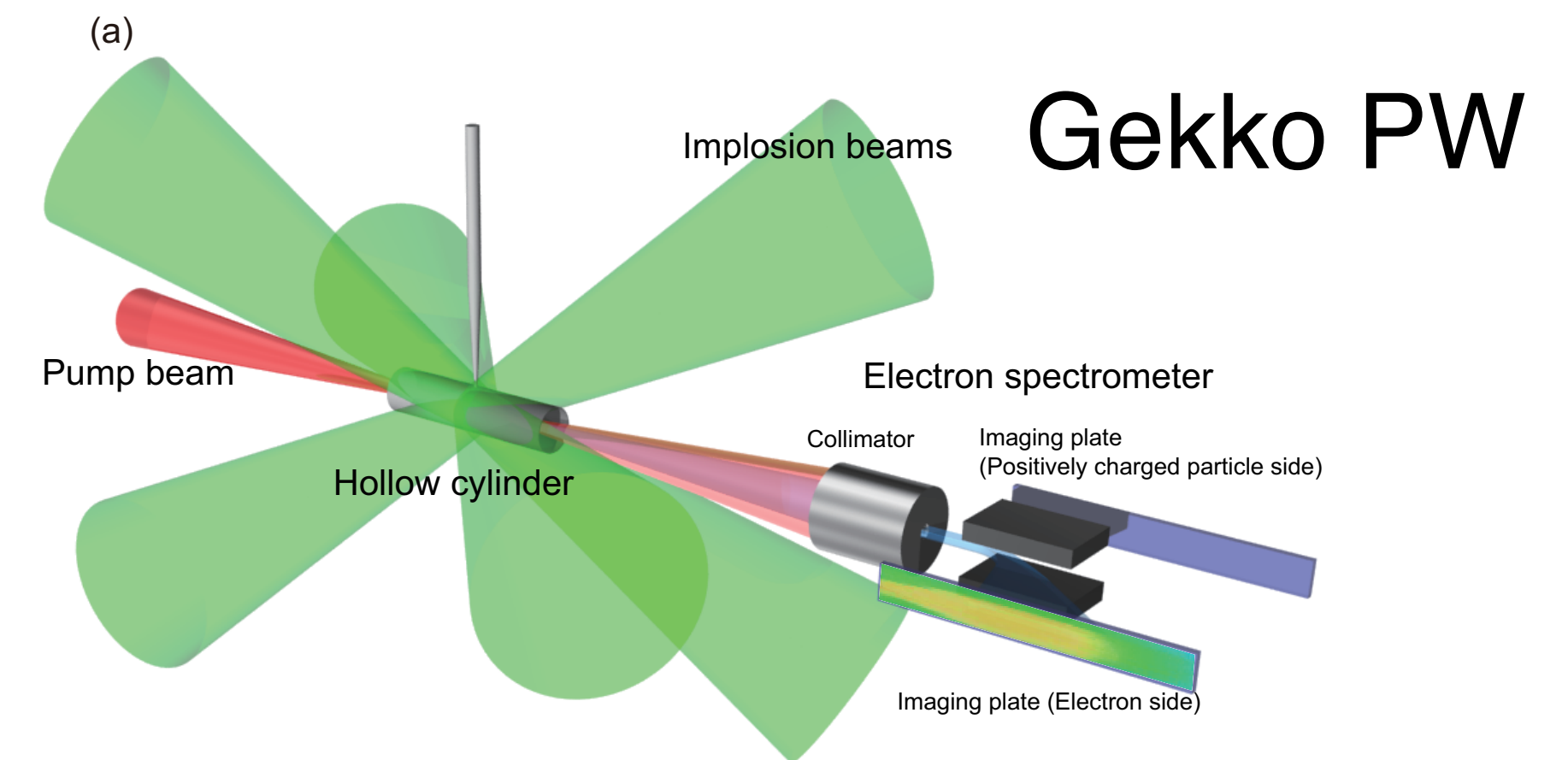
1. a large amplitude light pulse ($a > 1$)

➔ Gekko PW
(100 J, 700 fs, $a_0 \sim 1.9$)

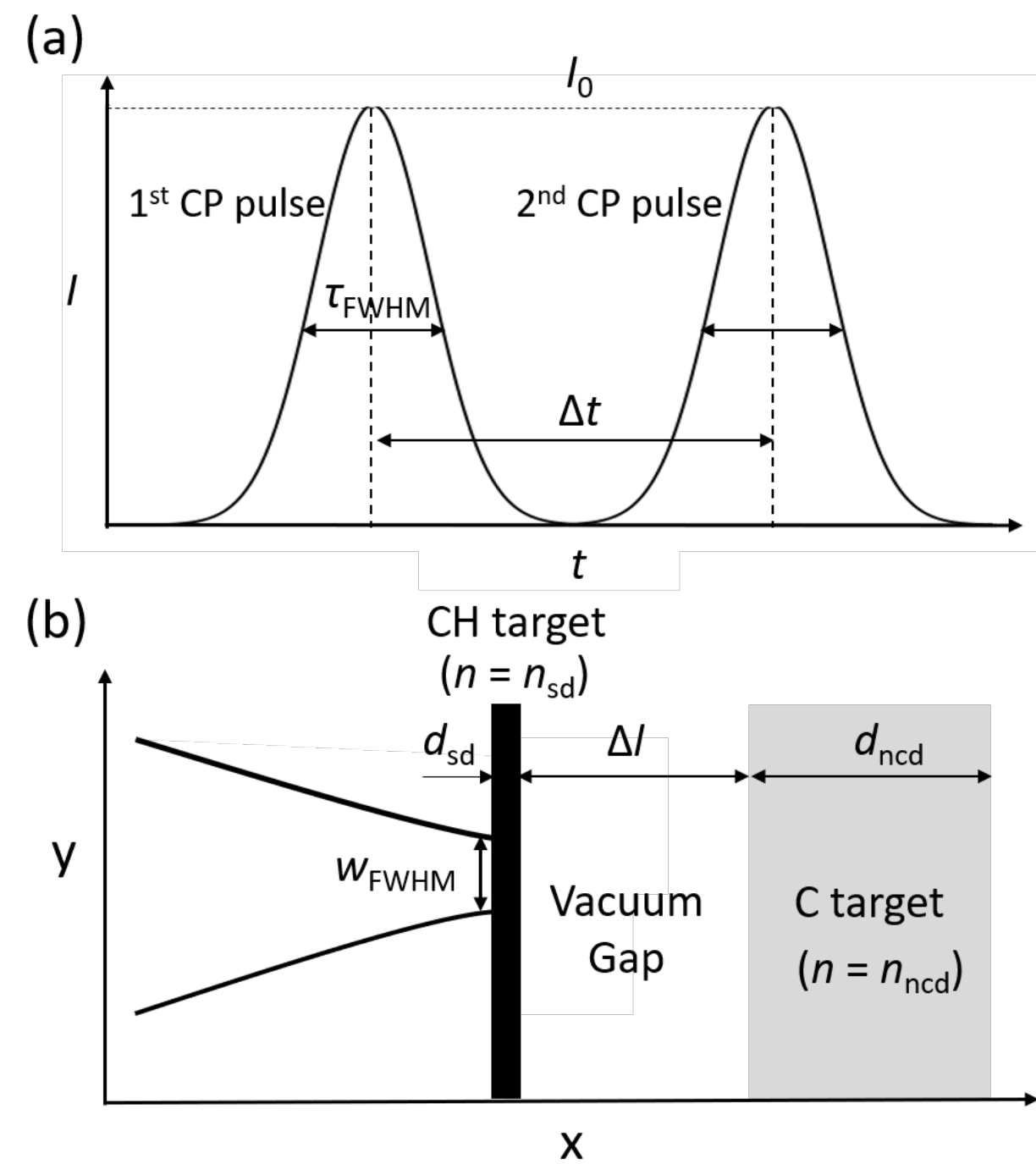
2. propagating in a plasma.

➔ Hollow cylinder implosion with Gekko XII

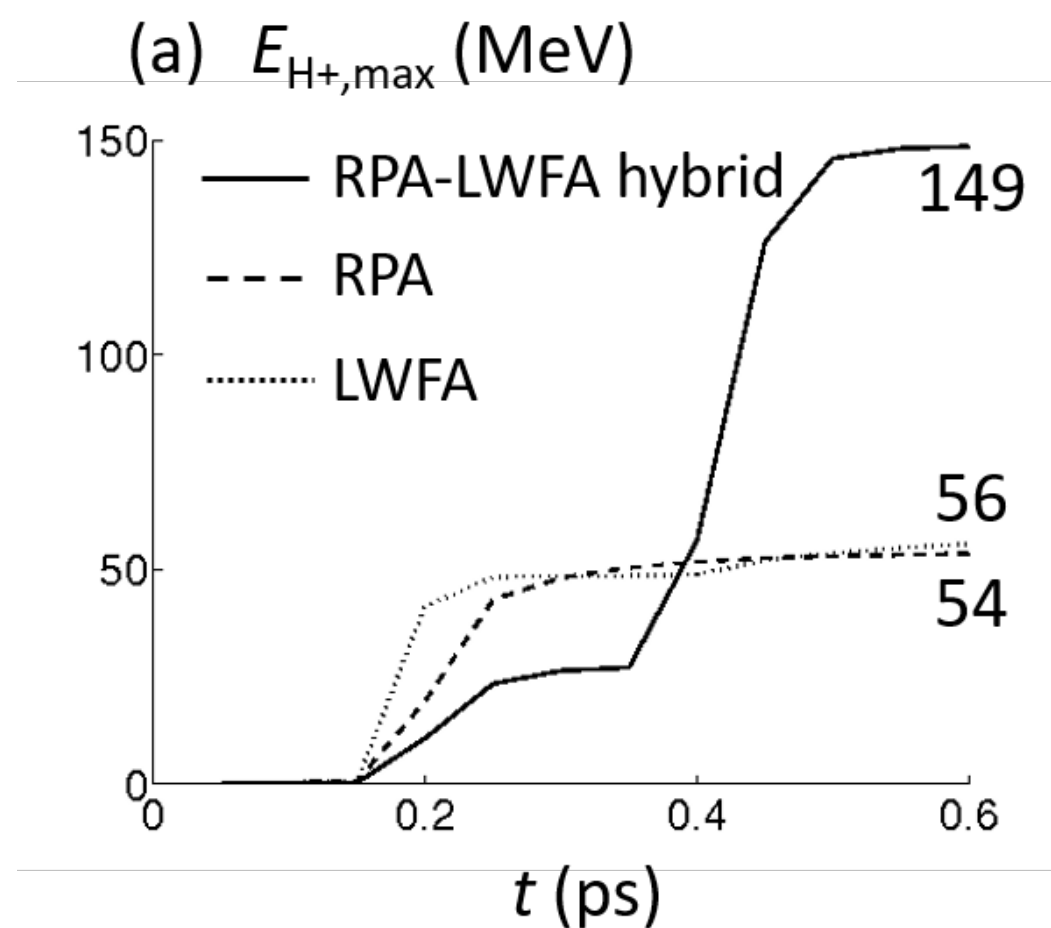
- Distribution functions of accelerated electrons are measured with electron spectrometer (ESM).
- Power law spectra independent of plasma density.



Relativistic ion acceleration



100 TW laser



- Graphene ion acceleration as the first stage
- Wakefield acceleration in the form target as the second stage
- Relativistic ion detectors
- Wakefield imaging with nonlinear Thomson scattering
- Machine learning on the detector

Controlled injection of energetic protons into the wakes

Frontiers

- Relativistic laboratory astrophysics
 - Energy Frontier
 - Relativistic ion acceleration
 - Relativistic ion detector
- Data science and Informatics
- Laser nuclear physics
- Extreme light field and plasmas
 - Induced Compton scattering from high brightness temperature radiation (13-5 Shuta Tanaka)
 - Thomson scattering of intense light from nonlinear plasmas (11-9 Kentaro Sakai)

Energy frontier in laser-driven ion
acceleration with large-area suspended
graphene with the aid of machine learning

Practical problems on ion acceleration experiment at relatively small laser facility

- radioactive contamination
- limited space and floor strength
- limited man power ...



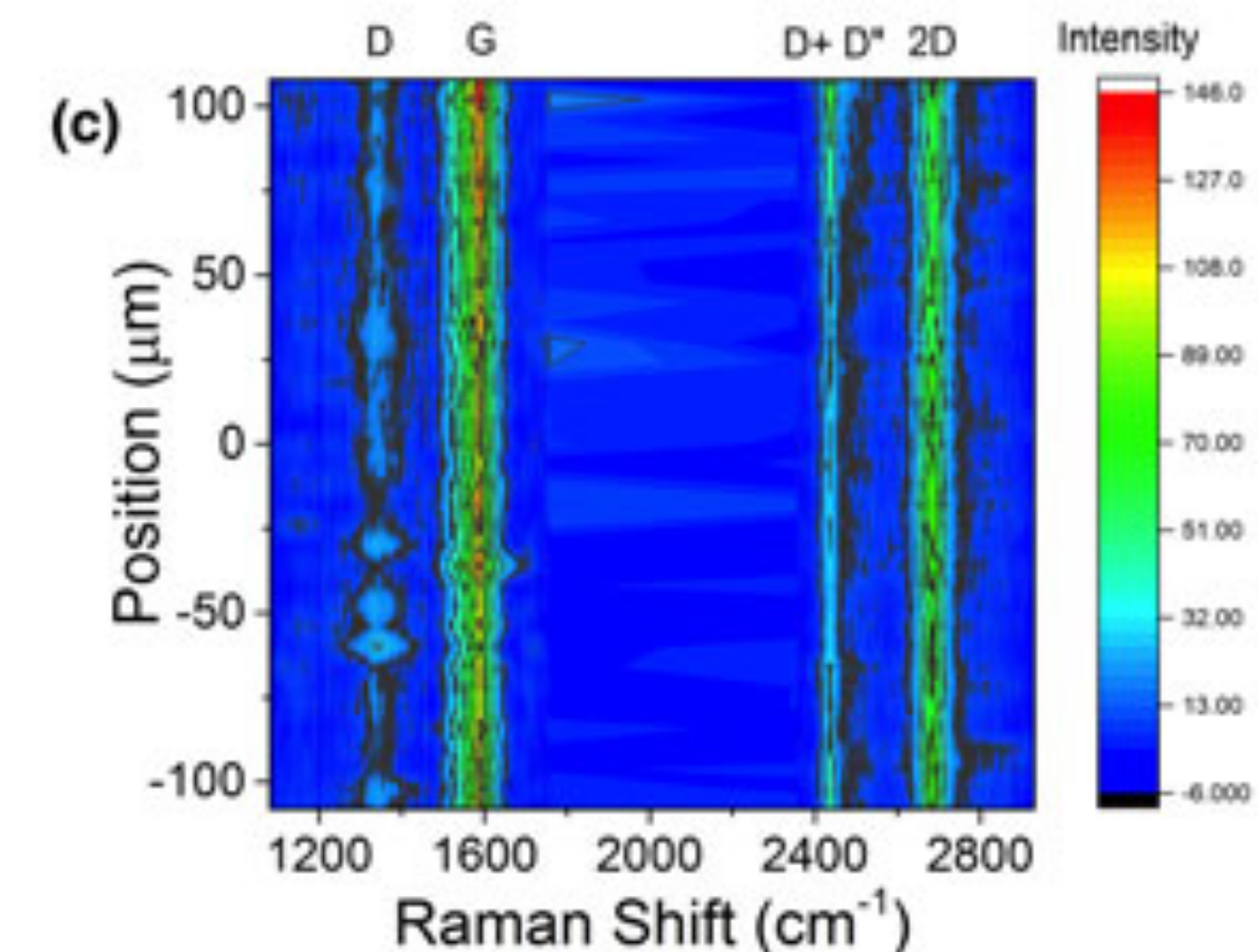
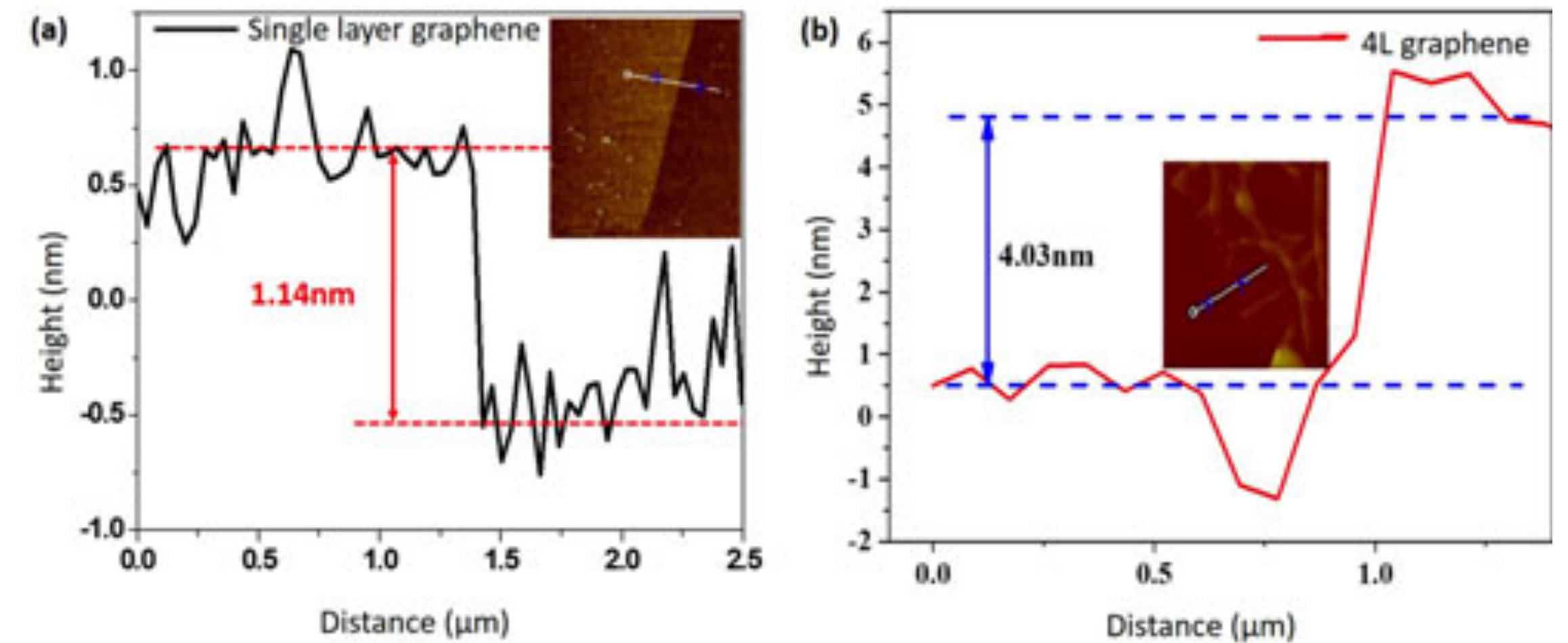
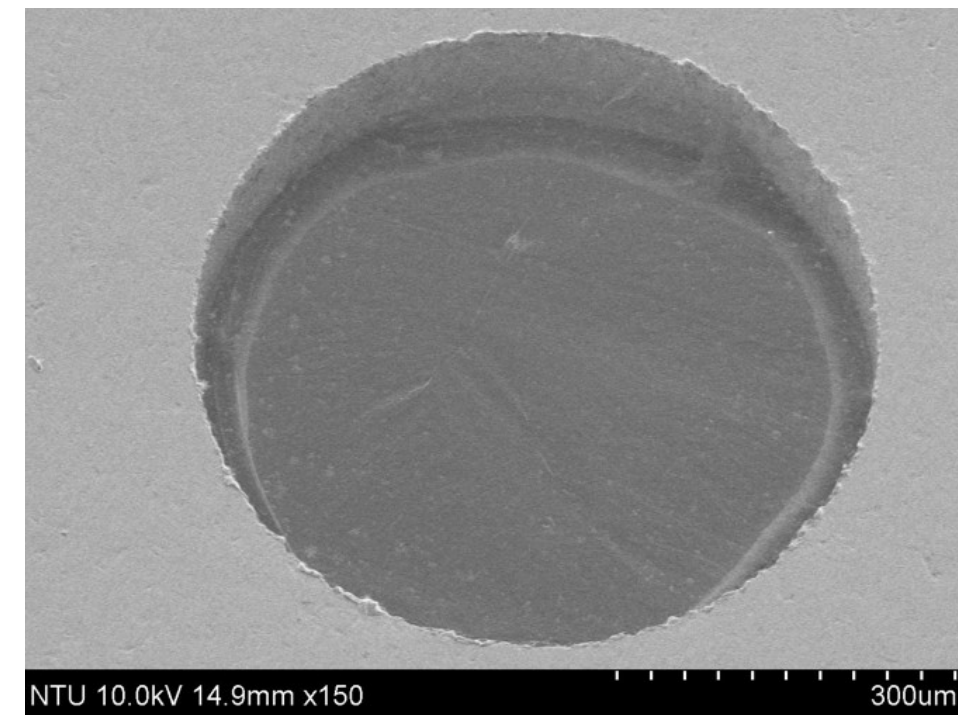
Need to suppress
radiation from
laser-matter interactions

These are also the case for medical applications

Large-area suspended graphene (LSG)

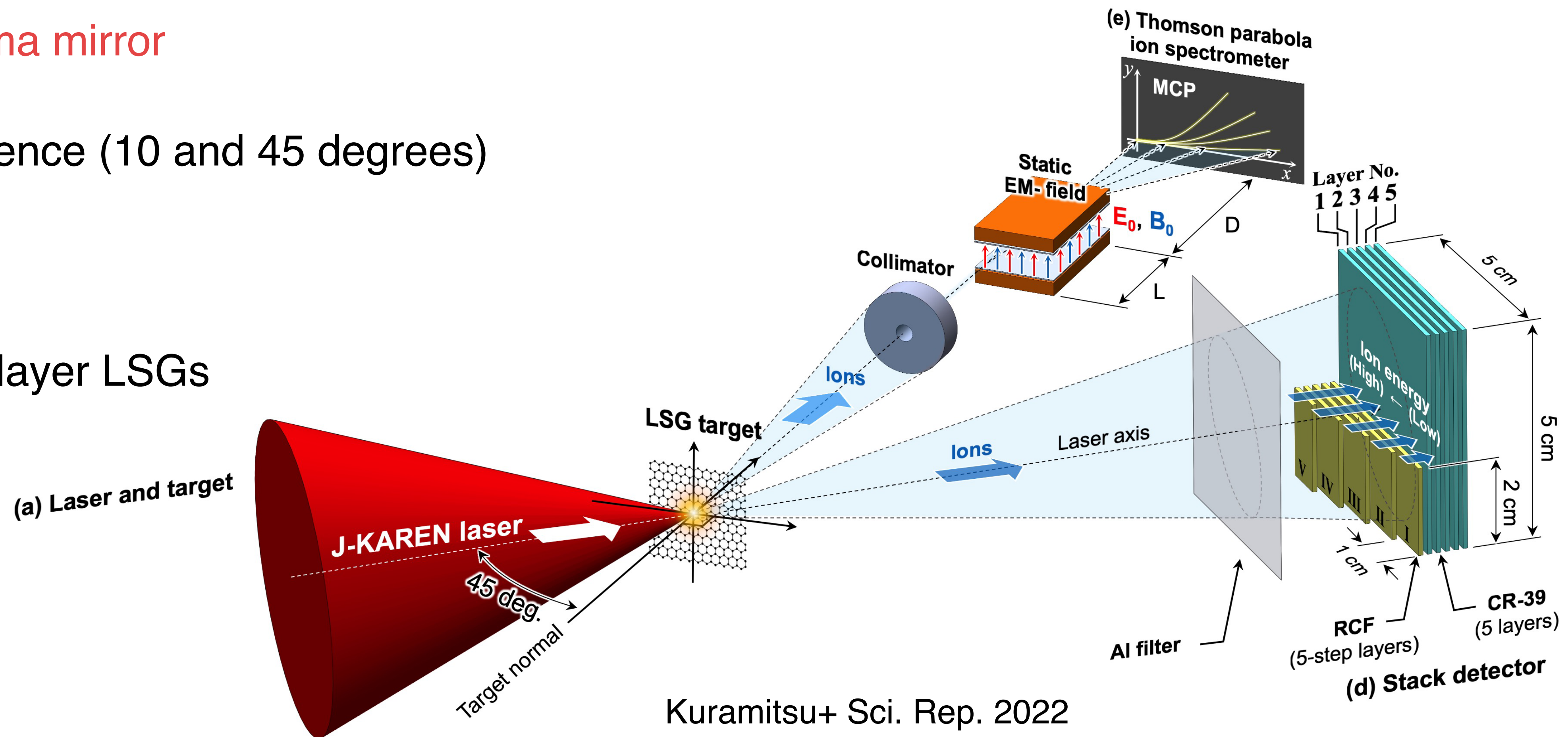
We have developed a large-area suspended graphene (LSG),

- Thinnest
- Lightest
- Strongest
- Transparent
- Thickness control at 1 nm by transferring layer by layer
- Extremely high thermal and electrical conductivity within the layer
- Reasonable



J-KAREN experiments

- 800 nm, 30 fs, 10J, 0.1 Hz, F/1.35, $5e21$ W/cm²
- Without plasma mirror
- Oblique incidence (10 and 45 degrees)
- Targets
 - 2, 4, and 8 layer LSGs



Best focus relativistic laser intensities

- Thomson parabola spectrometer with 8-layer LSGs

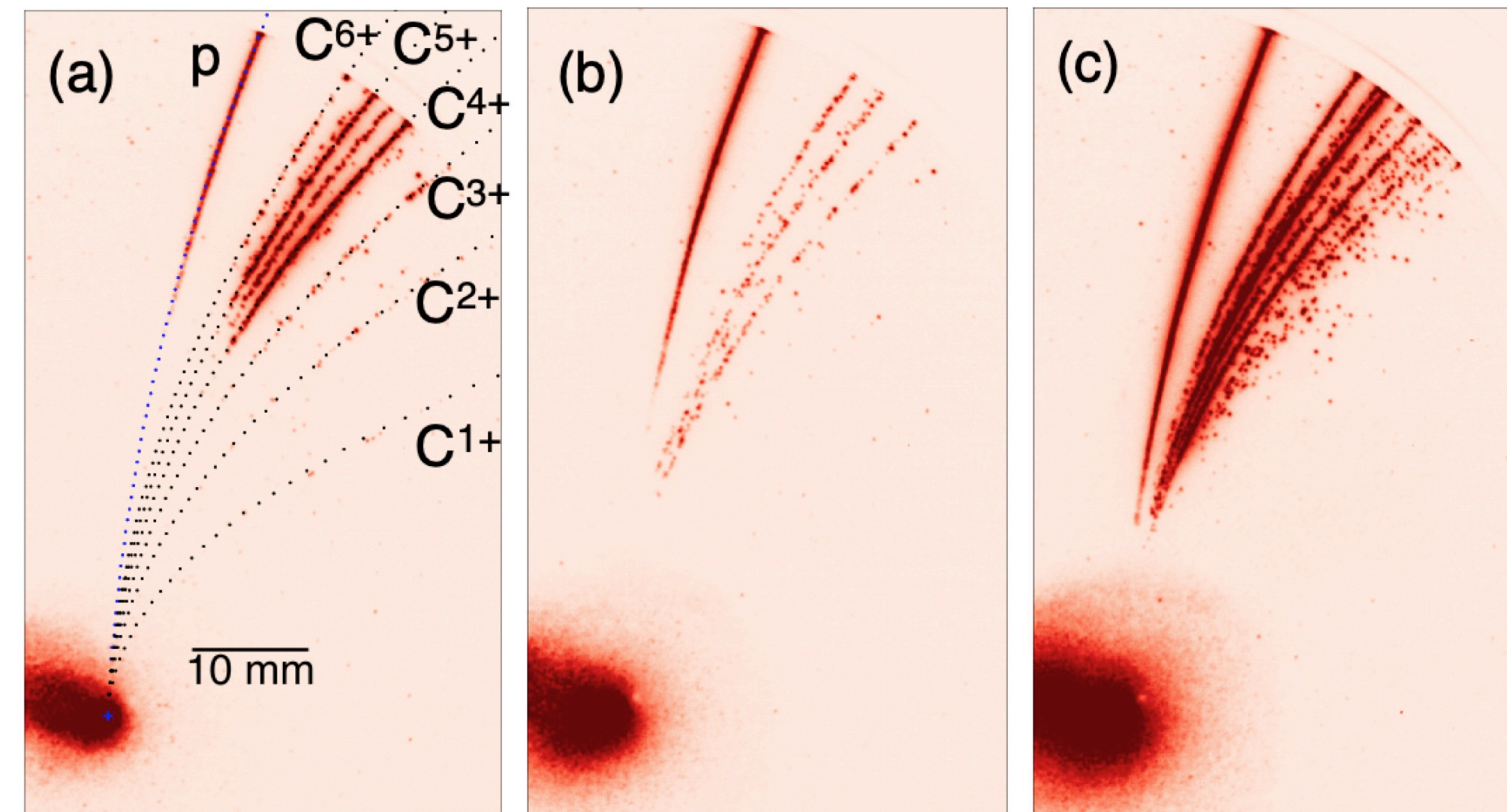
(a) $1.06 \text{ e}21 \text{ Wcm}^{-2}$

(b) $2.86 \text{ e}21 \text{ Wcm}^{-2}$

(c) $4.83 \text{ e}21 \text{ Wcm}^{-2}$

- $\sim 15 \text{ MeV}$ protons and $\sim 60 \text{ MeV}$ carbons

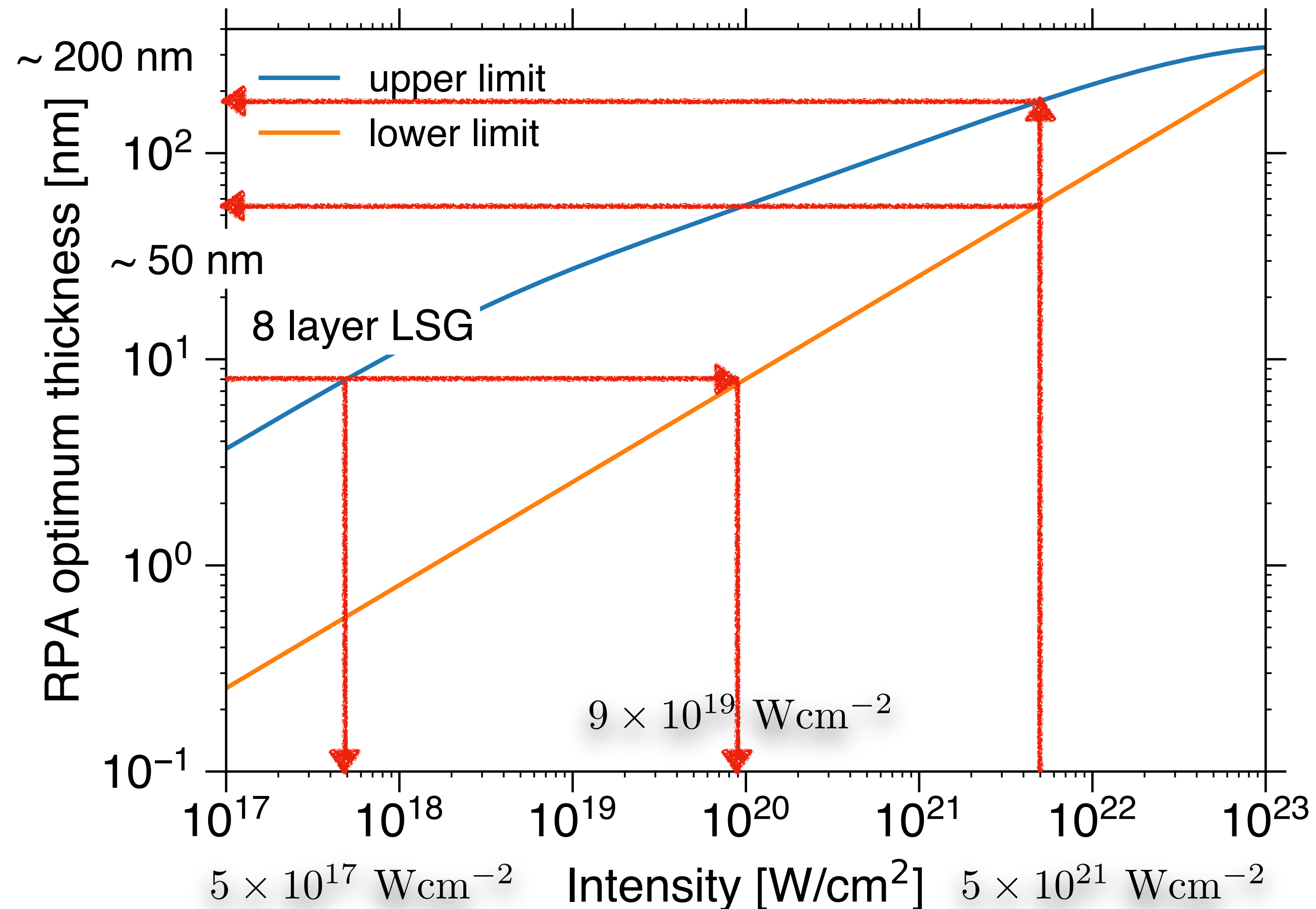
- Without plasma mirror



Kuramitsu+ Sci. Rep. 2022

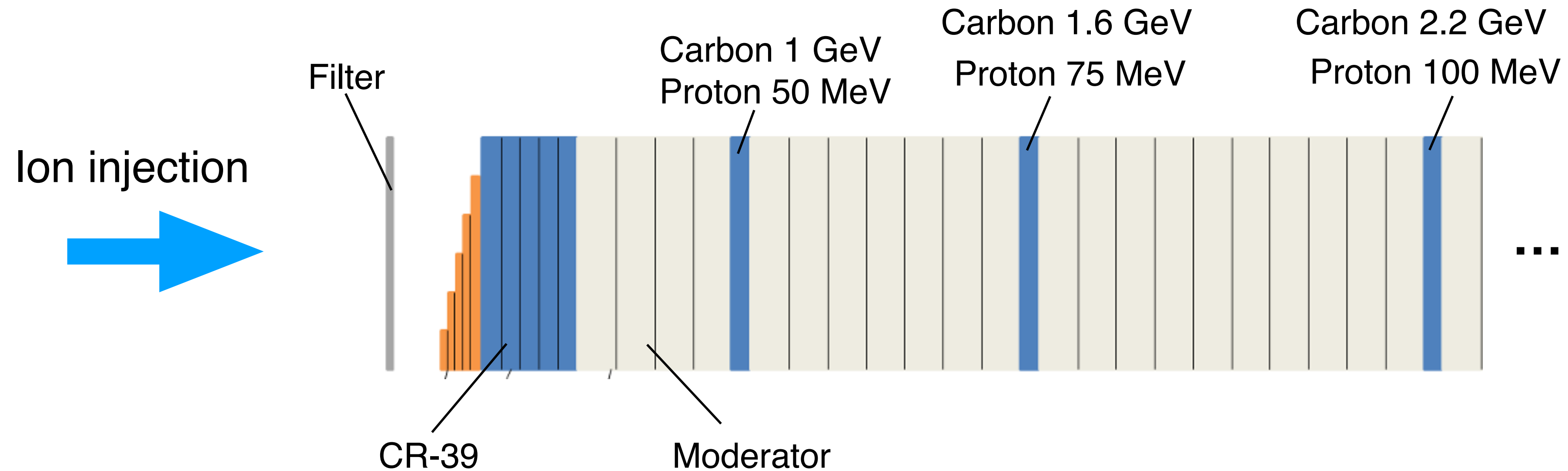
Irradiating the thinnest target by the highest intensity laser without plasma mirror to demonstrate robustness of LSG \Rightarrow Not optimized yet!

1. LSG optimization to J-KAREN laser
2. J-KAREN optimization to LSG



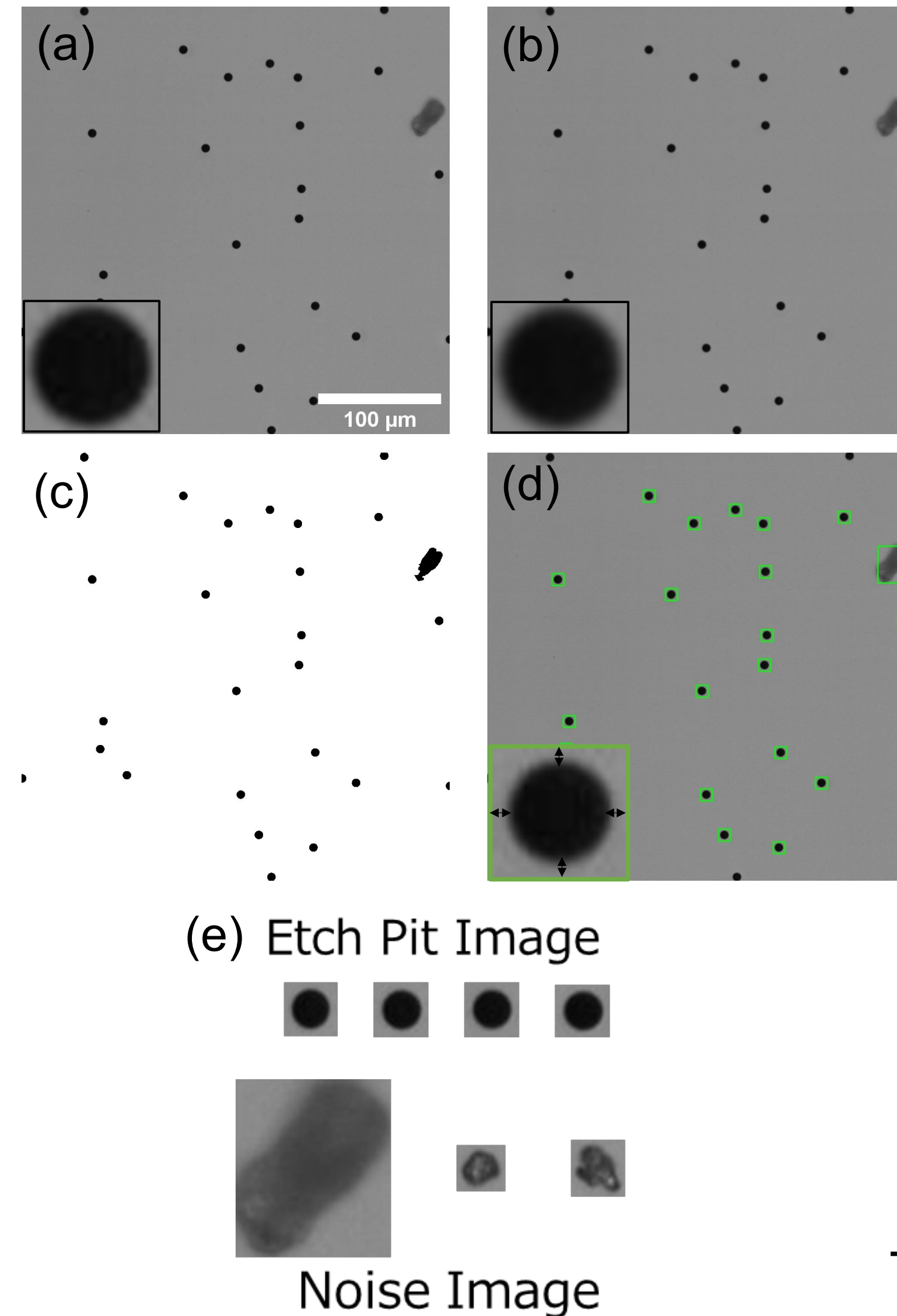
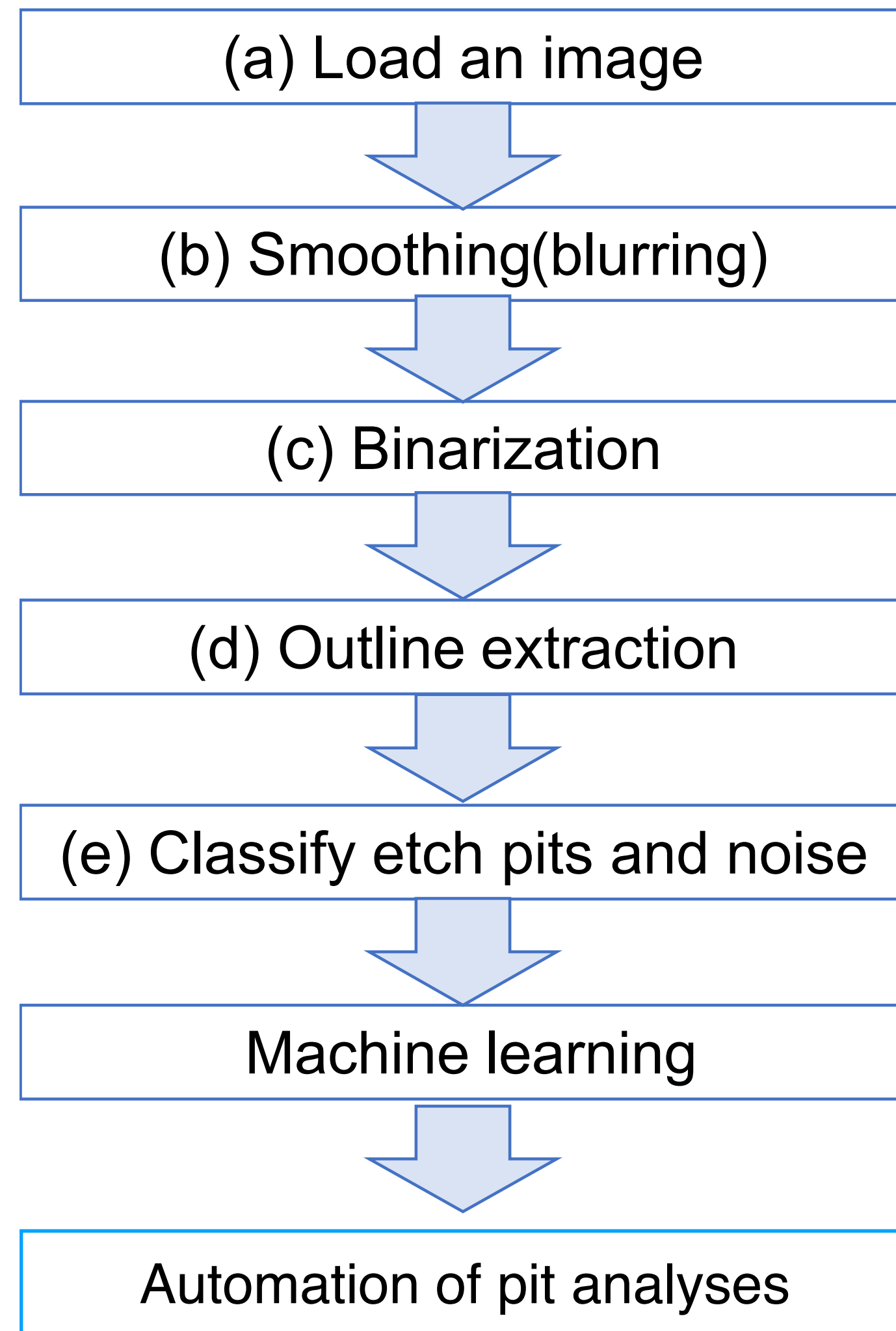
CR-39 stack

To resolve ion energy using CR-39

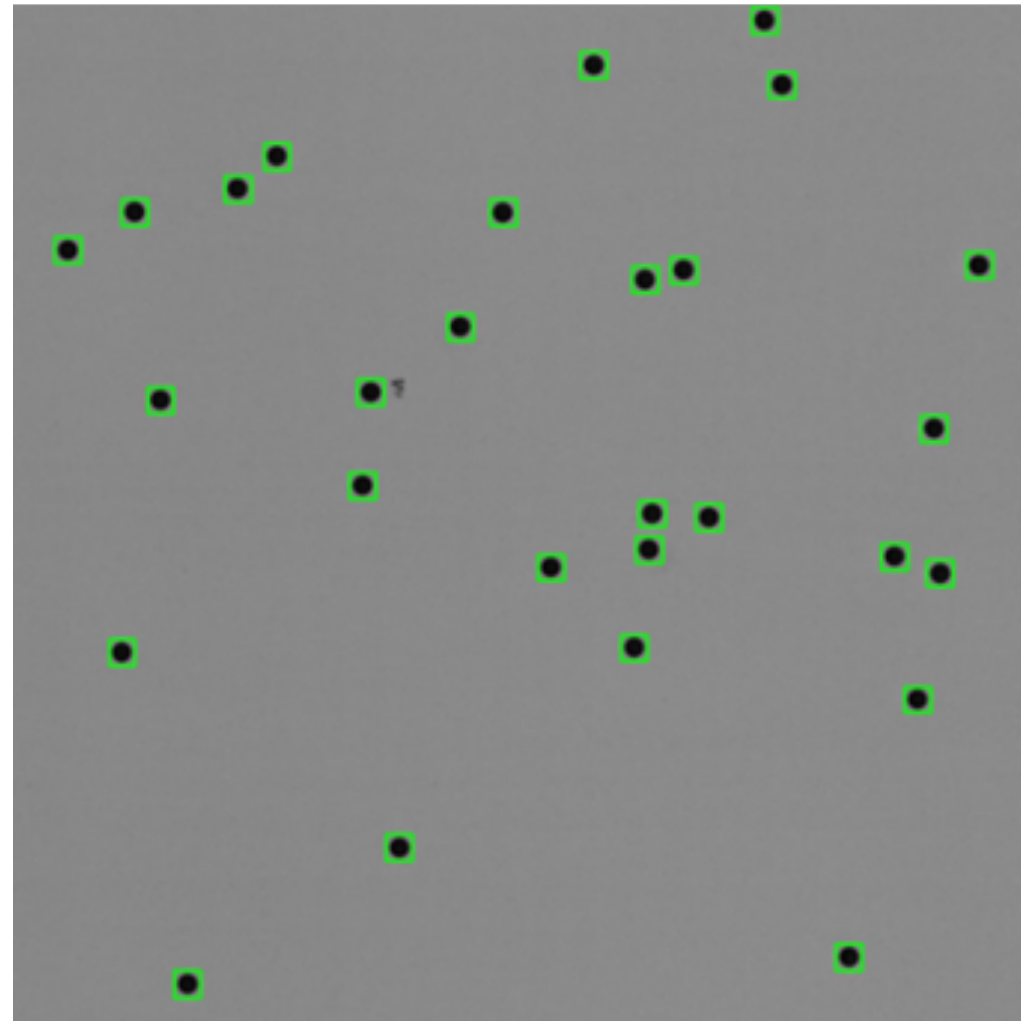


- To obtain ion spectra with CR-39 stack, it is required to find etch pits in large amounts of microscope images.
- ~10 CR-39 in 1 stack, ~10,000 microscope images in 1 CR-39 sheet
- **Millions of images should be analyzed in 1 experiment series.**

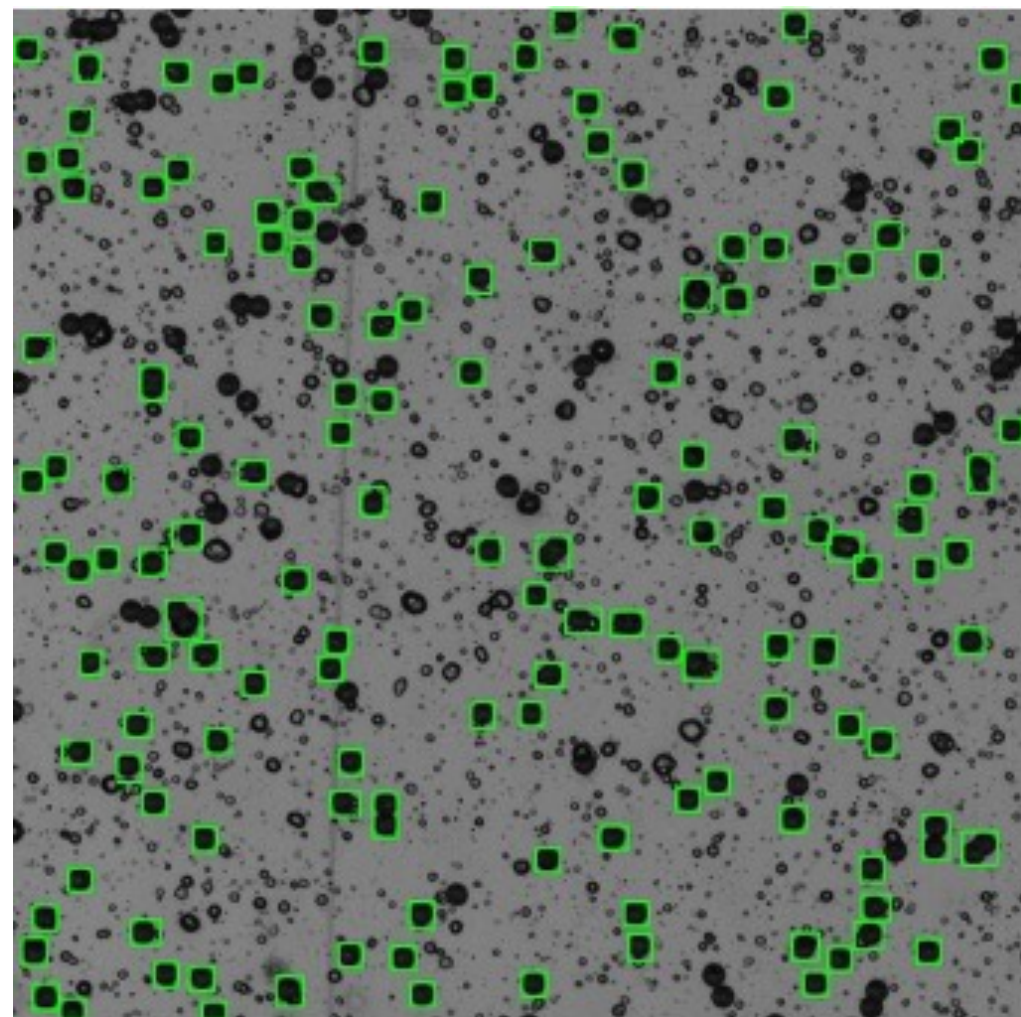
Automation of ion pit analyses with machine learning (ML)



ML (ExtraTreesClassifier)

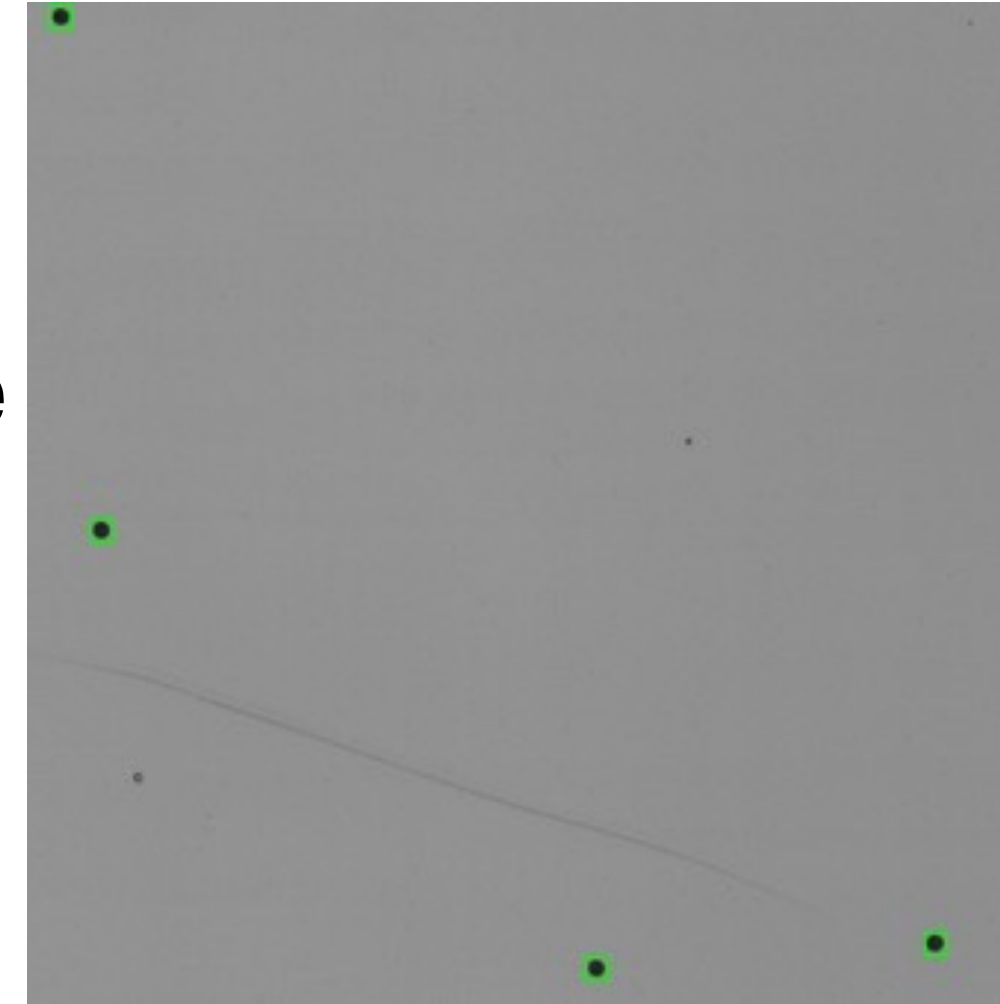


- HIMAC data
- Training data from HIMAC
- More than 5000 microscope images
- About 10^5 pit detection
- Precision: 98%
- Recall: 98%

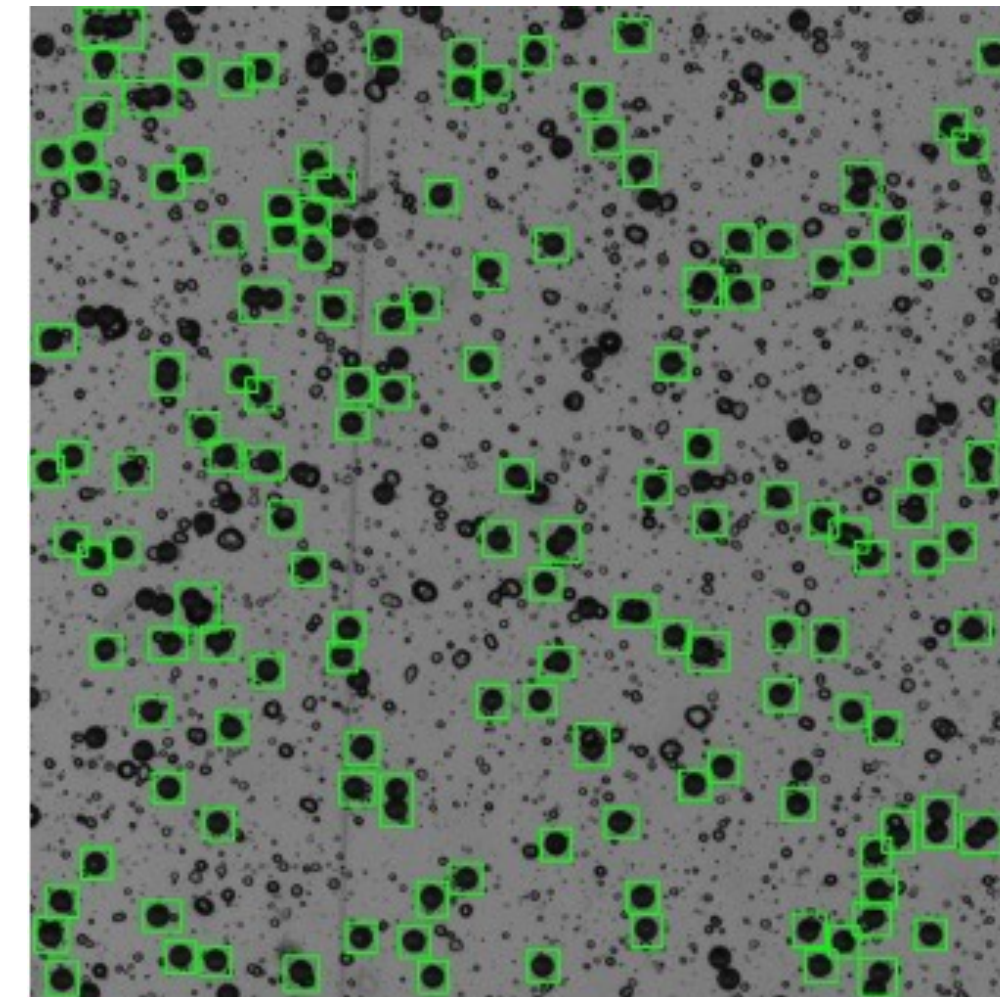


- LFEX data
- Training data from LFEX
- More than 17000 microscope images
- About 10^5 pit detection
- Cropping margin: 3 pixels
- Precision: 95%
- Recall: 76%

CNN (VGG16)



- LFEX data
- Training data from HIMAC
- All the ion pits are detected.



- LFEX data
- Training data from LFEX
- Cropping margin: 5 pixels
- Precision: 95%
- Recall: 83%

Summary 1

- We are exploring relativistic laboratory astrophysics aiming at relativistic ion acceleration relevant to cosmic rays.
- We have developed large-area suspended graphene as targets for laser-driven ion acceleration.
- We optimize the ion acceleration in two ways, LSG to laser and laser to LSG, and both successfully produce energetic protons and carbons.
- 132 MeV protons are accelerated with long (1.5 ps) and lower intensity laser ($\sim 10^{19}$ Wcm⁻²) and identified with machine-aided ion pit analyses.

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Core-to-Core Program



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